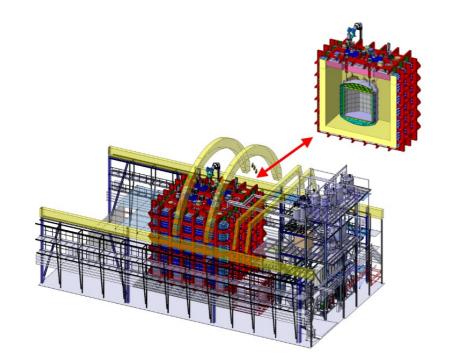




The DarkSide Program

(on behalf of the Global Argon Dark Matter Collaboration - GADMC) Science Fair @ GSSI – 06 of February 2024



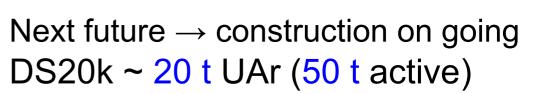


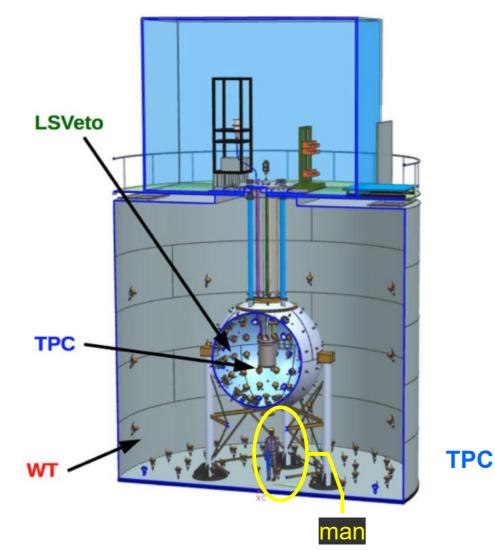
DarkSide Program: dark matter direct-detection experiments with double phase (U)Ar TPCs



The DS Program is staged with detectors of increasing mass.

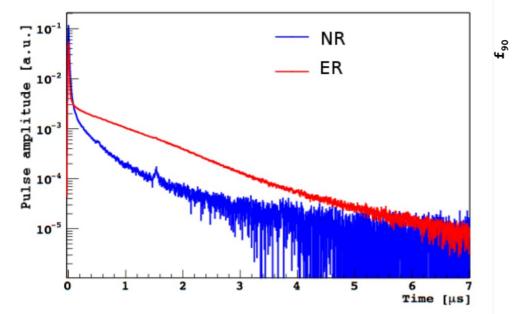
Last running experiment DarkSide-50 ~ 50 kg UAr





man

Why LAr? - Pulse Shape Discrimination (PSD) in LAr \rightarrow BG free WIMPs search

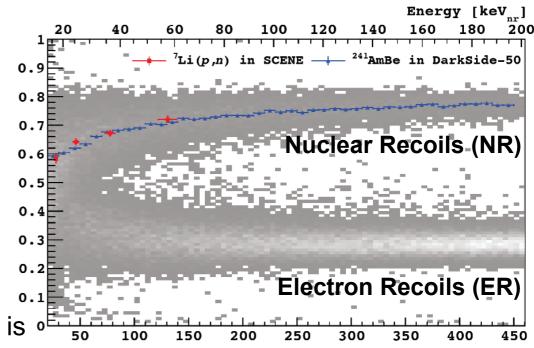


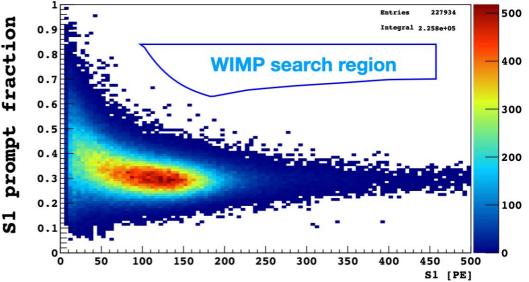
- The ratio of singlet/triplet excimer density is function on recoil type

- De-excitation: singlet 6 ns, triplet 1.5 us

$$f_{prompt} = \frac{\int_{t_0}^{t_{prompt}} S1(t)dt}{\int_{t_0}^{t_{tot}} S1(t)dt}$$

β/γ bg rejection up to factor 10⁹ demonstrated by DEAP-3600 *EPJC 81,823(2021)*



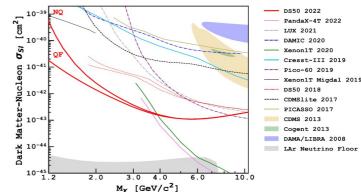


Still ongoing DarkSide50's data analysis - 2023 results

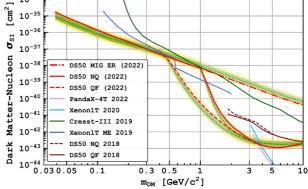
"Search for low-mass dark matter WIMPs with 12 ton-day exposure of DarkSide-50" PhysRevD.107.063001

- World's best limits for WIMPs with masses in the range (1.2 - 3.6) GeV/c²

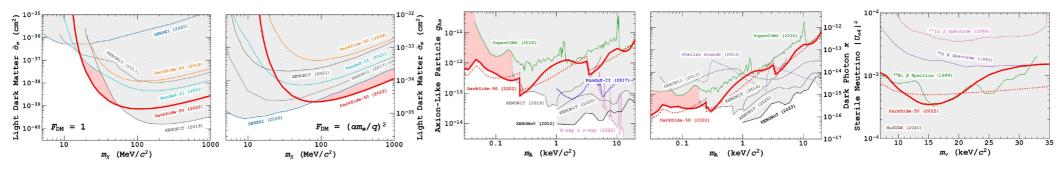
- One of the most downloaded Physical Review D papers of 2023



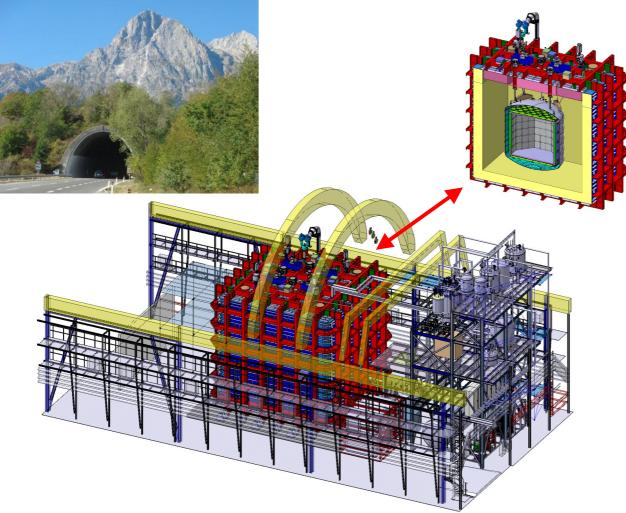
"Search for dark-matter--nucleon interactions via Migdal effect with DarkSide-50" PhysRevLett.130.101001



"Search for dark matter particle interactions with electron final states with DarkSide-50" PhysRevLett.130.101002



Next step → DarkSide-20k @ LNGS



- about 50 t of Underground Ar inside the TPC (20t fiducial)
- 10 years of foreseen activity
- Expected **0.1 background events in 200 t*yr exposure** [FV] → **still background free** [instrumental background]



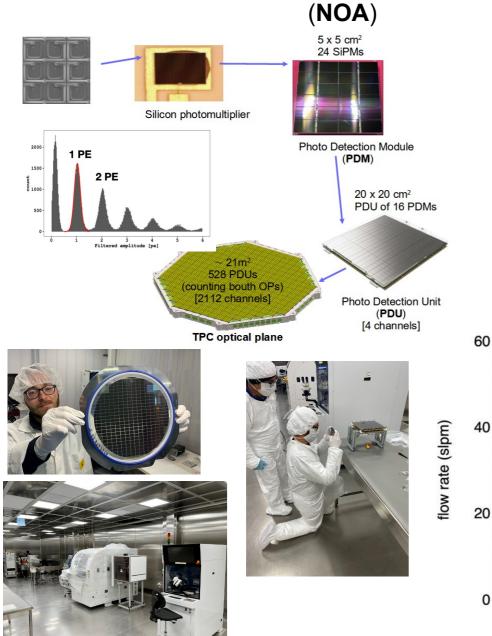
Construction on going



DarkSide @ GSSI - 1

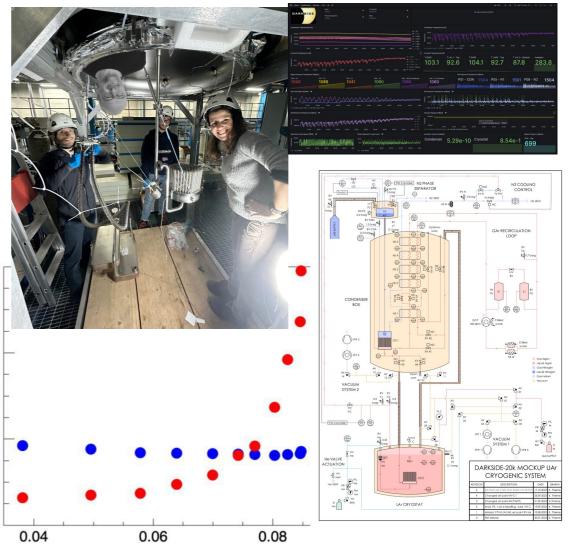
From PMTs to Silicon PhotoMultipliers

Photodetector production and characterisation @ Nuova Officina Assergi



UAr Cryogenics

Responsibilities in testing the DarkSide-20k UAr cryogenic system. Activities in Hall C - on going.



pressure difference (Bar)

0

DarkSide @ GSSI - 2 Analysis and sensitivity studies

Detector simulations

Leadership in simulation for:

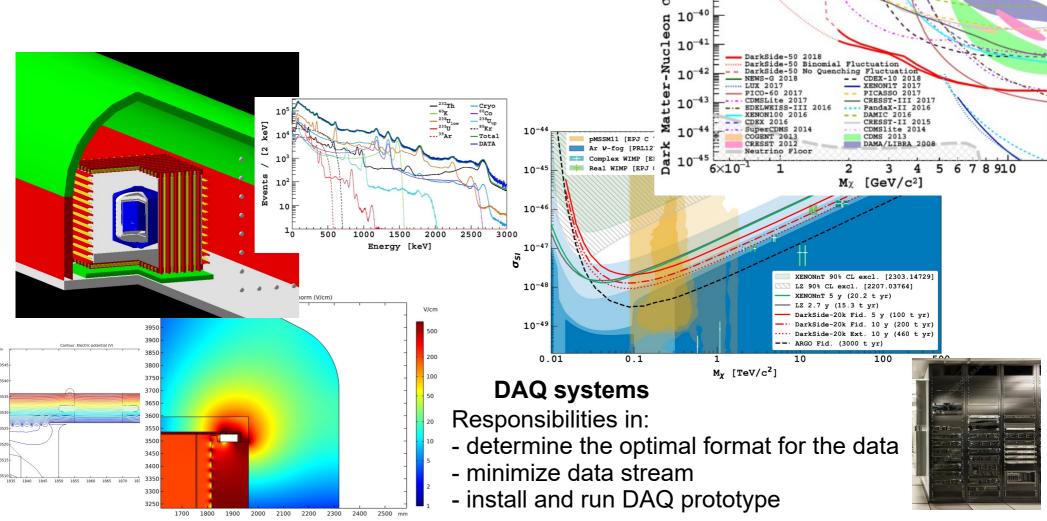
- Background levels (NR and ER) from the detector materials, the Hall C, cosmic rays
- Optimization of detector configuration
- Requirements on the DAQ
- Electric fields in the TPC

Leadership in analysis work for:

- DarkSide-50 data analysis
- Sensitivity studies

CH

— Performance of analysis and reconstruction algorithms



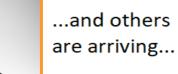
DarkSide @ GSSI - People











Mauro Caravati Pablo Kunzé Oscar Taborda Marek Walckzak Paolo Agnes



Backup

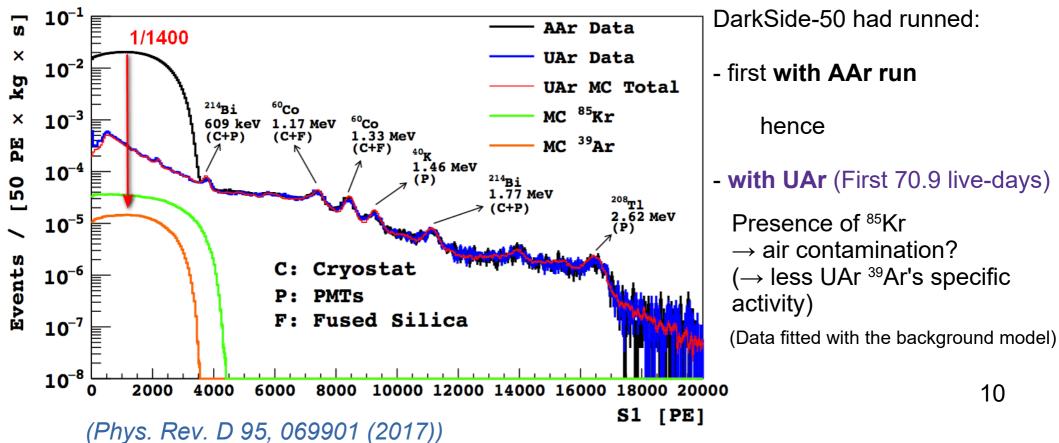
A key player: low radioactive Underground Argon (UAr)

- ³⁹Ar activity in atmospheric argon (AAr) is ~ 1 Bq/kg \rightarrow pileup problem w/ increasing mass target (β -decay; Q = 565 keV; t_{1/2} = 269 yr)
- ³⁹Ar activity sets the detection threshold at low energies (PSD is less effective or not usable)
- ³⁹Ar is produced by cosmic rays in the atmosphere

[as ${}^{37}Ar$: E.C. - (0.27 - 2.82) keV (L- K) -shell - $t_{1/2} \sim 35$ d]

 \rightarrow Argon from Underground (CO₂ well in Colorado)

DS50's UAr $\,^{39}\text{Ar}$ activity at 0.73 \pm 0.10 mBq/kg \Rightarrow ~1400 reduction factor wrt AAr



The Underground Argon journey in a nutshell

1) UAr extraction at the URANIA plant Industrial CO_2 extraction plant in Cortez (CO)

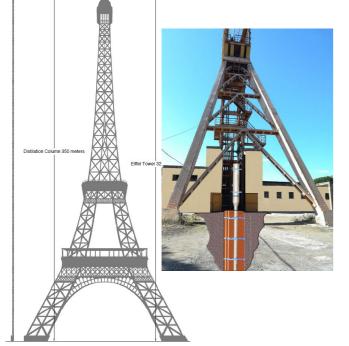
- Expected argon purity at outlet: 99.99%
- UAr extraction rate: 250-330 kg/day



3) Qualification at Canfranc (ES), DArT in ArDM

- Single-phase LAr detector
- Active volume ~1L
- Capable of measuring ³⁹Ar
 depletion factors of the order of
 1000 with 10% precision in weeks





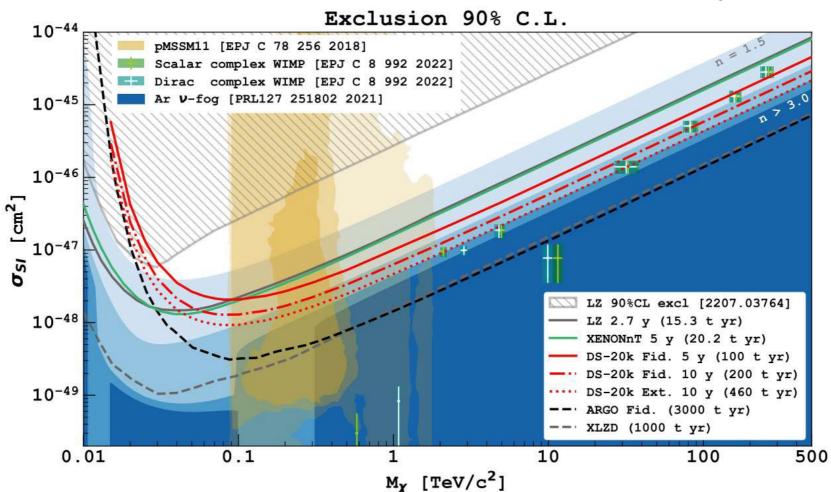
2) Cryogenic distillation at the ARIA facility

350 m tall criogenic distillation column installed in the shaft of a dismissed coal mine in Sardinia (It)

- Chemical purification rate 1 t/day

Prototype module actioned with nitrogen and argon showed the expected isotopic separation capability: 11 Expected a factor 10 of ³⁹Ar reduction per pass in the isotopic separation running configuration (~ some kg/day)

DarkSide-20k WIMP sensitivity



Sensitivity of DarkSide-20k to spin-independent WIMPs for different exposures, w/ or w/o fiducial cuts applied

With 20 tonnes fiducialization and 10 years running:

- Probing at 90% CL: $M_{\chi}~$ = 1 TeV/c² $~\sigma_{\text{SI}}\simeq 6.3~x~10^{\text{-}48}~cm^{2}$

Instrumental background < 0.1 neutrons in Rol (30~200 keV_{NR})
 in the neutrino fog