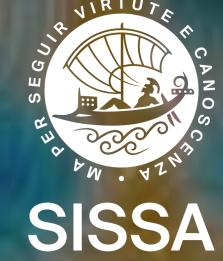


# Systematic effects and foreground cleaning for LiteBIRD satellite

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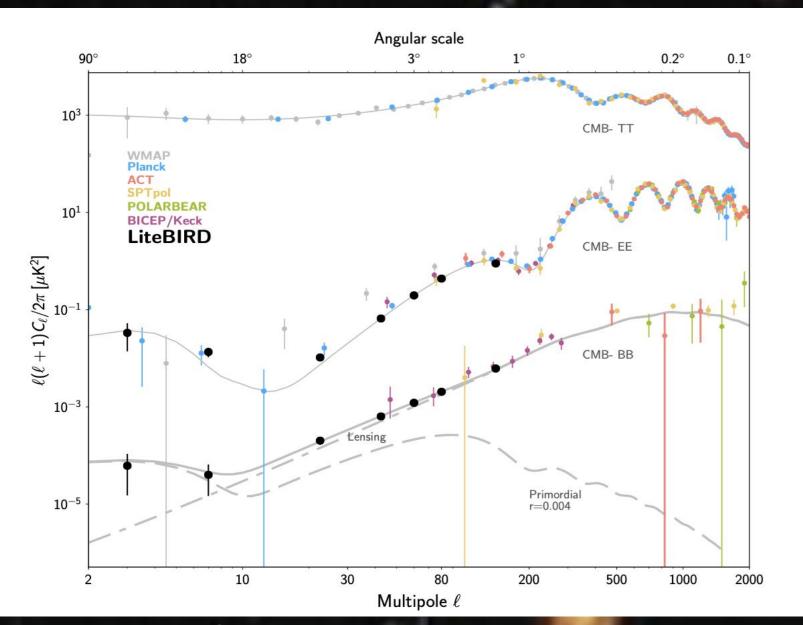
LiteBIRD is a space-borne experiment proposed by the Japanese Aerospace Exploration Agency (JAXA) and gathering a worldwide collaboration. This satellite aim at detecting CMB B-modes polarization which could contain an imprint of primordial gravitational waves generated by cosmic inflation. LiteBIRD will target the tensor-to-scalar ratio r, associated to the amplitude of primordial tensor perturbations (compared to scalar fluctuations) with a total uncertainty of  $\sigma_r < 10^{-3}$  by measuring both reionization and recombination bumps of the CMB B-modes power spectrum.

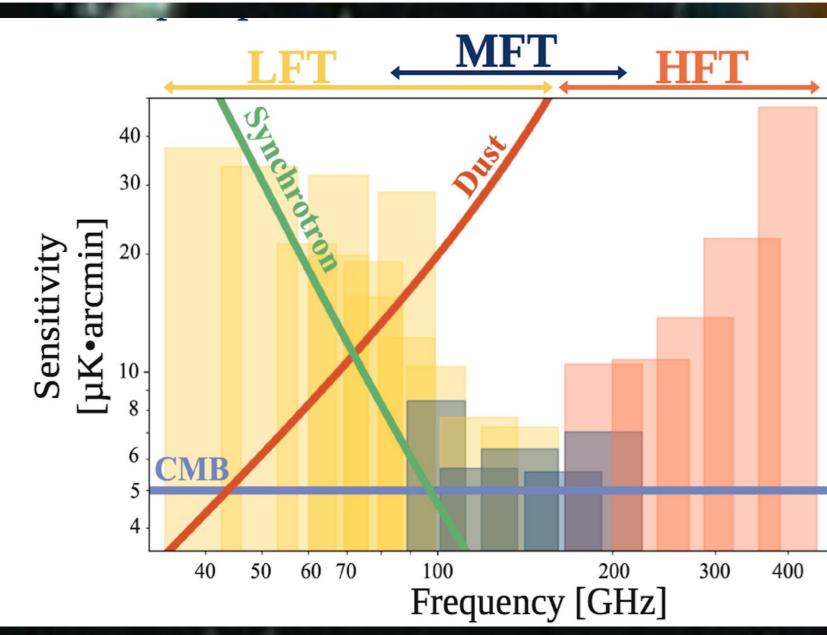
### Scientific target of LiteBIRD

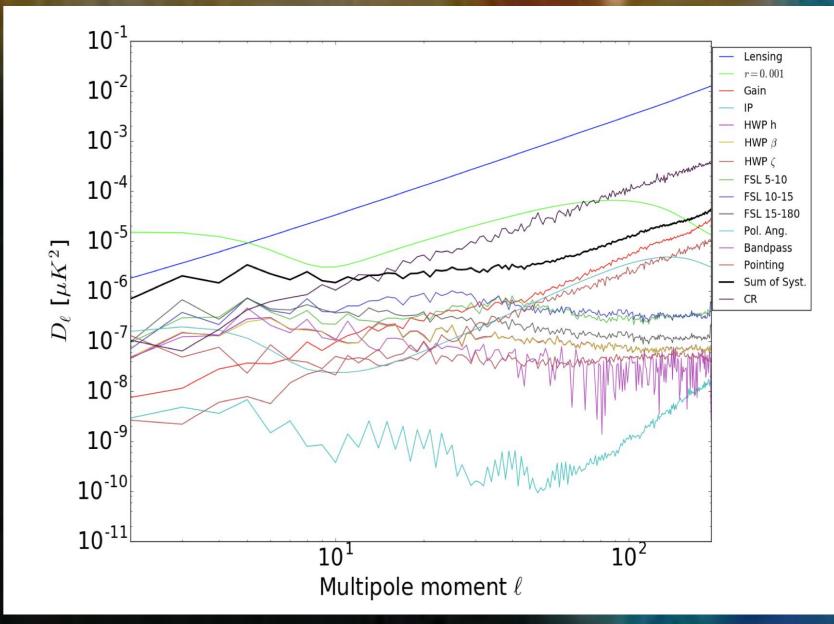
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# **Contamination by Galactic foregrounds**

### Systematic effects







TT, EE and BB CMB angular power spectra LiteBIRD will cover a multipole range up to l = 200to measure the reionization and recombination bumps of the CMB BB power spectrum.

Sensitivity of LiteBIRD over the frequency bands LiteBIRD will observe the sky in a frequency range from 34 to 448 GHz to characterize synchrotron (dominant at LF), dust emission (dominant at HF) and the CMB signal.

Angular power spectrum of the systematic effects The systematic effect arise from an imperfect characterization/calibration of the instrument. Oneway to control the impact of systematics is to set requirements on the calibration of a specific effect.

## Setting requirements on the gain calibration accuracy

- Simulate multi-frequency B-modes maps of CMB (r=0), synchrotron and dust + gaussian white noise
- Inject a gain mis-match at a specific channel and let the other channels unperturbed  $g = 1 + N(0, \Delta g)$
- Apply component separation on the set of maps containing the systematic effect and on the twin set of maps with ideal calibration (g=1)
- Estimate the tensor-to-scalar ratio value in both cases and determine the systematic bias
- Find the value of the gain calibration accuracy Δg for which the resulting mean bias is equal to the gain systematic budget per channel (6.5.10<sup>-6</sup>/22)

Foreground cleaning : Needlet Internal I Linear **Combination (NILC)** 

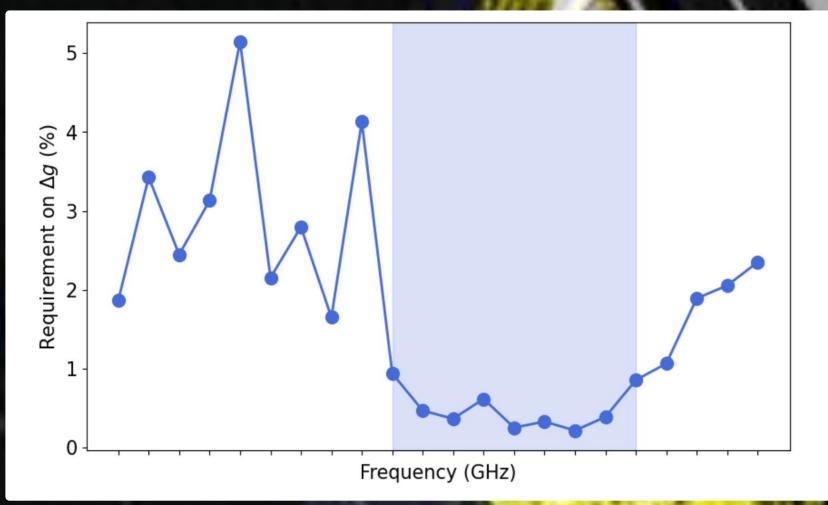
ILC-type methods consist in performing a linear of frequency data to minimize the variance of the output.  $X_{CMB} = \sum \omega_{\nu} X^{\nu} = \sum w_{\nu} (a_{CMB} X_{CMB} + X_{f\sigma}^{\nu} + X_{p}^{\nu})$ 

Likelihood estimation of the tensor-to-scalar ratio

Probability distribution function for an observed power spectrum  $C_{\rho}^{obs}$  to match a theoretical model  $C_{o}^{th}$ 

$$CIVIB = v = v \cdot CIVIB CIVIB Ig n'$$

$$\sum w_v \cdot a_{CMB} X_{CMB} = X_{CMB}$$

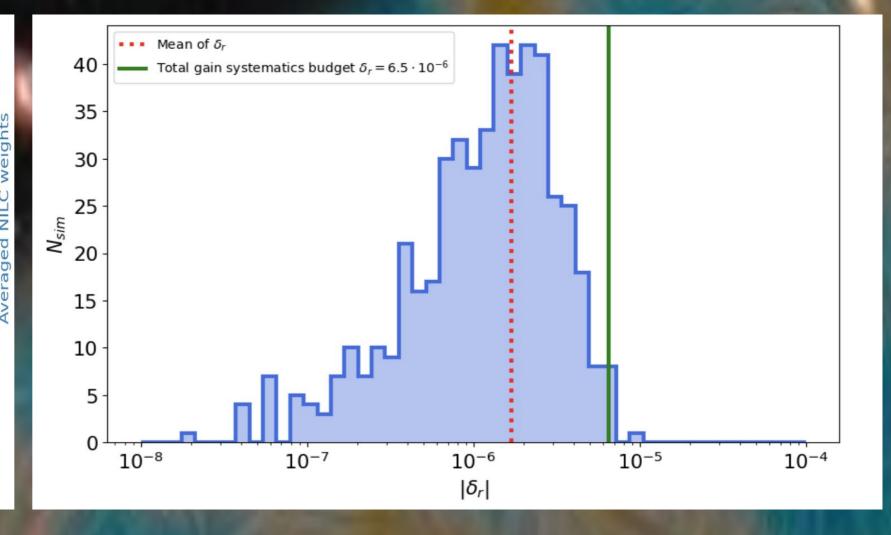


Requirements on the gain calibration accuracy The most sensitive frequency channels are t CMB frequencies (0.22% for the 166 GHz) while low and high frequency channels possess a requirement larger than 1%.

### Anti-correlation between the requirements on the gain calibration accuracy and the NILC weights

The frequency channels most affected by gain mis-calibration are those contributing the most to the CMB reconstruction process : characteristic of an analysis with ILC, the observation is different when employing parametric component separation.

$$\mathcal{L}(C_{\ell}^{\text{obs}}|r) = \sum_{\ell=1}^{\infty} 2\ell + 1 f_{sky} \left[ C_{\ell}^{\text{obs}} / C_{\ell}^{\text{in}} + \ln(C_{\ell}^{\text{in}}) - 2\ell - 1/2\ell + 1 \cdot \ln(C_{\ell}^{\text{obs}}) \right]$$



**Bias distribution obtained when all channels are** mis-calibrated simultaneously by their corresponding requirements

The resulting mean bias is below the total gain systematic budget by a factor of 3.5

F.Carralot, N.Krachmalnicoff, A.Carones et al. Requirements on the relative polarization gain calibration for LiteBIRD with blind component separation (in prep.) T.Ghigna, T.Mastumura, G.Patanchon . Requirements for future CMB satellite missions : photometric and band-pass response calibration (2020) M.Hazumi LiteBIRD: A Satellite for the Studies of B-Mode Polarization and Inflation From Cosmic Background Radiation Detection (2019)

#### A.Carones, M.Migliaccio, G.Puglisi et al. Multi-Clustering Needlet ILC for CMB B-mode component separation (2023)

#### J.Dick, M.Remazeilles, J.Delabrouille Impact of calibration errors on CMB component separation using FastICA and ILC (2009))