

Primordial Black Holes

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The work was supported by the RSF Grant 23-42-00066

Conference in memory of Veniamin Sergeyevich Berezinsky
Gran Sasso
Italy

October 1st to October 3rd, 2024

Black holes who are they?

The black hole existence was ingeniously predicted in 1783 by John Michell, an English country parson, famous for many other discoveries in physics.

He noticed that there could be stellar bodies having the second cosmic velocity larger than the speed of light. Since such objects neither shine nor reflect light, it would be impossible to observe them directly.

Michell called such, not emitting radiation stars as "dark stars". According to his idea a single dark star would be invisible, but if a double system of a dark and a usual star is formed, one may identify dark star observing the other one rotating around "nothing".

This is one of possible ways to observe black holes at the present time, among many others.

Observations of black holes

Nowadays many methods are invented to observe possible black holes.

BHs evaporate and shine (Hawking radiation), though nobody yet saw it.

Near-solar mass BHs are observed by X-rays from the accreting matter.

Gravitational lensing by BH, e.g. that's how $0.5M_{\odot}$ MACHOs are spotted.

Mass estimates via star motion around BH, as e.g. in our Galaxy.

The most powerful sources of radiation, quasars (QSO), are supermassive black holes, that radiate as thousands galaxies, $L_{QSO} = 10^{46} - 10^{47}$ erg/sec, i.e. $10^{13}L_{\odot}$, though they are practically point-like, their size is about $10^9 - 10^{10}$ km, i.e. smaller than the Solar System. The only known mechanism of QSO radiation is the process of ultrarelativistic particle collision in the process of matter accretion. QSOs shine until they consume all "food" around and remains almost in desert. After that the shining fades out as e.g. BH in the center of our Galaxy.

All these methods however only allow to determine the mass inside central volume. According to General Relativity there must be a BH inside.

Strictly speaking BH existence was not proven by "experiment".

LIGO/Virgo/KAGRA registration of gravitational waves from BH binaries demonstrates the first direct "picture" of the Schwarzschild metric and presents strong evidence that the source are PBHs.

General relativity and BHs

After creation of GR (Einstein, 1915) almost immediately exact solutions of GR equations describing all possible types of BHs have been found.

BHs have only three types of "hairs" that may be observed outside BH:

- 1. Gravitational field created by mass, M , asymptotically at large distances tending to the Newtonian limit.
- 2. Electric field created by charge Q , tending to the Coulomb one.
- 3. Field induced by rotation, created by spin J .

Four known exact solution describing all existing BHs.

Schwarzschild (1916), BH with $Q = J = 0$.

Reissner-Nordström (1916, 1918) $J = 0$, $Q \neq 0$.

Kerr, (1963), $Q = 0$, $J \neq 0$.

Kerr–Newman (1965) $J \neq 0$, $Q \neq 0$.

If photon mass is non-zero, no matter how tiny, electric field of BH would completely vanish. **No continuous limit to $m_\gamma = 0$.**

Discovery of charged BH could present the absolute upper bound on m_γ .

Schwarzschild metric

$$ds^2 = (1 - r_g/r)dt^2 - (1 - r_g/r)^{-1}dr^2 - r^2(d\theta^2 + \sin^2\theta d\phi^2),$$

where $r_g = 2M/m_{Pl}^2$.

System of units: $c = \hbar = k = 1$.

Gravitational coupling constant $G_N = 1/m_{Pl}^2$,

$m_{Pl} = 2.176 \times 10^{-5}g = 1.22 \times 10^{19} \text{ GeV} = 6.17 \times 10^{32} \text{ cm}^{-1}$.

Newtonian force: $F_N = G_N m_1 m_2 / r^2$, non-relativistic limit.

However, the source of gravity is not mass but the energy-momentum tensor. A lot of poor knowledge claims of antigravitating antimatter.

For the Sun: $r_g^\odot \approx 3 \text{ km}$.

For the Earth $r_g^\oplus \approx 3 \text{ cm}$. To make BH out of the Earth one has to compress it down to 3 cm.

Some unusual properties of space and time near BH horizon: from our position an object infalling into BH would never cross the BH horizon; from their point of view it happens during finite short time (look at the metric).

BH types by formation mechanisms

1. Astrophysical black holes,

created by the collapse of a star which exhausted its nuclear fuel. The expected masses should start immediately above the neutron star mass, i.e. about $3M_{\odot}$, but noticeably below $100M_{\odot}$. Instead we observe that the BH mass spectrum in the galaxy has maximum at $M \approx 8M_{\odot}$ with the width $\sim (1 - 2)M_{\odot}$. The result is somewhat unexpected but an explanation in the conventional astrophysical frameworks is possible.

Recently LIGO/Virgo discovered BHs with masses close to $100M_{\odot}$. Their astrophysical origin was considered **impossible due to huge mass loss in the process of collapse.**

Now some, quite exotic, formation mechanisms are specially invented. No problem if $100M_{\odot}$ BH is primordial.

BH created by accretion

2. BH formed by accretion to the mass excess in the galactic center. In any large galaxy there exists a supermassive BH (SMBH) at the center, with masses varying from a few millions M_{\odot} (e.g, Milky Way) up to almost hundred billions M_{\odot} . However, the conventional accretion mechanisms are not efficient enough to create such monsters during the universe life-time, $t_U \approx 14.6$ Gyr. At least 10-fold longer time is necessary, to say nothing about SMBH in 20 times younger universe, observed in impressive numbers by HST and JWST.

Primordial Black Holes

3. Primordial black holes (PBH) created during pre-stellar epoch

The idea of the primordial black hole (PBH) i.e. of black holes which could be formed the early universe prior to star formation was first put forward by Zeldovich and Novikov: "The Hypothesis of Cores Retarded During Expansion and the Hot Cosmological Model", *Astronomicheskij Zhurnal*, 43 (1966) 758, *Soviet Astronomy*, AJ.10(4):602–603;(1967).

According to their idea, the density contrast in the early universe inside the bubble with radius equal to the cosmological horizon might accidentally happen to be large, $\delta\rho/\rho \approx 1$, then that piece of volume would find itself inside its gravitational radius i.e. it became a PBH, decoupled from the cosmological expansion.

Elaborated later in S. Hawking, "Gravitationally collapsed objects of **very low mass**", *Mon. Not. Roy. Astron. Soc.* **152**, 75 (1971).

B. J. Carr and S. W. Hawking, "Black holes in the early Universe," *Mon. Not. Roy. Astron. Soc.* **168**, 399 (1974).

PBHs are mostly not accepted by astrophysical community. Why?!

BH types by masses

There is the following conventional division of black holes by their masses:

1. Supermassive black holes (SMBH): $M = (10^6 - 10^{11})M_{\odot}$.
2. Intermediate mass black holes (IMBH): $M = (10^2 - 10^5)M_{\odot}$.
3. Solar mass black holes: masses from a fraction of M_{\odot} up to $100M_{\odot}$.

The origin of most of these BHs is unclear in the conventional approach, except maybe of the BHs with masses of a few solar masses, that might be astrophysical.

Highly unexpected was great abundance of IMBH which are being observed at the present day universe during last few years in large numbers.

The suggestion that (almost) all black holes in the universe are primordial strongly reduce or even eliminate the tension.

However, in earlier works the emerged masses of PBH were quite low.

PBH and inflation

Cosmological inflation allows for formation of PBH with very large masses. Mass of perturbations inside cosmological horizon grows exponentially, that allows for creation even of supermassive PBHs.

It was first applied to PBH production in

A.Dolgov, J.Silk, PRD 47 (1993) 4244 (DS) "Baryon isocurvature fluctuations at small scale and baryonic dark matter".

a year later in: B.J. Carr, J.H. Hilbert, J.E. Lidsey, "Black hole relics and inflation: Limits on blue perturbation spectra", Phys.Rev.D 50 (1994) 4853, astro-ph/9405027;

and soon after in P. Ivanov, P. Naselsky, I. Novikov (May 10, 1994), "Inflation and primordial black holes as dark matter", PRD 50 (1994) 7173.

An avalanche of papers on inflationary formation of PBH nowadays.

Presently inflationary mechanism of PBH production is commonly used.

However, almost always the calculated mass spectrum is multi-parameter one and quite complicated.

Log-normal mass spectrum

The only exception is the log-normal mass spectrum predicted by Dolgov-Silk and further worked out by A.Dolgov, M. Kawasaki, N. Kevlishvili (DKK), NPB807 (2009) 229, "Inhomogeneous baryogenesis, cosmic antimatter, and dark matter":

$$\frac{dN}{dM} = \mu^2 \exp[-\gamma \ln^2(M/M_0)],$$

tested by LiGO/Virgo/Kagra with very good agreement, see below.

According to DS and DKK scenario initial large perturbations of baryonic number (isocurvature perturbations, since quarks are massless) are generated at inflationary stage, transformed into density perturbations when massless quarks turned into heavy protons and neutrons.

Hence M_0 should be equal to horizon mass at the QCD phase transition:

$M_0 \sim 10M_\odot$, as predicted, A.Dolgov, K.Postnov, "Why the mean mass of primordial black hole distribution is close to $10M_\odot$ ". JCAP 07 (2020) 063.

The horizon mass at QCD p.t. is close to $10M_\odot$, for $\mu = 0$. At larger chemical potential T_{pt} is smaller and M_{hor} is larger, namely, $\approx 17M_\odot$.

Seeding of galaxy formation by PBHs

Usually it is assumed that supermassive black holes (SMBHs) observed in large galaxies, are created by matter accretion to the galactic centers.

However, the necessary time is much larger than the universe age, even for the contemporary universe, with the age about 15 billion years, to say nothing of the 20 times younger universe at $z \sim 10$.

Contradiction between observations and canonical theory is neatly solved if the universe is populated by primordial black holes (PBH) that **seeded** formation of cosmic structures, as is suggested 30 years ago:

A.Dolgov, J.Silk, PRD 47 (1993) 4244 (DS), "Baryon isocurvature fluctuations at small scale and baryonic **dark matter**";

A.Dolgov, M. Kawasaki, N. Kevlishvili (DKK), NPB807 (2009) 229, "Inhomogeneous baryogenesis, **cosmic antimatter**, and dark matter"; and subsequent publications by our group.

Inverted mechanism of galaxy formation

In DS and DKK an **inverted** formation mechanism of galaxies and their central black holes is proposed. Namely, a primordial SMBH was formed firstly and later it **seeded** the galaxy formation.

The 30 year old DS idea of seeding, but of different kind, was rediscovered in several recent publications, under the pressure of the Hubble Space Telescope HST and the James Webb Telescope (JWST) observations of the early universe at redshifts $z = 5 - 15$.

Crisis in cosmology

Conventional Λ CDM cosmology encounters serious difficulties describing astronomical observations during **all the history** of the universe, **starting from our time with the universe age about 15 billion years, and back to the past down to ~ 300 million years discovered recently by HST, JWST, and ALMA at high redshifts $z \sim 10$.**

The inconsistencies discovered earlier are reviewed in 2018: A.D. "Massive and supermassive black holes in the contemporary and early Universe and problems in cosmology and astrophysics", Phys. Usp. 61 (2018) 2, 115.

The tension that existed during all life-time of the universe between the canonical theory and observations is neatly solved if the universe is populated by primordial black holes (PBH) that **seeded** formation of cosmic structures, **as is suggested in DS and DKK.**

Confirmation by "experiment"

Observational confirmation of DS/DKK mechanism:

- The calculated mass spectrum of PBH very well agrees with "experiment".
- Noticeable antimatter population of the Galaxy is anticipated and confirmed by the observations of **positrons, antinuclei, and antistars**.
- The early galaxy formation observed by HST and JWST is explained if galaxies are SEEDED by BHs, as it is also rediscovered in several papers of the last years.
- Discovery of IMBH, with $M \sim (10^3 - 10^5) M_{\odot}$ in dwarfs and globular clusters, predicted in AD & K. Postnov. "Globular Cluster **Seeding** by Primordial Black Hole Population", JCAP 04 (2017) 036, e-Print: 1702.07621 [astro-ph.CO].
- **Recent observation of antistar - gamma burster possible source star-antistar collision!!!**

Problems of the contemporary universe. Summary.

1. Supermassive BH (SMBH) in all large galaxies. Too short time (15 billion year) for their formation through the conventional accretion mechanism.
2. Huge SMBH in small galaxies and even in (almost) EMPTY space. No material for their creation. Pushed out of large galaxies? Wandering BHs?
3. Too old stars, older than the Galaxy and maybe one **older** than the universe?
4. MACHOs, non-luminous objects, $M \sim 0.5M_{\odot}$, observed through microlensing.
5. Problems with the BH mass spectrum in the Galaxy with $(M = 7.8 \pm 1.2)M_{\odot}$.
6. Origin and properties of the sources of the observed gravitational waves.
7. Origin of intermediate mass BH (IMBH) with masses $(10^3 - 10^5)M_{\odot}$. Plenty of them are observed everywhere in the universe, in particular in dwarfs and globular clusters, as predicted by AD & K. Postnov.
- 8. Discovery of BH with $M \approx 100M_{\odot}$, that is strictly forbidden but nevertheless observed by LIGO/Virgo.**
9. Strange stars in the Galaxy, too fast and with unusual chemistry.

Problems of the early universe

Serious problems, similar to those recently found by JWST, are known already for many years. HST discovered that the early universe, at $z = 6 - 7$ is too densely populated with quasars, alias SMBH, supernovae, gamma-bursters and it is very dusty. **No understanding how all these creature were given birth in such a short time exists in conventional cosmology.**

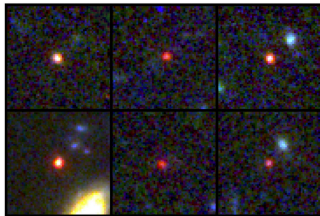
"Hubble" sees the universe up to $z = 6 - 7$, but accidentally a galaxy at $z \approx 12$ has been discovered for which both Hubble and Webb are in good agreement.

Huge BHs in small galaxies discovered in the early universe (as well as in the contemporary one). **Such huge BHs could not be created by the accretion of matter to galactic center, since the amount of material is too small.**

All the problems are neatly solved if the universe is populated by primordial black holes (PBH) and the astrophysical large bubbles with very high baryonic density, according to DS and DKK.

Impossible galaxies

I. Labbé et al, A population of red candidate massive galaxies 600 Myr after the Big Bang, Nature, published online 22.02.2023, Six candidate massive galaxies (stellar mass $> 10^{10}$ solar masses) at $7.4 \lesssim z \lesssim 9.1$ 500–700 Myr after the Big Bang, including one galaxy with a possible stellar mass of $\sim 10^{11} M_{\odot}$, too massive to be created in so early universe. According to the 'science' it is impossible to create so well developed galaxies. **NB: "May be they are supermassive black holes of the kind never seen before. That might mean a revision of usual understanding of black holes."** Well agrees with our predictions of PBHs.



The six candidate galaxies identified in the JWST data. (NASA, ESA, CSA, I. Labbé/Swinburne University of Technology)

Ultra-massive early QSO observed by ALMA

ALMA (Atacama Large Millimeter Array) confirmation of an obscured hyper-luminous radio-loud AGN at $z = 6.853$ associated with a dusty starburst in the 1.5 deg² COSMOS field,

R. Endsley et al, MNRAS 520, 3, , p 4609, (2023)

VIRCam and IRAC photometry perhaps suggests that COS-87259 is an extremely massive reionization-era galaxy with $M_* = 1.7 \times 10^{11} M_\odot$

Such a very high AGN luminosity suggests that this object is powered by $\sim 1.6 \times 10^9 M_\odot$ black hole if accreting **near the Eddington limit.**

BH mass is about 1% of the stellar mass, 100 times larger than usually.
Nearly impossible, but PBH could seed such monster.

Summarising: recent observations have found a large number of supermassive black holes already in place in the first few hundred million years after Big Bang. **The channels of formation and growth of these early, massive black holes are not clear, with scenarios ranging from heavy seeds to light seeds experiencing bursts of high accretion rate.**

Rediscovery of galaxy seeding by BH

Seeding was advocated also in A. Bogdan, A. Goulding, P. Natarajan, *et al*, "Evidence for heavy-seed origin of early supermassive black holes from a $z \approx 10$ X-ray quasar", *Nature Astron.* 8 (2024) 1, 126, 2305.15458 [astro-ph.GA] and A.D. Goulding, J.E. Greene, D. J. Setton, *et al*, "UNCOVER: The Growth of the First Massive Black Holes from JWST/NIRSpec—Spectroscopic Redshift Confirmation of an X-Ray Luminous AGN at $z = 10.1$ ", *Astrophys. J. Lett.* 955 (2023) 1, L24 • e-Print: 2308.02750 [astro-ph.GA].

It was postulated that seeds of the observed early galaxies and SMBHs could be either light BH with masses $(10 - 100)M_{\odot}$, or heavy ones, $M = (10^4 - 10^5)M_{\odot}$. According to the authors the light BHs could be remnants of the first stars and the heavy ones might be created by direct collapse of gas clouds, **assuming accretion at the Eddington limit.**

Observations of high-redshift quasars reveal that many supermassive black holes were in place less than 700 million years after the Big Bang.

However, the origin of the first BHs remains a mystery.

Rediscovery of quasar seeding by BH

The same paper, A. Bogdan, et al [Detection of an X-ray quasar in a gravitationally-lensed \$z = 10.3\$ galaxy suggests that early supermassive black holes originate from heavy seeds](#), 2305.15458 [astro-ph.GA]

The detection of an X-ray-luminous quasar powered by SMBH with the mass $\sim 4 \times 10^7 M_{\odot}$ in the galaxy identified by JWST at $z \approx 10.3$ is reported.

This mass is comparable to the inferred stellar mass of its host galaxy, in contrast to the usual examples from the local universe where mostly the BH mass is $\sim 0.1\%$ of the host galaxy's stellar mass. The combination of such a high BH mass and large BH-to-galaxy stellar mass ratio ~ 500 Myrs after the Big Bang is consistent with a picture wherein **such BHs originated from heavy seeds, accreting near Eddington limit, as is stated in the previous slide.**

Much simpler and easier if the seeds are primordial BH, suggested by DS and DKK.

Strong blow to the super-Eddington accretion

A dormant, overmassive black hole in the early Universe I. Juodzbalius, R. Maiolino, W.M. Baker, *et al*, 2403.03872: **Recent observations have found a large number of SMBHs already in place in the first few hundred million years after Big Bang. The channels of formation and growth of these early, massive black holes are not clear, with scenarios ranging from heavy seeds to light seeds experiencing bursts of high accretion rate.**

The detection, from the JADES survey, of broad H α emission in a galaxy at $z = 6.68$, which traces a black hole with mass of $\sim 4 \times 10^8 M_{\odot}$ and accreting at a rate of **only 0.02 times the Eddington limit.**

The black hole to stellar mass ratio is 0.4, i.e. 10^3 times above the local relation. Huge BH in small galaxies, see more examples below.

This object is most likely the tip of the iceberg of a much larger population of dormant black holes around the epoch of reionization. Its properties are consistent with scenarios in which short bursts of **super-Eddington accretion have resulted in black hole overgrowth and massive gas expulsion from the accretion disk; in between bursts, black holes spend most of their life in a dormant(?) state.**

Huge black holes in tiny galaxies at high z

Black holes in high z universe are too massive w.r.t. the expectations based on observations of BHs in contemporary large galaxies - PBHs solve the problem.

F. Pacucci, B. Nguyen, S. Carniani *et al* "JWST CEERS and JADES Active Galaxies at $z = 4 - 7$ Violate the Local $M_* - M_{BH}$ Relation at $> 3\sigma$: Implications for Low-mass Black Holes and Seeding Models", The Astrophysical Journal Letters, 957:L3 (10pp), 2023 November 1.

Black holes are overmassive by factor 10 – 100 compared to their low- z counterparts in galactic hosts of the same stellar mass.

M. Volonteri, M. Habouzit, M. Colpi. "What if young $z > 9$ JWST galaxies hosted massive black holes?" Monthly Notices of the Royal Astronomical Society, Volume 521, Issue 1, pp.241-250. Only MBHs overmassive relative to expected galaxy scaling relations, accreting at high Eddington rates, would be detectable. Their discovery would point to the presence of heavy MBH seeds, but care is needed to exclude the existence of lighter seeds as only overmassive MBHs.

Possible resolution: BH seeds operated too little time to create very massive galaxies as in the contemporary universe.

IMBH in tiny galaxies today

AD+K.Postnov prediction "Globular Cluster **Seeding** by Primordial Black Hole Population", JCAP 04 (2017) 036, e-Print: 1702.07621 [astro-ph.CO] is recently verified:

The observation of IMBH in Omega-Centauri (the most massive globular cluster of the Milky Way): M. Häberle, N. Neumayer, A. Seth *et al*, Nature, **631**, 285–288 (2024): $M_{IMBH} \gtrsim 8200 M_{\odot}$;

F. Peißker, M. Zajaček, M. Labaj, *et al*, "The Evaporating Massive Embedded Stellar Cluster IRS 13 Close to Sgr A*. II. Kinematic Structure".

Discovery of IMBH with $M \approx 3 \times 10^4 M_{\odot}$, 2024, ApJ **970** 74.

The origin of IMBH is mysterious, if they are not primordial.

Huge black holes in tiny galaxies today

Primordial IMBHs with masses of a few thousand solar mass explain formation of globular clusters (GCs) and dwarf galaxies, otherwise the formation is not well understood, even mysterious.

In the last several years several such IMBH inside GSs are observed. Similar IMBHs are observed in dwarf galaxies.

A striking example: discovery by the Hobby-Eberly Telescope at Texas's McDonald Observatory of a SMBH with $M_{BH} \approx 1.7 \cdot 10^{10} M_{\odot}$ i.e. 14% of the stellar mass of the galaxy.

Usually the mass of the central BH is about 0.1 % of the galaxy mass.

Huge BHs in dwarf galaxies

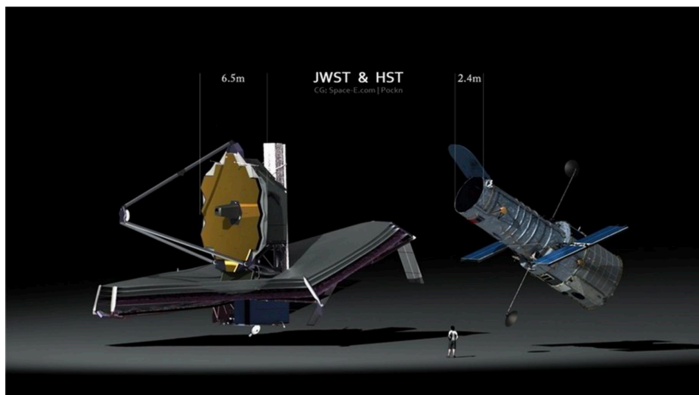
The seeding of dwarfs by intermediate mass BHs is confirmed by the recent data, e.g. in the dwarf galaxy SDSS J1521+1404 the BH is discovered with the mass $M \sim 10^5 M_{\odot}$.

Two Candidates for Dual AGN in Dwarf-Dwarf Galaxy Mergers, M. Mićić, et al, arXiv:2211.04609 [astro-ph.GA]. For the first time, astronomers have spotted evidence of a pair of dwarf galaxies featuring GIANT black holes on a collision course with each other. In fact, they haven't just found just one pair – they've found two.

Intermediate-mass black holes: finding of episodic, large-scale and powerful jet activity in a dwarf galaxy SDSS J090613.77+561015.2. Jun Yang et al, e-Print: 2302.06214 [astro-ph.GA,astro-ph.HE].

Discovery of an intermediate-mass black hole (IMBH) with a mass of $M_{BH} = 3.6_{-2.3}^{+5.9} \times 10^5 M_{\odot}$, **that surely cannot be created by accretion but might seed the dwarf formation.**

JWST infrared telescope and HST



Placing a telescope in space makes it possible to register electromagnetic radiation in the ranges in which the earth's atmosphere is opaque; primarily in the infrared range. Due to the absence of the influence of the atmosphere, the resolution of the telescope is 7-10 times greater than that of a similar telescope located on Earth.

Comparison of JWST and HST

HST: Distance: 570 km

Mirror 2.4 m

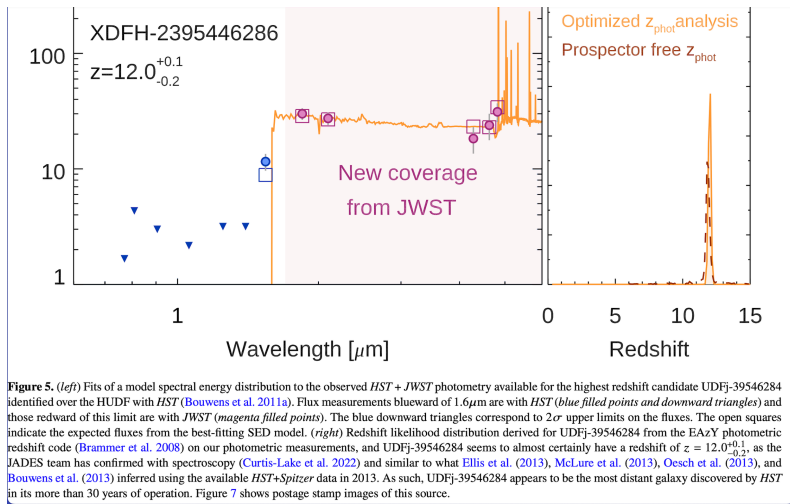
Wave length: optical, e.g. blue 450 nm and UV,
some IR: 0.8-2.5 microns;

JWST: Distance 1.5×10^6 km

Mirror: 6.5 m

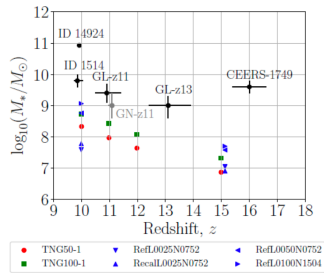
Wave length: 0.6 - 28,5 micron.

JWST and HST common galaxy

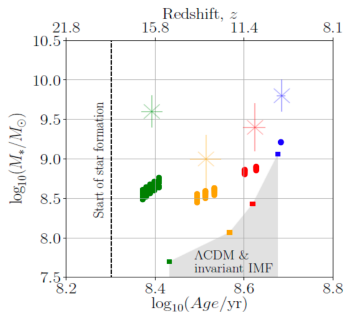


JWST and the conventional Λ CDM cosmology

Moritz Haslbauer et al, Has JWST already falsified dark-matter-driven galaxy formation? arXiv:2210.14915



Comparison of the size of the most massive galaxies, obtained in models of formation and growth of galaxies based on LCDM (colored dots) with JWST observations (black dots with errors) depending on the redshift of the observed galaxies.



Two orders of magnitude discrepancy.

Ultra-massive early QSO observed by ALMA

ALMA confirmation of an obscured hyper-luminous radio-loud AGN at $z = 6.853$ associated with a dusty starburst in the 1.5 deg² COSMOS field, R. Endsley et al, Monthly Notices of the Royal Astronomical Society, Volume 520, Issue 3, April 2023, Pages 4609–4620, Published: 24.02.2023

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Highly unexpected was abundance of IMBH which are appearing during last few years in huge numbers.

The assumption that (almost) all black holes in the universe are primordial strongly reduce or even eliminate the tension.

Predicted mass spectrum of PBH

The proposed mechanism is the first where inflation and Affleck-Dine baryogenesis are applied to PBH formation, repeated now in many works. The striking feature of it is the **log-normal** mass spectrum which is the only known spectrum **tested by "experiment" in a good agreement.**

$$\frac{dN}{dM} = \mu^2 \exp[-\gamma \ln^2(M/M_0)],$$

$M_0 \sim 10M_\odot$, is predicted, A.Dolgov, K.Postnov, "Why the mean mass of primordial black hole distribution is close to $10M_\odot$ ". JCAP 07 (2020) 063. The horizon mass at QCD p.t. is $10M_\odot$, for $\mu = 0$. At larger chemical potential the T_{pt} is smaller and M_{hor} is larger.

Gravitational waves from BH binaries

• GW discovery by LIGO strongly indicate that the sources of GW are PBHs. see e.g. S.Blinnikov, A.D., N.Porayko, K.Postnov, JCAP 1611 (2016), 036 "Solving puzzles of GW150914 by primordial black holes,"

1. Origin of heavy BHs ($\sim 30M_{\odot}$); there appeared much more striking problem of BH with $M \sim 100M_{\odot}$. See however, J. Ziegler, K. Freese, arXiv:2010.00254: DM annihilation inside stars
2. Formation of BH binaries from the original stellar binaries.
3. Low spins of the coalescing BHs .

To form so heavy BHs, the progenitors should have $M > 100M_{\odot}$. and a low metal abundance to avoid too much mass loss during the evolution. Such heavy stars might be present in young star-forming galaxies **but they are not observed in the necessary amount.** PBHs with the observed by LIGO masses may be easily created with sufficient density.

Chirp mass

Two rotating gravitationally bound massive bodies are known to emit **gravitational waves**. In quasi-stationary inspiral regime, the radius of the orbit and the rotation frequency are approximately constant and the GW frequency is twice the rotation frequency. **The luminosity of the GW radiation is:**

$$L = \frac{32}{5} m_{Pl}^2 \left(\frac{M_c \omega_{orb}}{m_{Pl}^2} \right)^{10/3},$$

where M_1 , M_2 are the masses of two bodies in the binary system and M_c is the so called chirp mass:

$$M_c = \frac{(M_1 M_2)^{3/5}}{(M_1 + M_2)^{1/5}},$$

and

$$\omega_{orb}^2 = \frac{M_1 + M_2}{m_{Pl}^2 R^3}.$$

Chirp mass distribution

A.D. Dolgov, A.G. Kuranov, N.A. Mitichkin, S. Porey, K.A. Postnov, O.S. Sazhina, I.V. Simkine [On mass distribution of coalescing black holes](#), JCAP 12 (2020) 017, e-Print: 2005.00892.

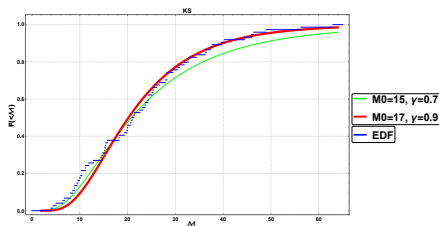
The available data on the chirp mass distribution of the black holes in the coalescing binaries in O1-O3 LIGO/Virgo runs are analyzed and compared with theoretical expectations based on the hypothesis that these black holes are primordial with log-normal mass spectrum.

The inferred best-fit mass spectrum parameters, $M_0 = 17M_\odot$ and $\gamma = 0.9$, fall within the theoretically expected range and shows excellent agreement with observations.

On the opposite, binary black hole formation based **on massive binary star evolution** require additional adjustments to reproduce the observed chirp mass distribution.

Chirp mass distribution

Model distribution $F_{PBH}(< M)$ with parameters $M_0 \approx 17M_\odot$ and $\gamma \sim 1$ for two best Kolmogorov-Smirnov tests. EDF= empirical distribution function.

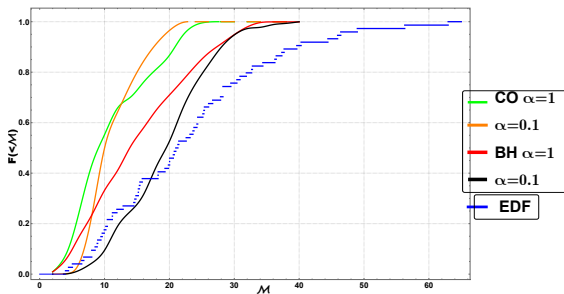


Similar value of the parameters are obtained in [M. Raidal et al, JCAP,2019. Feb. V. 2019, no. 2. P. 018. arXiv:1812.01930](#) and [L. Liu, et al arXiv:2210.16094](#).

See also [K. Postnov and N. Mitichkin, e-Print: 2302.06981](#).

Chirp mass distribution

Cumulative distributions $F(< M)$ for several **astrophysical** models of binary BH coalescences.



Conclusion: **PBHs with log-normal mass spectrum perfectly fit the data.**
Astrophysical BHs seem to be disfavoured.

PBH and inflation

In earlier works the predicted masses of PBH were quite low.

Inflation allows for formation of PBH with very large masses.

It was first applied to PBH production in DS paper,

a year later in: B.J. Carr, J.H. Hilbert, J.E. Lidsey, "Black hole relics and inflation: Limits on blue perturbation spectra", Phys.Rev.D 50 (1994) 4853, astro-ph/9405027;

and soon after in P. Ivanov, P. Naselsky, I. Novikov (May 10, 1994), Inflation and primordial black holes as dark matter, PRD 50 (1994) 7173.

Presently inflationary mechanism of PBH production is commonly used. It allows to create PBH with very high masses, but the predicted spectrum is multi-parameter one and quite complicated

The only exception is the log-normal spectrum of DS and DKK tested by observatons.

Black Dark Matter

The first suggestion PBH might be dark matter "particles" was made by S. Hawking in 1971 "Gravitationally collapsed objects of very low mass", *Mon. Not. R. astr. Soc.* (1971) 152, 75 and later by G. Chapline in 1975 who noticed that low mass PBHs might be abundant in the present-day universe with the density comparable to the density of dark matter. G.F. Chapline, *Nature*, 253, 251 (1975) "Cosmological effects of primordial black holes". Assumed flat mass spectrum in log interval:

$$dN = N_0(dM/M)$$

with maximum mass $M_{\max} \lesssim 10^{22}$ g, which hits the allowed mass range. The next one: DS (Mar 13, 1992), Baryon isocurvature fluctuations at small scales and **baryonic dark matter**, with more realistic masses. **first paper with inflation applied to PBH formation, so PBH masses as high as $10^6 M_{\odot}$, and even higher can be created, log-normal mass spectrum was predicted.**

Black Dark Matter

Constraints on PBHs - B.Carr, F. Kuhnel "Primordial Black Holes as Dark Matter: Recent Developments", arXiv:2006.02838, June 2020

Primordial black holes as dark matter candidates B. Carr, F. Kuhnel SciPost Phys.Lect.Notes 48 (2022), e-Print: 2110.02821 [astro-ph.CO]

For monochromatic mass spectrum of PBHs (model-dependent and have caveats).

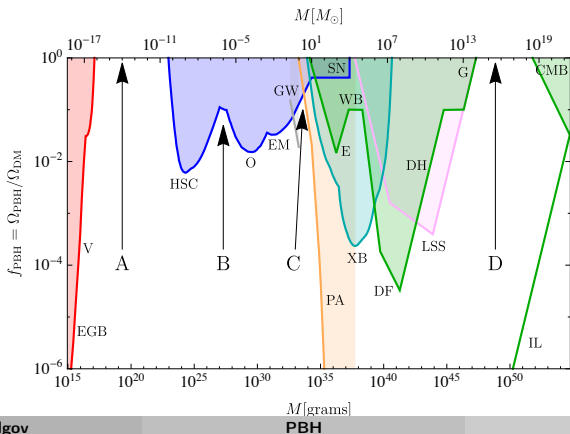


Figure caption

Constraints on $f(M)$ for a **monochromatic** mass function, from evaporations (red), lensing (blue), gravitational waves (GW) (gray), dynamical effects (green), accretion (light blue), CMB distortions (orange) and large-scale structure (purple). Evaporation limits from the extragalactic gamma-ray background (EGB), the Voyager positron flux (V) and annihilation-line radiation from the Galactic centre (GC). Lensing limits from microlensing of supernovae (SN) and of stars in M31 by Subaru (HSC), the Magellanic Clouds by EROS and MACHO (EM) and the Galactic bulge by OGLE (O). Dynamical limits from wide binaries (WB), star clusters in Eridanus II (E), halo dynamical friction (DF), galaxy tidal distortions (G), heating of stars in the Galactic disk (DH) and the CMB dipole (CMB). Large scale structure constraints(LSS). Accretion limits from X-ray binaries (XB) and Planck measurements of CMB distortions (PA). The incredulity limits (IL) correspond to one PBH per relevant environment (galaxy, cluster, Universe). **There are four mass windows (A, B, C, D) in which PBHs could have an appreciable density.**

BH clustering and DM

As is argued by S.G. Rubin, et al in "The Formation of Primary Galactic Nuclei during Phase Transitions in the Early Universe", Soviet Journal of Experimental and Theoretical Physics. 2001, V. 92, no. 6. 921. arXiv:hep-ph/0106187 **PBHs can be formed in clusters.**

Dynamical interactions in PBH clusters offers additional channel for the orbital energy dissipation thus increasing the merging rate of PBH binaries, and the **constraints on f_{PBH} obtained by assuming a homogeneous PBH space distribution can be weaker.** A recent analysis by Y. Eroshenko, V. Stasenko, "Gravitational waves from the merger of two primordial black hole clusters" arXiv:2302.05167 based on the PBH formation model M. Sasaki et al "Primordial Black Hole Scenario for the Gravitational-Wave Event GW150914", PRL. 2016. V. 117, no. 6. P. 061101, arXiv:1603.08338 and T. Nakamura, et al "Gravitational Waves from Coalescing Black Hole MACHO Binaries", ApJL 1997, V. 487, no. 2, P. L139, arXiv:astro-ph/9708060, **shows that even $f_{PBH} = 0.1 - 1$ is not excluded.**

BH clustering and DM

Surprisingly strong limits on BH number density in the Galaxy are presented in the paper P. Mroz, A. Udalski, M. K. Szymanski, *et al*, "No massive black holes in the Milky Way halo", *Nature* **632**, 749 (2024)

Based on the microlensing analysis the authors conclude that the compact objects with masses from $1.8 \times 10^{-4} M_{\odot}$ to $6.3 M_{\odot}$ could make not more than 1% into dark matter density and those with masses in the range $1.3 \times 10^{-5} M_{\odot} - 860 M_{\odot}$ less than 10%.

However, the bounds are obtained for delta-function mass spectrum, while the model predicts the log-normal one. For the latter the bounds are significantly weaker.

Moreover the clusterization of PBH, according to Y. Eroshenko, V. Stasenko, "Gravitational waves from the merger of two primordial black hole clusters", *Symmetry* **15**, 637 (2023), [arXiv:2302.05167 \[astro-ph.CO\]](https://arxiv.org/abs/2302.05167), absolutely excludes the bound of Mroz et al.

Intermediate summary and antimatter in the Galaxy

The mechanism of AD and DKK solves the problem of the observed population of the universe at high redshifts by SMBH (QSO), galaxies, SN, and of a large amount of dust.

The predicted log-normal spectrum of PBH is tested and confirmed by the observations (the only one existing in the literature).

The existence of IMBH in GCs is confirmed.

The crazy by-product of AD and DKK mechanism, namely prediction of antimatter in the Galaxy seems to come true as well.

Astronomical data of the several recent years present strong evidence in favour of noticeable antimatter population in our Galaxy including:

- Observation of gamma-rays with energy 0.511 MeV, which surely originate from electron-positron annihilation at rest.
- Very large flux of anti-helium nuclei, observed at AMS.
- **Several stars are found which produce excessive gamma-rays with energies of several hundred MeV which may be interpreted as indication that these stars consist of antimatter.**

Antimatter history

Search for galactic antimatter

B.P. Konstantinov, et al Cosmic Research, 4, 66 (1968);

B.P. Konstantinov, et al Bulletin of the Academy of Sciences of the USSR. Physical series, 33, No,11, 1820 (1969).

Antimatter int the univerese:

F. W. Stecker, et al Possible Evidence for the Existence of Antimatter on a Cosmological Scale in the Universe, Phys. Rev. Letters 27, 1469 (1971);

F. W. Stecker, Grand Unification and possible matter-antimatter domain structure in the the universe. Tenth Texas Symposium on Relativistic Astrophysics, p. 69 (1981),

Summary of the situation presented at 2002:

F. W. Stecker, "The Matter-Antimatter Asymmetry of the Universe (keynote address for XIVth Rencontres de Blois)" arXiv:hep-ph/0207323.

A.D. Dolgov, "Cosmological matter antimatter asymmetry and antimatter in the universe", keynote lecture at 14th Rencontres de Blois on Matter - Anti-matter Asymmetry • e-Print: hep- ph/0211260.

Antimatter history

Paul A.M. Dirac: “Theory of electrons and positrons”, Nobel Lecture, December 12, 1933: “It is quite possible that... these stars being built up mainly of positrons and negative protons. In fact, there may be half the stars of each kind. The two kinds of stars would both show exactly the same spectra, and there would be no way of distinguishing them by present astronomical methods.”

It seems that now we know ways to distinguish stars from an antistars by observations from the Earth. A.D. Dolgov, V.A. Novikov, M.I. Vysotsky, “How to see an antistar” JETP Lett. 98 (2013) 519, e-Print: 1309.2746

The spectra are not exactly the same, even if CPT is unbroken and the polarization of radiation could be a good indicator or the type of emitted neutrinos/antineutrinos from supernovae.

Antimatter history

Dirac was the second person to talk about antimatter. In 1898, 30 years before Dirac and one year after discovery of electron (J.J. Thomson, 1897) Arthur Schuster (another British physicist) conjectured that there might be other sign electricity, ANTIMATTER, and supposed that there might be entire solar systems, made of antimatter, INDISTINGUISHABLE from ours.

Schuster's wild guess: matter and antimatter are capable to annihilate and produce VAST energy.

He believed that they were gravitationally repulsive having negative mass. Two such objects on close contact should have vanishing mass!?

A. Schuster, Nature, 58 (1898) 367. Potential Matter. Holiday Dream.

"When the year's work is over and all sense of responsibility has left us, who has not occasionally set his fancy free to dream about the unknown, perhaps the unknowable?"

"Astronomy, the oldest and yet most juvenile of the sciences, may still have some surprises in store. May antimatter be commended to its case".

Antimatter in the Galaxy

Based on the conventional approach no antimatter object is expected to be in the Galaxy.

However, it was predicted in 1993 and elaborated in 2009 that noticeable amount of antimatter, even antistars might be in the Galaxy and in its halo:

A. Dolgov, J.Silk, PRD 47 (1993) 4244 "Baryon isocurvature fluctuations at small scale and baryonic dark matter."

A.Dolgov, M. Kawasaki, N. Kevlishvili, Nucl.Phys. B807 (2009) 229, "Inhomogeneous baryogenesis, **cosmic antimatter**, and dark matter".

Bounds on the density of galactic antistars are rather loose, because the annihilation proceeds only on the surface of antistars as analyzed in:

C.Bambi, A.D. Dolgov, "**Antimatter in the Milky Way**", Nucl.Phys.B 784 (2007) 132-150 • astro-ph/0702350,

A.D. Dolgov, S.I. Blinnikov, "**Stars and Black Holes from the very Early Universe**", Phys.Rev.D 89 (2014) 2, 021301 • 1309.3395,

S.I.Blinnikov, A.D., K.A.Postnov, "**Antimatter and antistars in the universe and in the Galaxy**", Phys.Rev.D 92 (2015) 023516 • 1409.5736.

Anti-evidence: cosmic positrons

Observation of intense 0.511 line, a proof of abundant positron population in the Galaxy. In the central region of the Galaxy electron–positron annihilation proceeds **at a surprisingly high rate**, creating the flux:

$$\Phi_{511 \text{ keV}} = 1.07 \pm 0.03 \cdot 10^{-3} \text{ photons cm}^{-2} \text{ s}^{-1}.$$

The width of the line is about 3 keV. Emission mostly goes from the Galactic bulge and at much lower level from the disk,

"Great Annihilator" in the Galactic bulge.

G. Weidenspointner *et al.*, *Astron. Astrophys.* **450**, 1013 (2006);

J. Knodlseder *et al.*, *Astron. Astrophys.* **441**, 513 (2005);

P. Jean *et al.*, *Astron. Astrophys.* **445**, 579 (2006).

Until recently the commonly accepted explanation was that e^+ are created in the strong magnetic fields of pulsars but the recent results of AMS probably exclude this mechanism, since the spectrum of \bar{p} and e^+ at high energies are identical. L'Aquila Joint Astroparticle Colloquium, 10th November, 2021 by S. Ting.

Anti-evidence: cosmic positrons

However, this conclusion is questioned in [astro-ph:1504.06472](#),
Signatures of a two million year old supernova in the spectra of cosmic ray
protons, antiprotons and positrons,
M. Kachelriess, A. Neronov, D.V. Semikoz, where it is shown that
these features are consistently explained by a nearby source which was
active ~ 2 Myr ago and has injected $(1 - 2) \times 10^{50}$ erg in cosmic rays.

Anti-evidence: cosmic antinuclei

Registration of anti-helium: In 2018 AMS-02 announced possible observation of six \overline{He}^3 and two \overline{He}^4 .

A. Choutko, AMS-02 Collaboration, "AMS Days at La Palma, La Palma, Canary Islands, Spain," (2018).

S. Ting, Latest Results from the AMS Experiment on the International Space Station. Colloquium at CERN, May, 2018.

Recent registration of more events L'Aquila Joint Astroparticle Colloquium, 10th November by S. Ting; and COSPAR 2022, 16-24 July:

7 \overline{D} ($\lesssim 15$ GeV) and 9 \overline{He} , (~ 50 GeV). **fraction $\overline{He}/He \sim 10^{-9}$, too high.**

Secondary creation of \overline{He}^4 is negligibly weak.

Nevertheless S. Ting expressed hope to observe \overline{Si} !!!

It is not excluded that the flux of anti-helium is even much higher because low energy \overline{He} may escape registration in AMS.

Deuterium/Helium problem

There is noticeable discrepancy between the large fraction of D with respect to He. In the case of the standard BBN this ratio should be smaller than unity, but the observed one is practically 1.

It is assumed that the abundances of D and He are determined by BBN with large β (or η). However if $\beta \sim 1$ there is no primordial D. On the other hand in our scenario formation of primordial elements takes place inside non-expanding compact stellar-like objects with fixed temperature. If the temperature is sufficiently high, this so called BBN may stop before abundant He formation with almost equal abundances of D and He. One can see that looking at abundances of light elements at a function of temperature. **Is it is so, antistars may have equal amount of \bar{D} and \bar{He} !!!**

Possible discovery of anti-stars in the Galaxy

S. Dupourqué, L. Tibaldo and P. von Ballmoos, Constraints on the antistar fraction in the Solar System neighborhood from the 10-year Fermi Large Area Telescope gamma-ray source catalog,
Phys Rev D.103.083016 103 (2021) 083016

We identify in the catalog 14 antistar candidates not associated with any objects belonging to established gamma-ray source classes and with a spectrum compatible with baryon-antibaryon annihilation.

Possible discovery of anti-stars in the Galaxy

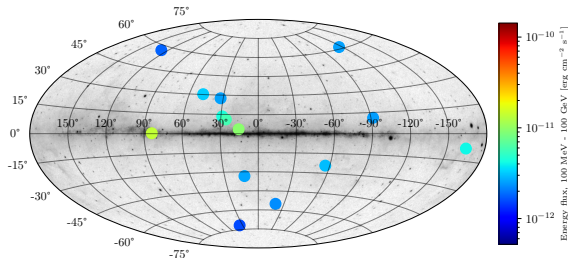


Figure: Positions and energy flux in the 100 MeV - 100 GeV range of antistar candidates selected in 4FGL-DR2. Galactic coordinates. The background image shows the Fermi 5-year all-sky photon counts above 1 GeV

X-ray signatures of antistars

X-ray signature of antistars in the Galaxy A.E. Bondar, S.I. Blinnikov, A.M. Bykov, A.D. Dolgov, K.A. Postnov e-Print: 2109.12699 [astro-ph.HE], JCAP, Sep 26, 2021,

In astrophysically plausible cases of the interaction of neutral atmospheres or winds from antistars with ionised interstellar gas, the hadronic annihilation **will be preceded by the formation of excited $p\bar{p}$ and $He\bar{p}$ atoms**. These atoms rapidly cascade down to low levels prior to annihilation giving rise to a series of narrow lines which can be associated with the hadronic annihilation gamma-ray emission. The most significant are L (3p-2p) 1.73 keV line (yield more than 90%) from $p\bar{p}$ atoms, and M (4-3) 4.86 keV (yield $\sim 60\%$) and L (3-2) 11.13 keV (yield about 25%) lines from $He^4\bar{p}$ atoms. These lines can be probed in dedicated observations by forthcoming sensitive X-ray spectroscopic missions XRISM and Athena and in wide-field X-ray surveys like SRG/eROSITA all-sky survey.

Antihelium and antistars

A.M. Bykov, K.A. Postnov, A.E. Bondar, S.I. Blinnikov, A.D. Dolgov, [Antistars as possible sources of antihelium cosmic rays, 2304.04623](#).

Possible sources of antinuclei in cosmic rays from antistars which are predicted in a modified Affleck-Dine baryogenesis scenario by DS (1993) are discussed. The expected fluxes and isotopic content of antinuclei in the GeV cosmic rays produced in scenarios involving antistars are estimated. It is shown that the flux of antihelium cosmic rays reported by the AMS-02 experiment can be explained by [Galactic anti-nova outbursts, thermonuclear anti-SN Ia explosions, a collection of flaring antistars, or an extragalactic source](#) with abundances not violating existing gamma-ray and microlensing constraints on the antistar population.

Gamma burster from star-antistar collision!?

Very powerful gamma-ray burster (GRB) was reported in M.E. Ravasio, O.S. Salafia, G. Oganessian, *et al*, SCIENCE, 25 Jul 2024, **385**, Issue 6707, p. 452; 2303.16223 [astro-ph.HE].

This event got the nickname: the Brightest Of All Time or the bf BOAT This extremely strong GRB occurred in October 2022. A bright megaelectronvolt emission line was observed, that appeared 280 seconds after the GRB began and then rapidly faded away while shifting to lower energies.

Usually the gamma-ray spectra of GRBs consist of a smooth continuum without absorption or emission lines. The authors interpret this line as having been produced by the annihilation of electron-positron pairs within the relativistic jet produced by the GRB possibly emerging from star-antistar annihilation !?

Star-antistar collision, was briefly mentioned in: A.D. Dolgov arXiv:2301.01365. Plenary talk at 36th Rencontres de Physique de la Vallée d'Aoste on Results and Perspectives in Particle Physics. It may be a quasi-periodic process of a star-antistar direct contact, explosion forcing them apart, and possible, but not necessary return to each other by gravitational attraction.

PBH Creation Mechanism

SUSY motivated baryogenesis, Affleck and Dine (AD).

SUSY predicts existence of scalars with $B \neq 0$. Such bosons may condense along flat directions of the quartic potential:

$$U_\lambda(\chi) = \lambda|\chi|^4 (1 - \cos 4\theta)$$

and of the mass term, $U_m = m^2\chi^2 + m^{*2}\chi^{*2}$:

$$U_m(\chi) = m^2|\chi|^2[1 - \cos(2\theta + 2\alpha)],$$

where $\chi = |\chi| \exp(i\theta)$ and $m = |m|e^\alpha$. If $\alpha \neq 0$, C and CP are broken.

In GUT SUSY baryonic number is naturally non-conserved - non-invariance of $U(\chi)$ w.r.t. phase rotation.

Creation Mechanism

Initially (after inflation) χ is away from origin and, when inflation is over, starts to evolve down to equilibrium point, $\chi = 0$, according to Newtonian mechanics:

$$\ddot{\chi} + 3H\dot{\chi} + U'(\chi) = 0.$$

Baryonic charge of χ :

$$B_\chi = \dot{\theta} |\chi|^2$$

is analogous to mechanical angular momentum. χ decays transferred baryonic charge to that of quarks in B-conserving process.

AD baryogenesis could lead to baryon asymmetry of order of unity, much larger than the observed 10^{-9} .

Creation Mechanism

If $m \neq 0$, the angular momentum, B , is generated by a different direction of the quartic and quadratic valleys at low χ . If CP-odd phase α is small but non-vanishing, both baryonic and antibaryonic domains might be formed with possible dominance of one of them.

Matter and antimatter objects may exist but globally $B \neq 0$.

Affleck-Dine field χ with CW potential coupled to inflaton Φ (AD and Silk; AD, Kawasaki, Kevlishvili):

$$U = g|\chi|^2(\Phi - \Phi_1)^2 + \lambda|\chi|^4 \ln\left(\frac{|\chi|^2}{\sigma^2}\right) + \lambda_1(\chi^4 + h.c.) + (m^2\chi^2 + h.c.).$$

Coupling to inflaton is the general renormalizable one.

When the window to the flat direction is open, near $\Phi = \Phi_1$, the field χ slowly diffuses to large value, according to quantum diffusion equation derived by Starobinsky, generalized to a complex field χ .

Creation Mechanism

If the window to flat direction, when $\Phi \approx \Phi_1$ is open only during a short period, cosmologically small but possibly astronomically large bubbles with high β could be created, occupying a small fraction of the universe, while the rest of the universe has normal $\beta \approx 6 \cdot 10^{-10}$, created by small χ . The mechanism of massive PBH formation quite different from all others. The fundament of PBH creation is build at inflation by making large isocurvature fluctuations at relatively small scales, with practically vanishing density perturbations.

Initial isocurvature perturbations are in chemical content of massless quarks.
Density perturbations are generated rather late after the QCD phase transition.
The mechanism is very much different from other conventional ones.

The emerging universe looks like a piece of Swiss cheese, where holes are high baryonic density objects occupying a minor fraction of the universe volume.

Results

- PBHs with log-normal mass spectrum - confirmed by the data!
- Compact stellar-like objects, as e.g. cores of red giants.
- Disperse hydrogen and helium clouds with (much) higher than average n_B density. Strange stars with unusual chemistry and velocity.
- β may be negative leading to creation of (compact?) antistars which could survive annihilation with the homogeneous baryonic background.
- Extremely old stars would exist even, "older than universe star" is found; the older age is mimicked by the unusual initial chemistry. Several such stars are observed.

The mechanism of PBH creation pretty well agrees with the data on the mass spectrum and on existence of antimatter in the Galaxy, especially of antistars. So we may expect that it indeed solves the problems created by HST and JWST.

The DS/DKK model removes the tension between Λ CDM cosmology and astronomical data during all the history of the universe and predicts and explains the origin of the observed antimatter population of the Milky Way