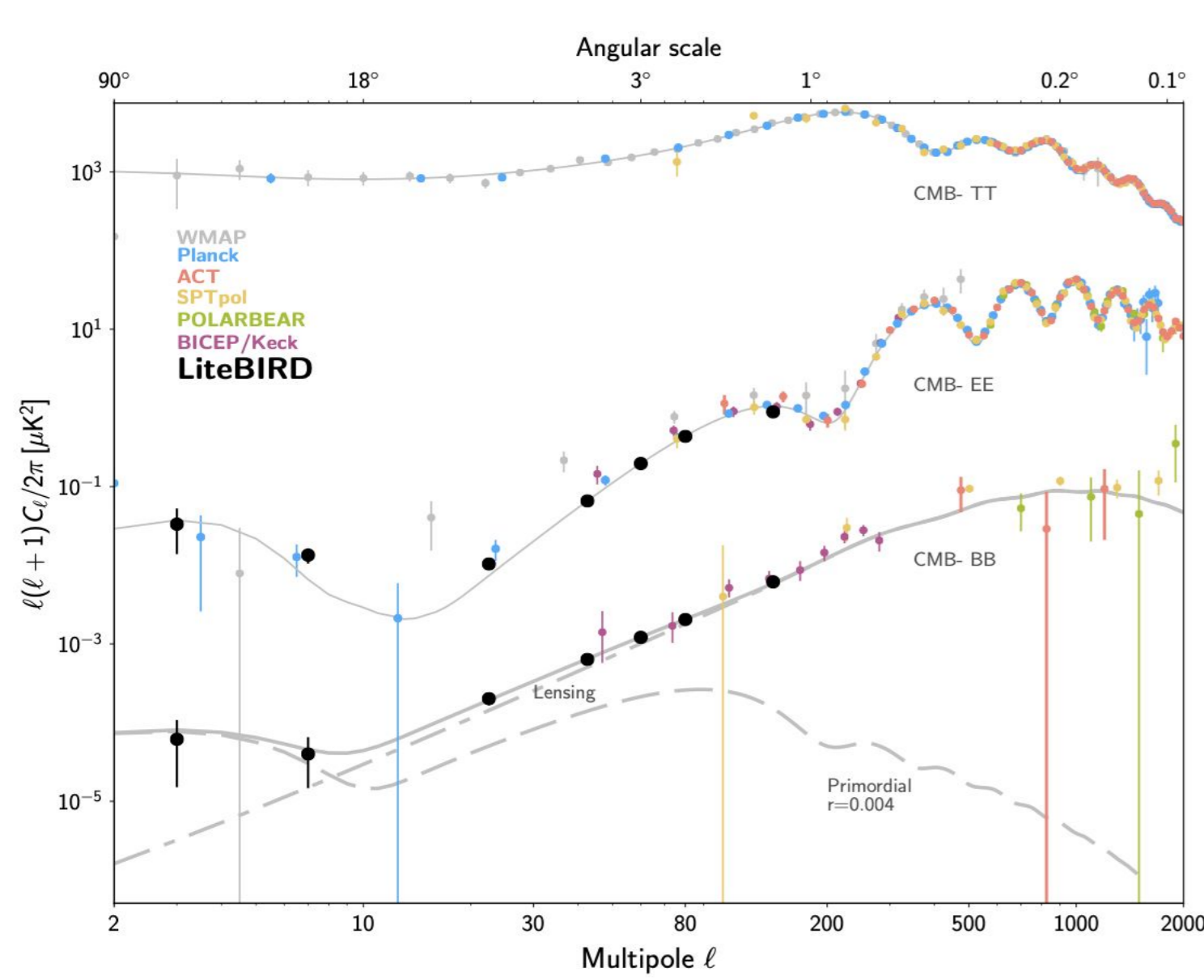


LiteBIRD is a space-borne experiment proposed by the Japanese Aerospace Exploration Agency (JAXA) and gathering a worldwide collaboration. This satellite aim at detecting CMB B-modes polarization which could contain an imprint of primordial gravitational waves generated by cosmic inflation. LiteBIRD will target the tensor-to-scalar ratio  $r$ , associated to the amplitude of primordial tensor perturbations (compared to scalar fluctuations) with a total uncertainty of  $\sigma_r < 10^{-3}$  by measuring both reionization and recombination bumps of the CMB B-modes power spectrum.

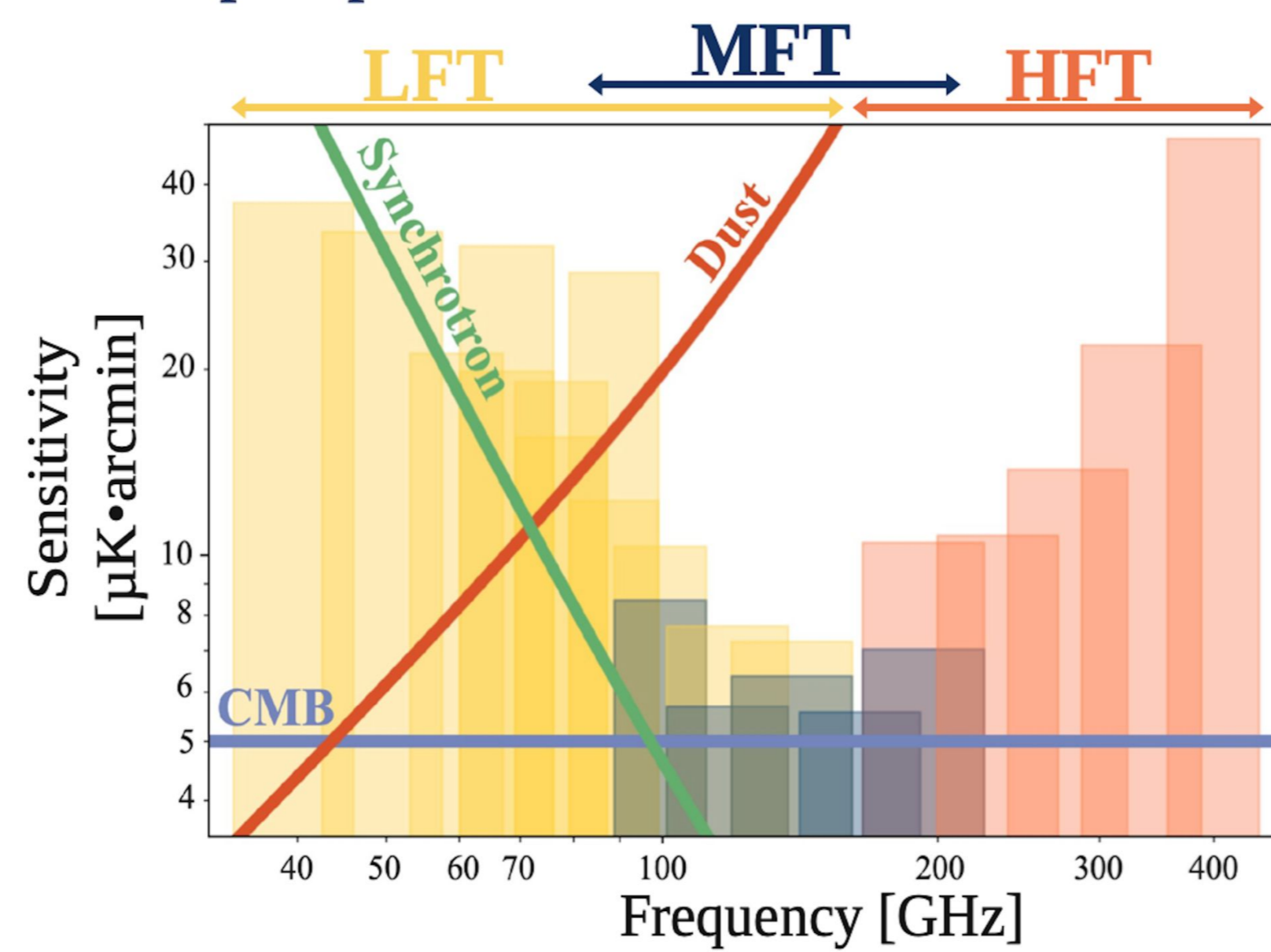
## Scientific target of LiteBIRD



### TT, EE and BB CMB angular power spectra

LiteBIRD will cover a multipole range up to  $\ell = 200$  to measure the reionization and recombination bumps of the CMB BB power spectrum.

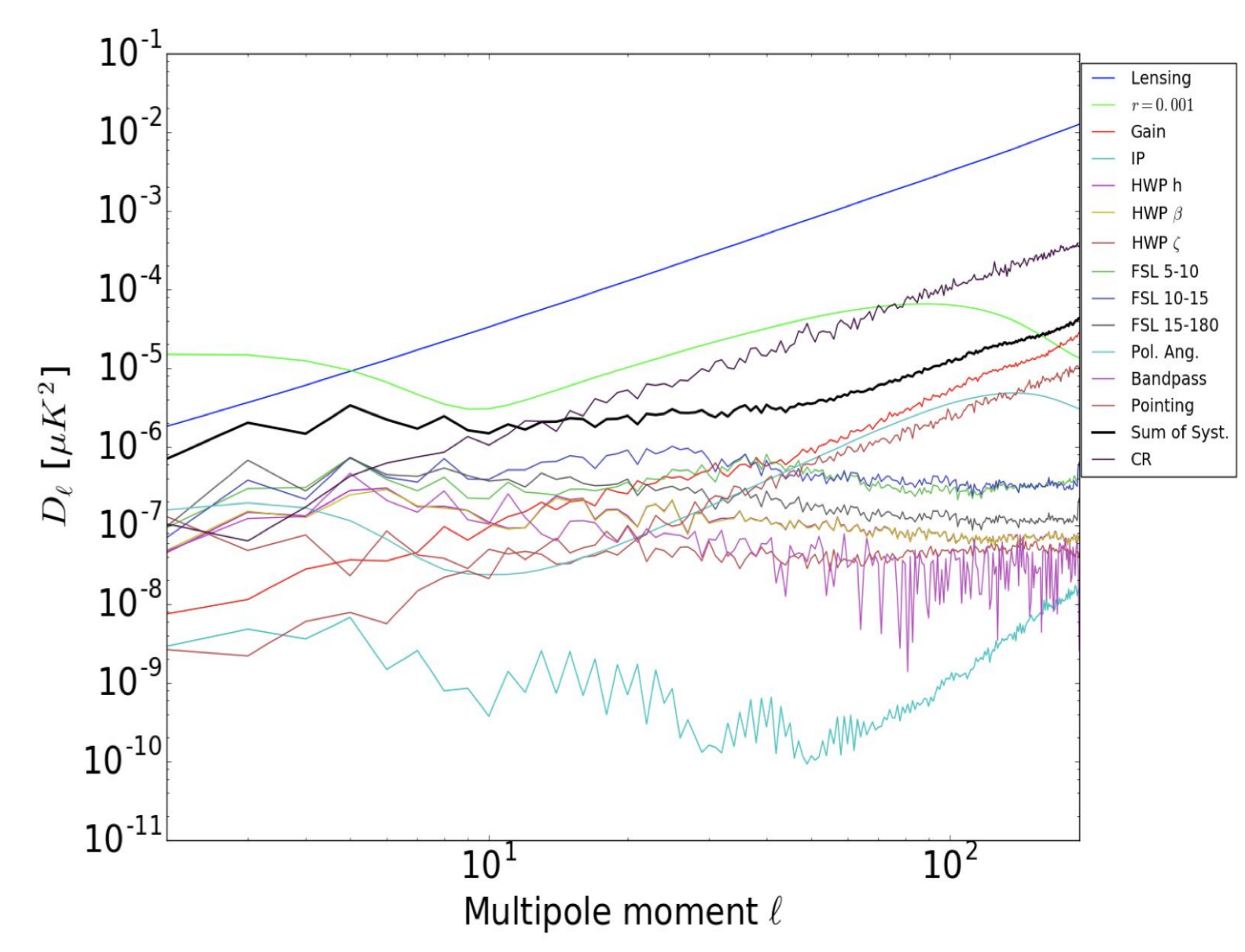
## Contamination by Galactic foregrounds



### Sensitivity of LiteBIRD over the frequency bands

LiteBIRD will observe the sky in a frequency range from 34 to 448 GHz to characterize synchrotron (dominant at LF), dust emission (dominant at HF) and the CMB signal.

## Systematic effects



### Angular power spectrum of the systematic effects

The systematic effect arise from an imperfect characterization/calibration of the instrument. Oneway to control the impact of systematics is to set requirements on the calibration of a specific effect.

## Setting requirements on the gain calibration accuracy

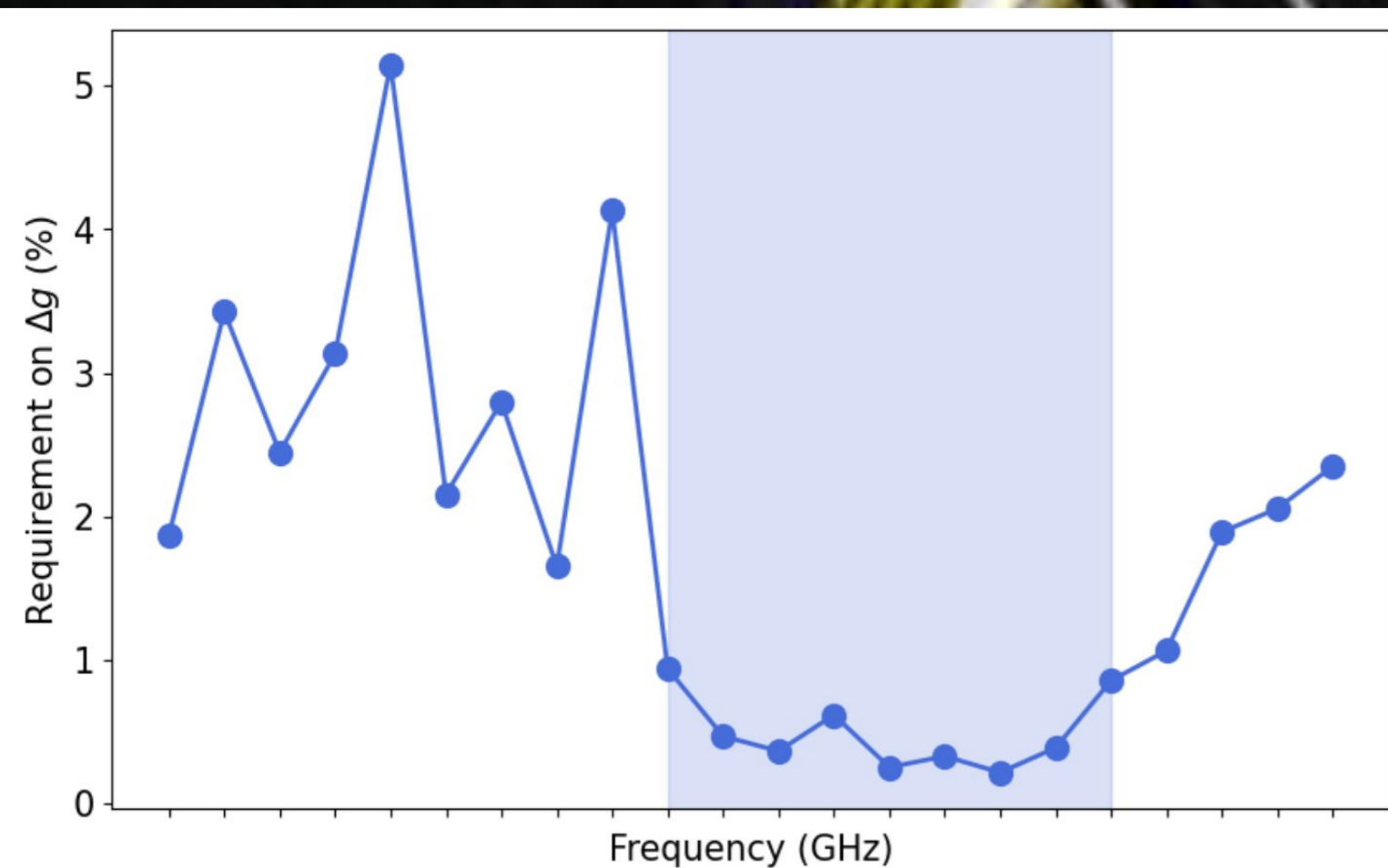
- Simulate multi-frequency B-modes maps of CMB ( $r=0$ ), synchrotron and dust + gaussian white noise
- Inject a gain mis-match at a specific channel and let the other channels unperturbed  $g = 1 + N(0, \Delta g)$
- Apply component separation on the set of maps containing the systematic effect and on the twin set of maps with ideal calibration ( $g=1$ )
- Estimate the tensor-to-scalar ratio value in both cases and determine the systematic bias
- Find the value of the gain calibration accuracy  $\Delta g$  for which the resulting mean bias is equal to the gain systematic budget per channel ( $6.5 \cdot 10^{-6}/22$ )

## Foreground cleaning : Needlet Internal Linear Combination (NILC)

ILC-type methods consist in performing a linear of frequency data to minimize the variance of the output.

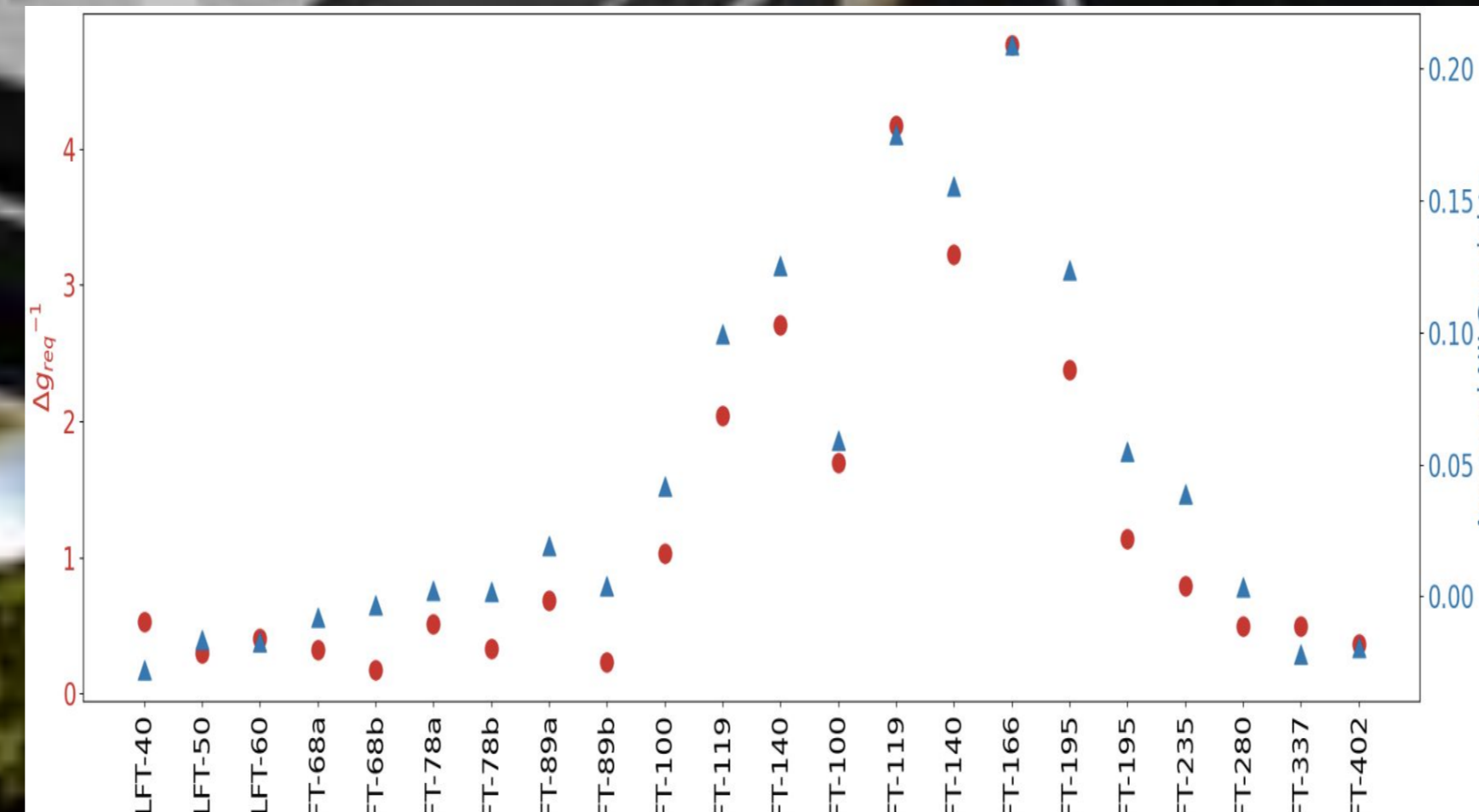
$$X_{\text{CMB}} = \sum w_v \cdot X^v = \sum w_v \cdot (a_{\text{CMB}} X_{\text{CMB}} + X_{\text{fg}}^v + X_n^v)$$

$$\sum w_v \cdot a_{\text{CMB}} X_{\text{CMB}} = X_{\text{CMB}}$$



### Requirements on the gain calibration accuracy

The most sensitive frequency channels are those at CMB frequencies (0.22% for the 166 GHz) while low and high frequency channels possess a requirement larger than 1%.



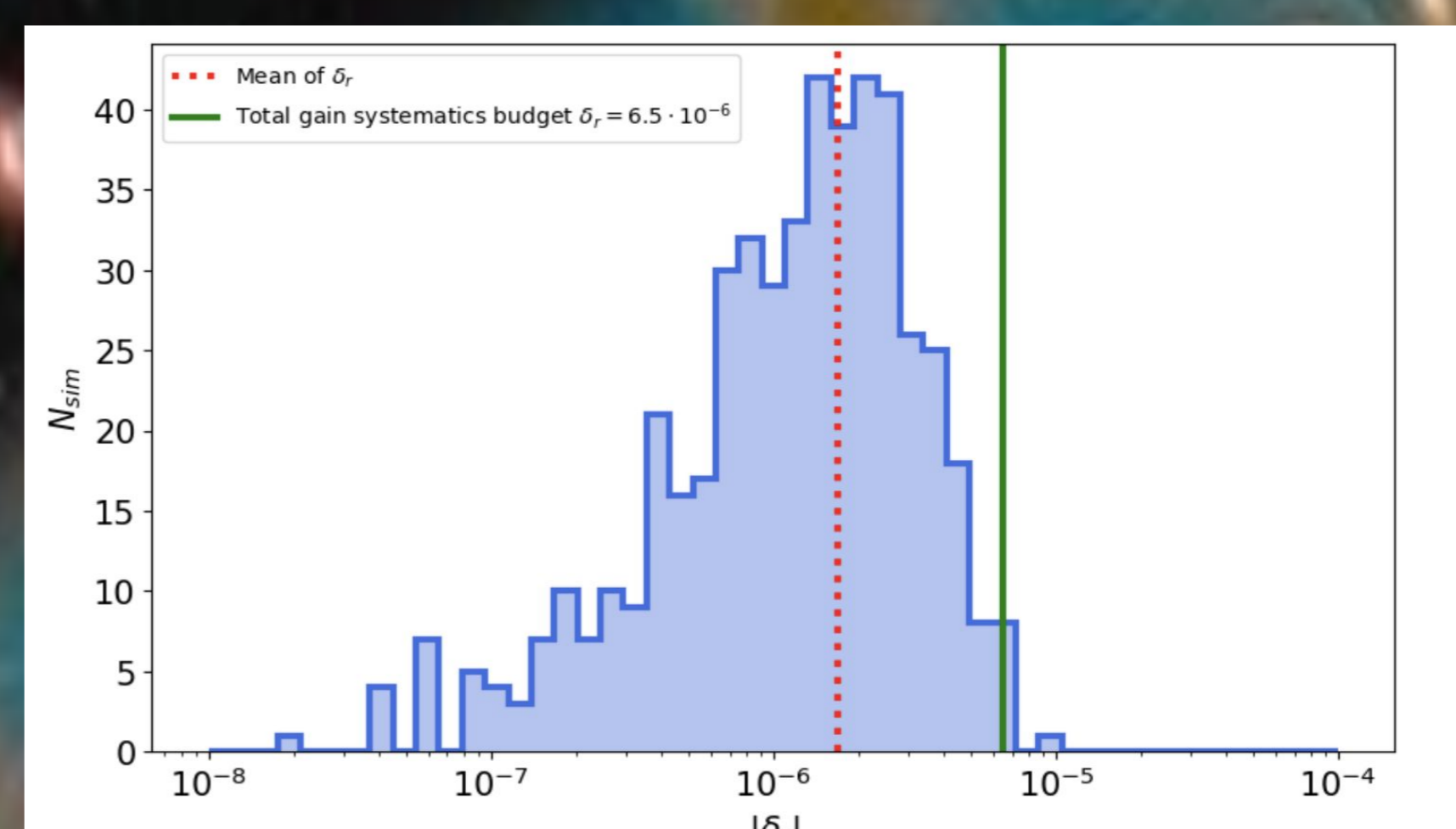
### Anti-correlation between the requirements on the gain calibration accuracy and the NILC weights

The frequency channels most affected by gain mis-calibration are those contributing the most to the CMB reconstruction process : characteristic of an analysis with ILC, the observation is different when employing parametric component separation.

## Likelihood estimation of the tensor-to-scalar ratio

Probability distribution function for an observed power spectrum  $C_\ell^{\text{obs}}$  to match a theoretical model  $C_\ell^{\text{th}}$

$$\mathcal{L}(C_\ell^{\text{obs}} | r) = \sum 2\ell+1 f_{\text{sky}} [C_\ell^{\text{obs}}/C_\ell^{\text{th}} + \ln(C_\ell^{\text{th}}) - 2\ell-1/2\ell+1 \cdot \ln(C_\ell^{\text{obs}})]$$



### Bias distribution obtained when all channels are mis-calibrated simultaneously by their corresponding requirements

The resulting mean bias is below the total gain systematic budget by a factor of 3.5