



Directional Dark Matter searches with the CYGNO/INITIUM detector

***3D tracking with the
CYGNO/INITIUM experiment***

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Ph.D. 3rd year report - Astroparticle Physics

Gran Sasso Science Institute - XXXVI cycle

October, 2023

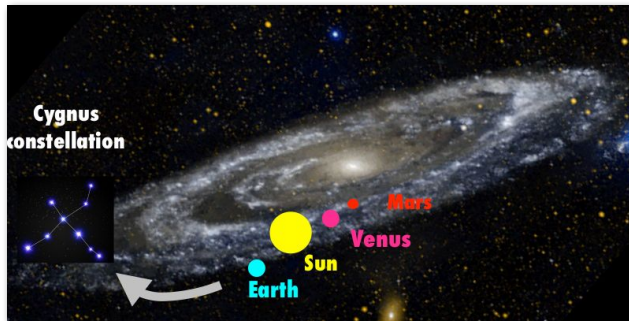
WIMPs - How to see them?

DM forms a halo within our galaxy.

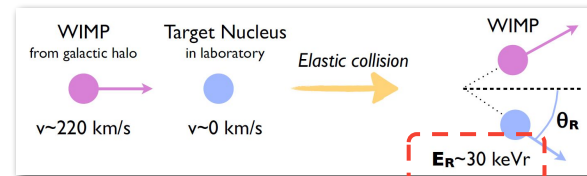
+

Solar system rotates around galaxy towards Cygnus constellation

Earth susceptible to an apparent WIMP wind from Cygnus direction!



...from WIMP scattering kinematics...

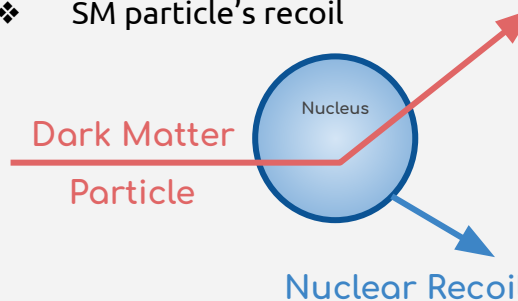


...the nuclear recoil is non-relativistic, of energies in the range 1 - 100 keV



Direct detection

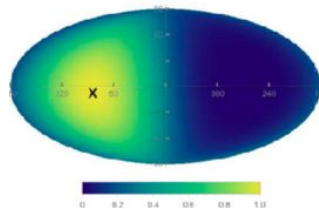
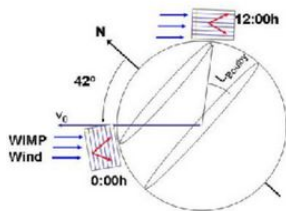
- ❖ $SM + \chi \rightarrow SM + \chi$
- ❖ SM particle's recoil



WIMPs - Directionality and beyond the neutrino floor

Exploring the **DIRECTION** dependency results in a characteristic effect - **anisotropy in the angular distribution of nuclear recoils**

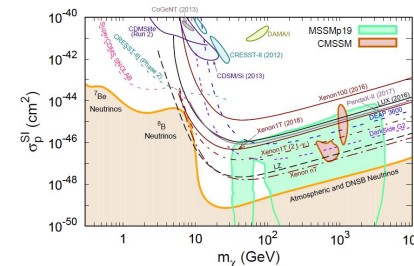
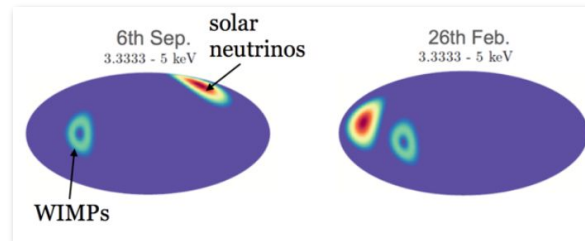
No background can mimic



The **CEvNS** produces NRs identical to the DM-induced ones. To **search DM at smaller cross-sections**, experiments need to **somehow venture into the neutrino fog**

Below $10 \text{ GeV}/c^2 \rightarrow$ Mostly **solar neutrinos**

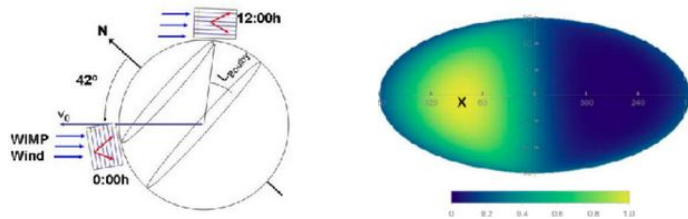
In galactic coords., the **Sun and Cygnus are never superimposed!**



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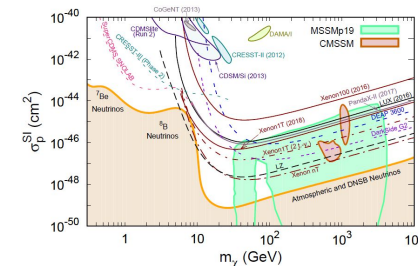
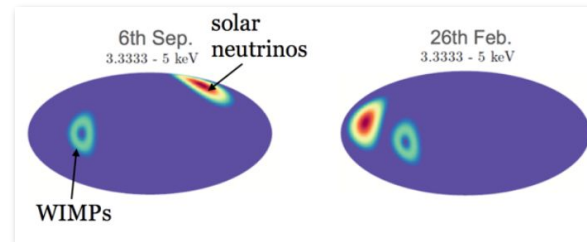


Where other experiments struggle to **prove / disprove DM**, **directional discrimination** strikes as the only way to **positively prove the existence of Dark Matter!**

The **CEvNS** produces NRs identical to the DM-induced ones. To **search DM at smaller cross-sections**, experiments need to **somehow venture into the neutrino fog**

Below $10 \text{ GeV}/c^2 \rightarrow$ Mostly **solar neutrinos**

In galactic coords., the **Sun and Cygnus are never superimposed!***



- Searching **beyond** the **neutrino floor**
- Properties of the **solar neutrino flux and DM halo**

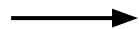


The logo for the CYGNO project. It features the word "CYGNO" in a bold, red, sans-serif font. The letter "Y" is replaced by a black silhouette of a bird in flight, with its wings spread wide.

A CYGNus tpc module
with Optical readout

CYGNO - What's the setup?

Time
Projection
Chamber



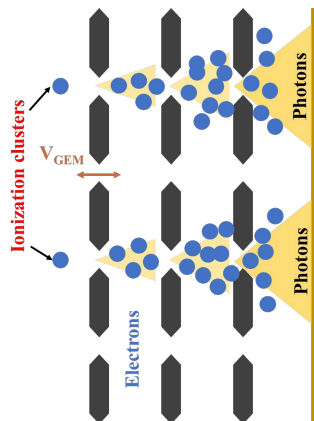
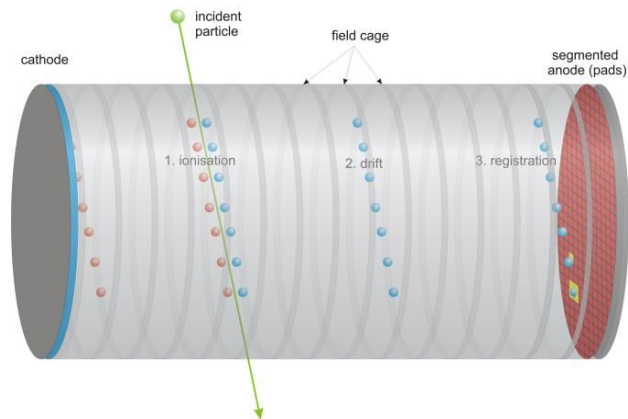
Triple GEM

Charge
amplification
& light production



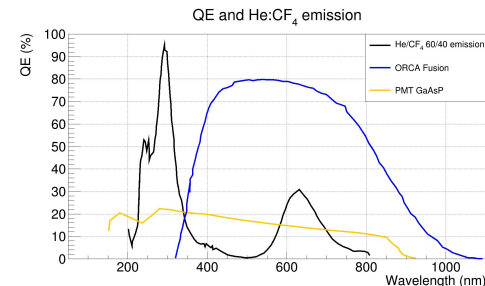
Camera & PMT

Optically read the **light produced by the de-excitation of the gas molecules** during electron multiplication.



Carbon tetrafluoride (CF₄)

→ Significant light yield at the camera's QE peak



CYGNO - What's the setup?

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Projection
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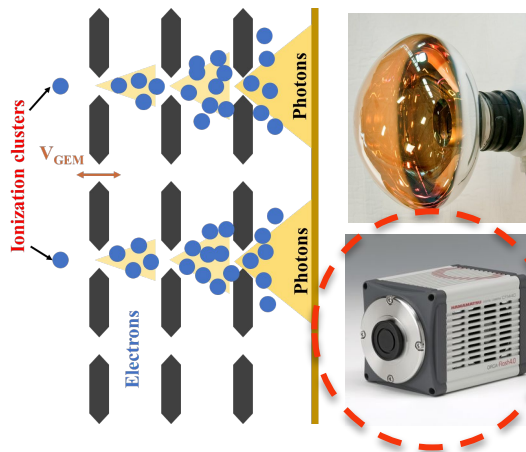
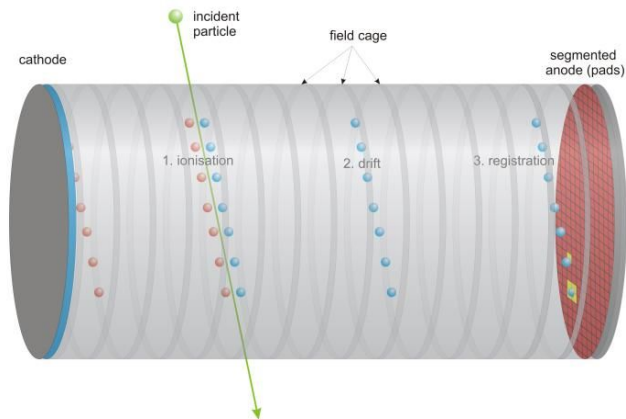
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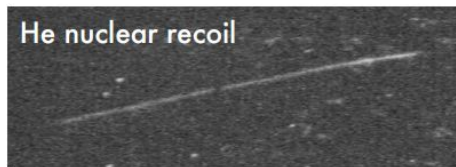


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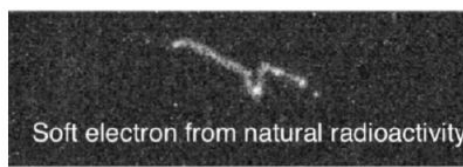
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Using the camera's high granularity, we can measure the **energy** & **X & Y coordinates**



He nuclear recoil



Soft electron from natural radioactivity

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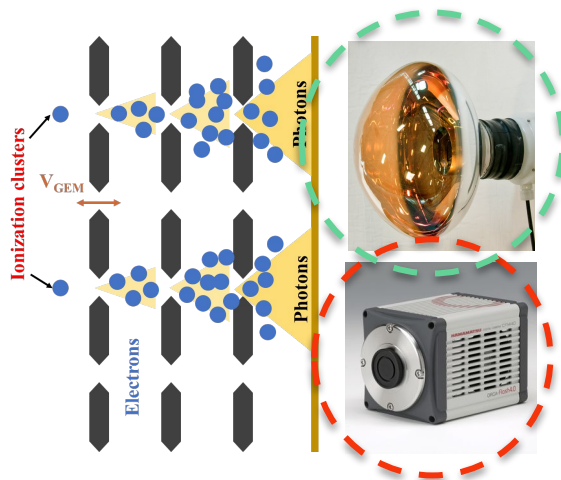
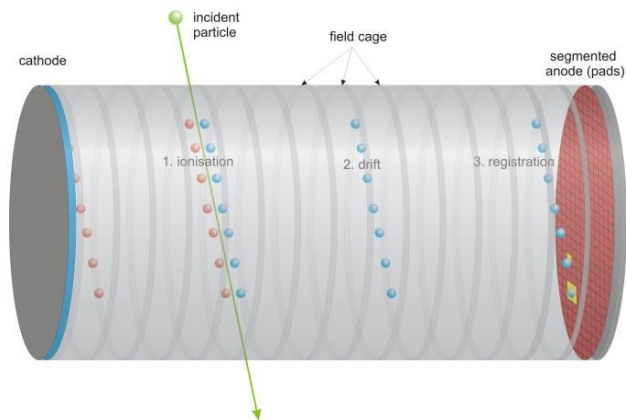
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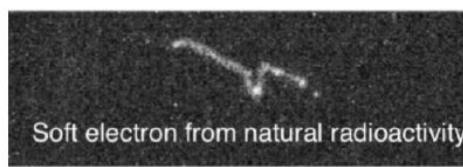
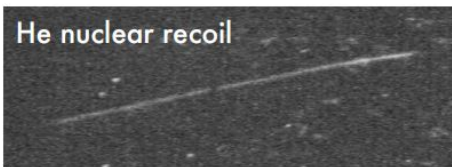
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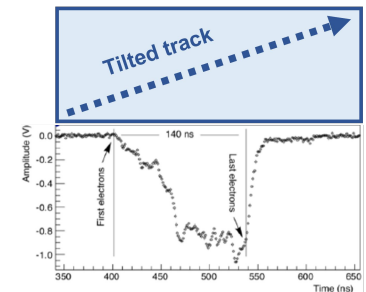
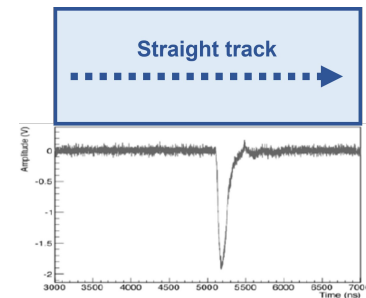
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1. Measure integrated energy.
2. Charge carriers' times of arrival → **dZ coordinate** (track's tilt)



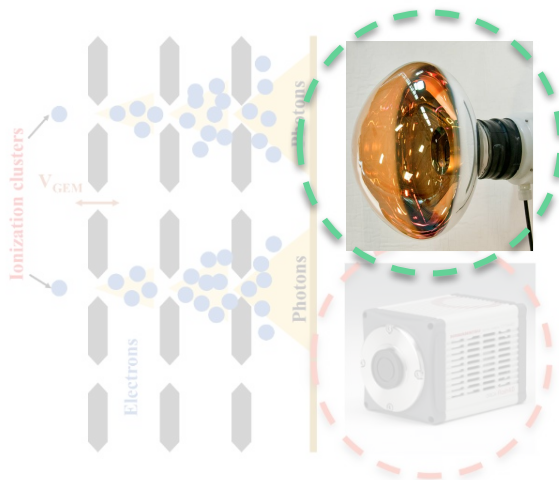
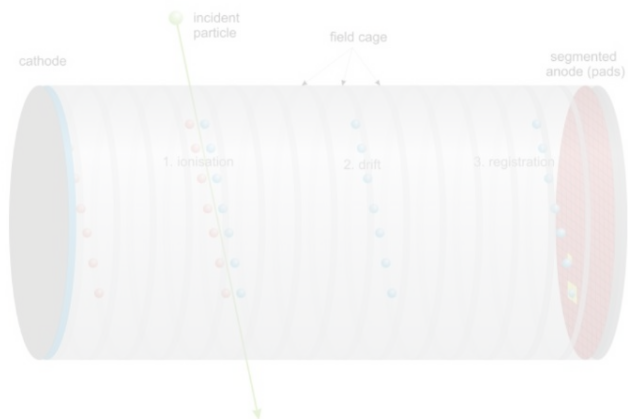
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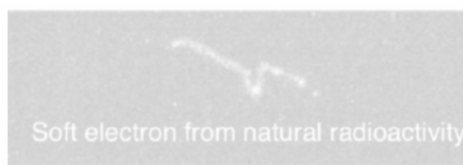
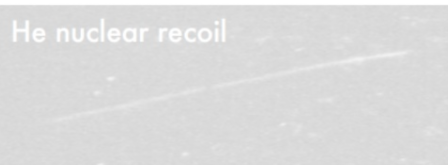
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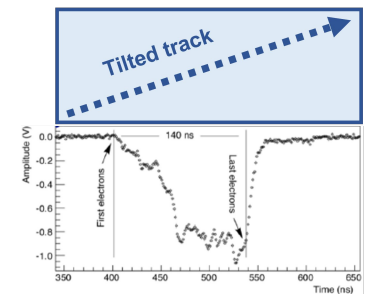
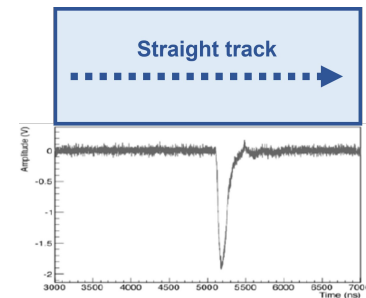
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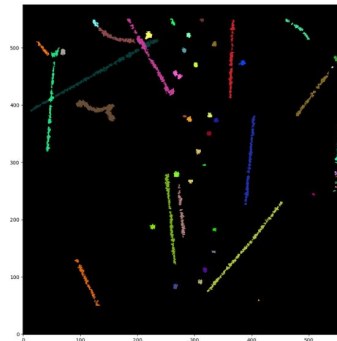
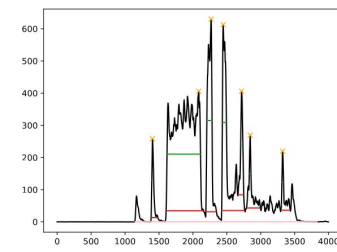
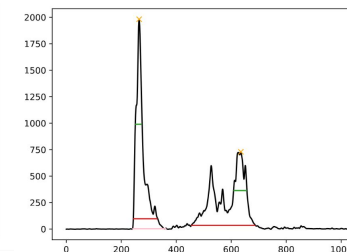
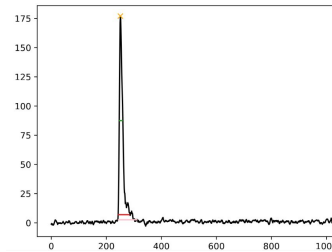
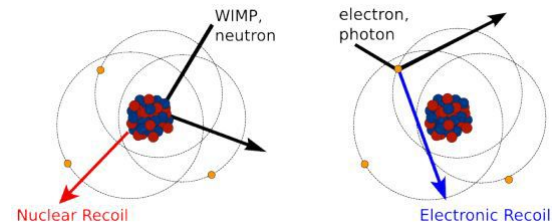
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A particle interacts inside our gas, He:CF₄ →
We look for the **ionization signal** → Camera pictures and **PMT waveforms** are recorded.

Reconstruction of variables of interest from PMT waveforms.

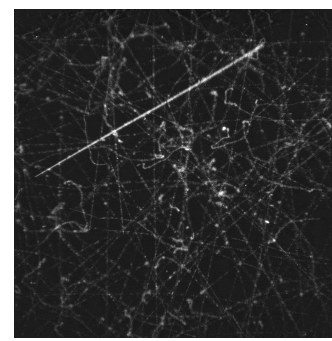
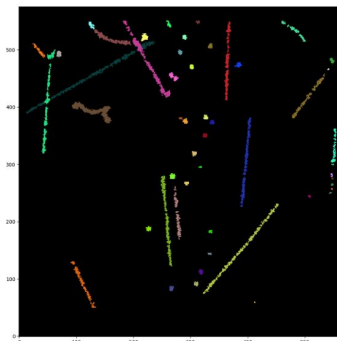
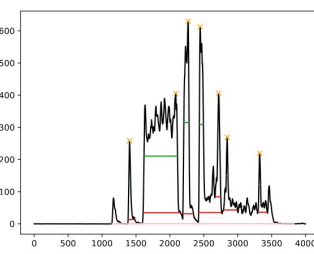
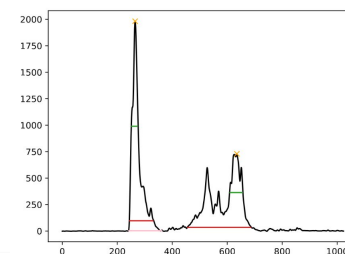
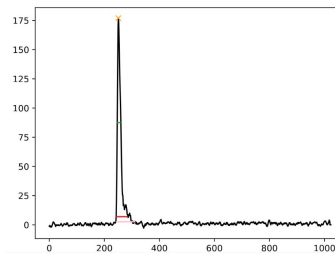
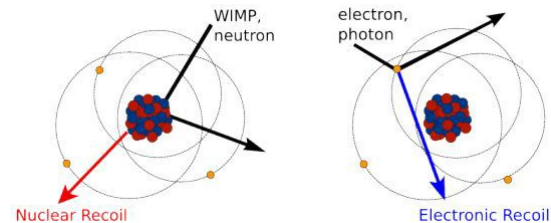
Merging with variables obtained with the sCMOS camera



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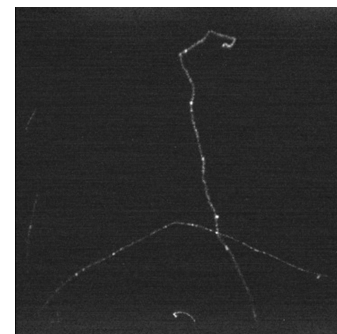
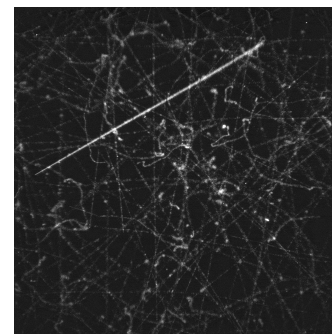
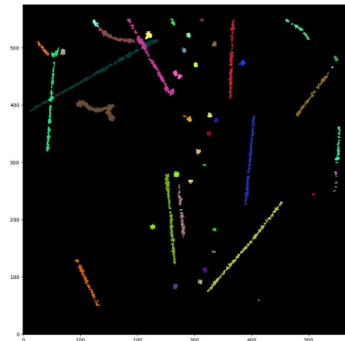
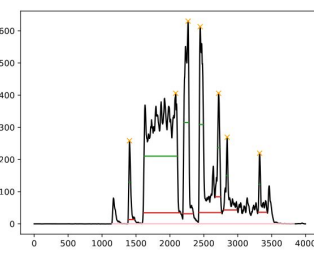
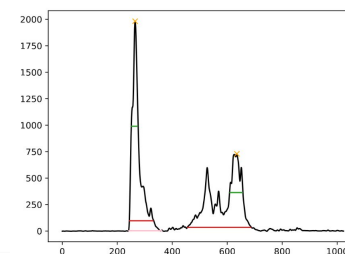
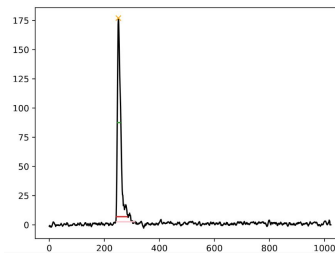
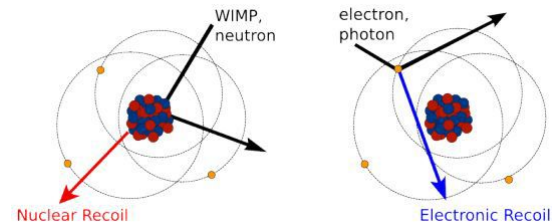
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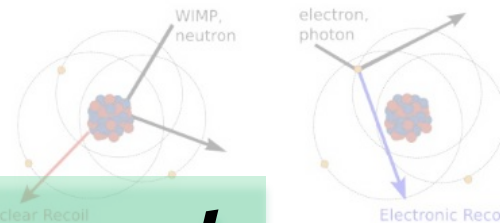
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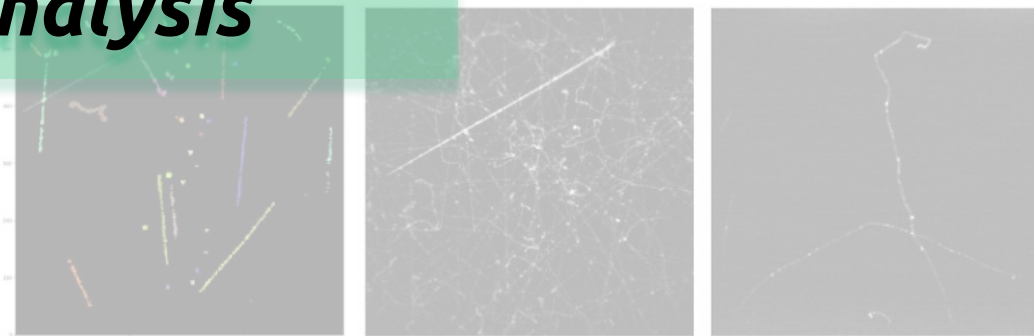
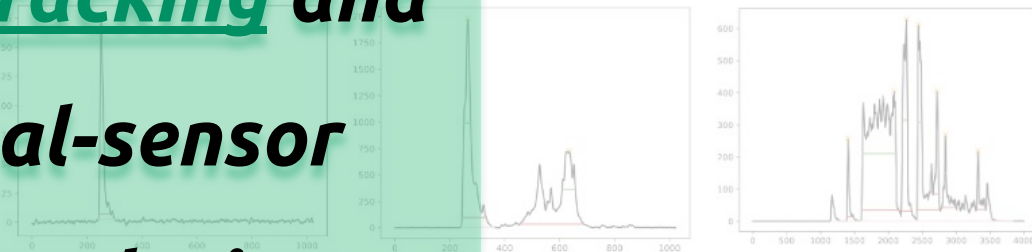
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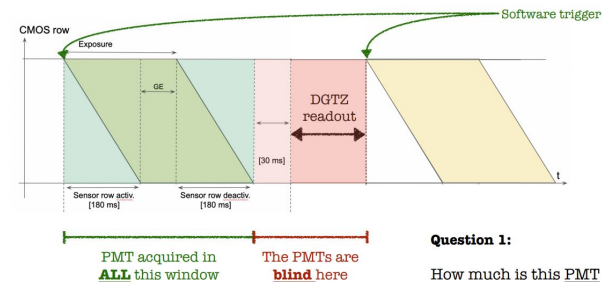
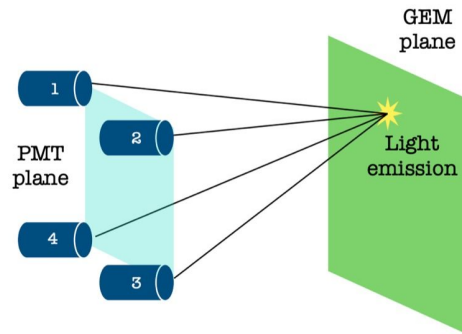
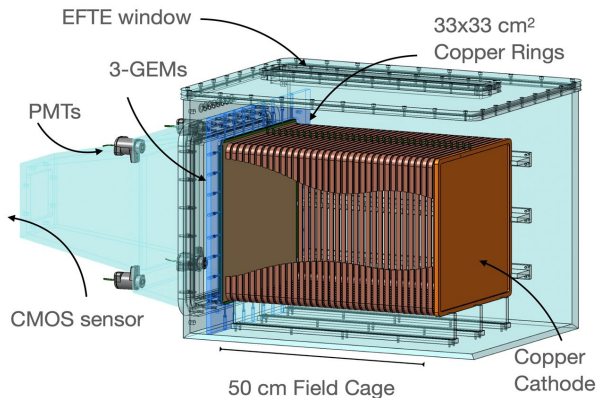
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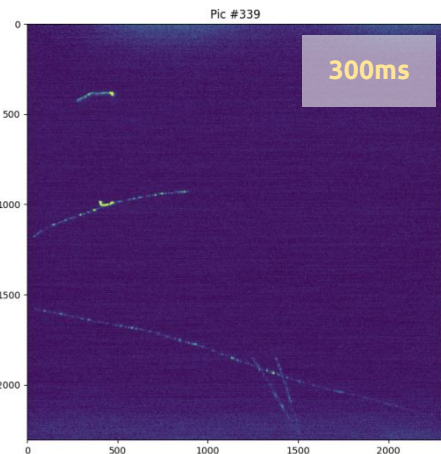
3D tracking and **dual-sensor** **analysis**



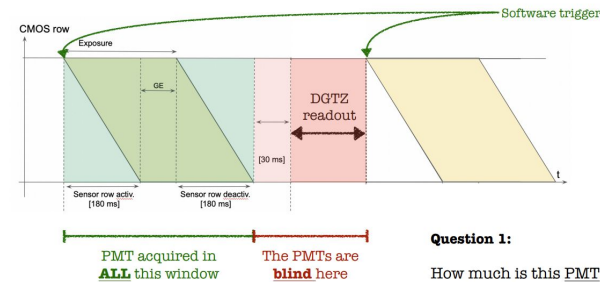
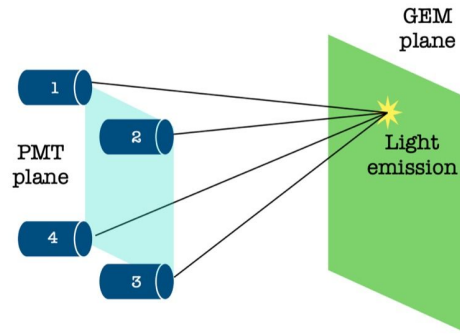
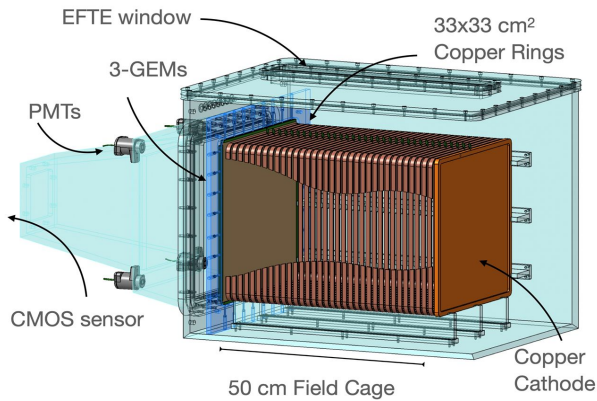
Ph.D. with CYGNO - How does LIME work?



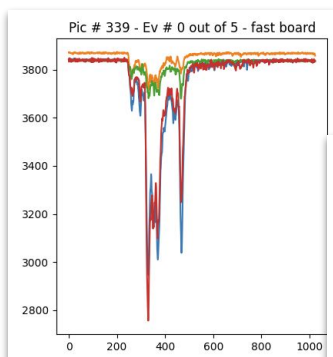
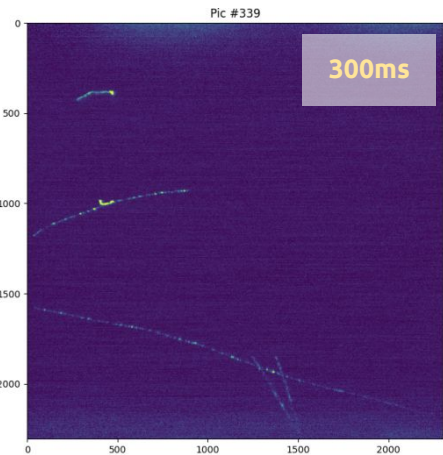
Question 1:
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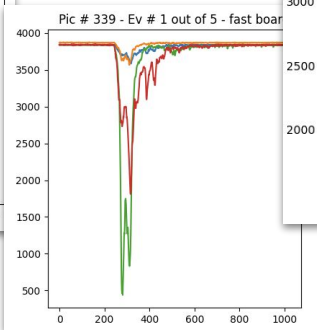
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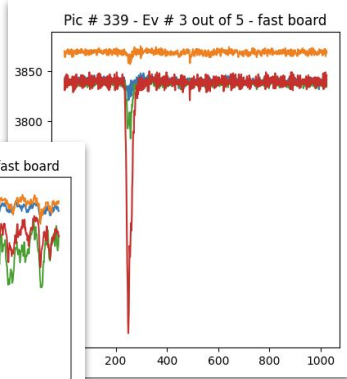
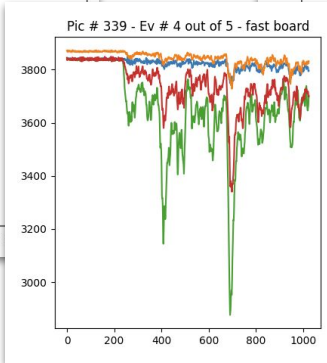
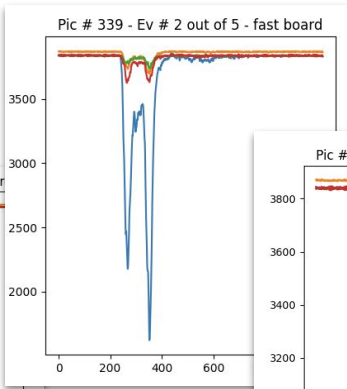
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1.5us
16us



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16us



1. PMT reconstruction

- a. Full integration of **waveforms information** within CYGNO's data framework

2. (L,x,y) Bayesian fit

- a. Simultaneously fits **ionization X - Y and Light**
 - i. This because the **light** seen by each PMT **changes with distance** ($L \propto 1/R^\alpha$)

3. Association of PMT & camera events

- a. A picture might contain **many events**. These must be **disentangled** for connection with PMT.
 - i. **Closest neighbour approach** under study with comparison with camera variables

4. 3D reconstruction

- a. Compute track angle / ΔZ
 - i. **Single peaks** vs. **Time over Threshold (ToT)**

5. Cross-analysis with camera's info.

- dE/dx
- NR vs ER
- PID

6. **Longitudinal diffusion** from ^{55}Fe spots **(+ NID)**

7. Fully or partial **integration of (x,y,L) fit** in PMT reconstruction.

8. **PMT Simulation**

9. Study of the **α parameter in $L \propto 1/R^\alpha$**

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- These are the **points of interest for my work**
- I have also **actively participated and chaired PMT meetings** where these topics are discussed.
- I have also met my co-workers in Rome Sapienza many times and well as **cross-checked and gave suggestions** in the results of other students

8. **PMT Simulation**

9. Study of the **α parameter in $L \propto 1/R^\alpha$**

□□ First CYGNO PMT analysis □□

1. PMT reconstruction → Full integration of **waveforms information** within CYGNO's data framework

→ Goal:

- ◆ Create a **TTree** with the PMT reconstructed variables.
- ◆ Run PMT reconstruction together with Camera reconstruction on the cloud.
- ◆ Create a **framework** that allows the *analysis of PMT events stand-alone and together with camera*.
 - (L,x,y), Z coordinate, Z diffusion, 3D reco, NR vs ER, PID, etc.

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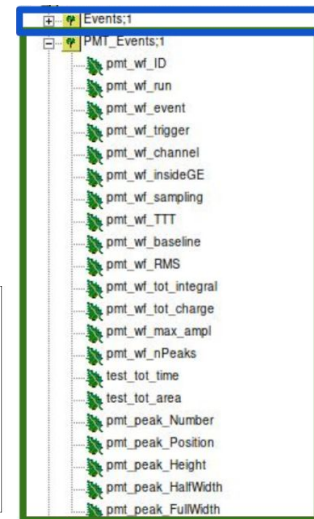
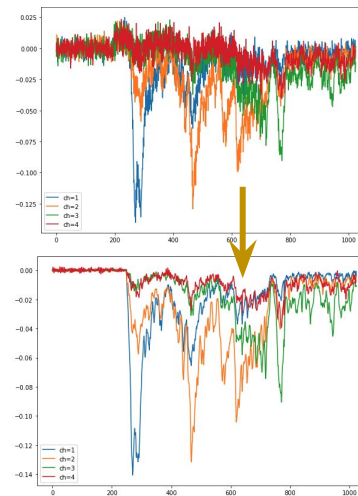
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→ Steps already taken:

- ◆ **Retrieve data** from DAQ through **cygno_libs**
 - All four PMTs and all triggers associated with a given picture.
 - Waveforms **automatically corrected**.
- ◆ Tree created with basic variables.
 - pmt_wf_[run/event/trigger/channel/sampling]
 - pmt_peak_[Position/Height/HalfWidth/FullWidth]
 - pmt_tot_[integral/charge]
 - pmt_max_ampl
 - pmt_nPeaks
 - pmt_[baseline/RMS]



- (L,x,y) Bayesian fit → Simultaneously fits **ionization X - Y and Light**
- Association of PMT & camera events
- Fully or partial **integration of (x,y,L) fit** in PMT reconstruction.

Method (1) 

Measure L_{i-4}
Infer x, y, L

Reconstruction performed with a Bayesian analysis:

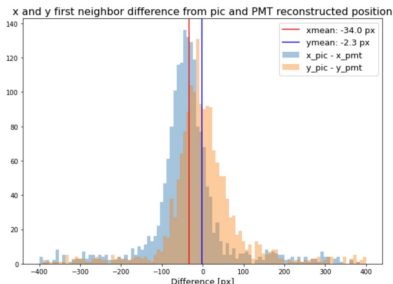
$$p(\theta|\{x_i\}) = \frac{p(\{x_i\}|\theta) \cdot \pi(\theta)}{p(\{x_i\})}$$

Labels: Likelihood, Prior pdf, Posterior pdf, Normalization factor

Where:
 $R_i = \sqrt{x_i^2 + y_i^2 + z_i^2}$
 $\mu_i = \frac{L}{R_i^4}$

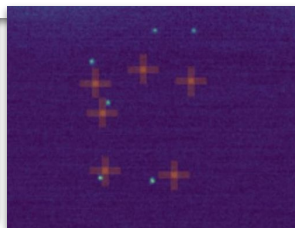
Likelihood: $p(\{x_i\}|\theta) = \prod_i \mathcal{N}(\{x_i\}|\mu_i(\theta))$

Closest neighbour (PMT waveform assigned to the closest cluster found)




• **Performance:**

- 37% within 1cm
- 73% within 2cm



- (L,x,y) Bayesian fit → Simultaneously fits **ionization X - Y and Light**
- Association of PMT & camera events
- Fully or partial **integration of (x,y,L) fit** in PMT reconstruction.

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Measure L_{1-4}
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$$p(\theta|\{x_i\}) = \frac{p(\{x_i\}|\theta) \cdot \pi(\theta)}{p(\{x_i\})}$$

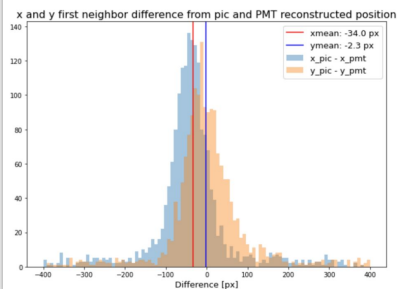
Labels: Likelihood, Prior pdf, Posterior pdf, Normalization factor

Where:
 $R_i = \sqrt{x_i^2 + y_i^2 + z_i^2}$

Likelihood: $p(\{x_i\}|\theta) = \prod_i \mathcal{N}(\{x_i\}|\mu_i(\theta)) \rightarrow \mu_i = \frac{L}{R_i^4}$

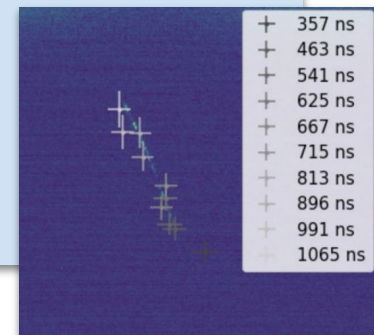
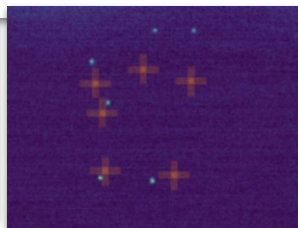
- Fit allows to retrieve event's **energy**
 - ◆ I used it to calculate dE/dx by merging it with it **Time-over-Threshold**
- Fit allows position comparison with camera picture
 - ◆ I'm working on a process to do the **1-to-1 coincidence** between ionization cluster and PMT waveform
 - Closest position neighbour
 - Closest energy neighbour
- Given it's importance I'm **currently integrating** of the fit in the overall CYGNO analysis framework.

Closest neighbour (PMT waveform assigned to the closest cluster found)



• **Performance:**

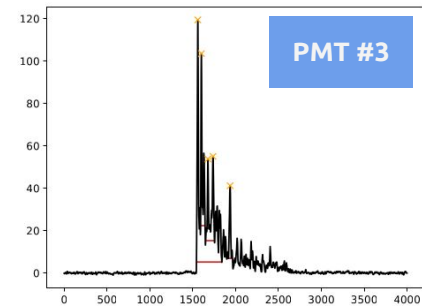
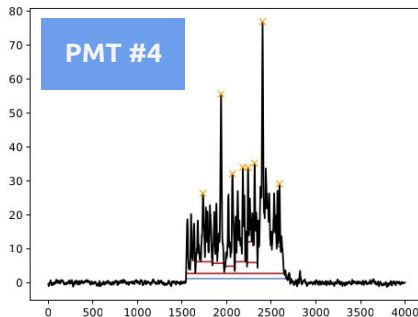
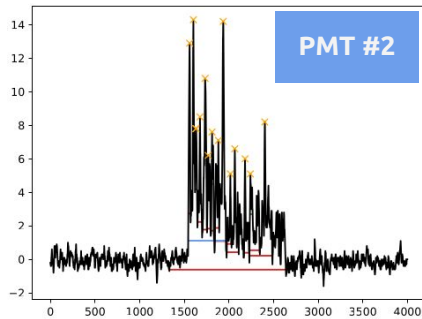
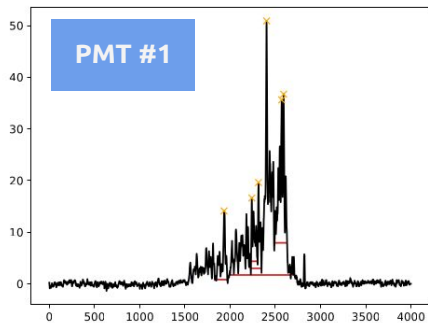
- 37% within 1cm
- 73% within 2cm



4. 3D reconstruction

9. Cross-analysis with camera's info

⇒ Time over threshold (TOT)



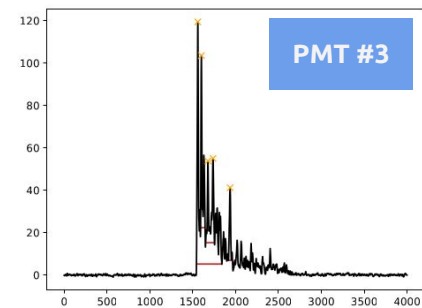
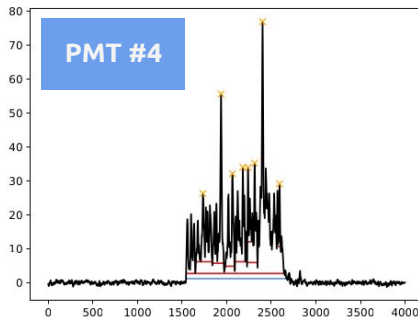
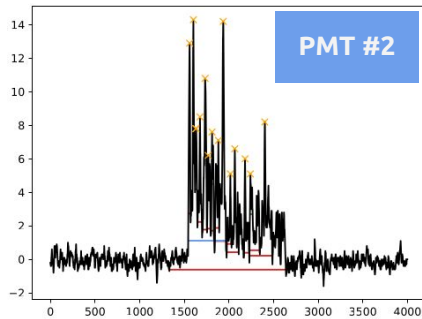
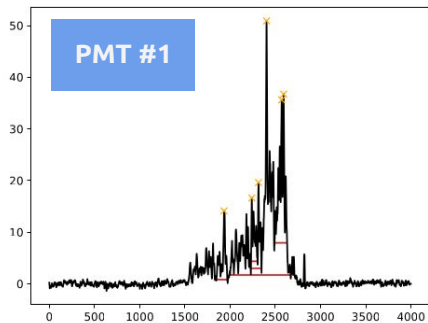
→ Measurement of the time length of the signal which is above a given threshold.

- ◆ Not trivial when each PMT sees a different signal intensity and tracks can have very complicated paths
- ◆ I do a **weighted average based on waveform's SNR** ⇒ *Only correct for timing purposes*

4. [3D reconstruction](#)

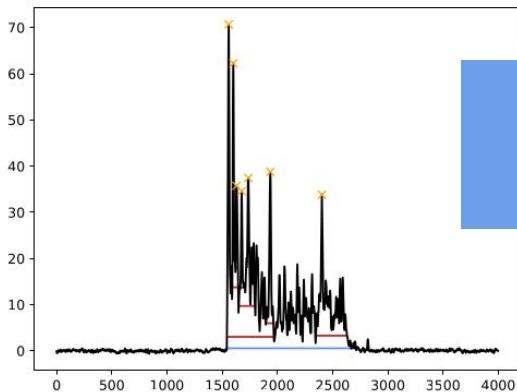
9. [Cross-analysis with camera's info](#)

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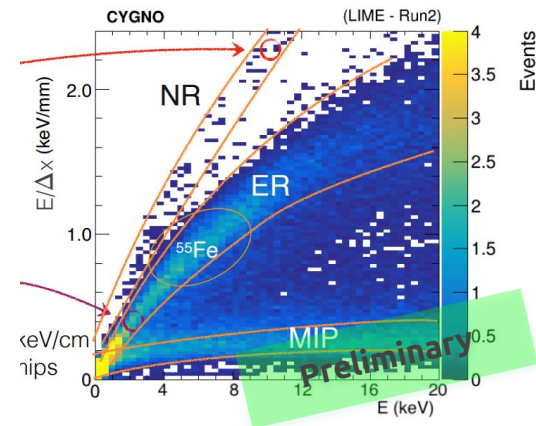
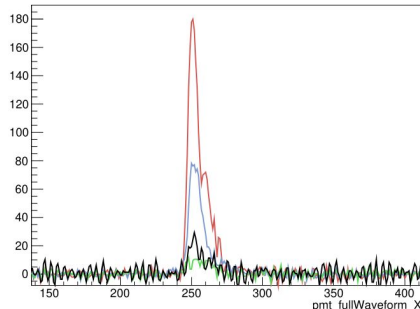
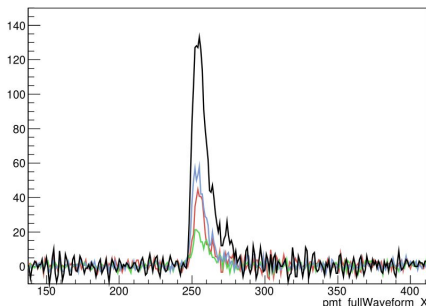
Weighted
average
waveform

4. [3D reconstruction](#)

9. [Cross-analysis with camera's info](#)

⇒ Time over threshold (TOT)

1. I tried to replicate sCMOS sensor **dE/dx vs. E** plot
 - a. Each PMT sees a **different intensity** but approximately the same time extension ⇒ Results in a different dE/dx.
 - i. Solved by using Bayesian fit **L (fitted luminosity)** of a given trigger mixing the 4 PMT signals.



For the ***length*** I used the ***Time over Threshold***

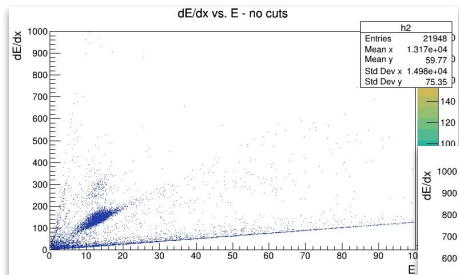


Thus, in fact I'm plotting the ***dE/dz***

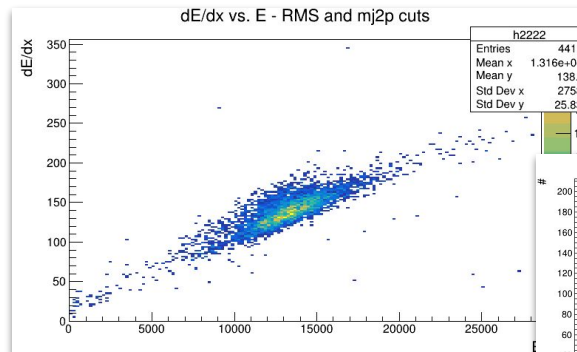
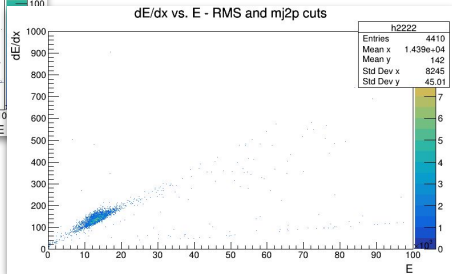
- 4. [3D reconstruction](#)
- 9. [Cross-analysis with camera's info](#)

⇒ Time over threshold (TOT)

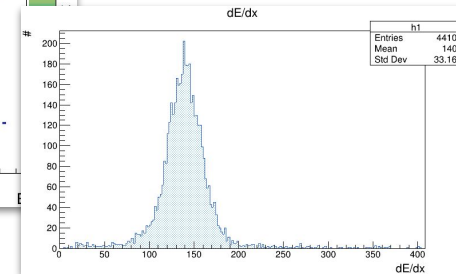
1. I calculated the dE/dx of calibration runs with ^{55}Fe
2. This is calculated using the the **fitted L** and **time over threshold**
3. Different **cuts** were studies: **RMS** and **mj2p**



Applying cuts



Zoomed in

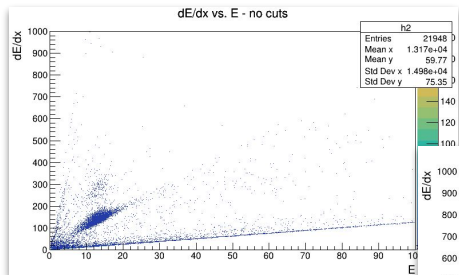
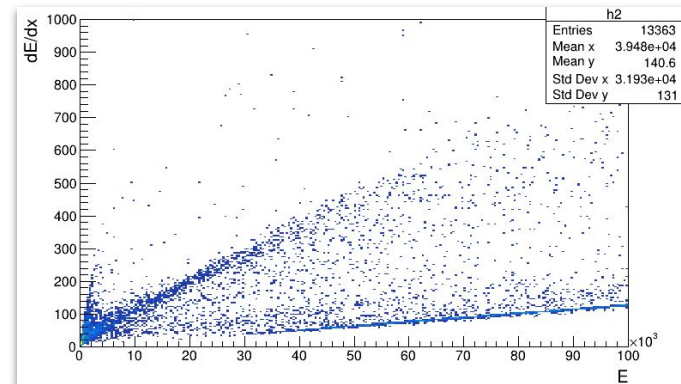


- 4. [3D reconstruction](#)
- 9. [Cross-analysis with camera's info](#)

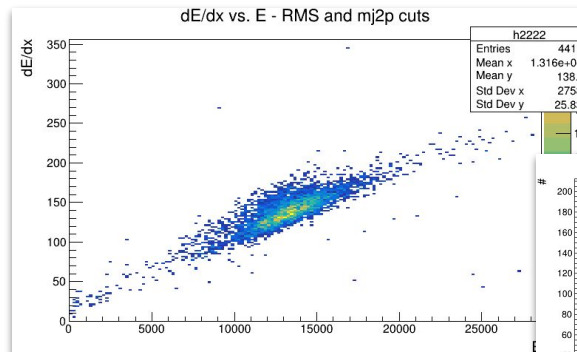
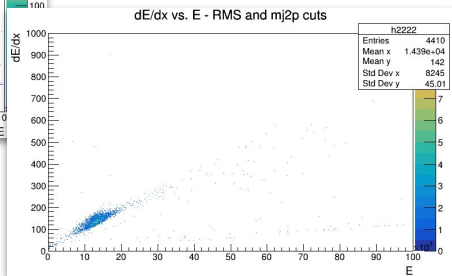
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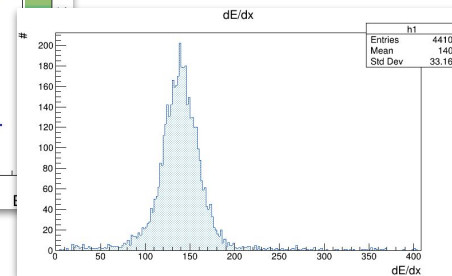
Comparing to
BG-only run



Applying cuts



Zoomed in

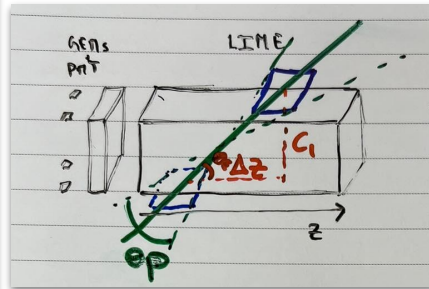
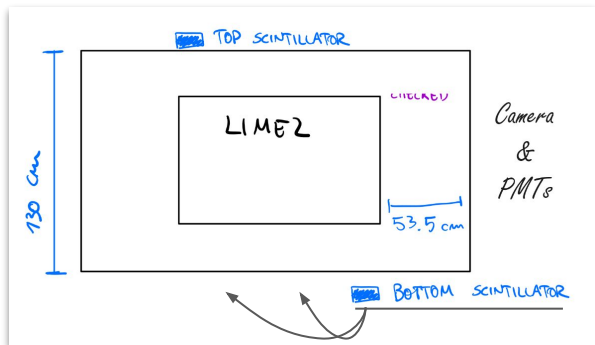


- 4. [3D reconstruction](#)
- 9. [Cross-analysis with camera's info](#)

⇒ Next studies ⇒ Tilted cosmic rays analysis

Setup:

- Two scintillator bars are placed on top and bottom of LIME
- LIME DAQ triggered by coincidence of two scintillators
 - ◆ Only certain angles (by geometry) are possible
 - ◆ 3 different scintillator position were used

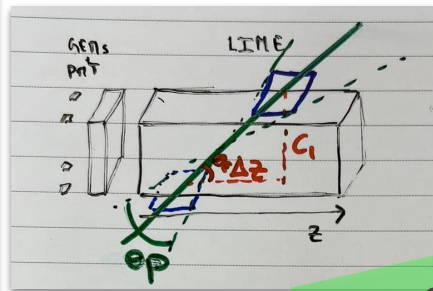
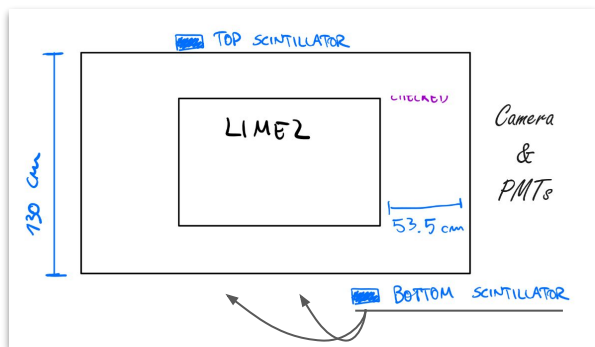


4. [3D reconstruction](#)
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 - ◆ Only certain angles (by geometry) are possible
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Ongoing...

Motivation & Math:

- *This measurement presents a clear dataset with track with well-defined orientation and energy deposit (MIP)*
 - We have a given **range of possible angles** of entering LIME (given by geometry of LIME + scintillators)
 - We can **superimpose** it with the **cosmic muons angle distribution** at ground ($\propto \cos^2(\Theta)^*$) to get the **theoretical angle distribution** (Θ_{teo}).
 - PMT measures the Time over Threshold
 - Multiplied by $v_{drift e^-}$ gives the Δz
 - Height of LIME (c_1) is known (33 cm)
 - The tracks inclination (α) will be $\tan^{-1}(c_1/\Delta z)$
 - We can compare Θ_{teo} with α .
 - **First CYGNO 3D analysis (on a distribution)**
 - **Gives us a measurement of PMT Reco / ToT efficiency**

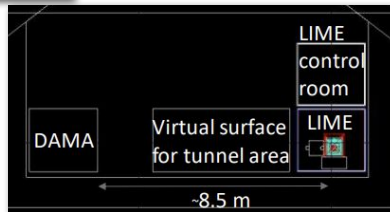
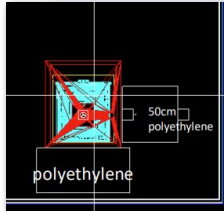
*<https://dx.doi.org/10.1088/1475-7516/2023/04/025>

- 4. [3D reconstruction](#)
- 9. [Cross-analysis with camera's info](#)

⇒ Next studies ⇒ Nuclear recoil data set

Setup:

- *AmBe* source placed near LIME
- *LIME was shielded* with Polyethylene, water and copper to block external radiation, thus creating a *clean dataset*

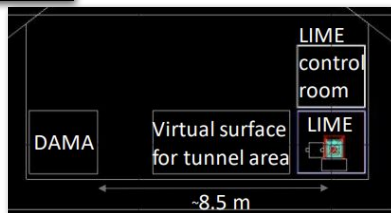
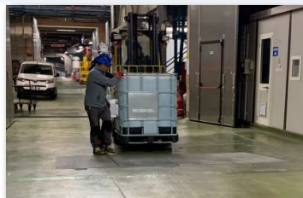
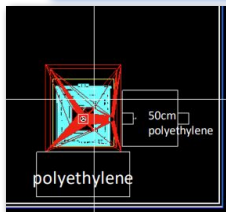


- 4. [3D reconstruction](#)
- 9. [Cross-analysis with camera's info](#)

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Setup:

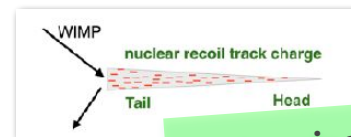
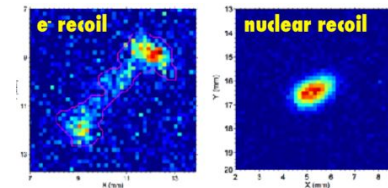
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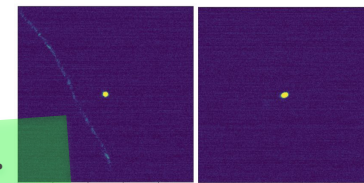
Motivation & Expectations:

- Clean **dataset of NR** obtained with a AmBe source → *WIMP-like signals*
 - 3D from ToT since they're straight tracks
 - Dual-sensor analysis (sCMOS + PMT)
 - dE/dx analysis → PID for NR vs. ER
 - Track sense/direction from head-tail asymmetry

■ **Paramount for CYGNO directional searches!**



Ongoing...



Parallel work with MANGO



a Multipurpose Apparatus for Negative ion studies with GEM Optical readout

Parallel work with MANGO



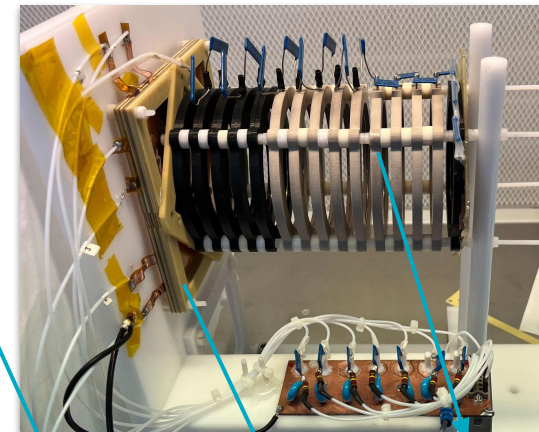
a Multipurpose Apparatus for Negative ion studies with GEM Optical readout

→ Enhancement of light yield

- Different types/configuration of **GEMs**
- ITO vs. Mesh

→ Negative Ion Drift

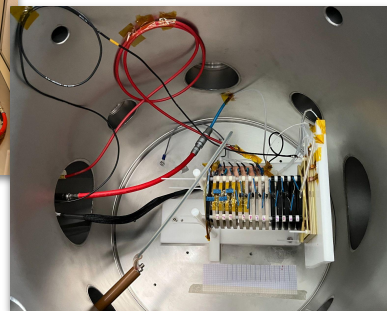
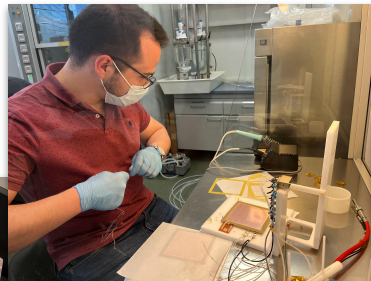
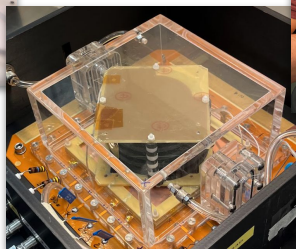
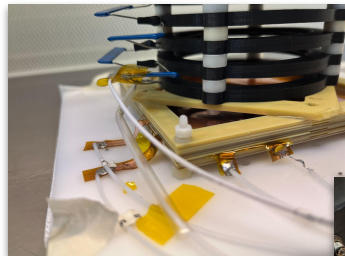
- Optimal **gas pressure, composition & amplification** config.
- **Transverse & Longitudinal Diffusion**
- **Ion mobility**



15 cm drift field cage

3 GEM stack:
50 μm thick
140 μm pitch
 \varnothing 70 μm holes

GEMs facing **SCMOS** and **PMT**



Parallel work with MANGO



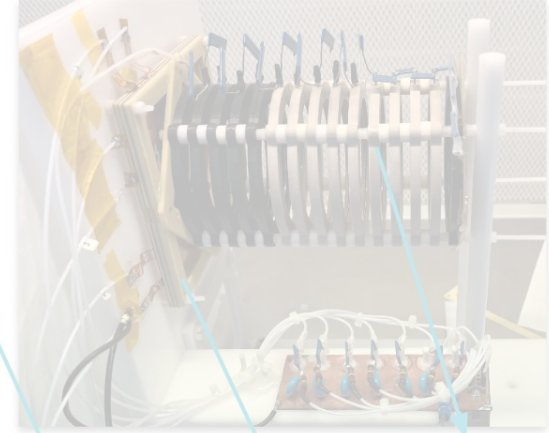
a Multipurpose Apparatus for Negative ion studies with GEM Optical readout

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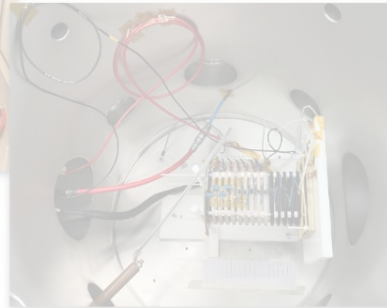
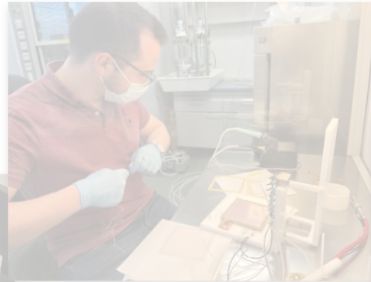
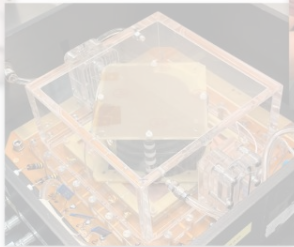
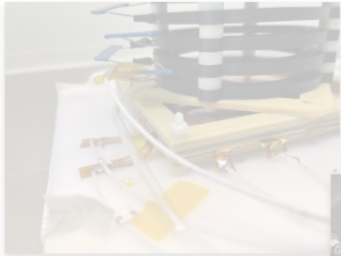
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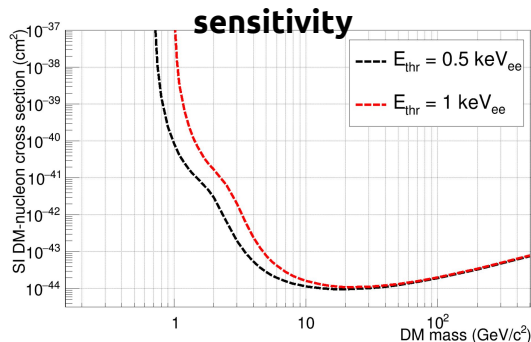
3 GEM stack:
50 μm thick
140 μm pitch
 $\varnothing 70 \mu\text{m}$ holes

GEMs facing **sCMOS** and **PMT**



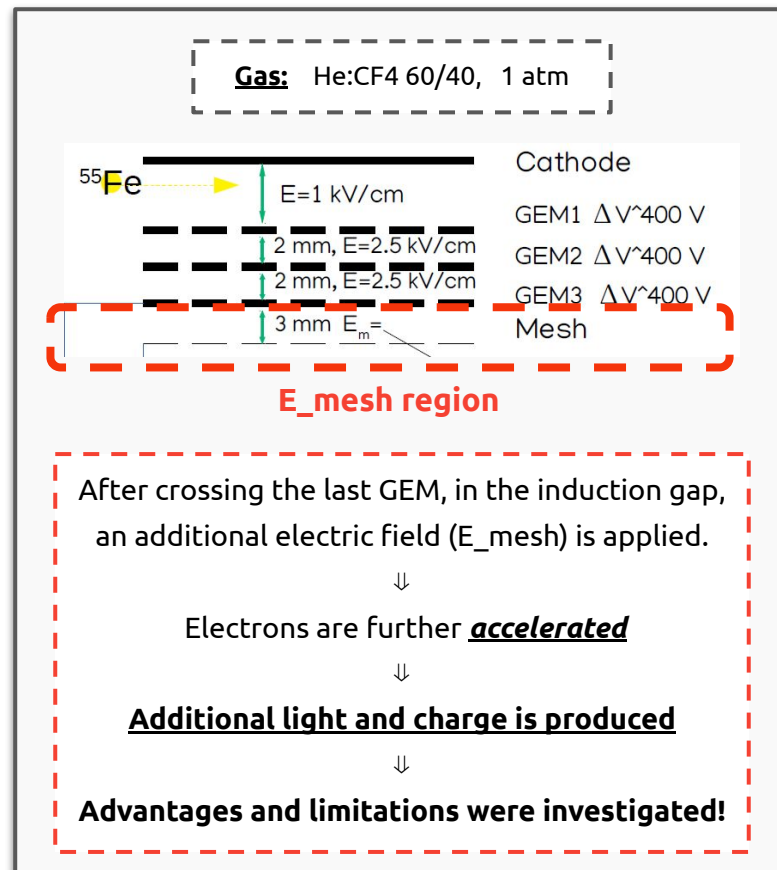
In CYGNO, the energy threshold is proportional to the amount of photons collected.

□ Light ⇒ □ Energy Threshold ⇒ □ DM

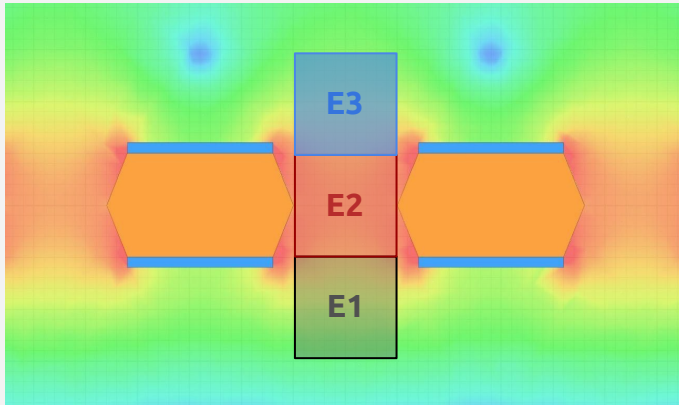


Options:

- Increase V_{GEM} ⇒ □ Sparks ✗
- Increase #GEMs ⇒ □ Diffusion/saturation ✗
- Applying a stronger electric field after the last GEM ⇒ □ Light + ~ spatial/energy res. ✓

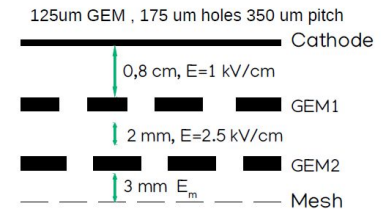
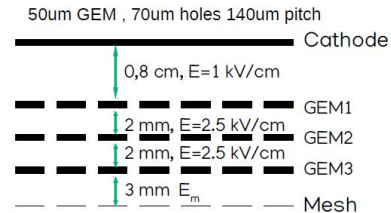


- ❑ To better quantify these effects, we used the 3 **contiguous** regions where the electric field profile was studied:
 - ❑ Below (E1), inside (E2) and above (E3) GEM.
- ❑ Box length = length of internal hole diameter, for **thin** and **Thick**.
- ❑ Obtain the electric field in matrix in these squares (1 μm res).
- ❑ Average over all points.
- ❑ Average over 10 holes.

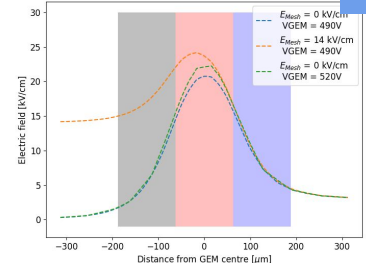
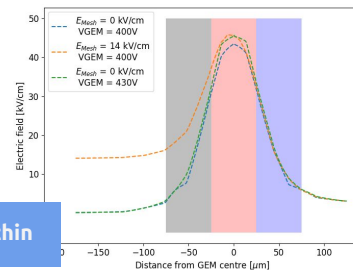
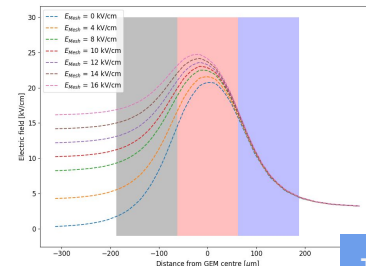
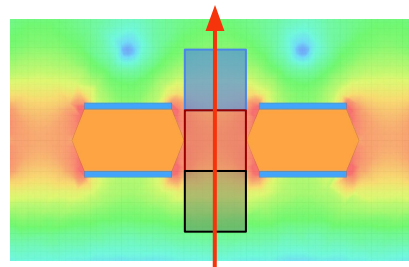
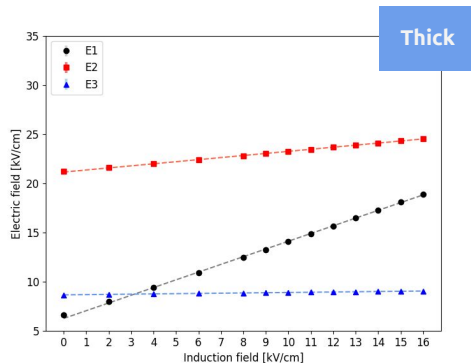
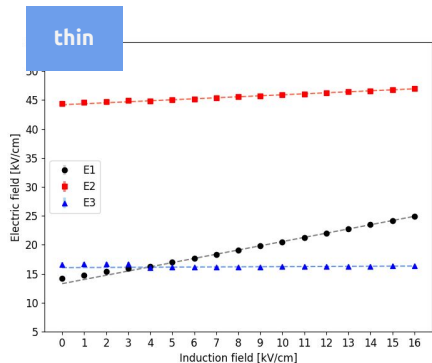


Tests performed:

- **Scan V_mesh**, with fixed V_GEMs
- **Scan on V_GEMs**, with a fixed V_mesh
- 2 different configurations:
 - ◆ **thin** (t - t - t)
 - ◆ **Thick** (T - T)



Simulation - Average field - Expectations



- ★ E_{mesh} **doesn't** change field **above** GEM.
- ★ E_{mesh} **increases** field **inside** GEM.
 - Linear increase of the electric field. ✓
- ★ E_{mesh} **greatly** affects field **below** GEM.
 - Reaches values enough to produce amplification. ✓
 - TT configuration works with lower fields thus we expected the light yield increase to be more notorious with the thick GEMs. ✓

- The addition of the induction field **changes the shape** of the profile:
- ★ The field's maximum **increases in amplitude and shifts** towards the **bottom** part of the GEM hole.
 - ★ Generates a **large region below the GEM where amplification and/or light production can happen.**
 - ★ For Thick GEMs, the influence is **relatively much larger.** ✓

Scenario	GEM conf.	E_mesh [kV/cm]	Light integral	ΔE [%]	Diffusion [μm]
Maximizing induction (E_mesh) field	ttt	15 ± 0.3	33500 ± 140	13.8 ± 0.3	388 ± 5
	TT	14 ± 0.3	58800 ± 300	25.7 ± 0.5	356 ± 5
	Tt	12.8 ± 0.2	11830 ± 50	26.8 ± 0.5	280 ± 4

- When maximizing E_mesh, the **ttt maintains a good ΔE** when compared with the other configurations.
- For the **TT case**, maximizing E_mesh allows for a **very high increase of light yield**, with similar intrinsic diffusion to the ttt (but losing on ΔE).
- Finally, the **Tt configuration excelled in intrinsic diffusion** in all the tests performed (check back-up), but with an overall worse ΔE .

Paper under final collaboration review!

★ In sum, the choice boils down to the experiment's requirements:

• ttt \Rightarrow Energy resolution

• TT \Rightarrow Light yield

• Tt \Rightarrow Intrinsic diffusion

Parallel work with MANGO



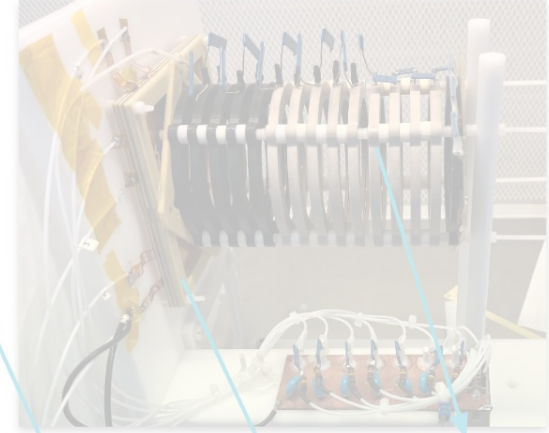
a Multipurpose Apparatus for Negative ion studies with GEM Optical readout

→ Enhancement of light yield

- Different types/configuration of GEMs
- ITO vs. Mesh

→ Negative Ion Drift

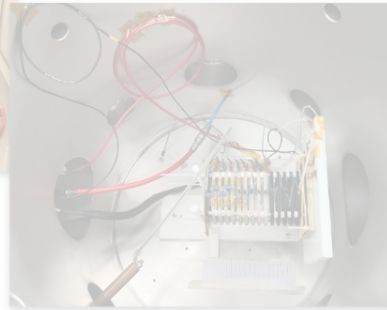
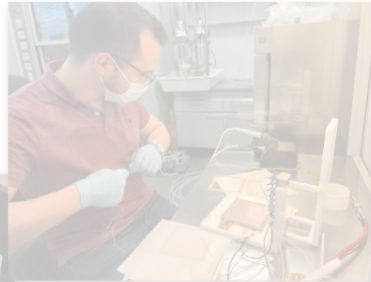
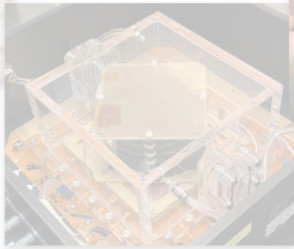
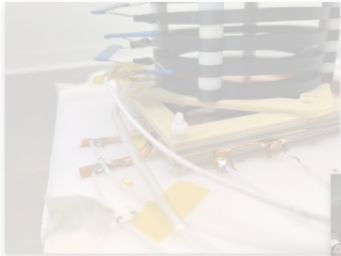
- Optimal gas pressure, composition & amplification config.
- Transverse & Longitudinal Diffusion
- Ion mobility



15 cm drift field cage

3 GEM stack:
50 μm thick
140 μm pitch
 $\varnothing 70 \mu\text{m}$ holes

GEMs facing
sCMOS and
PMT



Negative Ion Drift - INITIUM concept



an Innovative Negative Ion Time projection chamber for Underground dark Matter searches | INITIUM | Project | Fact sheet | H2020 | CORDIS | European Commission



Advantages:

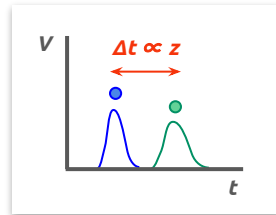
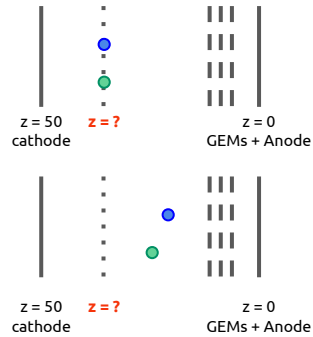
Reduced diffusion

Longitudinal and transverse **diffusion reduced to thermal limit**

$$\sigma_D = \sqrt{\frac{4\varepsilon L}{eE}}$$

Better spatial resolution!

Multiple charge carriers



Absolute Z from Δt between minority charge carriers

Negative Ion Drift - INITIUM concept



an Innovative Negative Ion Time projection chamber for Underground dark Matter searches | INITIUM | Project | Fact sheet | H2020 | CORDIS | European Commission



Advantages:

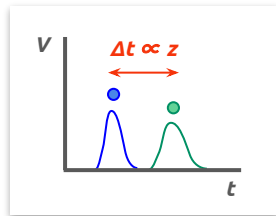
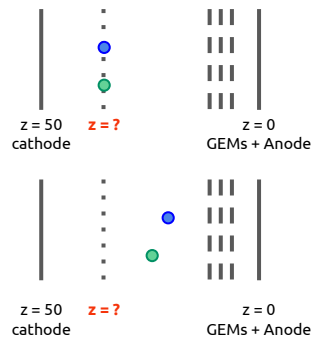
Reduced diffusion

Longitudinal and transverse **diffusion reduced to thermal limit**

$$\sigma_D = \sqrt{\frac{4\epsilon L}{eE}}$$

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Multiple charge carriers



Absolute Z from Δt between minority charge carriers

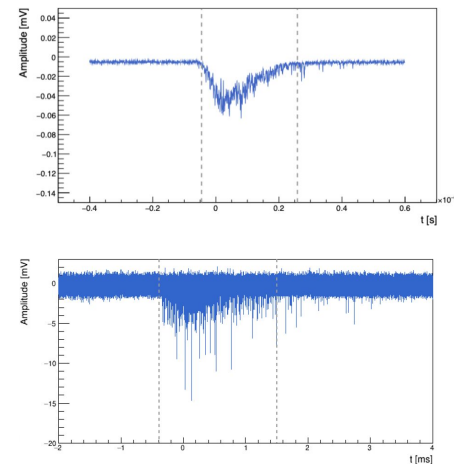
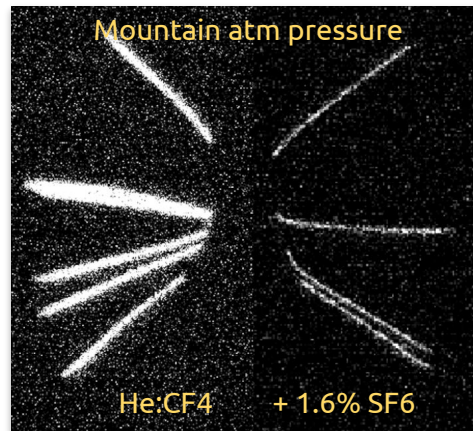
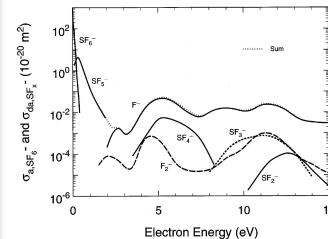


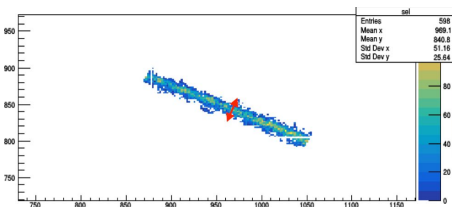
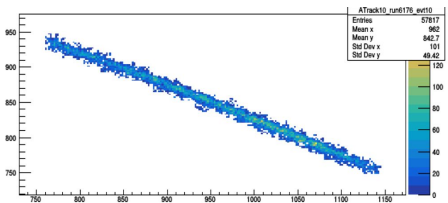
Table 2
Collisional processes that originate the SF_6^- ion species of interest for this work. Adapted from Ref. [46].

Reaction	Process	Energy
$e^- + SF_6 \rightarrow SF_6^{+*}$	Electron attachment	<1 meV
$SF_6^{+*} \rightarrow SF_6 + e^-$	Autodetachment	(Metastable: > 1 μs)
$SF_6^{+*} + SF_6 \rightarrow SF_6^+ + SF_6$	Collisional stabilization	
$e^- + SF_6 \rightarrow SF_5^- + F$	Dissociative electron attachment	0-2 eV
$e^- + SF_6 \rightarrow SF_4^- + 2F$		3-8 eV
$e^- + SF_6 \rightarrow F^- + SF_5$		1-14 eV
$e^- + SF_6 \rightarrow F_2^- + SF_4$		1-14 eV
$SF_6^{+*} + SF_6 \rightarrow SF_5^+ + SF_6 + e^-$	Collisional detachment	>90 eV
$SF_6^+ + SF_6 \rightarrow SF_6 + SF_6^+$	Charge transfer	
$SF_6^+ + SF_6 \rightarrow SF_5^+ + F + SF_6$	Dissociative charge transfer	>1 eV

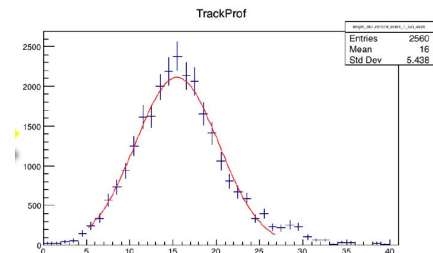


<https://doi.org/10.1016/j.nima.2022.1661416>

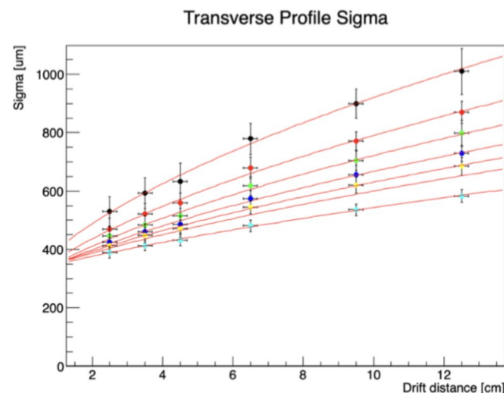
Negative Ion Drift - sCMOS analysis



Transverse profile of track fitted with Gaussian to estimate diffusion



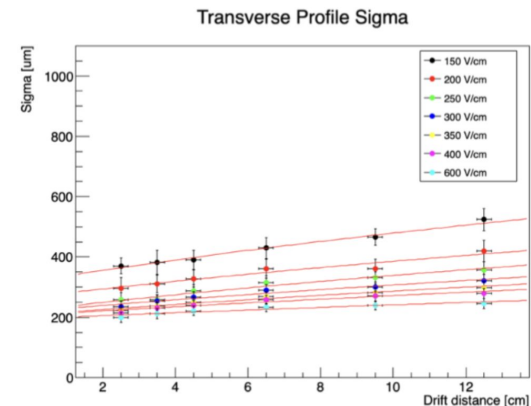
He:CF₄ 60:40 (Electron drift)



$$\sigma_{meas} = \sqrt{\sigma_0^2 + \sigma_T^2 L}$$

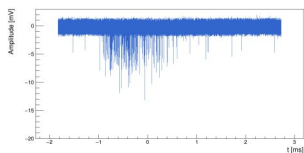
Drift field [V/cm]	σ_0^{ED}	σ_T^{ED}	σ_0^{NID}	σ_T^{NID}
150	300 ± 100	280 ± 20	320 ± 30	110 ± 10
200	290 ± 60	230 ± 10	260 ± 30	88 ± 20
250	284 ± 60	210 ± 10	220 ± 20	81 ± 10
300	300 ± 40	190 ± 10	220 ± 20	68 ± 10
350	300 ± 40	170 ± 10	210 ± 20	62 ± 10
400	310 ± 30	160 ± 10	210 ± 20	56 ± 9
600	320 ± 22	140 ± 10	200 ± 20	45 ± 10

He:CF₄:SF₆ 59:39.6:1.6 (Negative ion drift)

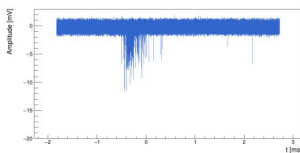


- Negative Ion Drift operation!
- **Transverse diffusion** reduced by **3 times!**
- **What about the PMT signals?**

→ I focused on the **PMT analysis**



(a) 300 V/cm.

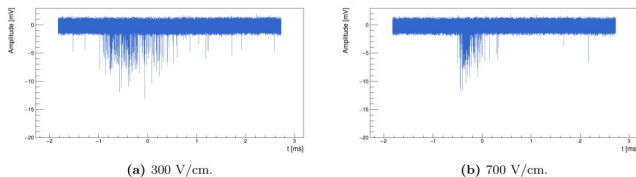


(b) 700 V/cm.

Peculiar signal:

- Thousands of small peaks (\sim ns width) over large time span (\sim few ms) \Rightarrow **primary ionization cluster counting?**
 \sim Perfect resolution
- Features visible \Rightarrow Note the enlargement of signal with lower drift field.
- **First optical observation of NID** \Rightarrow
 \Rightarrow **Few or none literature** on this

→ I focused on the **PMT analysis**



Peculiar signal:

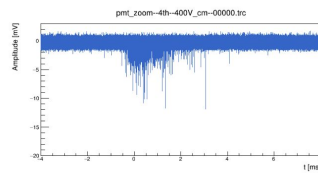
→ Thousands of small peaks (~ns width) over large time span (~ few ms) ⇒ **primary ionization cluster counting?**
~> **Perfect resolution**

→ Features visible ⇒ Note the enlargement of signal with lower drift field.

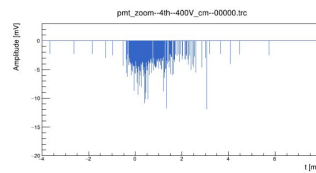
→ **First optical observation of NID** ⇒
⇒ **Few or none literature** on this

□ □ **First ever PMT analysis for optical NID signals** □ □

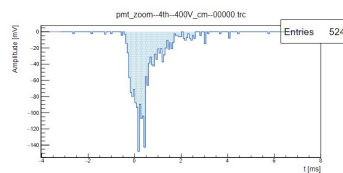
- Initial approach ⇒ **Time rebinning**
 - Threshold** ⇒ Only peaks above $6 \times \text{RMS}$ are taken into account
 - Rebin** ⇒ Selected points are put into histogram
 - Delimitation** ⇒ Start (end) when 2 bins are above (below) 10 mV
 - Systematics** ⇒ Varying #bins & threshold voltage (**reworked analysis**)



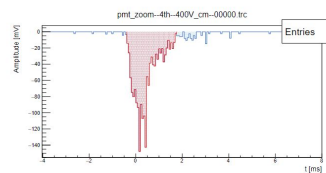
(a) Original signal.



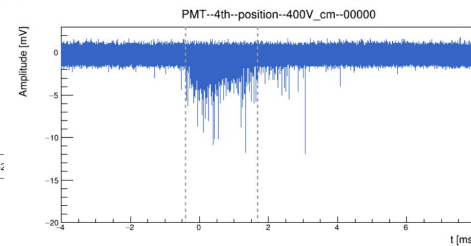
(b) Only peaks over noise threshold (2.5 mV).



(c) Rebinned waveform.



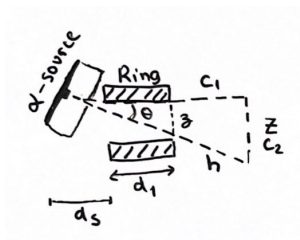
(d) Reconstructed signal in red.



(e) Final boundaries in the original signal.

Final measurement: **Ion Mobility** → **Updated version!**

1. Tilted ED alpha tracks
 - a. Distribution → Get mean value (Δt)
 - b. Knowing electrons' velocity in gas (v)
 - i. **Z travelled by track** → **1.5 cm**



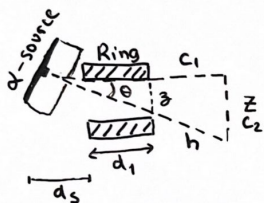
2. Tilted NID alpha tracks
 - a. Average time window (Δt)
 - i. **Z / Δt = v_{ion}**

Negative Ion Drift - PMT analysis

Final measurement: **Ion Mobility** → **Updated version!**

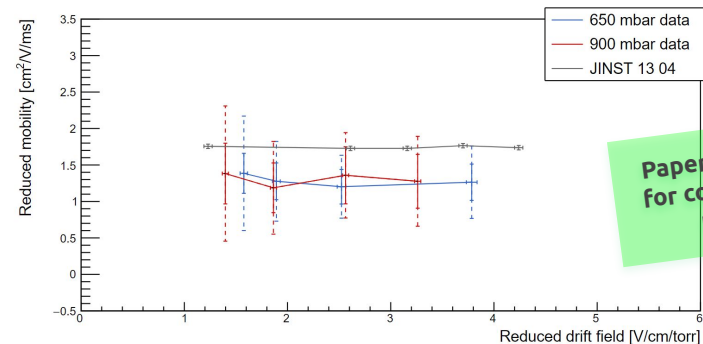
1. Tilted ED alpha tracks

- Distribution → Get mean value (Δt)
- Knowing electrons' velocity in gas (v)
 - Z travelled by track** → 1.5 cm



2. Tilted NID alpha tracks

- Average time window (Δt)
 - Z / $\Delta t = v_{ion}$**

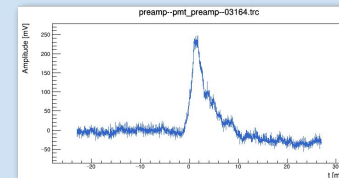


Paper submitted
for collaboration
review!

- Data **self-consistent** with previous year's data (900 mbar)
- Charge carriers' **mobility** consistent with **SF6**- in literature
- PMT signal undoubtedly proved NID with optical readout!**

Future:

→ New data with pre-amp to be analyzed.



- ◆ Could help understanding the signal's structure
- ◆ **Diffusion measurement to complement sCMOS analysis**

Thank you for your attention!

Scientific Communications (November 2022 - October 2023):

1. Papers:

- 1.1. **LIME — A gas TPC prototype for directional Dark Matter search for the CYGNO experiment**
Amaro, F. D.; Baracchini, E.; Benussi, L. et al
NIM A, Mar. 2023
- 1.2. **50 litres TPC with sCMOS-based optical readout for the CYGNO project**
Mazzitelli, G.; Amaro, F. D.; Baracchini, E. et al
NIM A, Jan. 2023
- 1.3. **Noise assessment of CMOS Active Pixel sensors for the CYGNO Experiment**
Almeida, B. D.; Amaro, F. D.; Antonietti, R. et al
Meas. Sci. Technol., Sep. 2023 (accepted)
- 1.4. **Dual-Polarity Ion Drift Chamber: Experimental results with Xe-SF₆ mixtures**
Marques, A.P.; Marques, D.J.G.; Duarte, N.G.S. et al
NIM A, Jan. 2023
- 1.5. **Directional IDBSCAN to detect cosmic-ray tracks for the CYGNO experiment**
Amaro, F. D.; Antonietti, R.; Baracchini, E. et al
Meas. Sci. Technol., Sep. 2023
- 1.6. **A 50 liter CYGNO prototype overground characterization**
Amaro, F. D.; Baracchini, E.; Benussi, L. et al
EPJ C (submitted)
- 1.7. **Secondary scintillation yield from GEM electron avalanches in He-CF₄ and He-CF₄-isobutane for CYGNO – Directional Dark Matter search with an optical TPC**
Amaro, F. D.; Baracchini, E.; Benussi, L. et al
Phys. Letter B (submitted)

2. Conference Communications:

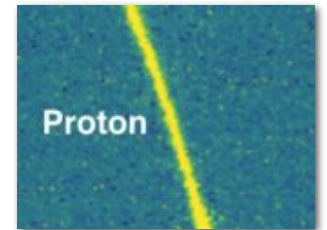
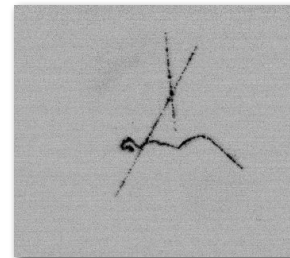
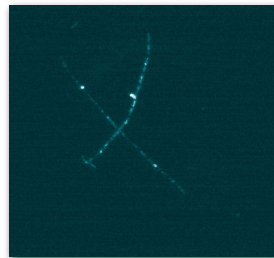
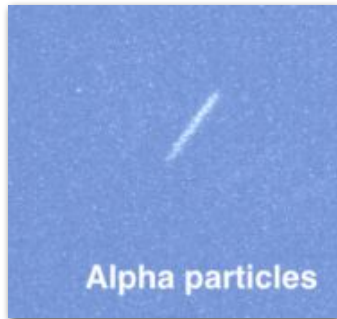
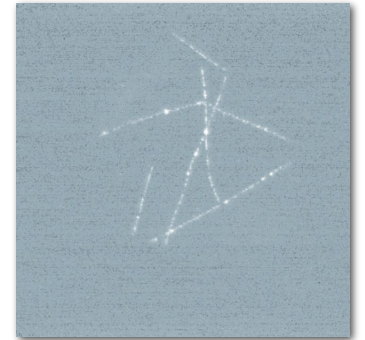
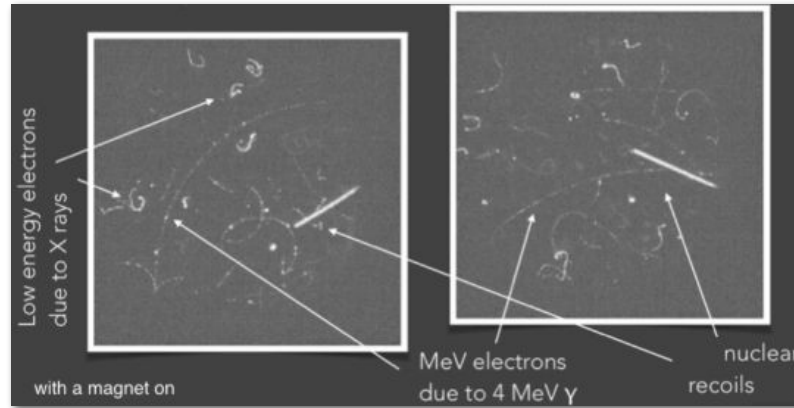
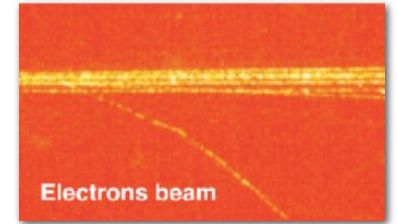
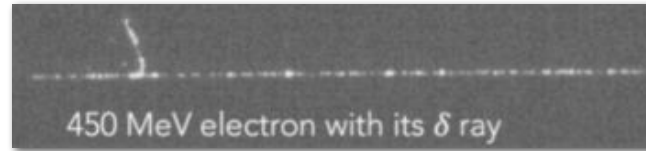
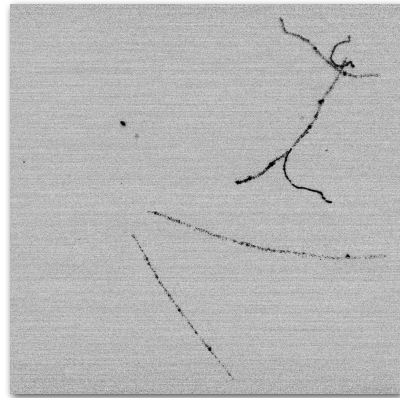
- 2.1. **The CYGNO experiment, a directional optically readout detector for Dark Matter searches**
Oral presentation
CPAD 2022, the Coordinating Panel for Advanced Detectors, 29/11 - 2/12, 2022, Stony Brook, US
- 2.2. **CYGNO, a directional Dark Matter TPC optically readout**
Oral presentation
LIDINE 2023, Light Detection in Noble Elements, 20 - 22 Sep, 2023, Madrid, Spain



Backup

& more details

CYGNO - *Some pictures!*



→ Given a setup with **different materials** and **voltages** applied, Maxwell calculates the **electric field** within a defined region. *Different GEM configurations were designed and studied.*

Source

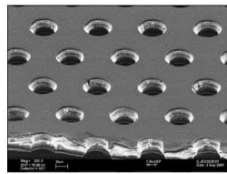


Figure 6. Electron microscope picture of a "standard" GEM foil, with 70 µm holes at 140 µm pitch in a triangular pattern.

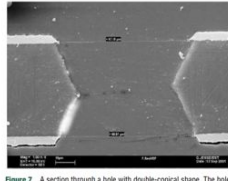
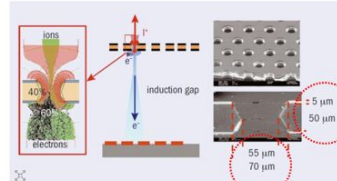


Figure 7. A section through a hole with double-conical shape. The hole diameter at the metal surface is 70 µm and the opening in the center of the polymer is 50 µm.

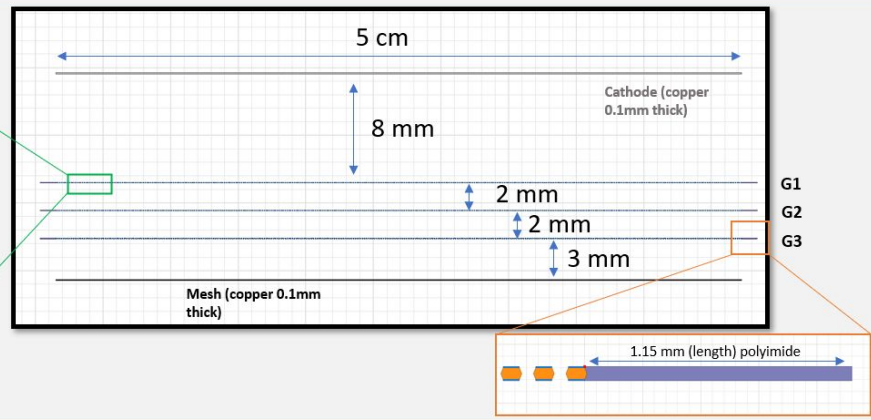
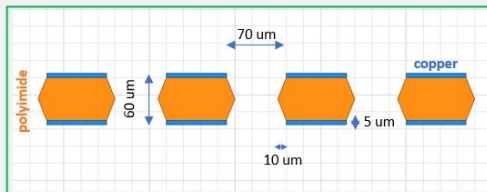


A gas electron multiplier (GEM) detector: schematic (centre) and micrographs of a real device from above (top right) and showing the profile of the walls of the holes (lower right), which give the field lines and amplification illustrated on the left.

<https://cerncourier.com/a/the-continuing-rise-of-micropattern-detectors/>

https://indico.cern.ch/event/346614/contributions/813299/attachments/683646/939073/GEM_at_CERN_seminar.pdf

Outcome



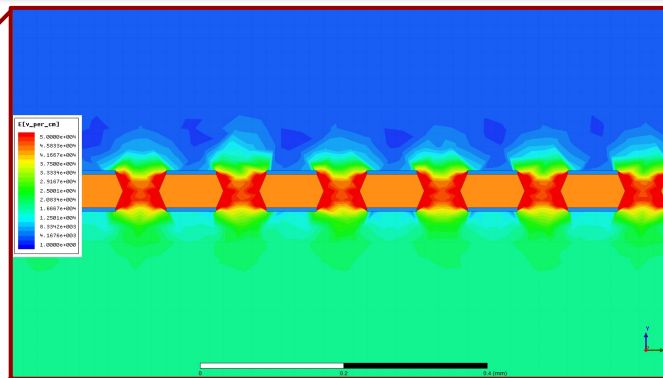
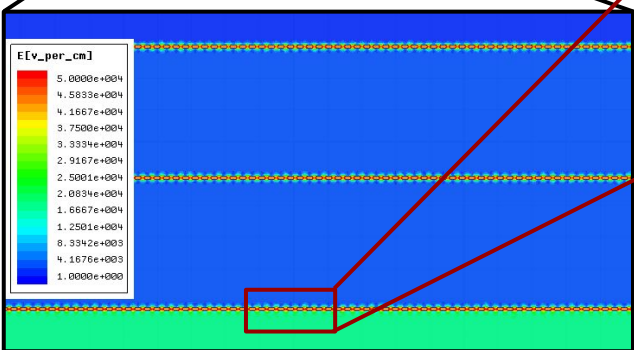
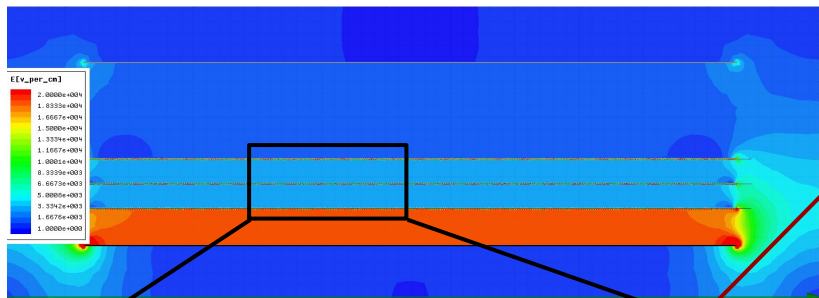
GEMs studied:

- Thin:
 - 50 µm Kapton
 - 70 µm holes
 - 140 µm pitch

- Thick:
 - 125 µm Kapton
 - 175 µm holes
 - 350 µm pitch

Example of the calculated electric field in the t-t configuration.

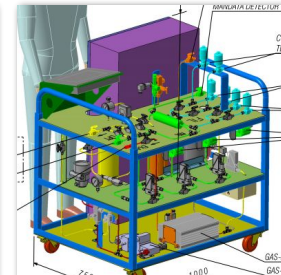
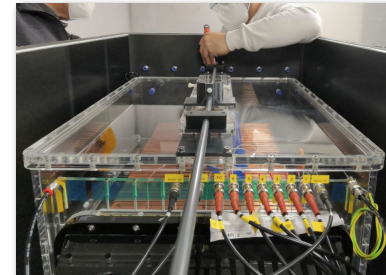
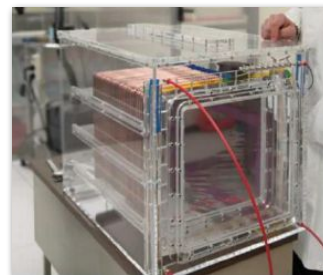
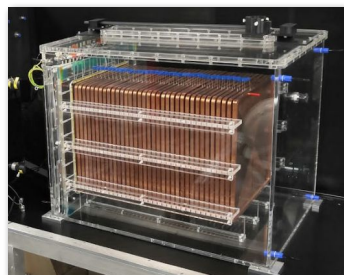
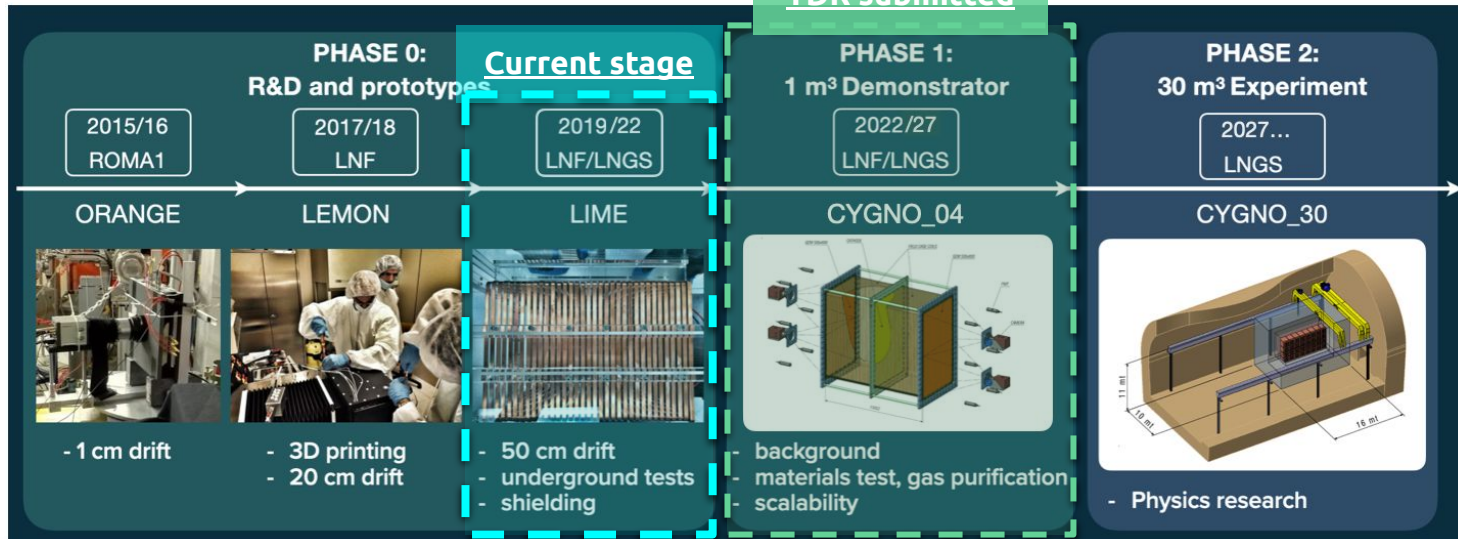
- V_GEMs : 400V
- TFs: 2.5 kV/cm, in 2 mm
- DF: 1kV/cm, in 8mm
- **Induction field: 18.3 kV/cm, in 3 mm**



- We focus on the **middle** to avoid border effects.
- **Macro fields constant** (parallel plates)
 - ◆ We focus on the fields close to the GEMs
- Induction field innocuous to GEM 1 and GEM 2
 - ◆ We focus on the effects observed in GEM 3

Several ongoing efforts in different fronts:

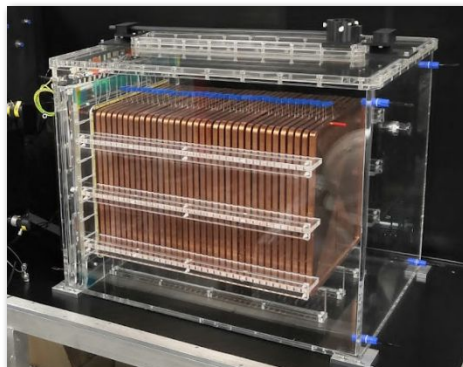
- Sensitivity
- 3D reconstruction
- Directionality
- ER vs. NR discrimination (ML)
- Shielding optimization
- Data vs. MC



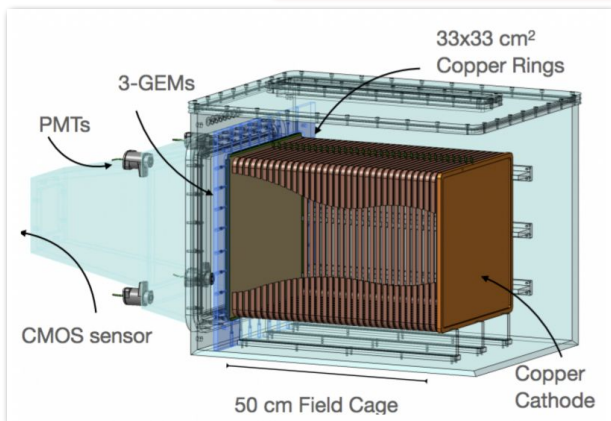
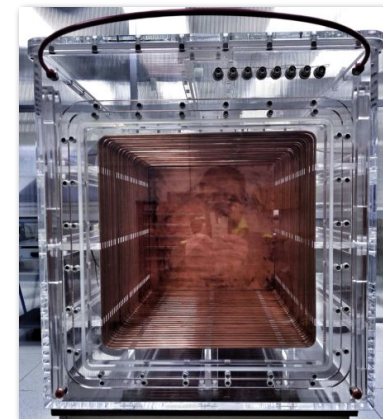
...for more info:

[CYGNO – Directional Dark Matter Search](#)
<https://www.facebook.com/cygno.experiment>





- Single-sided cathode, **50 L gaseous TPC**
- At **atm pressure**, room temperature and **He:CF₄**
- **Triple 33x33 cm² GEM stack** for amplification
- **Optical readout**
 - ◆ 4 PMTs
 - ◆ 1 sCMOS camera (ORCA Fusion)
- Copper ring field cage, **50 cm drift**



ORCA-Fusion CAMERA SPECS

LOW NOISE AND EXCEPTIONAL
READOUT NOISE UNIFORMITY

READOUT NOISE
0.7 electrons rms
Ultra-quiet Scan

DSNU
0.3 electrons rms

PRNU
0.06 % rms
At 7500 electrons

HIGH SPEED
100 frames/s
At 2304 × 2048 ROI

PIXEL SIZE
6.5 μm × 6.5 μm

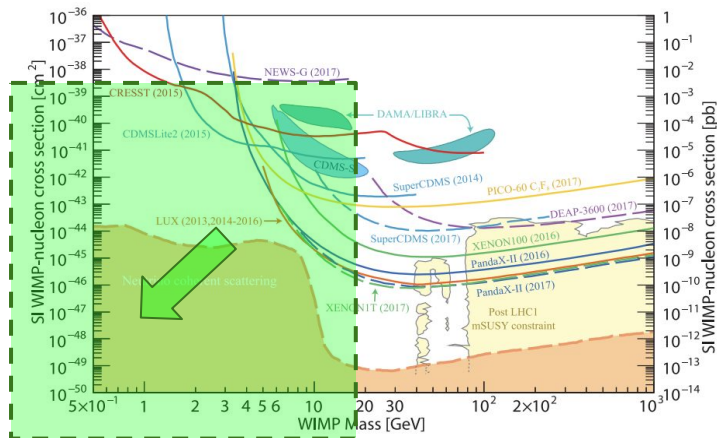
DYNAMIC RANGE
21 400:1

HIGH RESOLUTION
2304 × 2304
5.3 Megapixels

PEAK QE
80 %



CYGNO Dark Matter exploration region



Low Density @ atm pressure

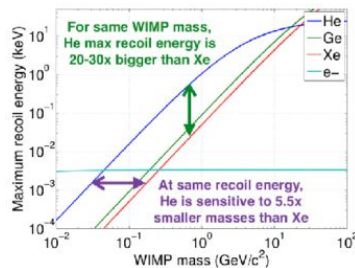
- Allows tracks of several millimeters at few keV without compromising exposure.

$\leq 10 \text{ GeV}/c^2$

- To observe lower WIMP masses:
 - ◆ Lower thresholds are necessary since lower $m\chi$ originate lower energy recoils.
 - ◆ Light nuclei used to maximize energy transfer.

Helium (He)

- Light target for SI in low mass range.



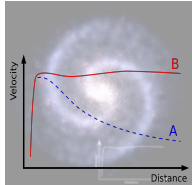
Fluorine (F)

- Heavier target to intermediate WIMP masses.
- Also Sensitive to SD coupling since $A = 19$ (odd).

Dark Matter - What, why and where?

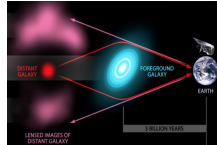
→ In the past few years, several **gravitational** anomalies have been found that **support the existence of a new type of matter.**

1. Galaxy rotation curves



$$v = \sqrt{\frac{GM}{r}}$$

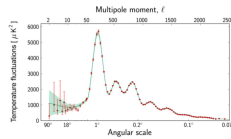
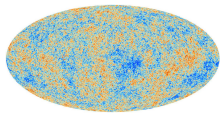
2. Gravitational lensing



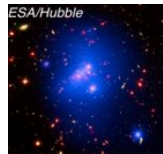
$$\Delta\Phi = \frac{4MG}{bc^2}$$

Universe's weight seems inconsistent with observations....

3. Planck's CMB measurement



4. Motion of galaxies inside clusters



$$\langle v^2 \rangle \approx \frac{GM}{\langle r \rangle}$$

This “matter” dominates the universe and only interacts **gravitationally...**



Commonly called **Dark Matter**

Best explanation (?)

WIMPs

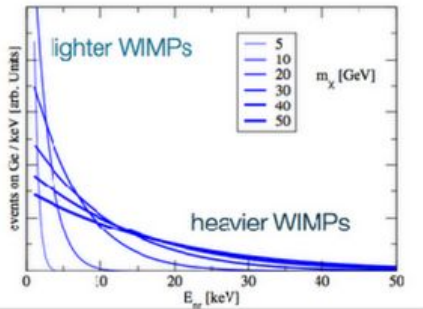
($m\chi \sim \text{GeV to TeV}$)

Highly justified theory independently predicted by **extensions** of the **Standard Model** at the weak-scale and **Cosmology!**

WIMPs - What dependency to explore?

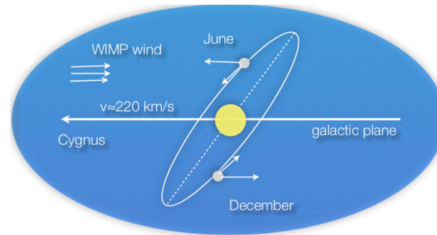
Increasing reliability but increasing difficulty in the experimental technique.

1. Exploring the **ENERGY** dependency

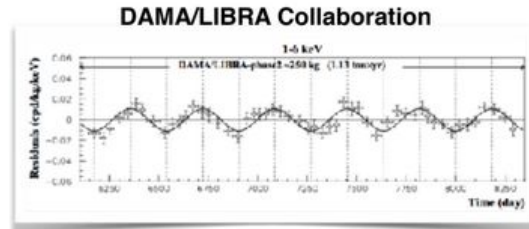


Results in a falling exponential with no peculiar features. The background has a similar spectrum.

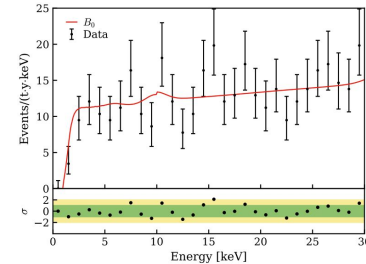
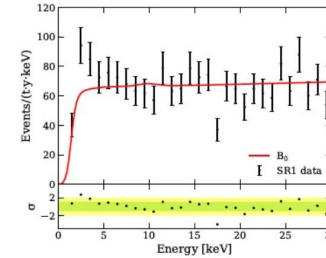
2. Exploring the **TIME** dependency



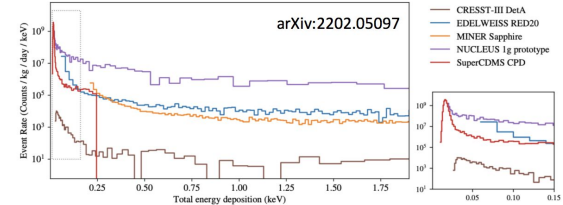
Results in a few % annual modulation.



[The Annual Modulation Signature for Dark Matter: DAMA/LIBRA-Phase1 Results and Perspectives](#)



Exponentially rising background towards lower energies

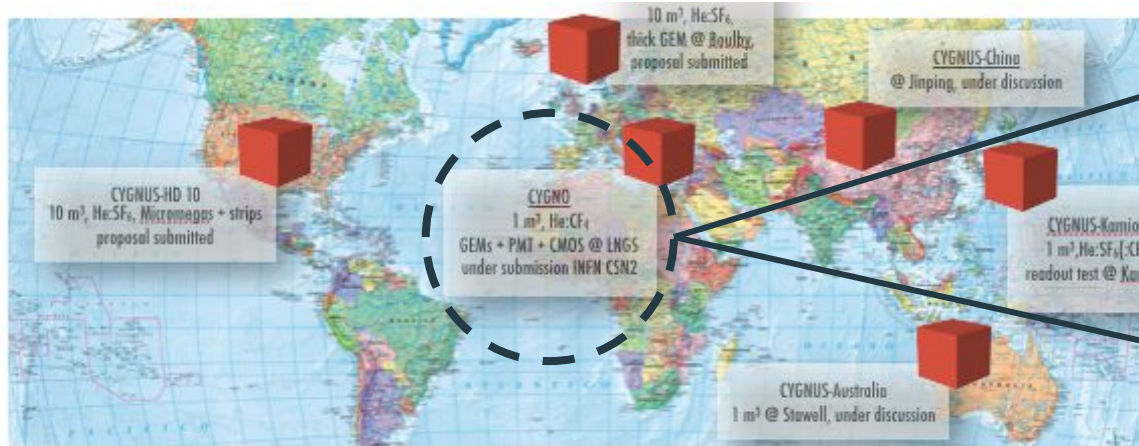


Currently limiting the sensitivity globally!
Origin still unknown, but a lot of R&D is going on ...

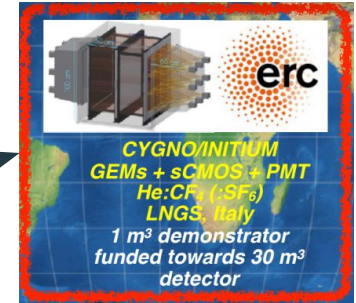
In all of these, it's hard to prove / disprove DM.

The CYGNO project

CYGN is part of a proto-collaboration, CYGNUS, focused on establishing a **Galactic Directional Recoil Observatory** that could test and study DM hypothesis beyond the neutrino floor.



<https://inspirehep.net/literature/1813839>



Within the CYGNUS collaboration, several approaches are being studied. The Italian group, CYGN, is developing a **gaseous TPC** based on the setup:

GEMs + sCMOS + PMT to test **Optical Readout**