



Earliest MACRO Data Contained Evidence for Neutrino Oscillations

Neil D Pignatano
(RTX ret., ex-Caltech, ex-NASA/JPL)

When you have eliminated the impossible, whatever remains, however improbable, must be the truth. – Sherlock Holmes



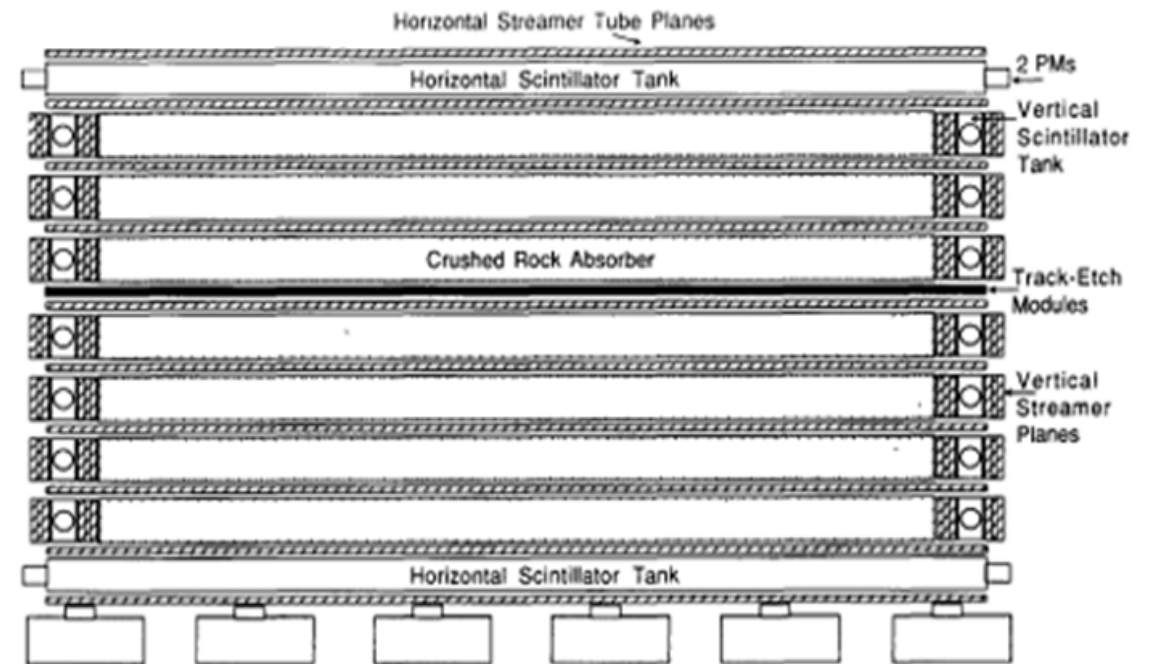
Monopole, Astrophysics, and Cosmic Ray Observatory

ABSTRACT

The earliest data collected by MACRO was analyzed for upward going muons, indicative of neutrino interactions in the ground a few meters below the detector. The sources of these neutrinos were expected to be from pion/kaon decay due to cosmic ray interactions, and, at much lower intensities, galactic (e.g., X-ray binaries) or extra-galactic (e.g., Active Galactic Nuclei) sources. When data from the first super-module's inaugural run was analyzed, a lower number of upward going muons was observed than expected, albeit not up to the standard of statistical significance required, and was dismissed as a statistical anomaly. Other aspects of the analysis however, lent credence to the veracity of the finding. This analytical approach is described in detail.

First Supermodule

- Central and bottom scintillator planes only
 - ERP muon trigger
- Streamer tubes and strips
 - Bari ST trigger
- Data set
 - Runs 5 – 440, 22 Feb – 30 May 89
 - Runs 673 – 4162, 10 Oct 89 – 15 Nov 91
 - Live time: 557d 01h 45m
- Custom analysis software
 - Calibration
 - Tracking
 - Prediction/expectation





Monopole, Astrophysics, and Cosmic Ray Observatory

Timeline

- October 1990
 - Capri Collaboration Meeting
 - Low up μ event rate discussed
 - Detailed analysis on detector efficiencies as source of low rate
- October 1991
 - Bologna Collaboration Meeting
 - Low up μ event rate looks real
 - Detailed analysis of other potential reasons for low rate
 - Statistical fluke, hitherto undiscovered detector inefficiency, MC error, *et al.*
- March 1993
 - Caltech Collaboration Meeting
 - Low up μ event rate must be real

It's All About the Calibration

- Mechanical (geometry)
 - Location of streamer tubes and strips
 - Survey provided
 - Location of scintillator tanks
 - Surveyed *in situ*
- Electrical
 - TDC and ADC calibration
 - Pedestals
 - Timing offsets
 - Gains
 - 13 calibrations performed (~ one per live month)
- TDC Calibration
 - Used sample of through-going muons
 - One ERP and Stream Tube muon trigger
 - One reconstructed track in each streamer tube view
 - One and only one scintillator channel hit in each layer
 - Streamer tube track must pass through both tanks with a minimum 19 cm path length
 - About 40% of all through-going muons in sample

Scintillator Position and Time of Flight

- Position in Scintillator

- $R_{SC} = k_0 T_0 - k_1 T_1 + C_R$
 - Where k_s are conversion gains (cm/digital clock tick)
 - T_s are TDC data
 - C_R is a delay offset
 - R_{SC} is the distance from “0” tank end
 - Gains and offset determined by linear least squares minimization

- Time of Flight

- $tof = 0.5v^{-1}[(k_0 T_0 + k_1 T_1)_C - (k_0 T_0 + k_1 T_1)_B] + C_{tof}$
- Where v is speed of light in scintillator
- C_{tof} is a timing offset
- tof is the time of flight measured by streamer tube track (path length between scintillators divided by speed of light in vacuum)
- Corrections due to PMT pulse height variations (“time walk”) were performed



Monopole, Astrophysics, and Cosmic Ray Observatory

Event Selection

- Main criteria: ERP \wedge Bari \wedge 1 reconstructed track in each view
 - 1 388 584 total events
- From these, select only those events which have at least one ERP hit for the bottom and central scintillators layers only
 - 868 616 events
- Additional cuts
 - Track must pass through one “hit” tank in each layer
 - ADC/TDC data valid
 - $|\text{TDC pos} - \text{Track pos}| \leq 75 \text{ cm}$
 - $0.5 \leq E/\langle E \rangle \leq 5$ (pulse height consistency)
 - $(E_{\text{max}}/E_{\text{min}})_{\text{tank end}} \leq 1.5$
 - Vert face channels < 2
 - **→ 778 448 events in final data sample**

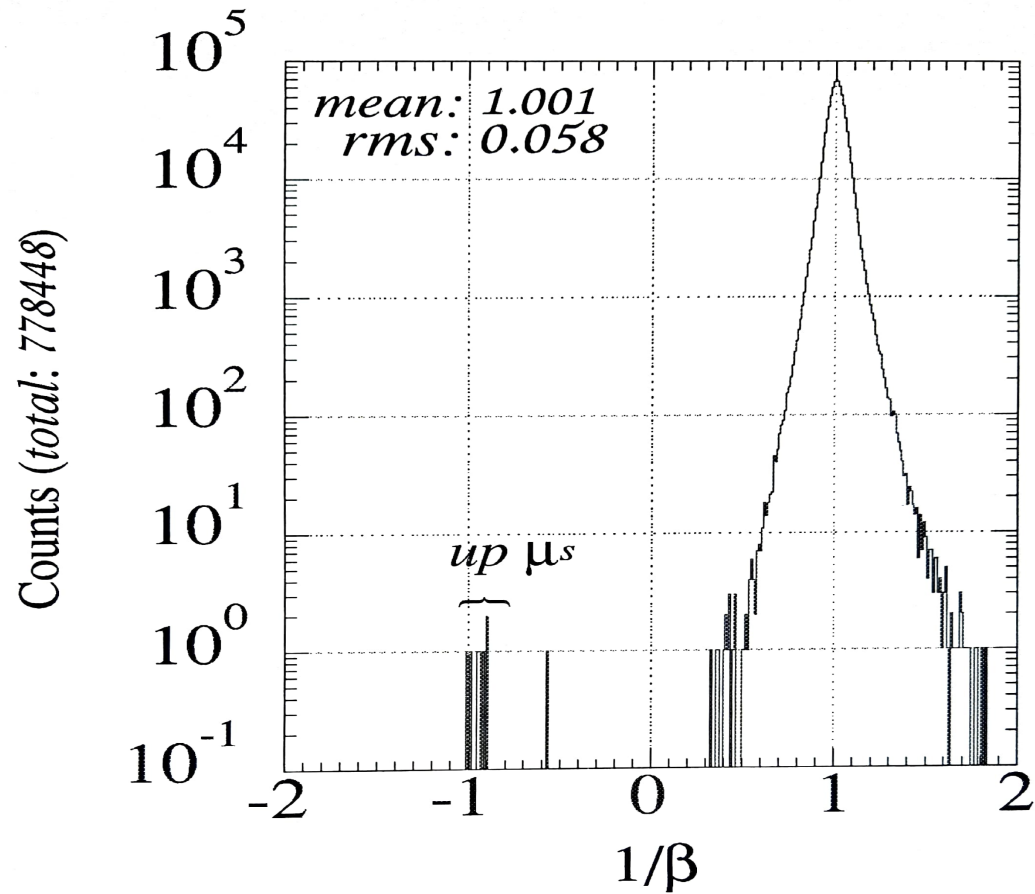


Monopole, Astrophysics, and Cosmic Ray Observatory

Detection Efficiency

- Cut efficiency is calculated
 - $\epsilon_{\text{cut}} = 778448 / 868616 = 0.896$
- Determine trigger efficiency by studying (see my Capri '90 talk for full details)
 - $N_{\text{obs}} = \epsilon_{\text{ERP}} \epsilon_{\text{Bari}} \epsilon_{\text{Tr}} N$
 - $\epsilon_{\text{ERP}} = 0.975$ (ERP trigger w/ hit in each horizontal plane)
 - $\epsilon_{\text{Bari}} = 0.995$ (Bari streamer tube trigger)
 - $\epsilon_{\text{Tr}} = 0.960$ (track reconstruction)
- Net efficiency
 - $\epsilon = \epsilon_{\text{ERP}} \epsilon_{\text{Bari}} \epsilon_{\text{Tr}} \epsilon_{\text{cut}} = \underline{\underline{0.835}}$

Down and Up μ Counts





Monopole, Astrophysics, and Cosmic Ray Observatory

What is Expected?

- Downward muons
 - Produced from π and K decay
 - Use Gaisser (1989) absolute flux
 - Propagate through Gran Sasso rock overhead to produce muon flux at MACRO
 - Monte Carlo simulation provides event rate
 - And “calibrates” MACRO’s effective area

Use muon and neutrino fluxes from Gaisser (1989):

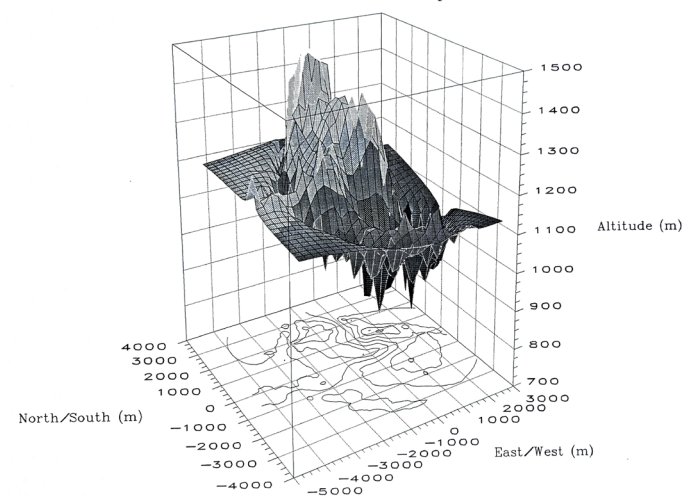
$$\frac{d^2 N_\mu}{dE d\Omega} = N(E) \left[\frac{Q_{\pi\mu}}{1 + \frac{B_{\pi\mu} E \cos\theta}{E_\pi}} + \frac{0.635 Q_{K\mu}}{1 + \frac{B_{K\mu} E \cos\theta}{E_K}} \right]$$

$$\frac{d^2 N_\nu}{dE d\Omega} = N(E) \left[\frac{Q_{\pi\nu}}{1 + \frac{B_{\pi\nu} E \cos\theta}{E_\pi}} + \frac{0.635 Q_{K\nu}}{1 + \frac{B_{K\nu} E \cos\theta}{E_K}} \right]$$

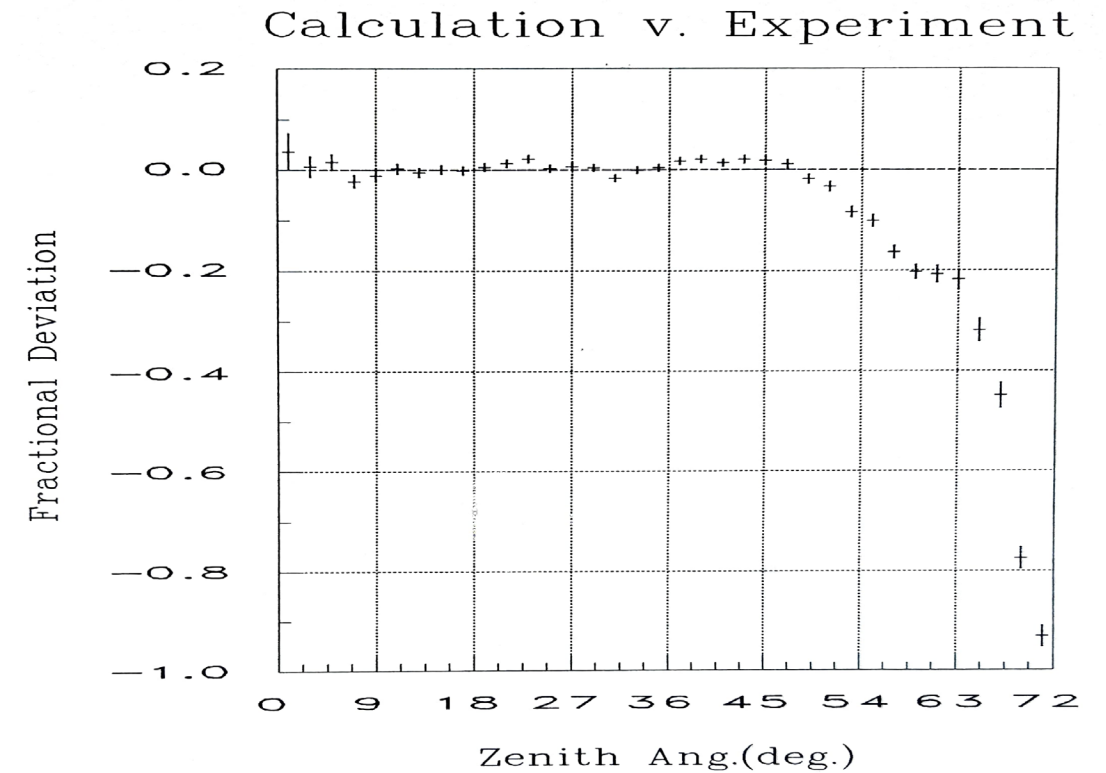
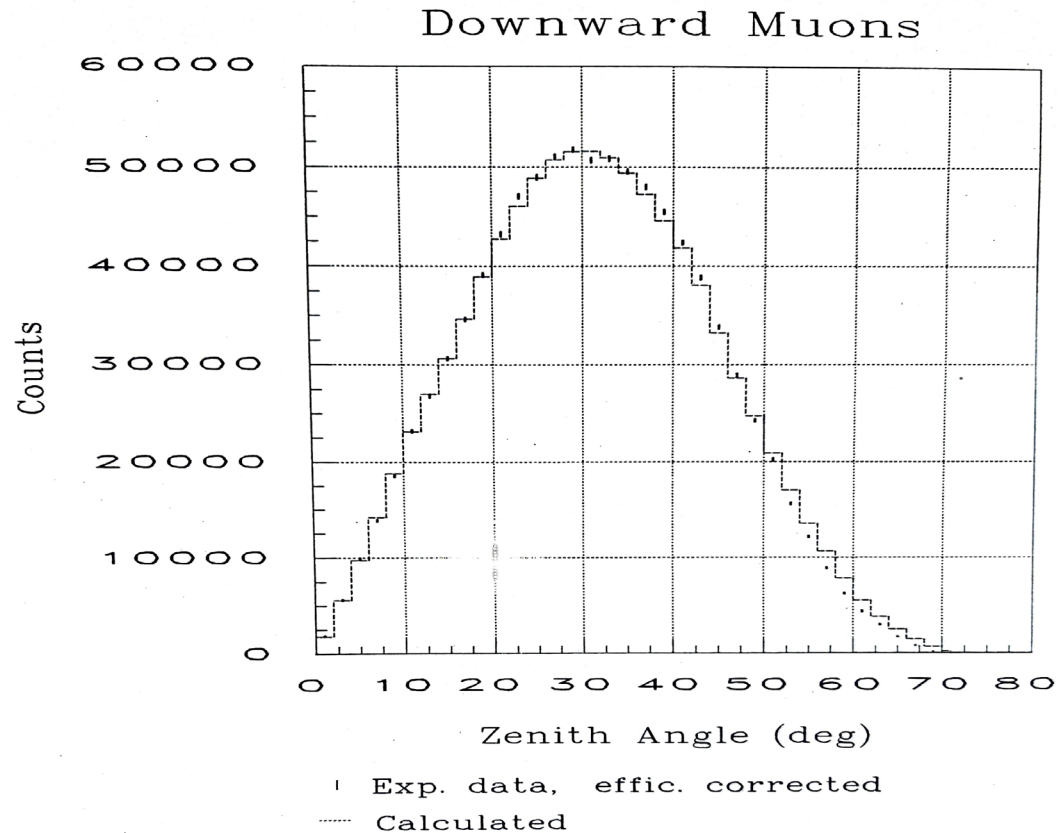
$$N(E) = 1.8 E^{-2.7} \left. \vphantom{N(E)} \right\} \text{Primary flux}$$

NB: This simple forms are valid only for $\theta \leq 60^\circ$

Gran Sasso Rock Map



Down μ : Experiment v Theory





Monopole, Astrophysics, and Cosmic Ray Observatory

Upward μ

- Compute differential cross section using suitable set of structure functions
- Do MC simulation to determine flux of upward μ that reach MACRO from ν 's interacting with earth
- Compute event rate with detector MC

Everything is in place except for $\frac{d\sigma}{dE_\mu}$. Start w/ doubly differential cross section:

$$\frac{d^2\sigma}{dx dy} = \frac{G_F^2 M_p E_\nu}{\pi} \frac{M_W^4}{[2ME_\nu xy + M_W^2]^2}$$

$$\times \left\{ \left(1-y + \frac{M_{xy}}{2E_\nu}\right) F_2^{\nu\ell}(x, Q^2) + y^2 2x F_3^{\nu\ell}(x, Q^2) \pm \left(y - \frac{y^2}{2}\right) x F_3^{\nu\ell}(x, Q^2) \right\}$$

Structure function definitions: D.W. Duke & J.F. Owens, Phys. Rev. D30, 49(1984).

$Q_0 := 2$ Q^2 dependence threshold, GeV

$\Lambda := 2$ Low Energy cutoff, GeV

$$C_1 := \ln\left(\frac{Q_0^2}{\Lambda^2}\right)$$

$$z(x, E_\nu, E_\mu) := Q^2 \left[2Mx(E_\nu - E_\mu) - Q_0^2 \right] \ln \left[\frac{2M(E_\nu - E_\mu)x}{\Lambda^2} \right] \frac{1}{C_1}$$

Structure fn. parameter

$$\eta_1(z) := 0.419 + 0.004z - 0.007z^2 \quad \eta_2(z) := 0.763 - 0.237z + 0.026z^2$$

$$\eta_3(z) := 3.46 + 0.724z - 0.066z^2 \quad \eta_4(z) := 4.00 + 0.627z - 0.019z^2$$

$$\gamma_{ud}(z) := 4.40 - 4.86z + 1.33z^2 \quad \gamma_{d\ell}(z) := 0.421z + 0.033z^2$$

$$N_{ud}(z) := 3 \left[\frac{\Gamma(\eta_1(z)) \cdot \Gamma(\eta_2(z) + 1)}{\Gamma(\eta_1(z) + \eta_2(z) + 1)} \left(1 + \frac{\gamma_{ud}(z) \eta_1(z)}{\eta_1(z) + \eta_2(z) + 1} \right) \right]^{-1}$$

$$N_{d\ell}(z) := \left[\frac{\Gamma(\eta_3(z)) \cdot \Gamma(\eta_4(z) + 1)}{\Gamma(\eta_3(z) + \eta_4(z) + 1)} \left(1 + \frac{\gamma_{d\ell}(z) \eta_3(z)}{\eta_3(z) + \eta_4(z) + 1} \right) \right]^{-1}$$

$$V(x, z) := N_{ud}(z) x^{\eta_1(z)} (1-x)^{\eta_2(z)} (1+\gamma_{ud}(z)x)$$

Valence quark structure fn

$$d(x, z) := N_{d\ell}(z) x^{\eta_3(z)} (1-x)^{\eta_4(z)} (1+\gamma_{d\ell}(z)x)$$

Down quark str. fn. for proton

$$\Lambda_g(z) := 1.265 - 1.132z + 0.293z^2 \quad \Lambda_d(z) := 0.135z - 0.075z^2$$

$$\Lambda_G(z) := 1.56 - 1.71z + 0.638z^2$$

$$a_g(z) := -0.372z - 0.029z^2$$

$$a_d(z) := -0.036 - 0.222z - 0.058z^2$$

$$a_G(z) := -0.949z + 0.325z^2$$

$$b_g(z) := 8.05 + 1.59z - 0.153z^2$$

$$b_d(z) := 6.35 + 3.26z - 0.909z^2$$

$$b_G(z) := 6.0 + 1.44z - 1.05z^2$$

$$c_g(z) := 6.31z - 0.273z^2$$

$$c_d(z) := 3.03z + 1.50z^2$$

$$c_G(z) := 9.0 - 7.19z + 0.255z^2$$

$$f_g(z) := 10.5z - 3.17z^2$$

$$f_d(z) := 17.4z - 11.3z^2$$

$$f_G(z) := 16.5z + 10.9z^2$$

$$\gamma_g(z) := 14.7z + 9.80z^2$$

$$\gamma_d(z) := 17.9z + 15.6z^2$$

$$\gamma_G(z) := 15.3z - 10.1z^2$$

$$S(x, z) := \frac{1}{6} \left[\Lambda_g(z) x^{a_g(z)} (1-x)^{b_g(z)} (1 + a_g(z)x + b_g(z)x^2 + \gamma_g(z)x^3) \right]$$

Sea Quarks

$$c(x, z) := \left[\Lambda_d(z) x^{a_d(z)} (1-x)^{b_d(z)} (1 + a_d(z)x + b_d(z)x^2 + \gamma_d(z)x^3) \right]$$

Charmed Quarks

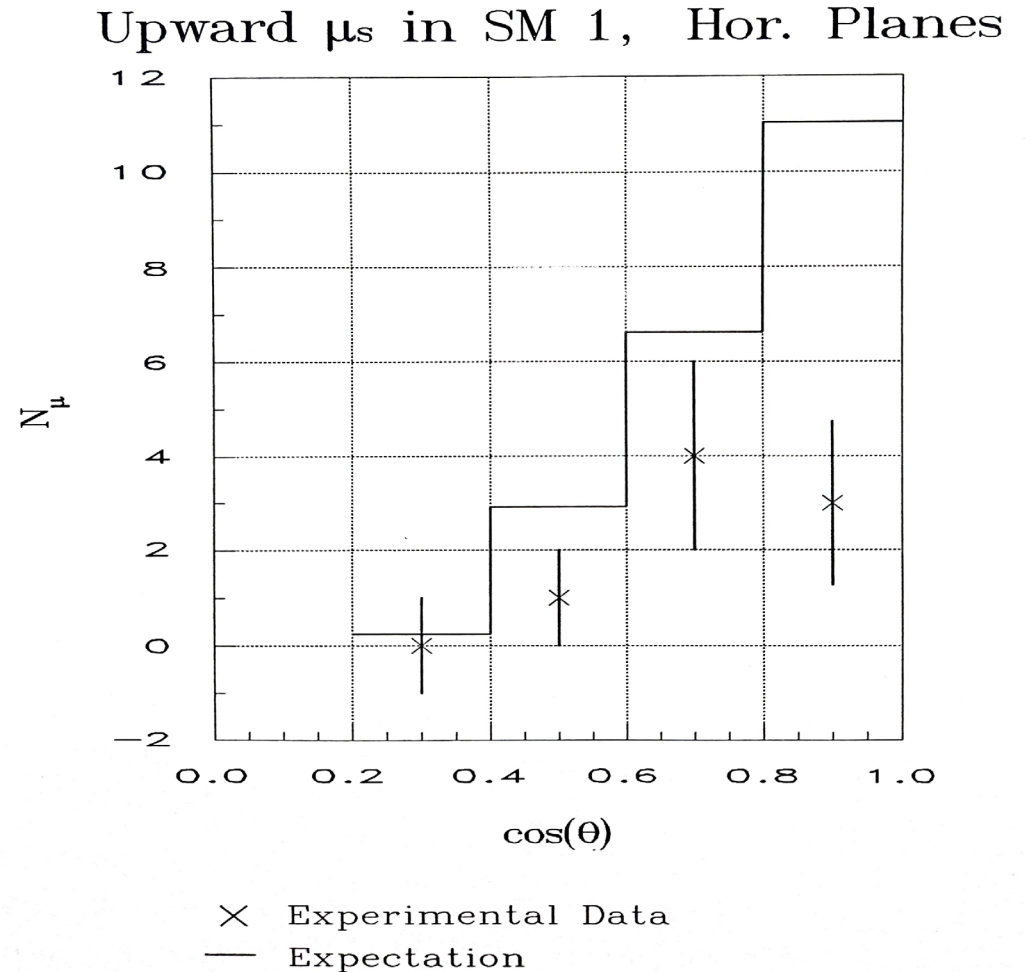
$$G(x, z) := \left[\Lambda_G(z) x^{a_G(z)} (1-x)^{b_G(z)} (1 + a_G(z)x + b_G(z)x^2 + \gamma_G(z)x^3) \right]$$

Gluons

$$N_{d\ell}(z) := \int_0^1 V(x, z) dx + 6 \int_0^1 S(x, z) dx + \int_0^1 G(x, z) dx + 2 \int_0^1 c(x, z) dx \quad x_0 := .25 \quad N_{d\ell}(x_0) = 1$$

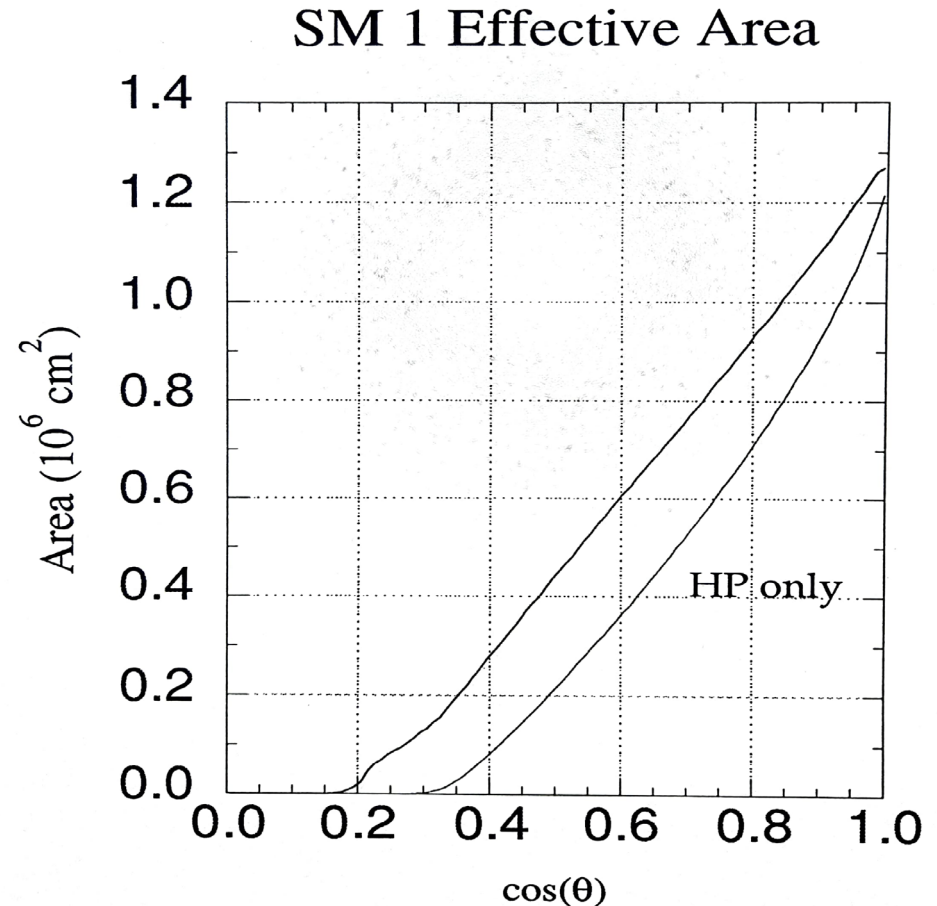
Observation v Expectation

- Reasonably good agreement with theory
 - EXCEPT for $0.8 < \cos(\theta) \leq 1.0$
- Possible explanations
 - Just unlucky
 - Detector efficiency less at small nadir angles
 - It's a **real** deficit



Eliminate the Impossible

- Just unlucky
 - Statistical argument is weak since MACRO has largest effective area at small zenith/nadir angles
 - Demonstrated by down muon observation
 - Effective area goes down by factor of two across histogram bins
- Detector efficiency
 - Looking at down muons, detector efficiency agrees well with calculated value at small zenith
 - Detector symmetry precludes up v down preference in efficiency



More Impossibilities

- Other “plausible” sources of the deficit were proffered, but all turned out to be wild goose chases (caccia ai fantasmi)
 - Upward muon spectrum too soft to traverse detector, i.e., below energy threshold required to pass through detector
 - Track reconstruction efficiency at low energy ($E < 3 \text{ GeV}$)
 - Error in cross section calculation
 - Does not explain anomaly in single nadir angle bin, i.e., all nadir angles should have been affected.
 - Model of the earth used in neutrino propagation MC contained errors

Whatever Remains, However Improbable...

- **The upward μ deficit is real!**
 - MACRO is symmetric with respect to up v down
 - No preferred direction
 - Largest effective area at small zenith/nadir angles
 - Area at 45° is 58% of that at 0°
 - There should be many more events near 0° than 45°
 - This is the simplest explanation as required by Occam's Razor (although not guaranteed)

...Must Be the Truth!