

laser frequency & intensity noise requirement in ET

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Beam path in the interferometer

Laser noise in signal path:

1. Bright port → PRC field , after circulating in coupled PRC-Arm.
2. PRC field → SRC field, due to contrast defect, Schnupp asymmetry.
3. SRC field → dark port, after circulating in coupled SRC-Arm.
4. dark port → readout.

Laser noise in Local oscillator path:

Balanced homodyne readout:

1. Bright port → PRC field , after circulating in coupled PRC-Arm.
2. PRC field → readout.

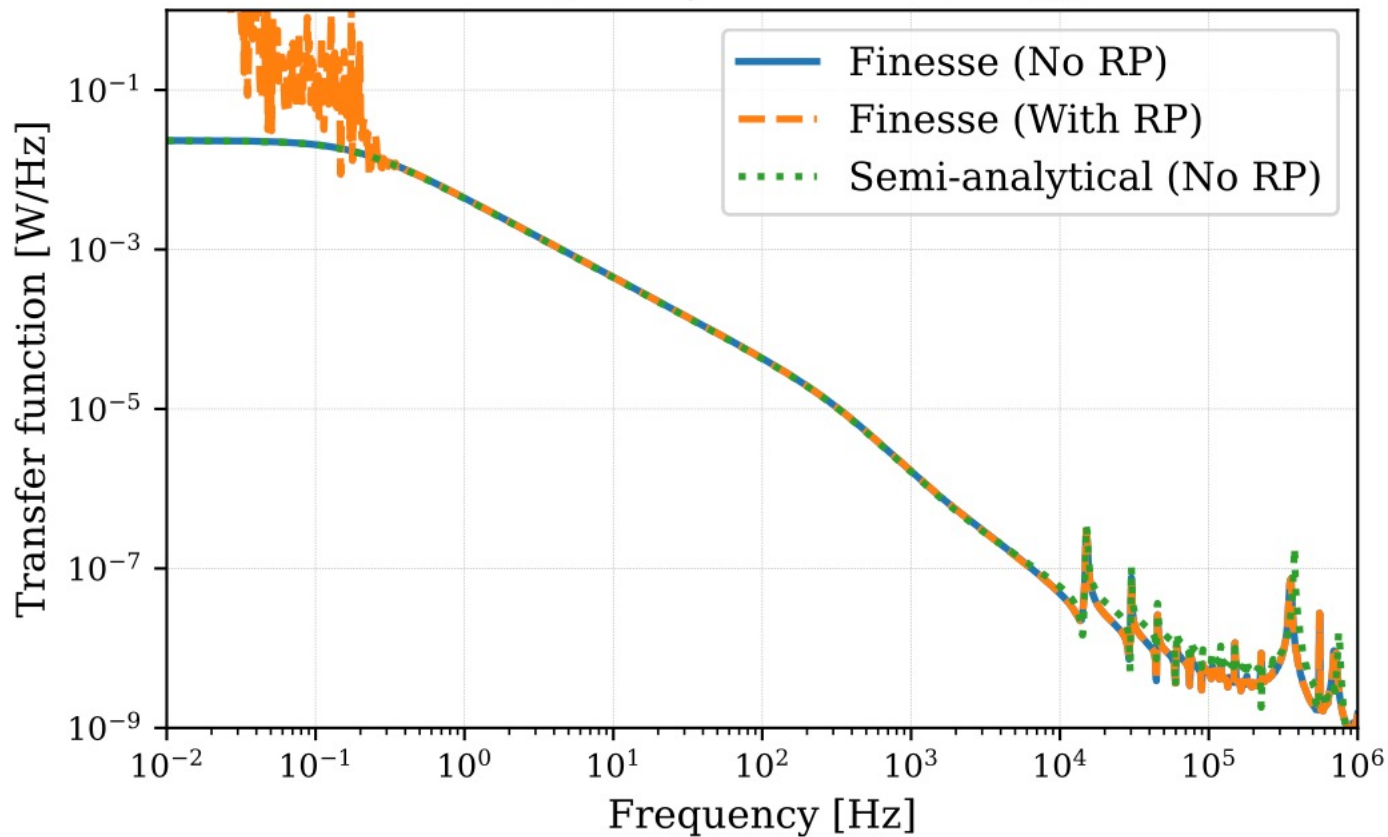
Parameters

Finesse	
ITM T	0.007
ITM T asymmetry	1%
Mirror loss	37.5e-6
Loss asymmetry	10%
LO power	100mW
Dark port power	1.41mW

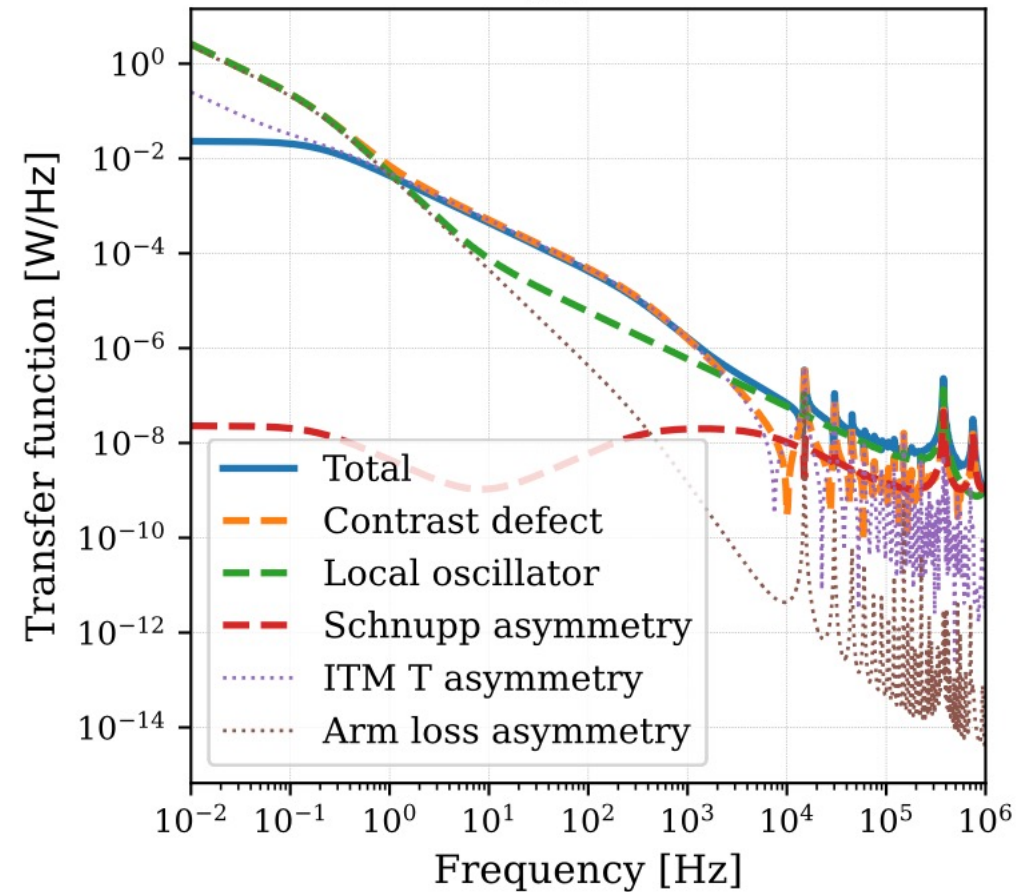
Semi-analytical	
ITM T	0.007
ITM T asymmetry	1%
Mirror loss	37.5e-6
Loss asymmetry	10%
LO power	100mW
Dark port power	1.56mW

Frequency noise TF in W/Hz

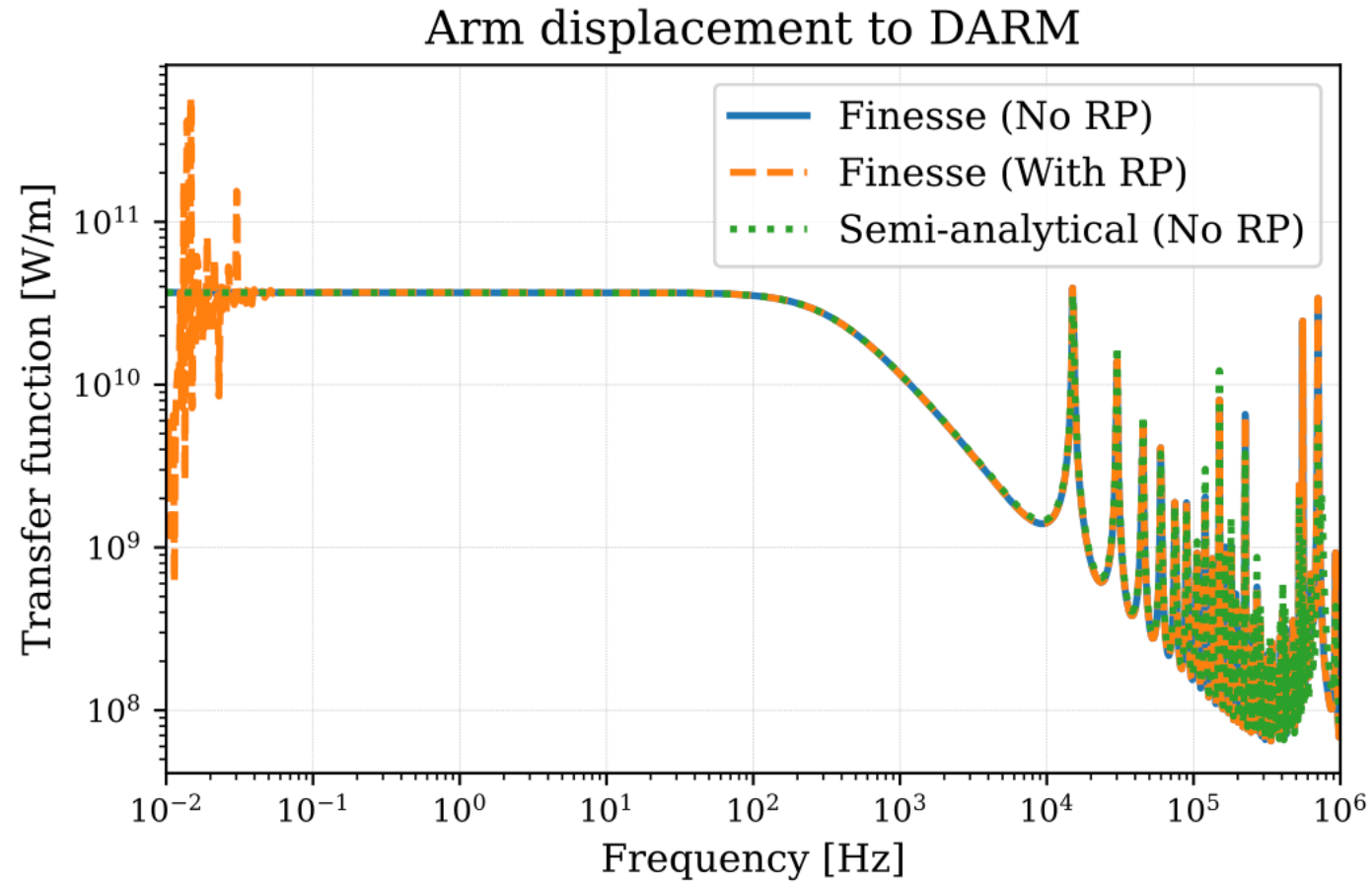
Laser frequency to DARM, $\delta T_{itm} = 1\%$



$\delta T_{itm} = 1\%$

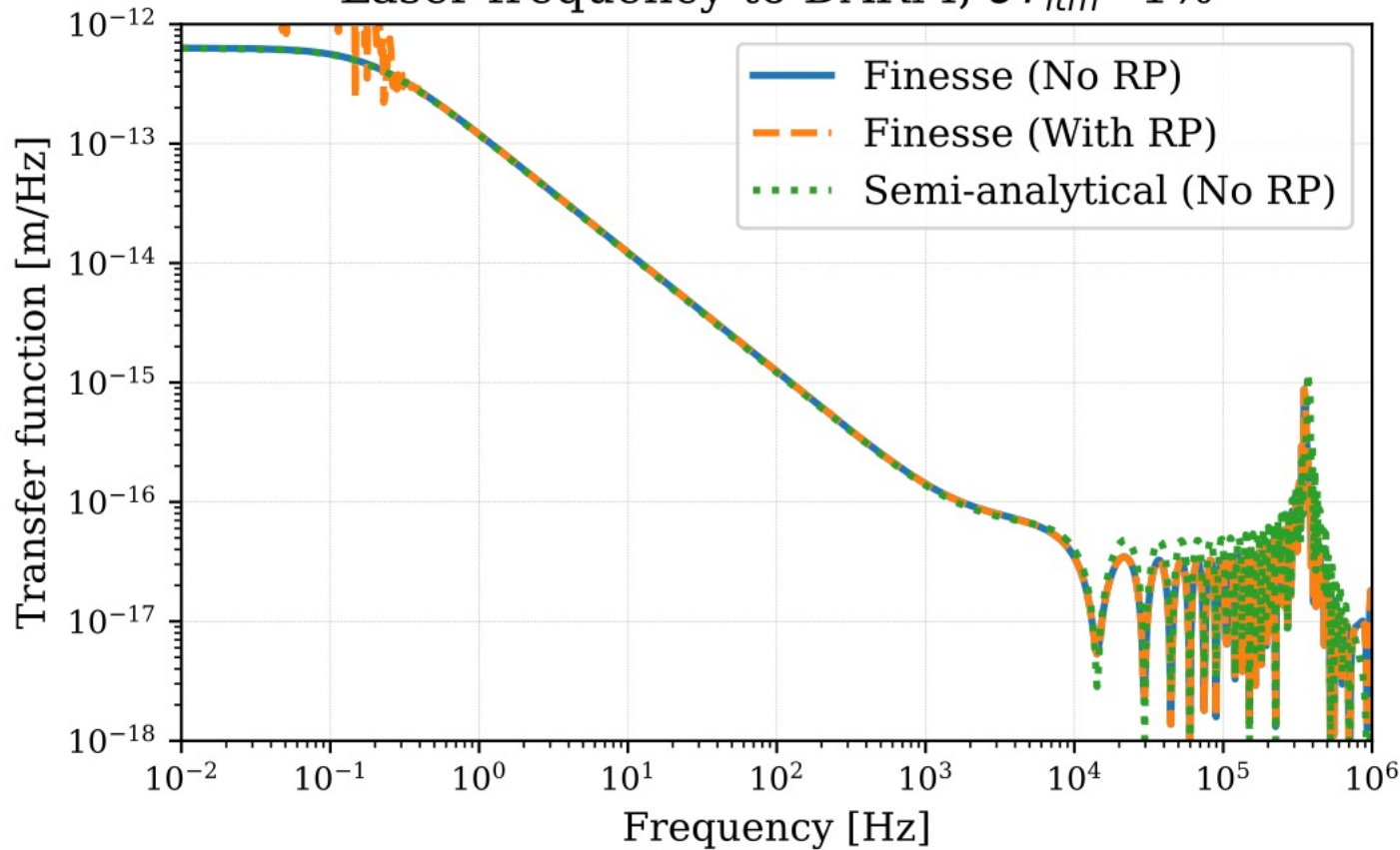


Darm motion TF in W/m

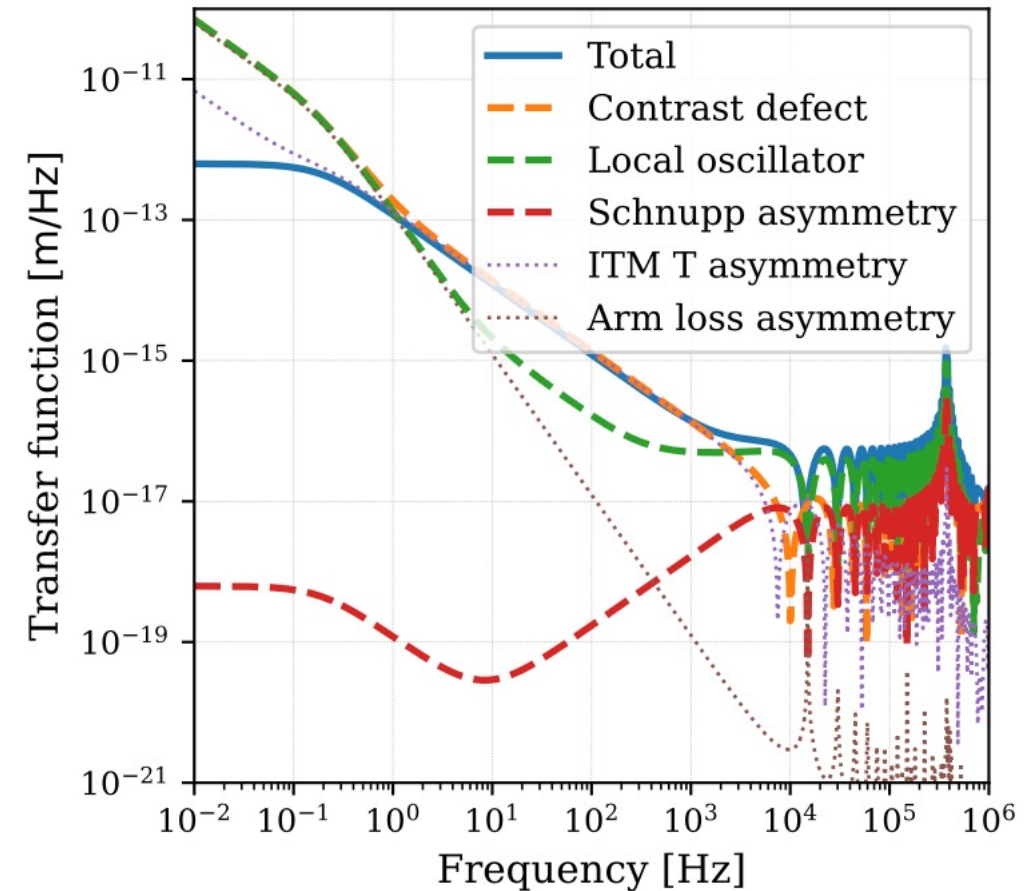


Calibrate frequency noise to equivalent darm motion in $\frac{m}{\text{Hz}} = \frac{W}{\text{Hz}} / \frac{W}{m}$

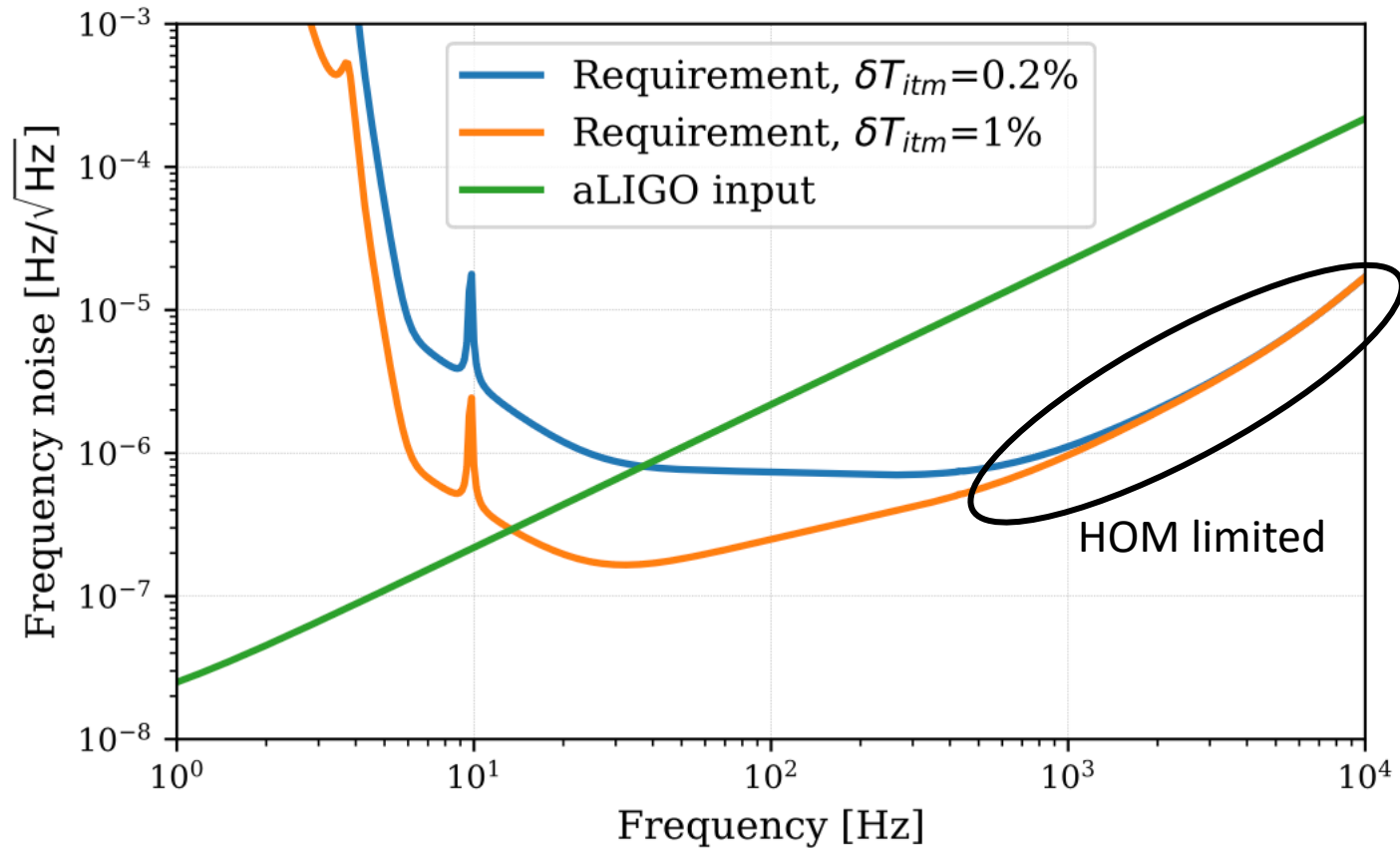
Laser frequency to DARM, $\delta T_{itm} = 1\%$



$\delta T_{itm} = 1\%$



Requirement (result from finesse including RP)



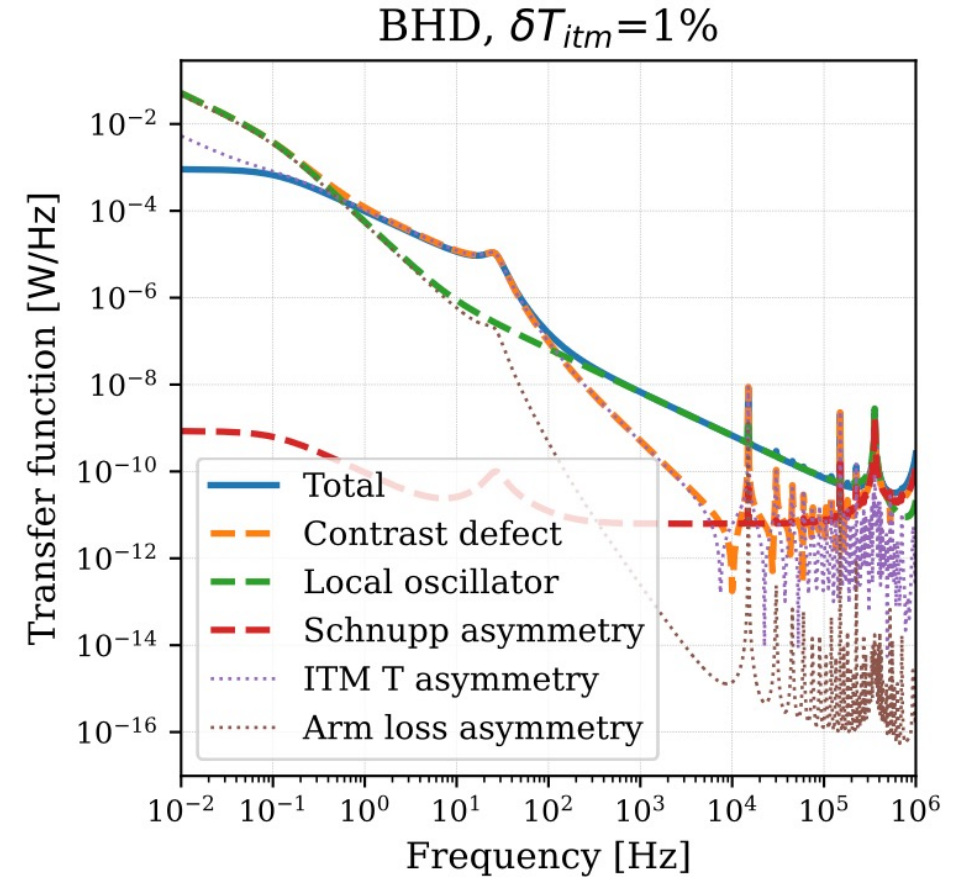
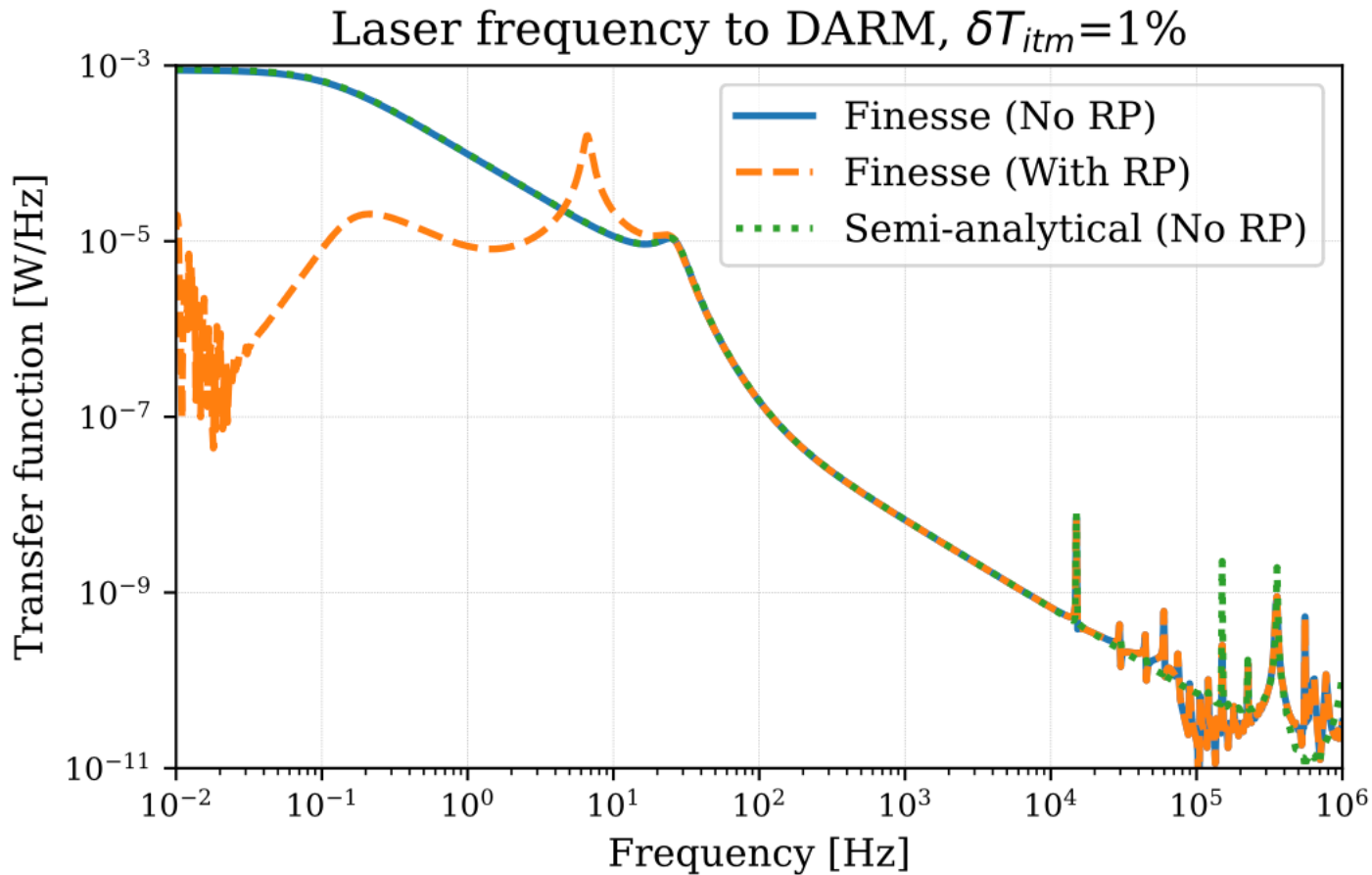
Including a factor of 10 safe margin and 0.4e-15m/Hz constant noise from HOM.

Parameters

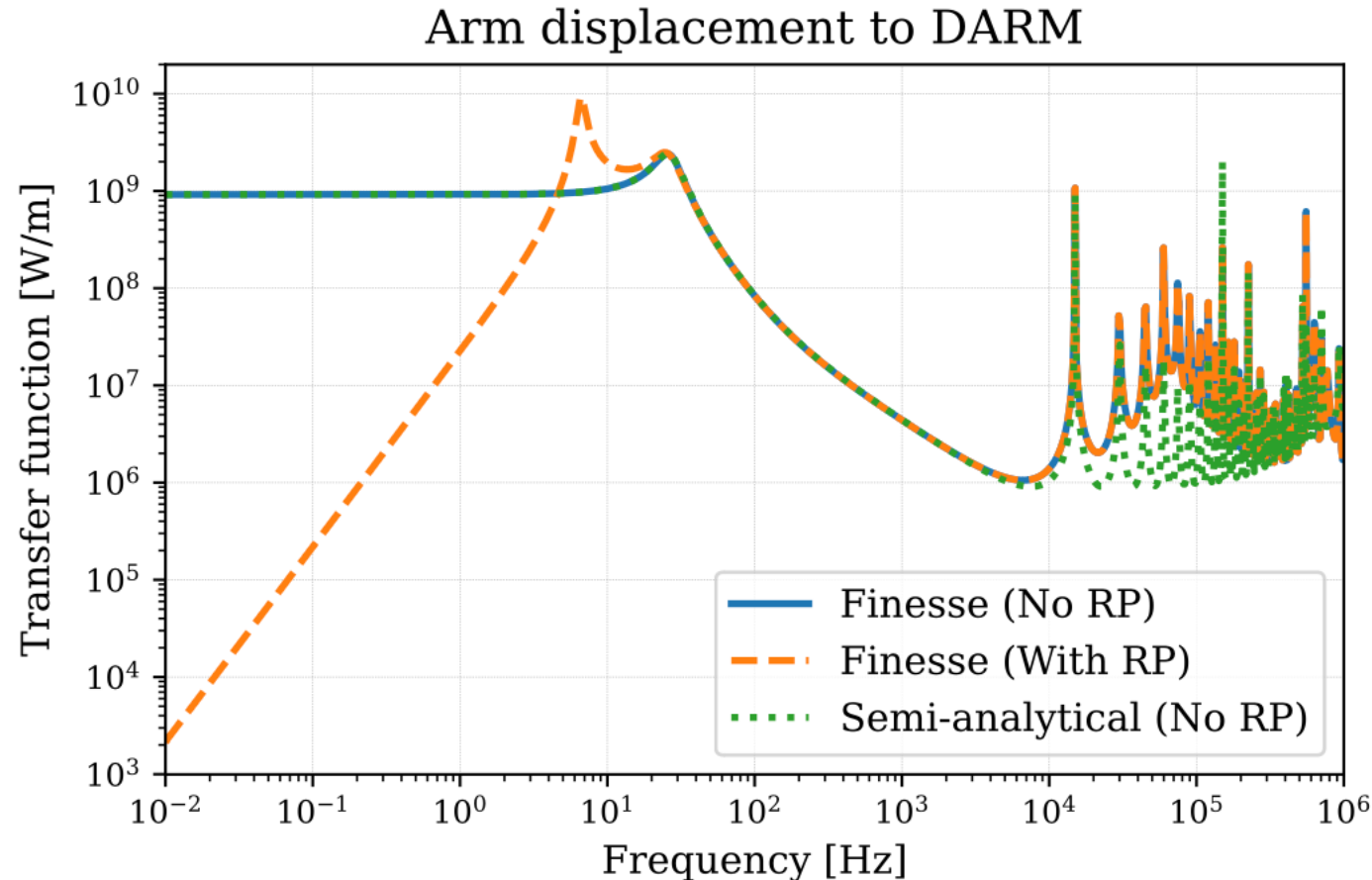
Finesse	
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Frequency noise TF in W/Hz



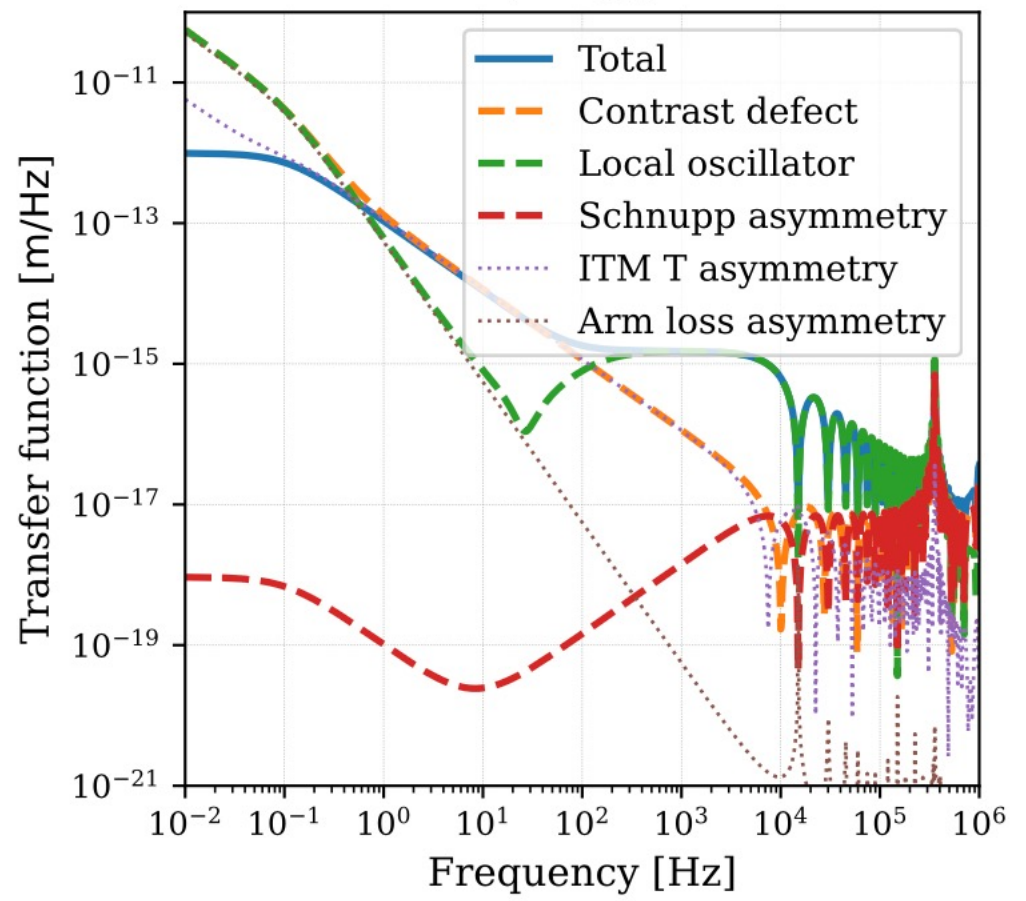
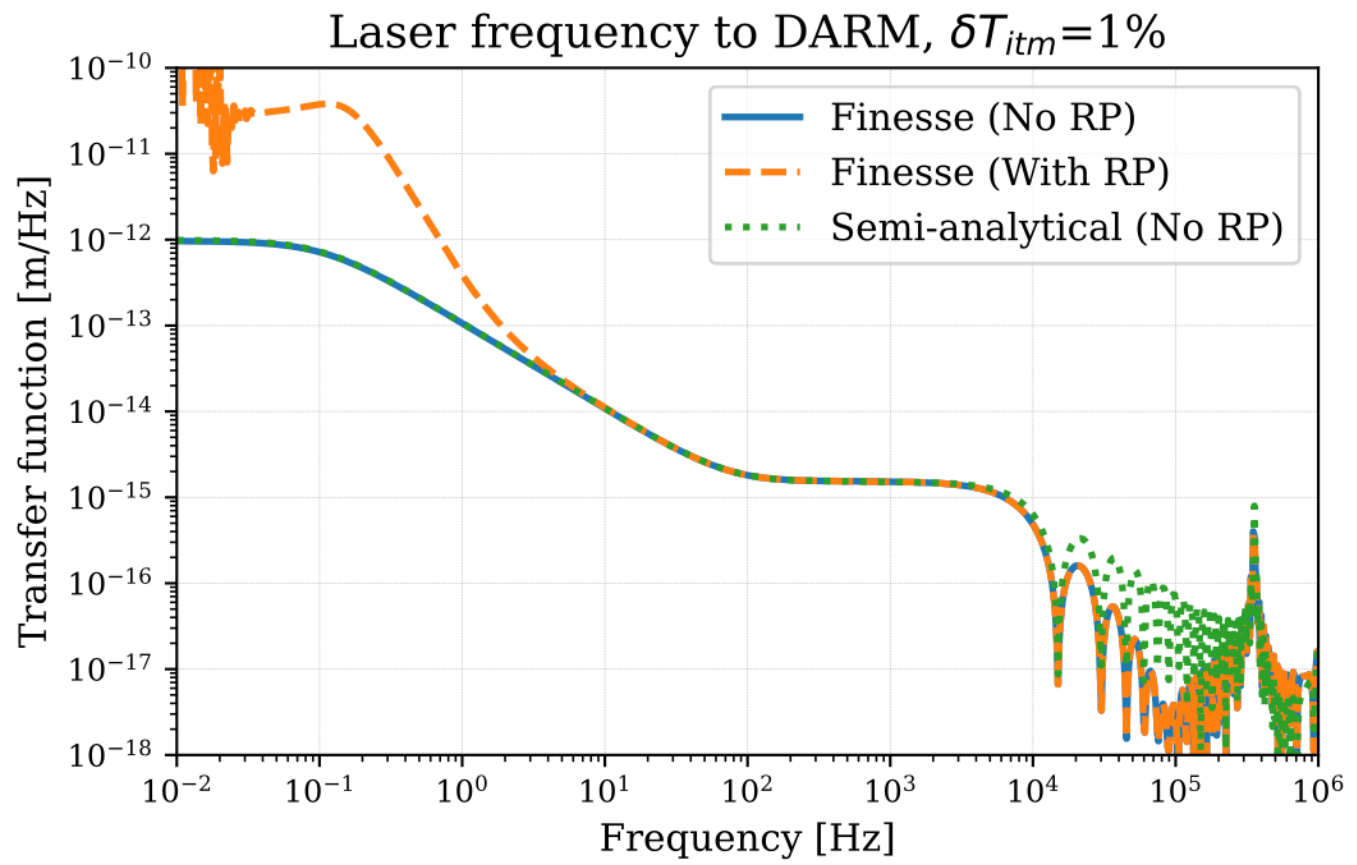
Darm motion TF in W/m



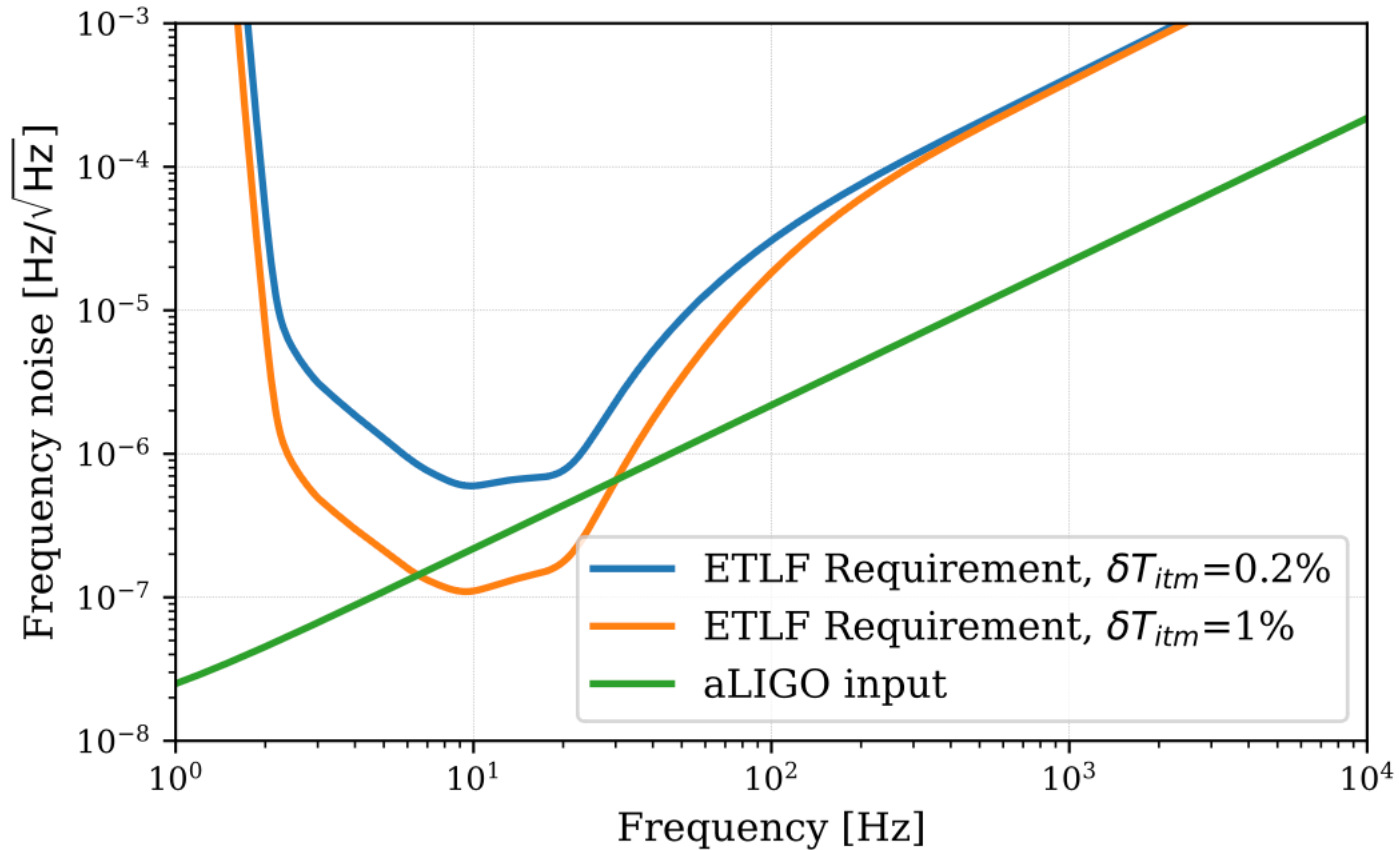
In Finesse, note that to measure the phase quadrature with SRM 34.2° detuned, the LO phase is set to -17.2° .

Calibrate frequency noise to equivalent darm motion in $\frac{m}{Hz} = \frac{W}{Hz} / \frac{W}{m}$

BHD, $\delta T_{itm} = 1\%$



Requirement (result from finesse including RP)



Including a factor of 10 safe margin and 0.4×10^{-15} m/Hz constant noise from HOM.

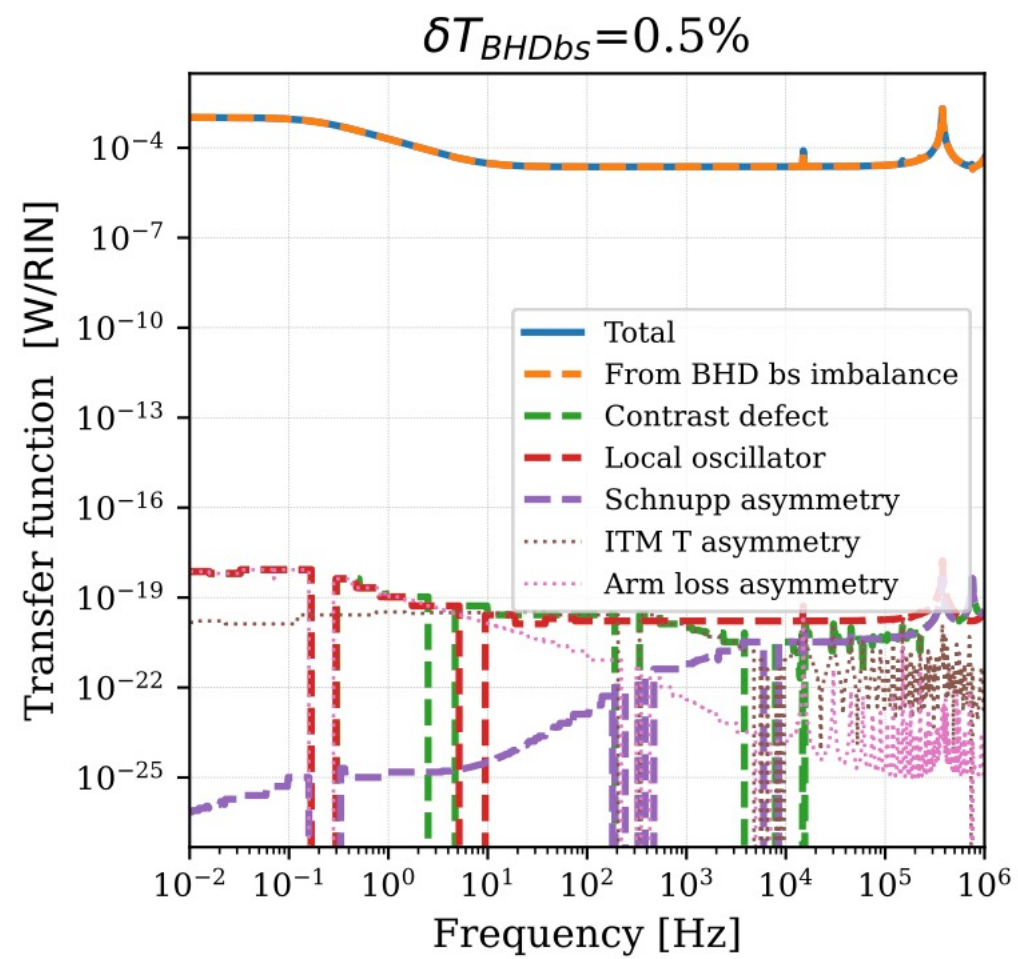
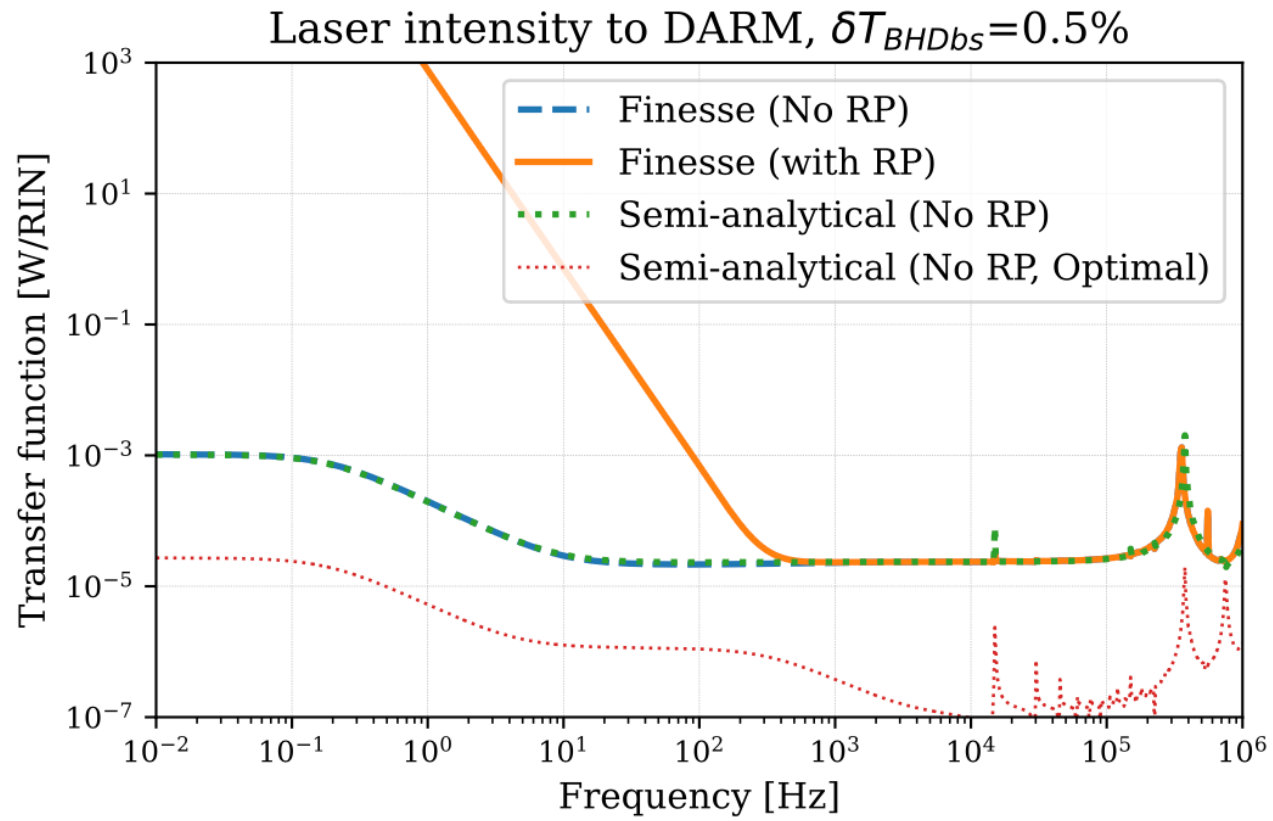
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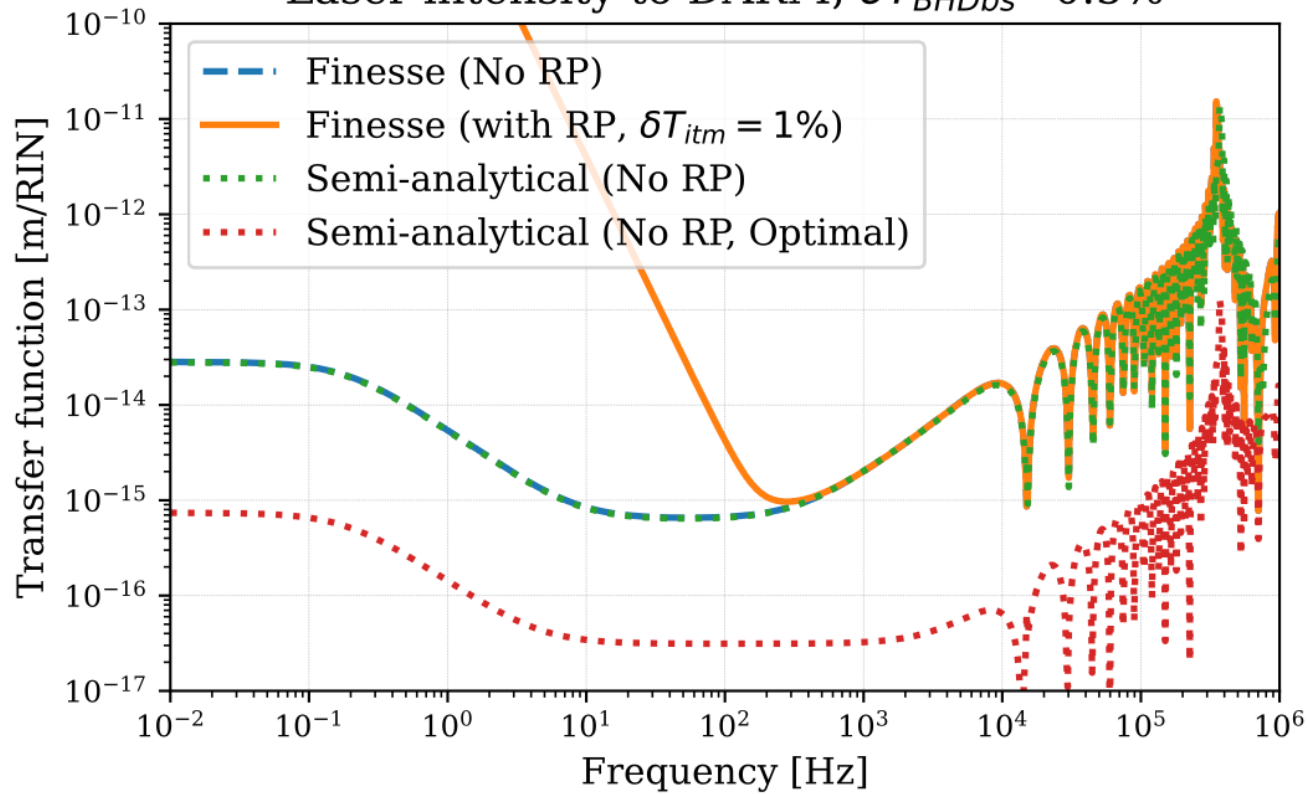
The LO static field (on phase quadrature) is orthogonal to the amplitude noise in signal beam. Here an imperfection on BHD beamsplitter can introduce the noise from LO itself. : LO static field x the local noise.

Intensity noise TF in W/RIN

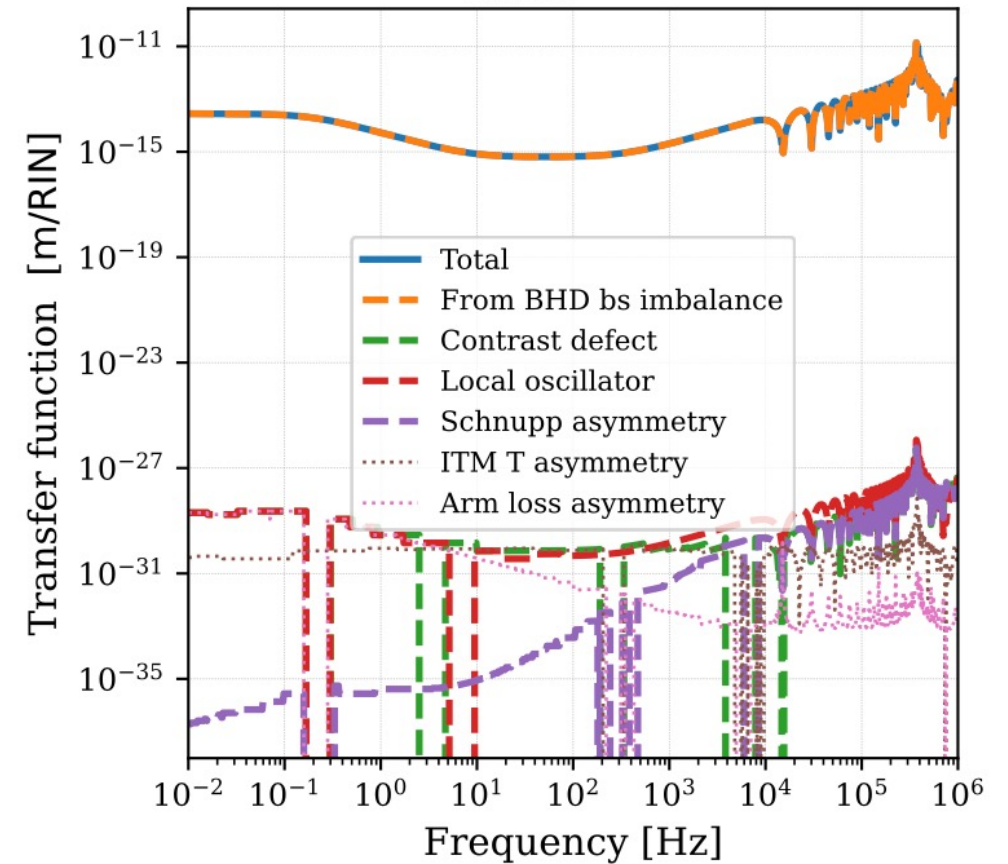


Calibrate intensity noise to equivalent darm motion in $\frac{m}{RIN} = \frac{W}{RIN} / \frac{W}{m}$

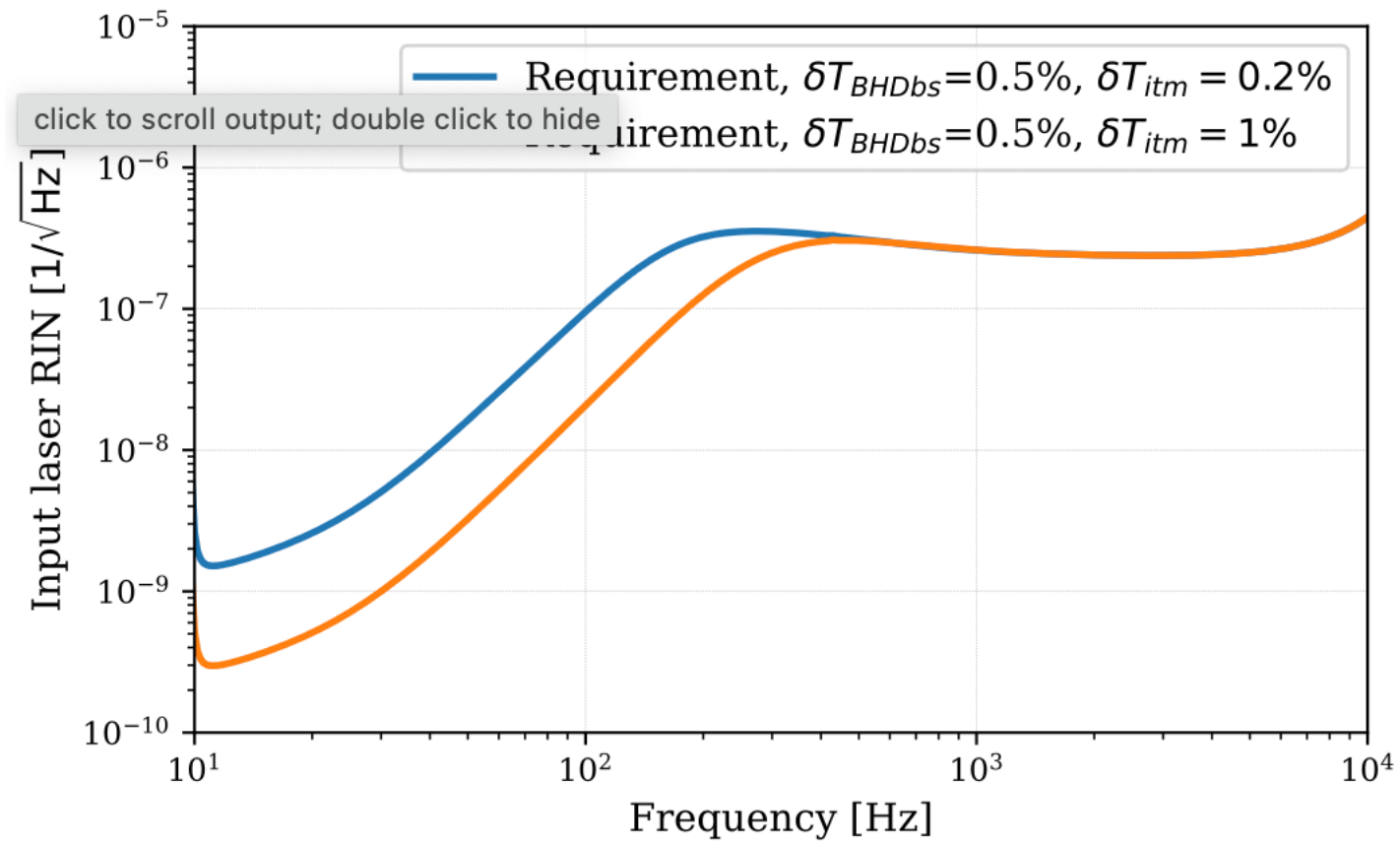
Laser intensity to DARM, $\delta T_{BHDbs}=0.5\%$



$\delta T_{BHDbs}=0.5\%$



Requirement



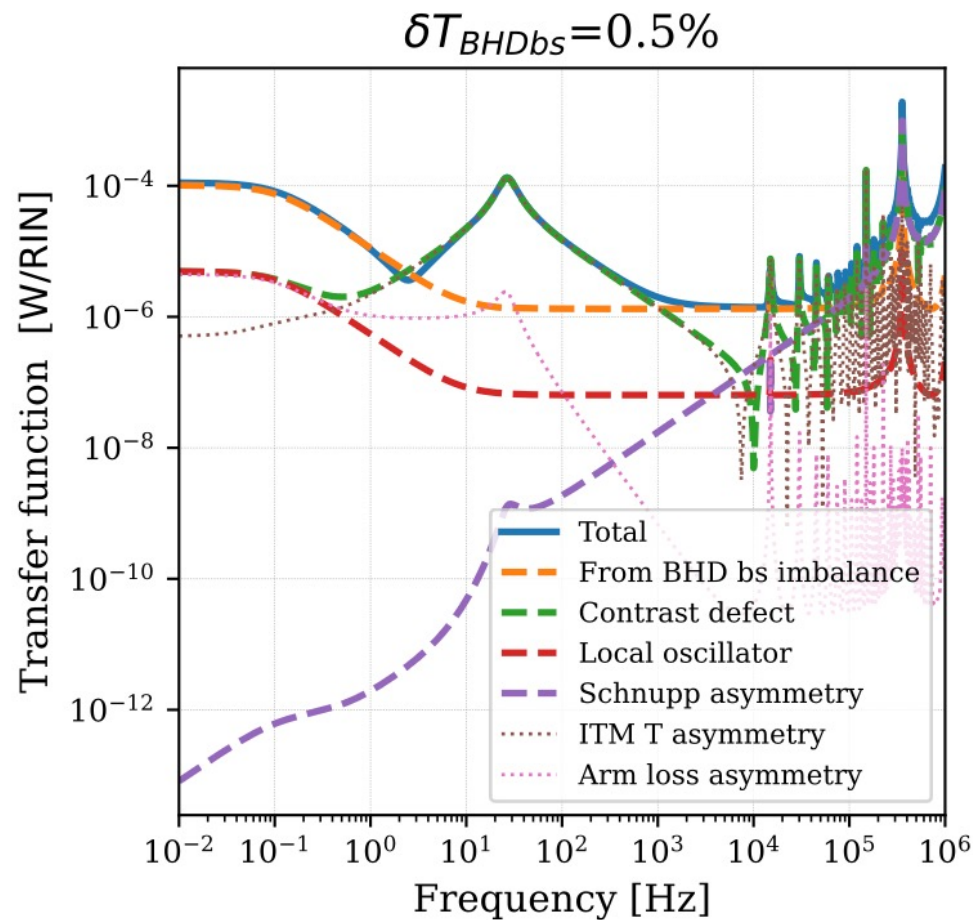
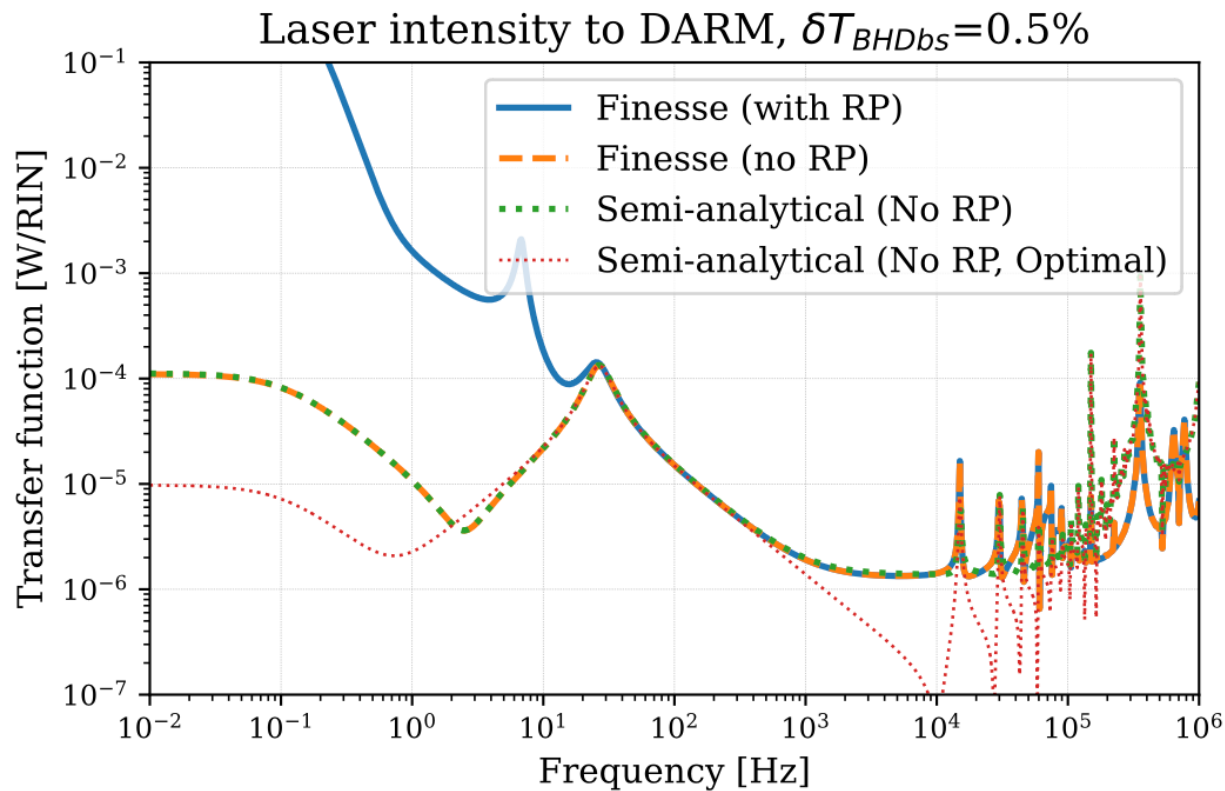
Including a factor of 10 safe margin.

Parameters

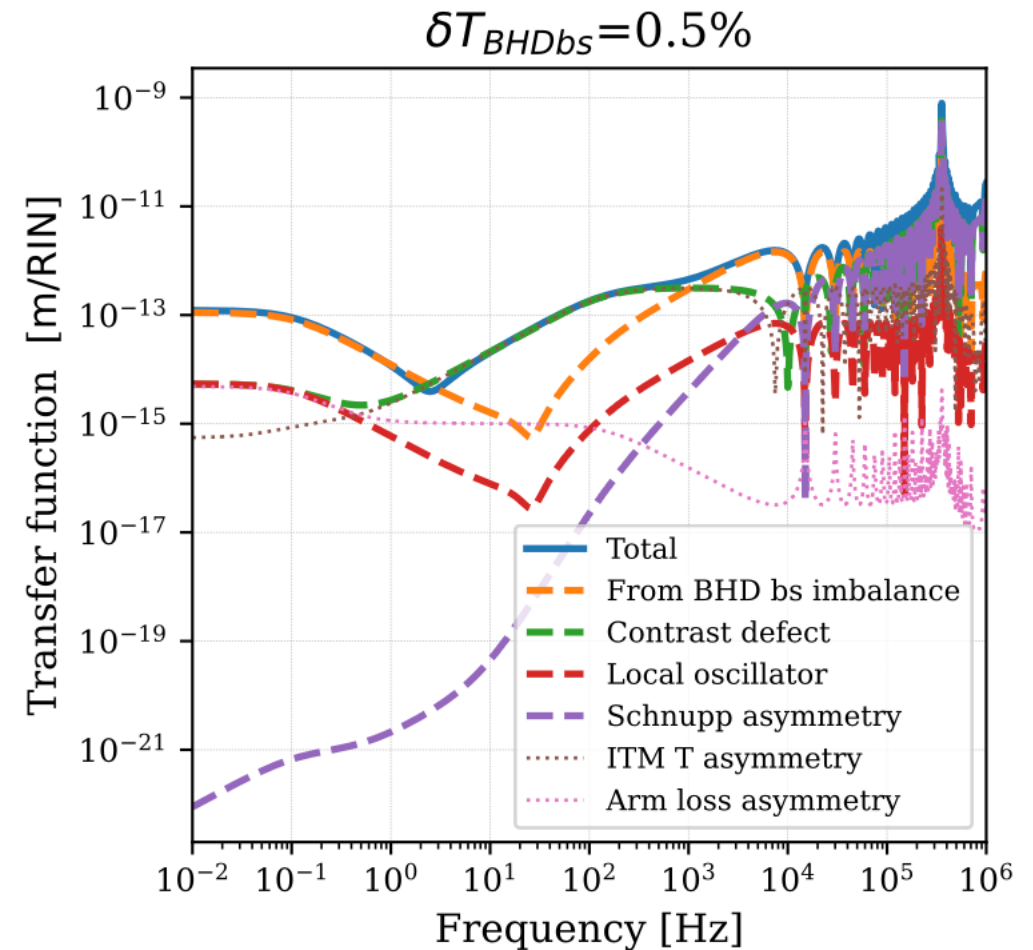
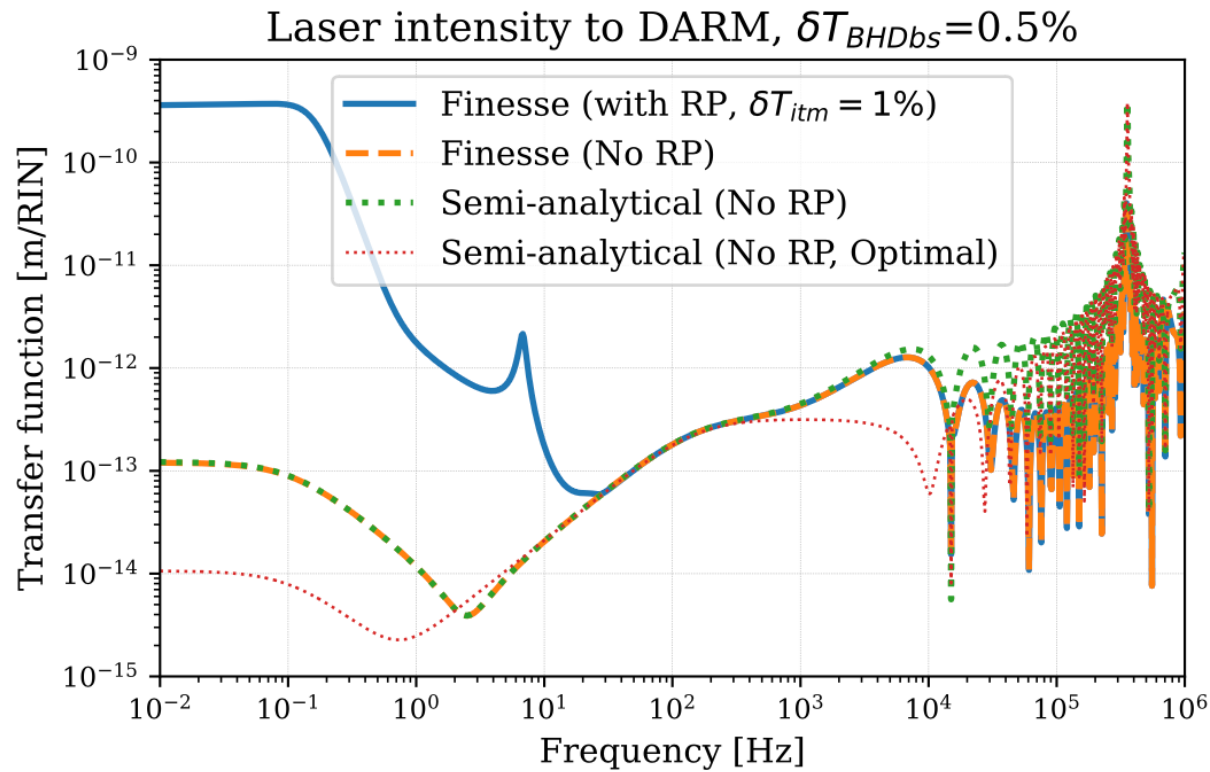
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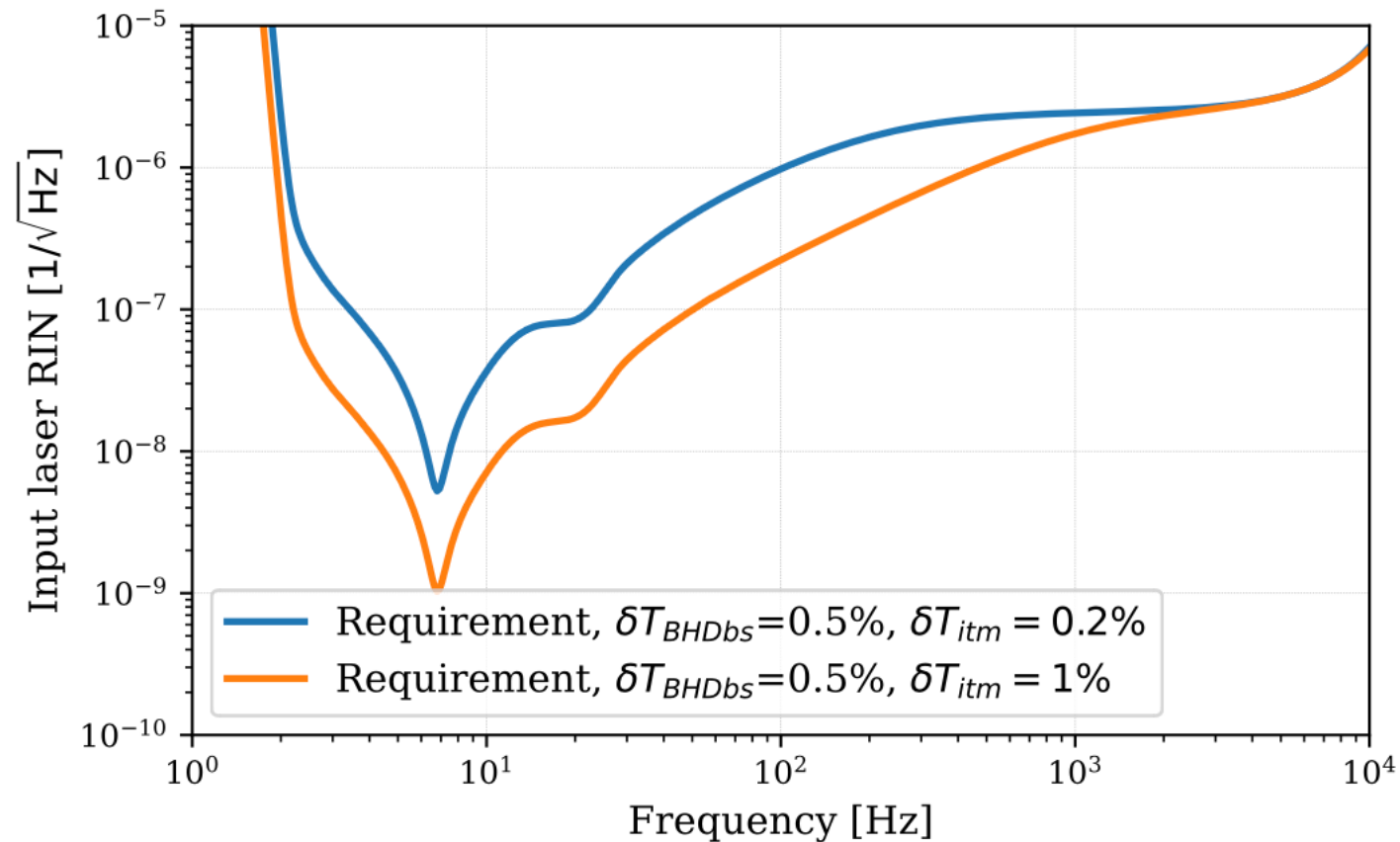
Intensity noise TF in W/Hz



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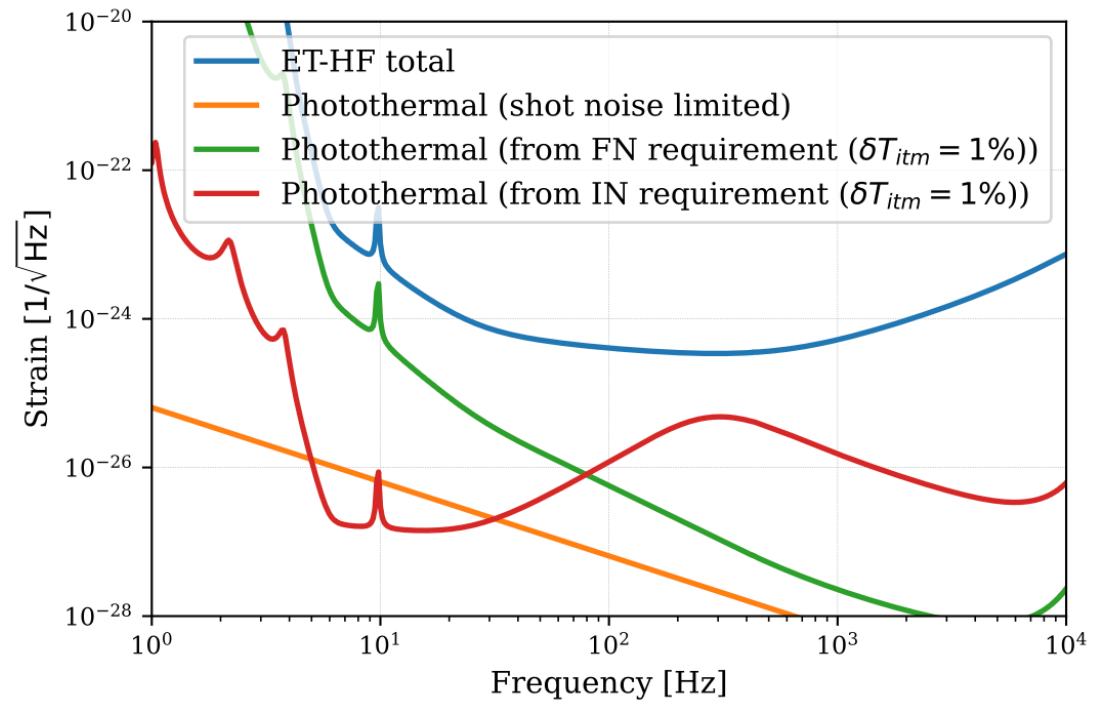


Intensity noise requirement

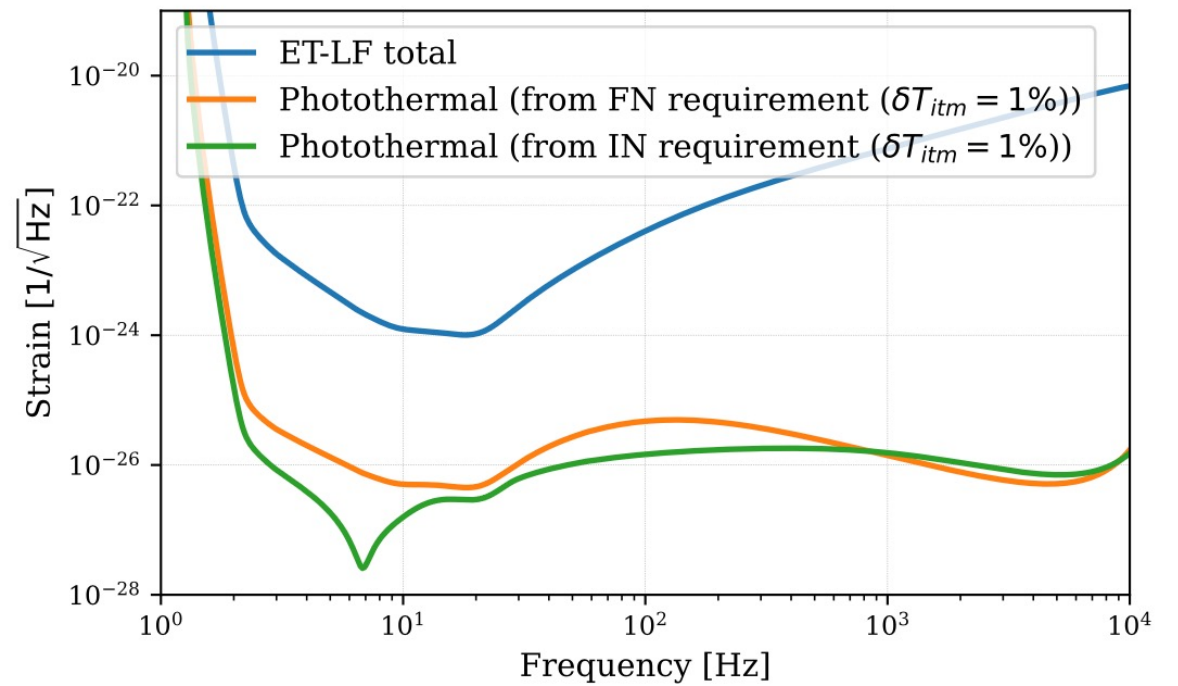


Including a factor of 10 safe margin.

Photothermal noise



Assuming 0.1 ppm absorption



Assuming 5 ppm absorption

Activities at AEI and Artemis

(Debanjan Adhikari , Benno Willke, Marina Trad Nery)

Almost done:

- Analytical calculation (in parallel with Teng) of the laser noise requirements at IFO input
Cross check our calculation with Teng's one
- Goals: bring intuitive understanding of the TFs and of the requirements to the laser groups
make a well documented Python code available for the community (good for newcomers)

Future plans:

- Include beam jitter requirements at IFO input
- Laser requirements at PSL-IO interface (in synergy with input optics team) using Finesse