# Lunar Gravitational Wave Antenna

Jacopo Tissino Supervisor: Jan Harms

### **Detector concept**

- Inertial sensing of seismic motion
- The lunar seismic background is much quieter than the Earth's
- Permanently shadowed region at the pole,  $T \sim 40$  K
- Cryogenic operation at  $T \sim 4 K$
- **SQUID** or interferometric readout

# **Science targets**

- Multibanding for BNS (and BBH)
- Intermediate mass BBH (horizon at  $z\gtrsim 50$  for  $10^3M_\odot \lesssim M \lesssim 10^4M_\odot$ )
- Extreme and intermediate mass ratio inspirals
- Neutron star-white dwarf and double white dwarf binaries
- Lunar science: formation history, geologic models...

## Forecasting

The ingredients required to make a forecast:

- Displacement **sensitivity** estimates
- Lunar response to GWs
- Seismic **background** model
- Fisher matrix approach (assuming a matched-filtering search)
- Lunar motion in the Solar System

Work within this poster is preliminary: the LGWA collaboration is writing the whitepaper in these months. The mission is planned for the 2030s.



Extremely sensitive seismometers, deployed in a crater at the Moon's pole, for deci-Hertz gravitational wave science.





# Horizon for IMBH binaries





# Sensitivity curve and multibanding



## Payload



Horizon for equal-mass BBH