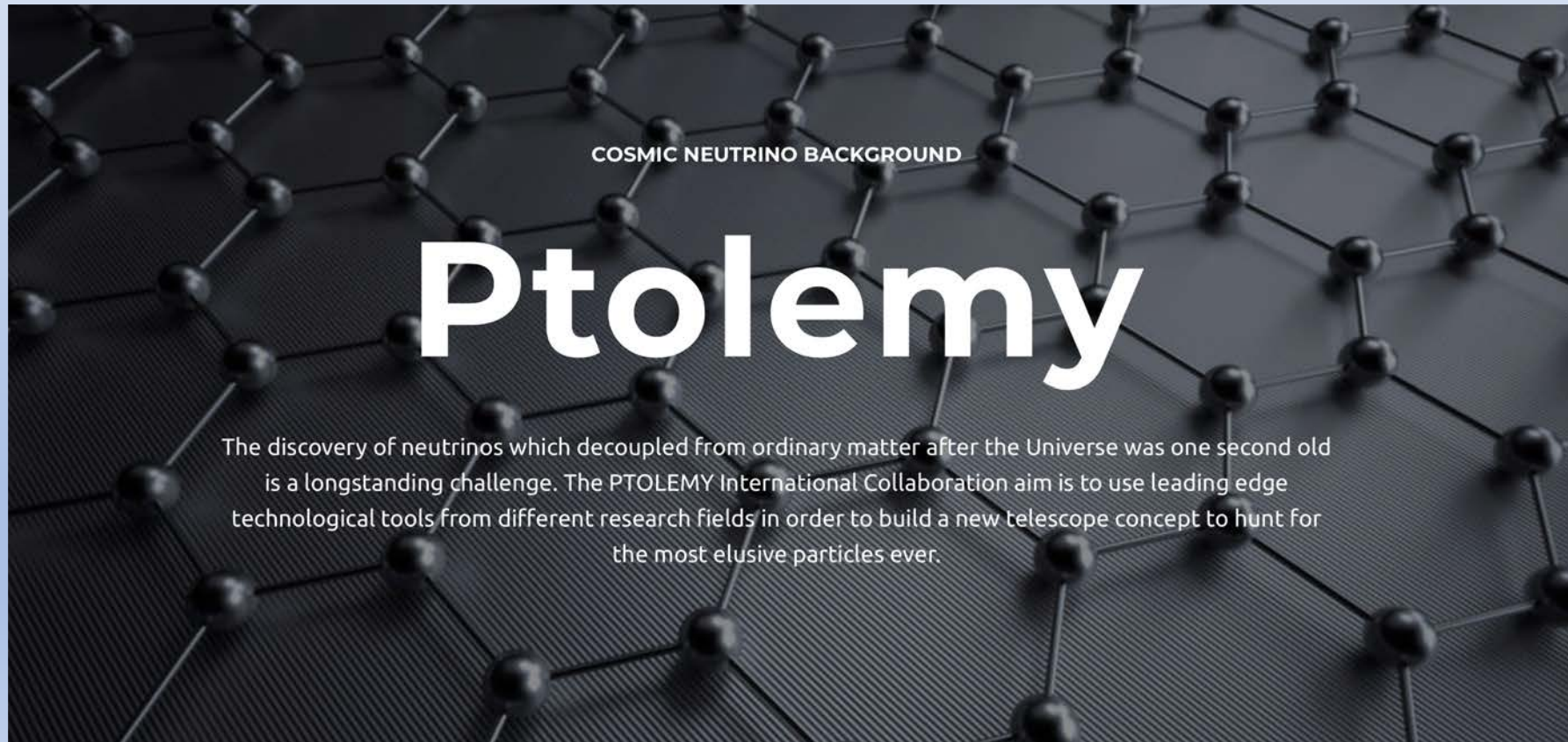


PTOLEMY: A detector for the oldest neutrinos in the Universe

Chris Tully (Princeton)

26 April 2023

L'Aquila Joint Astroparticle Colloquium
Gran Sasso Laboratory



COSMIC NEUTRINO BACKGROUND

Ptolemy

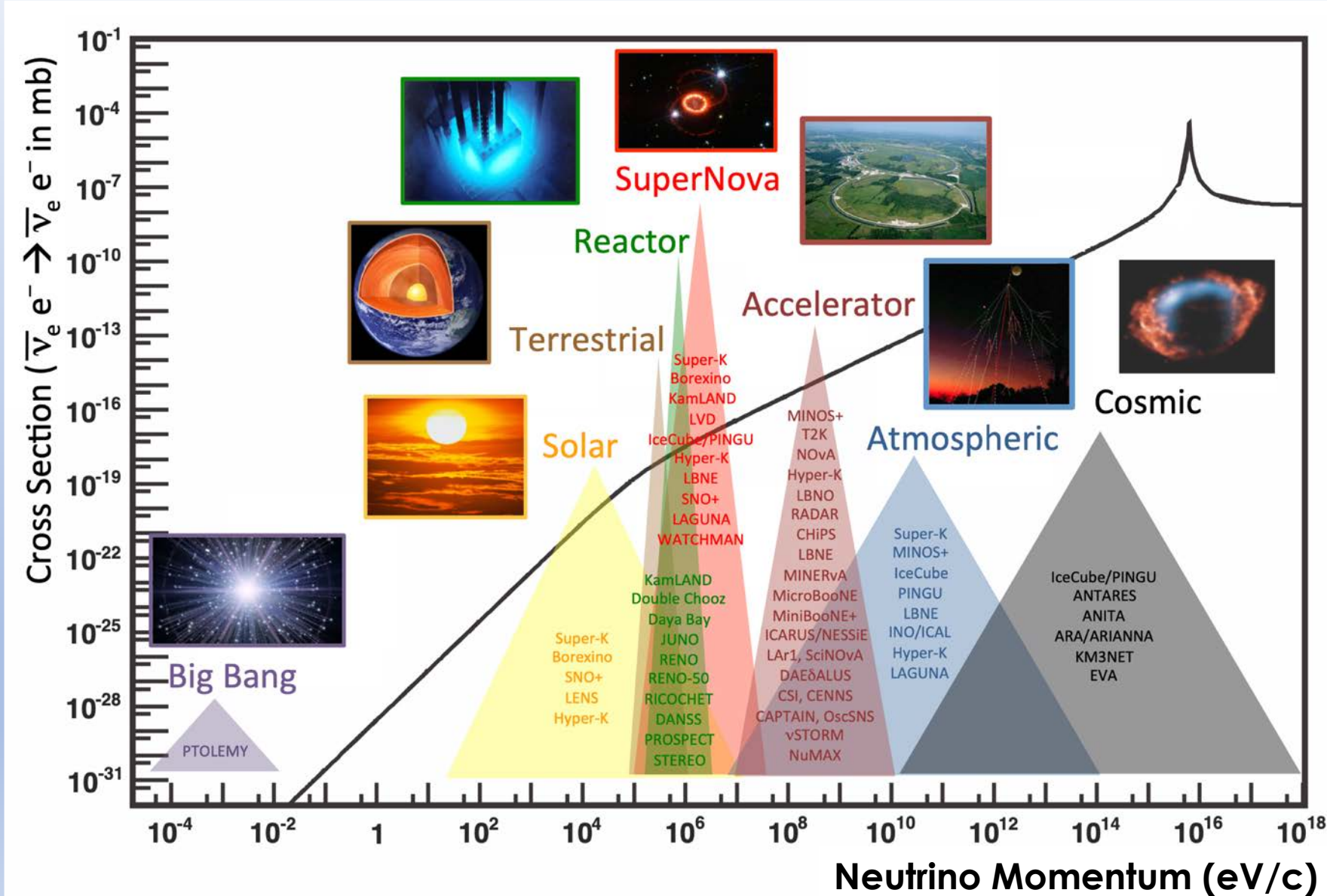
The discovery of neutrinos which decoupled from ordinary matter after the Universe was one second old is a longstanding challenge. The PTOLEMY International Collaboration aim is to use leading edge technological tools from different research fields in order to build a new telescope concept to hunt for the most elusive particles ever.

<http://ptolemy.lngs.infn.it>

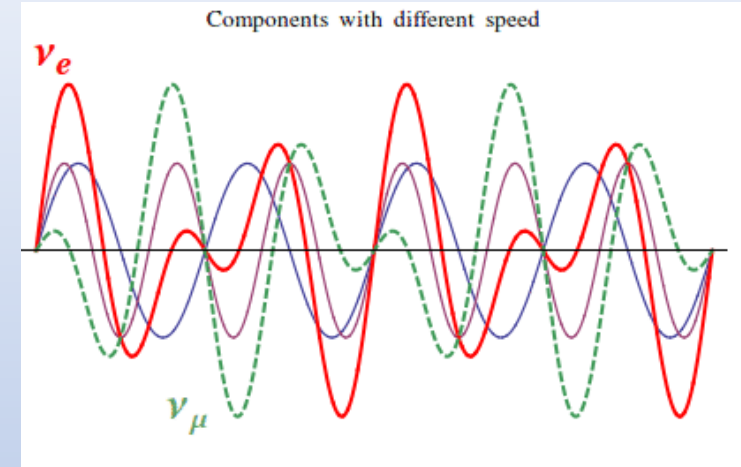
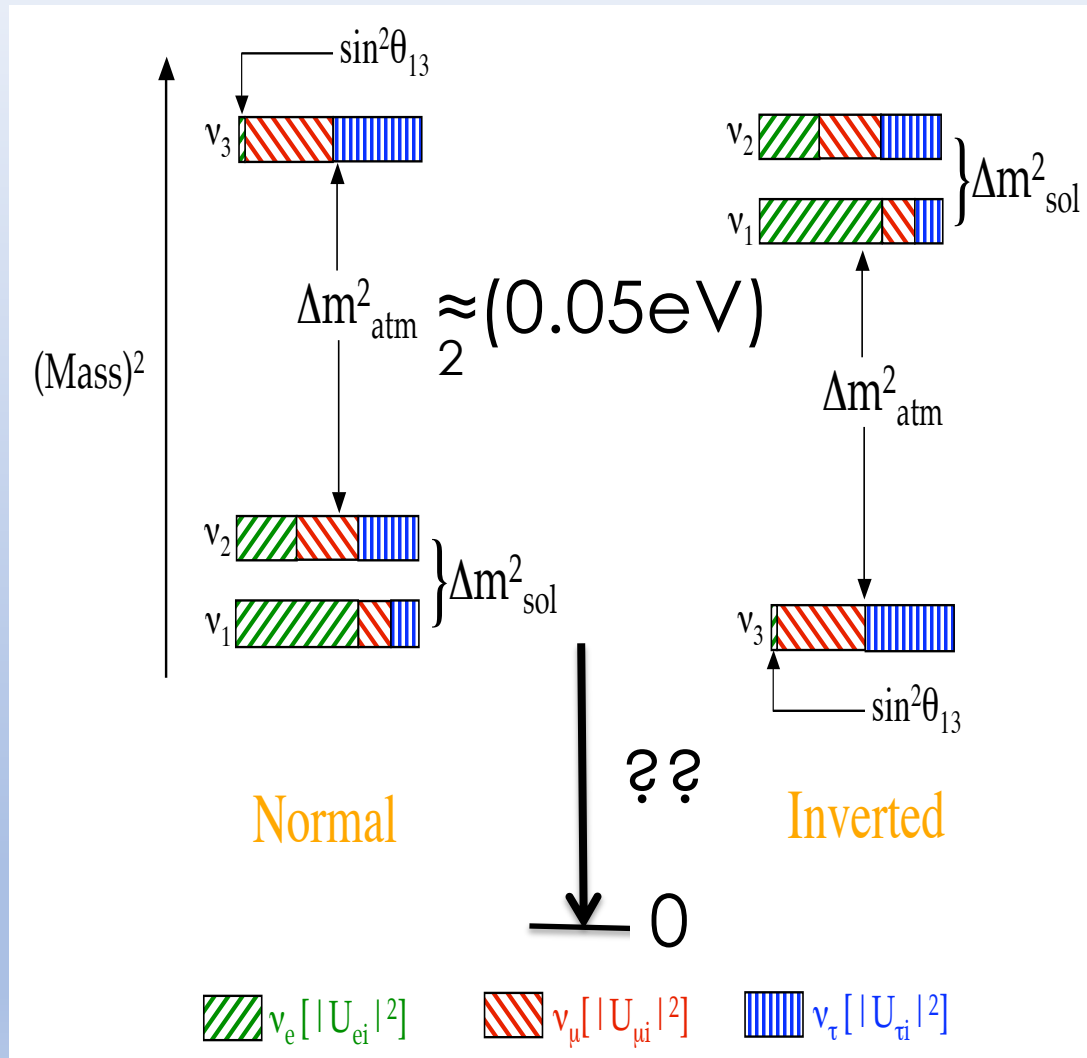


SIMONS FOUNDATION

Neutrinos sources across the Cosmos

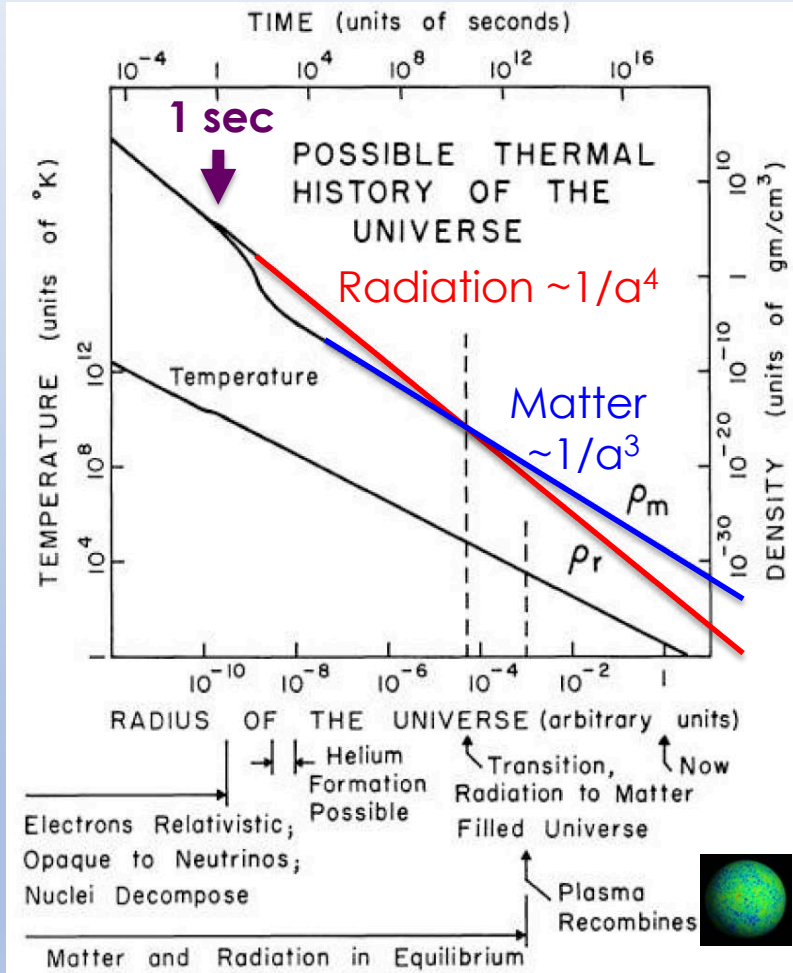


Neutrino Masses from Oscillations



3 mass eigenstates
 \times
 3 flavors
 (electron, muon, tau)

Cosmic Neutrino Background



Neutrino number density:

$$n_\nu = 112/\text{cm}^3$$

Temperature:

$$T_\nu \sim 1.95\text{K}$$

Time of decoupling:

$$t_\nu \sim 1 \text{ second}$$

neutron/proton ratio
@start of nucleosynthesis

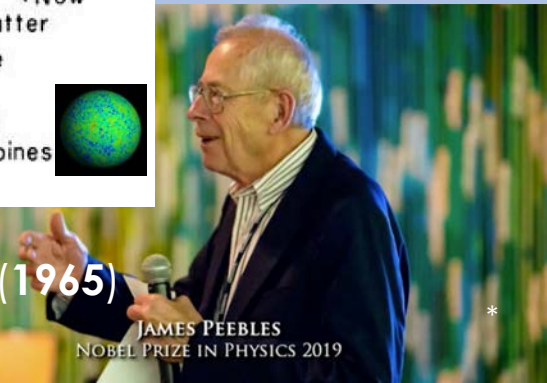
Velocity distribution:

$$\langle v_\nu \rangle \sim T_\nu / m_\nu$$

Non-linear distortions
Villaescusa-Navarro et al
(2013)

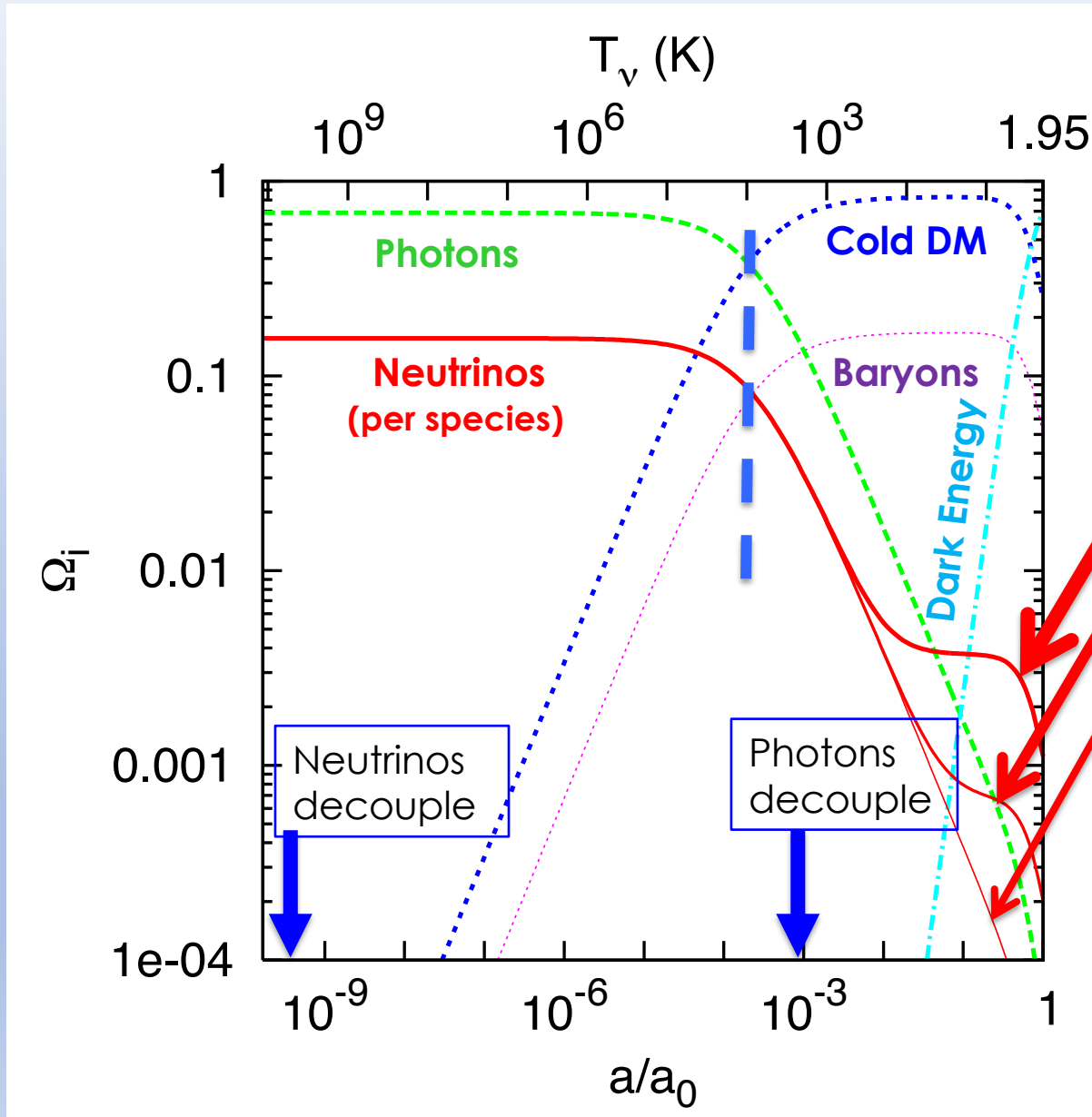
Dicke, Peebles*, Roll, Wilkinson (1965)

[Cosmology's Century \(2020\)](#)



JAMES PEEBLES
NOBEL PRIZE IN PHYSICS 2019

Cosmic Elements

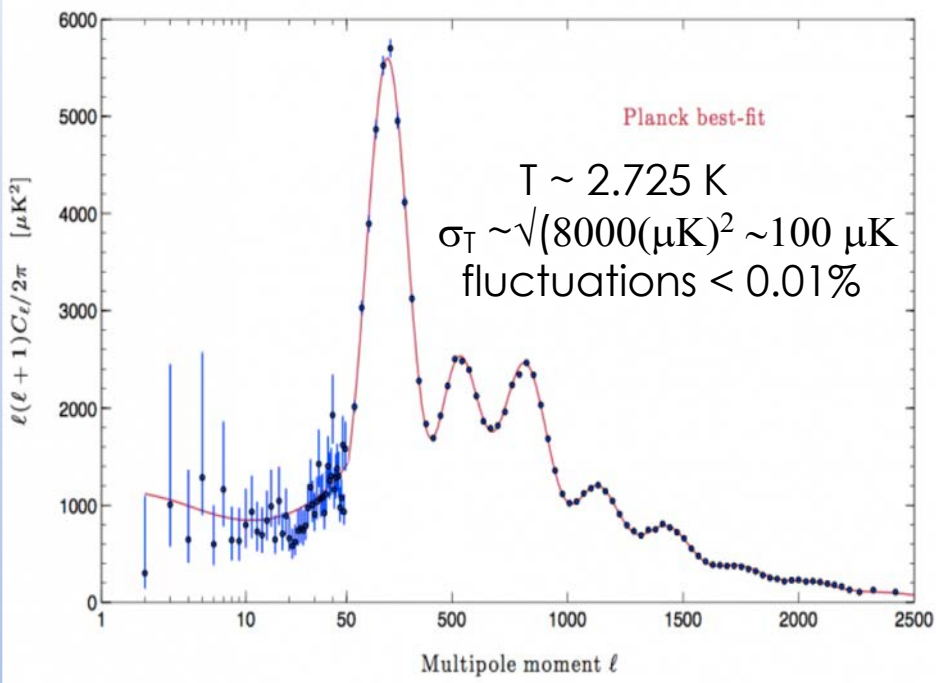


J. Lesgourgues

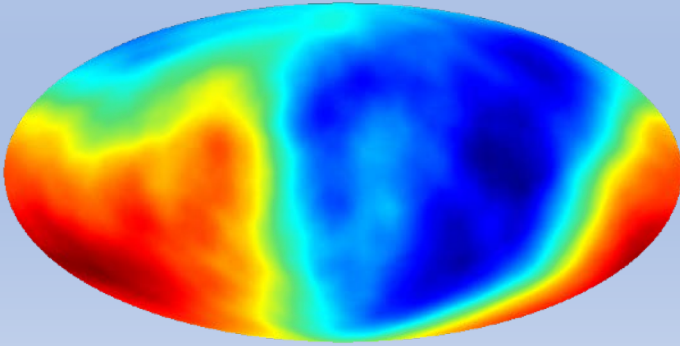
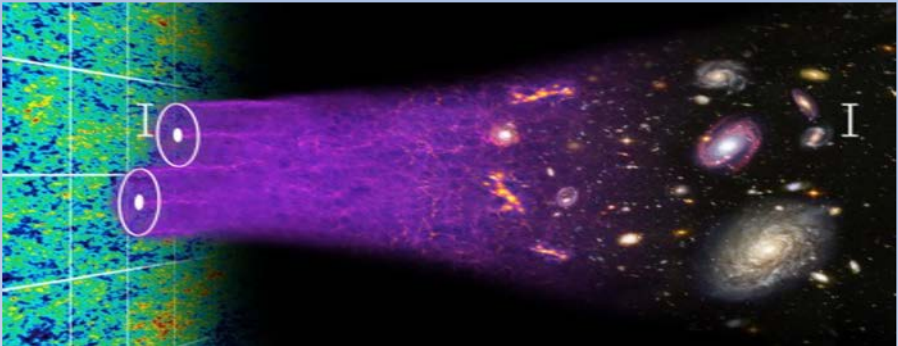
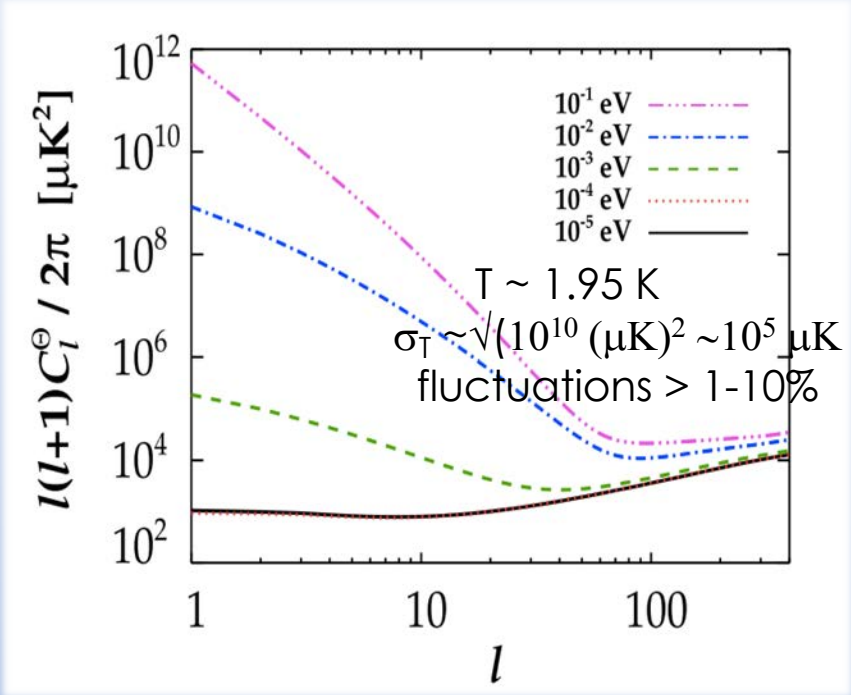
Individual neutrino contributions assuming Normal Hierarchy and $m_3 = 0.05$ eV, $m_2 = 0.009$ eV, $m_1 = 0$

Neutrino Flux on the Sky

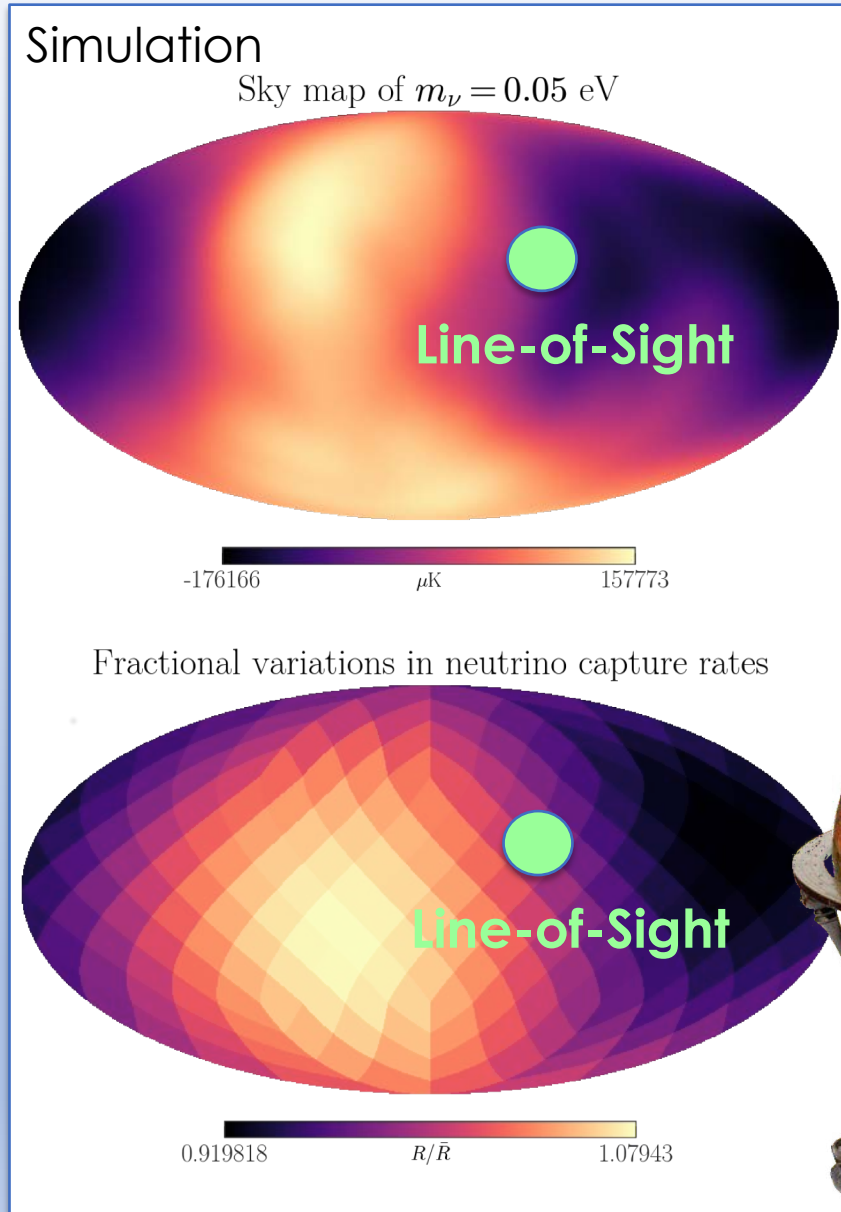
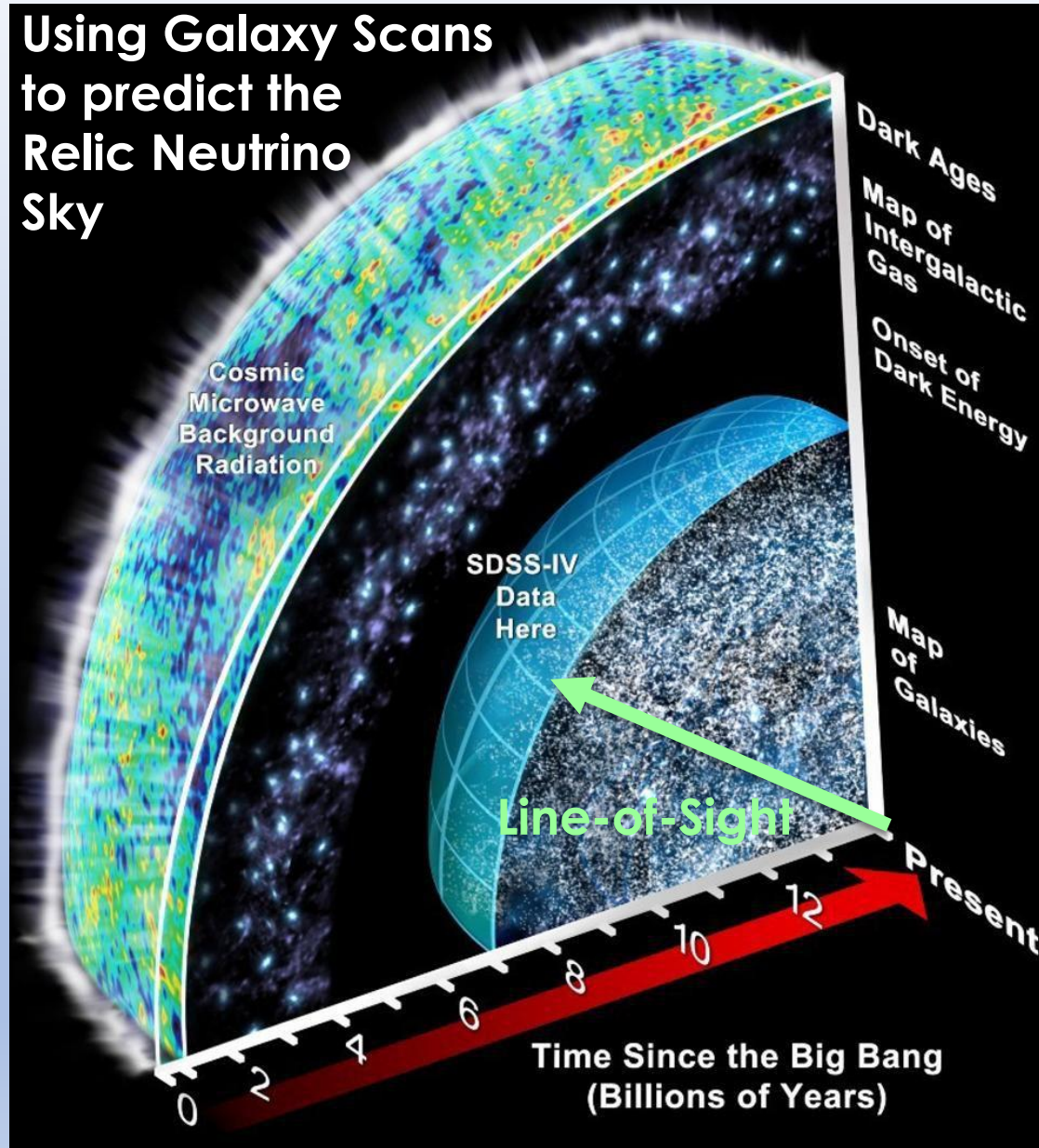
Cosmic Microwave Background (CMB)



Cosmic Neutrino Background (CNB)



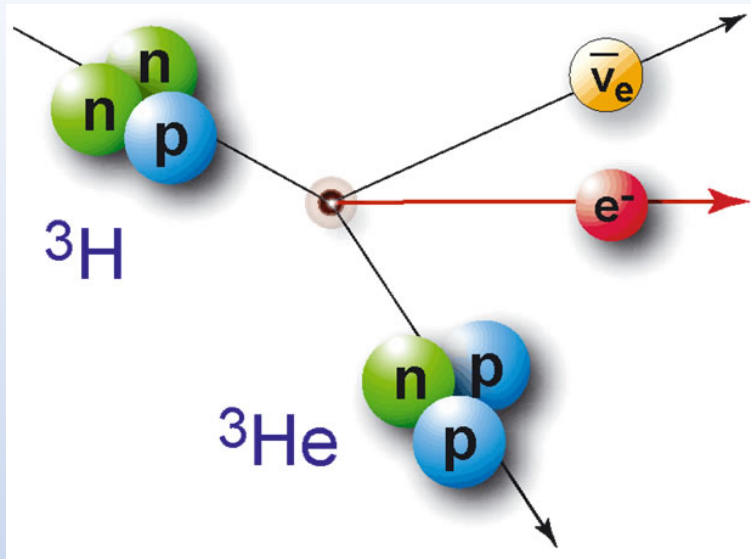
PTOLEMY: Experiment to measure relic neutrinos from the Big Bang



50meV
neutrino mass
→ CNB w/
~10%
anisotropies



Neutrino capture on Tritium

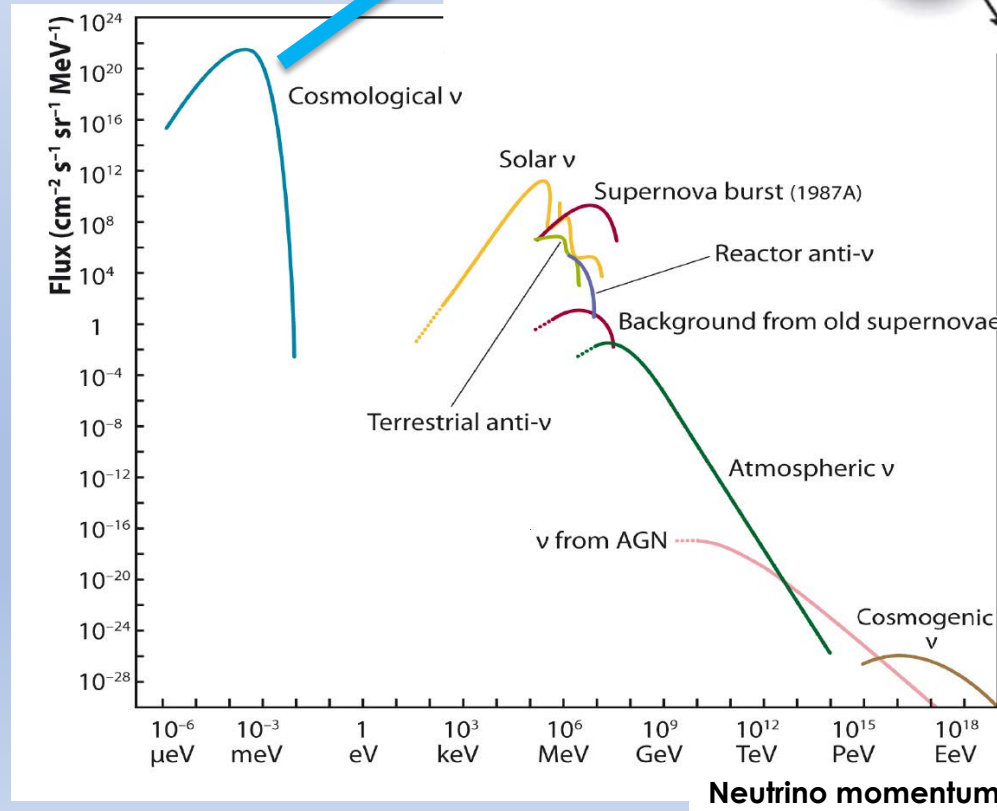
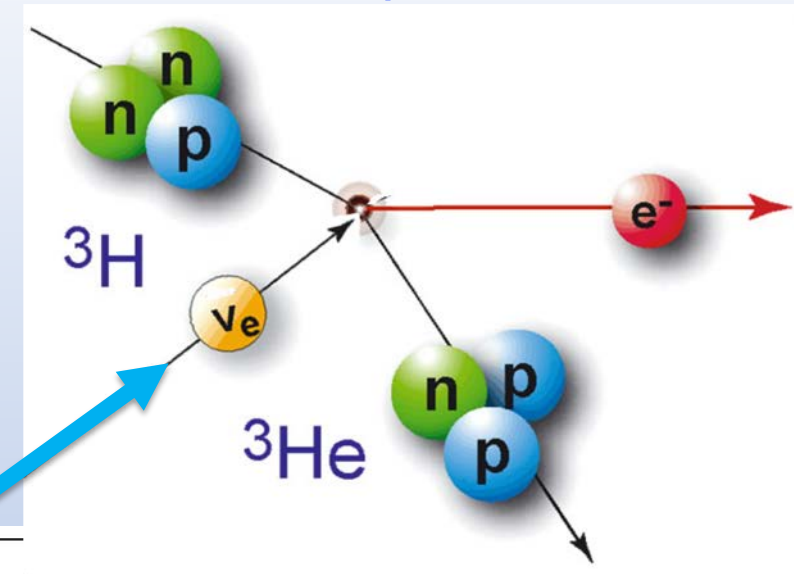


Tritium β -decay
(12.3 yr half-life)

Relic neutrino momentum ~ 0.17 meV

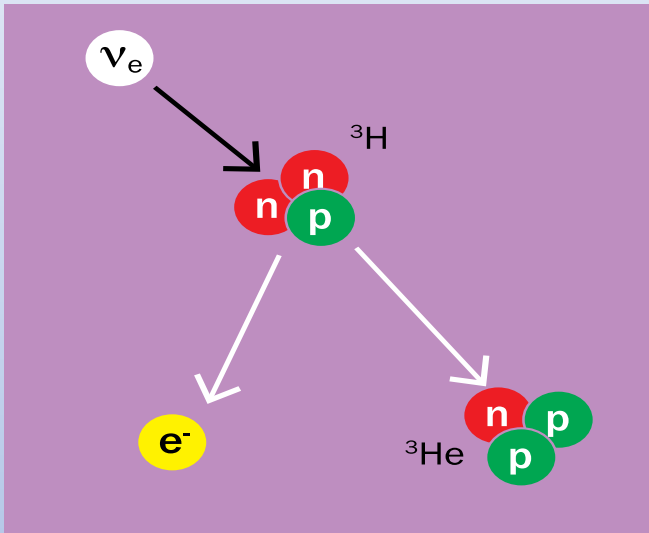
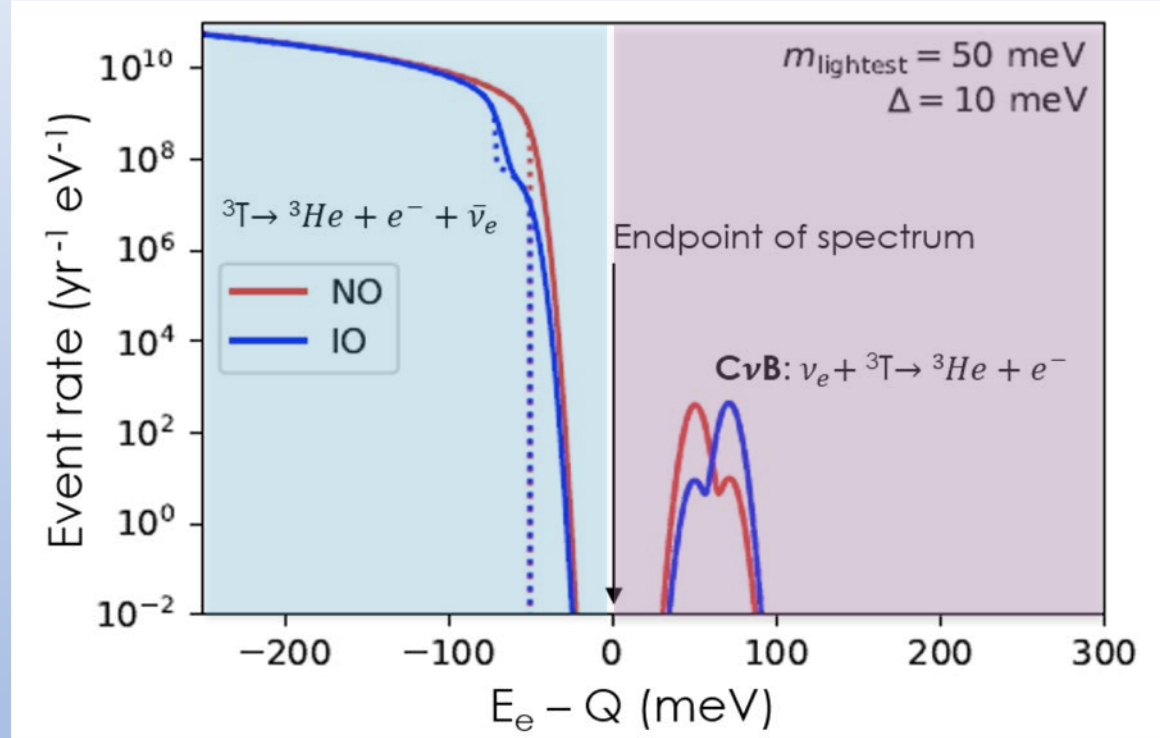
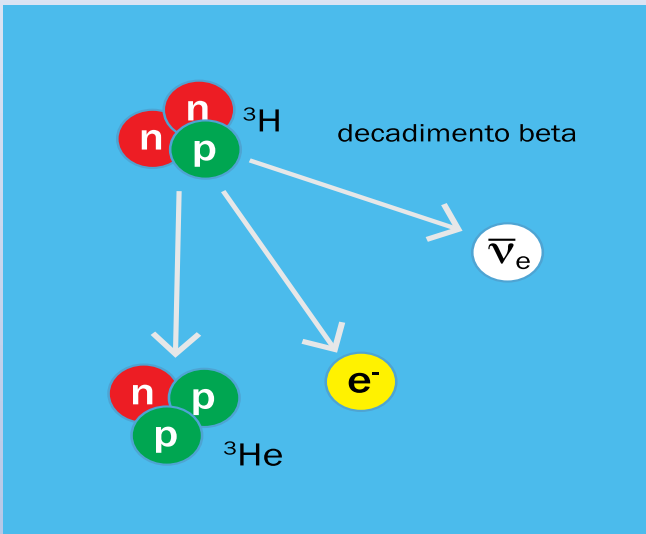
For $m_\nu = 50$ meV,
 $KE = p^2/2m$
 $= 0.17$ meV (0.17 meV/100 meV)
 $= 0.3$ μ eV

Ultra-Cold!



Detection Concept: Neutrino Capture

- Basic concepts for relic neutrino detection were laid out in a paper by Steven Weinberg in **1962** [*Phys. Rev.* 128:3, 1457] applied for the first time to massive neutrinos in **2007** by Cocco, Mangano, Messina [[DOI: 10.1088/1475-7516/2007/06/015](https://doi.org/10.1088/1475-7516/2007/06/015)] and revisited in **2021** by Cheipesh, Cheianov, Boyarsky [<https://arxiv.org/abs/2101.10069>]



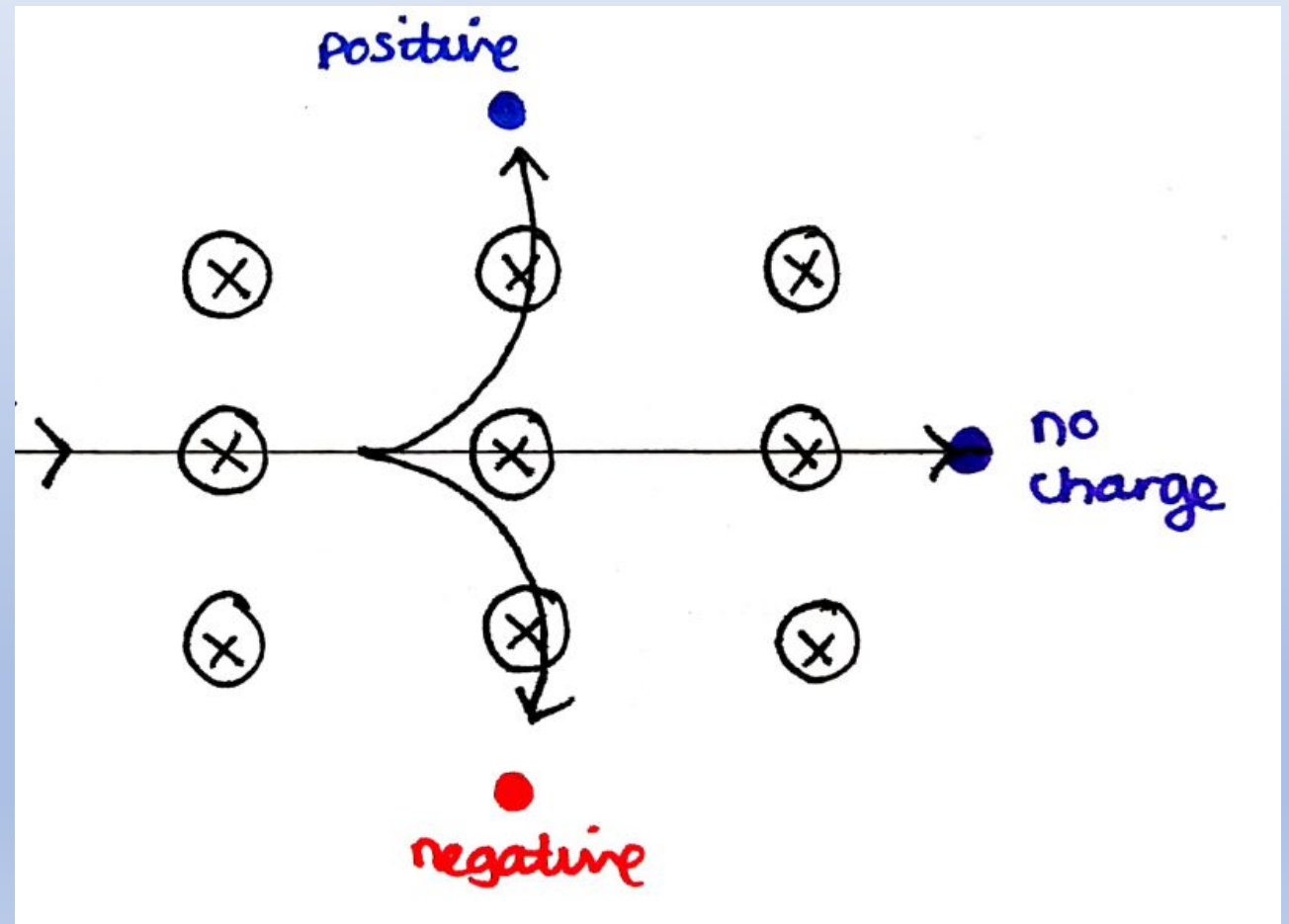
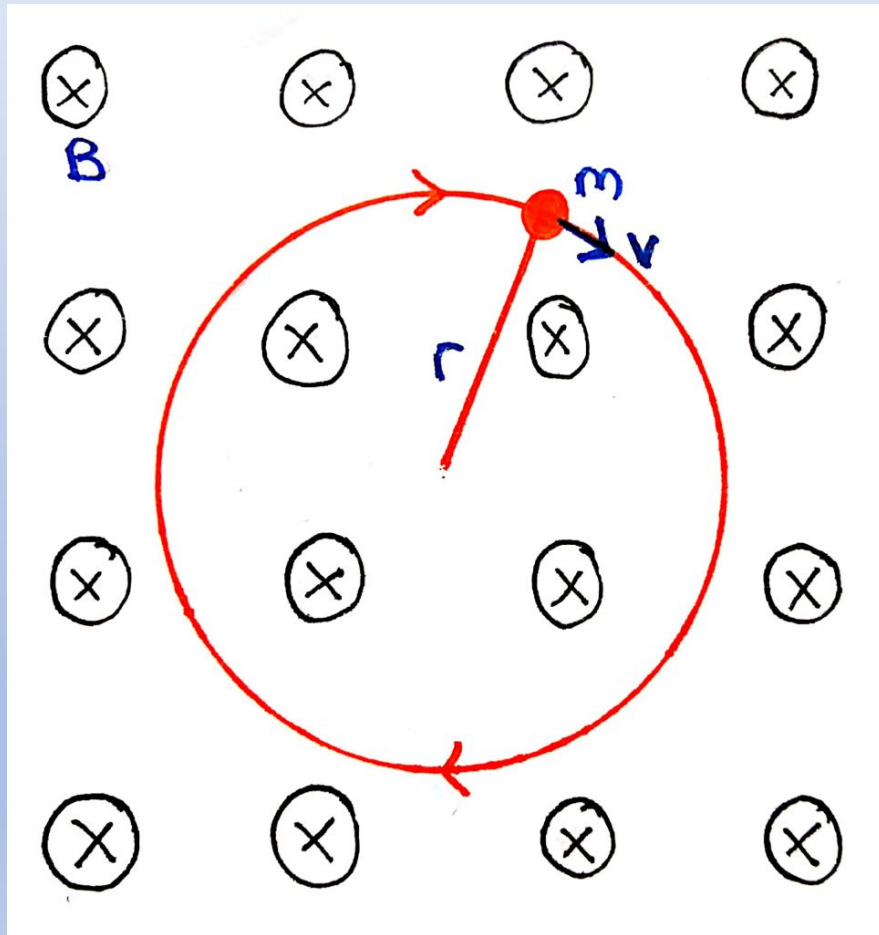
What do we know?
 Gap (2m) constrained to
 $m < \sim 200\text{meV}$
 from precision cosmology
 Electron flavor expected with
 $m > \sim 50\text{meV}$
 from neutrino oscillations

CvB Detection Requires:

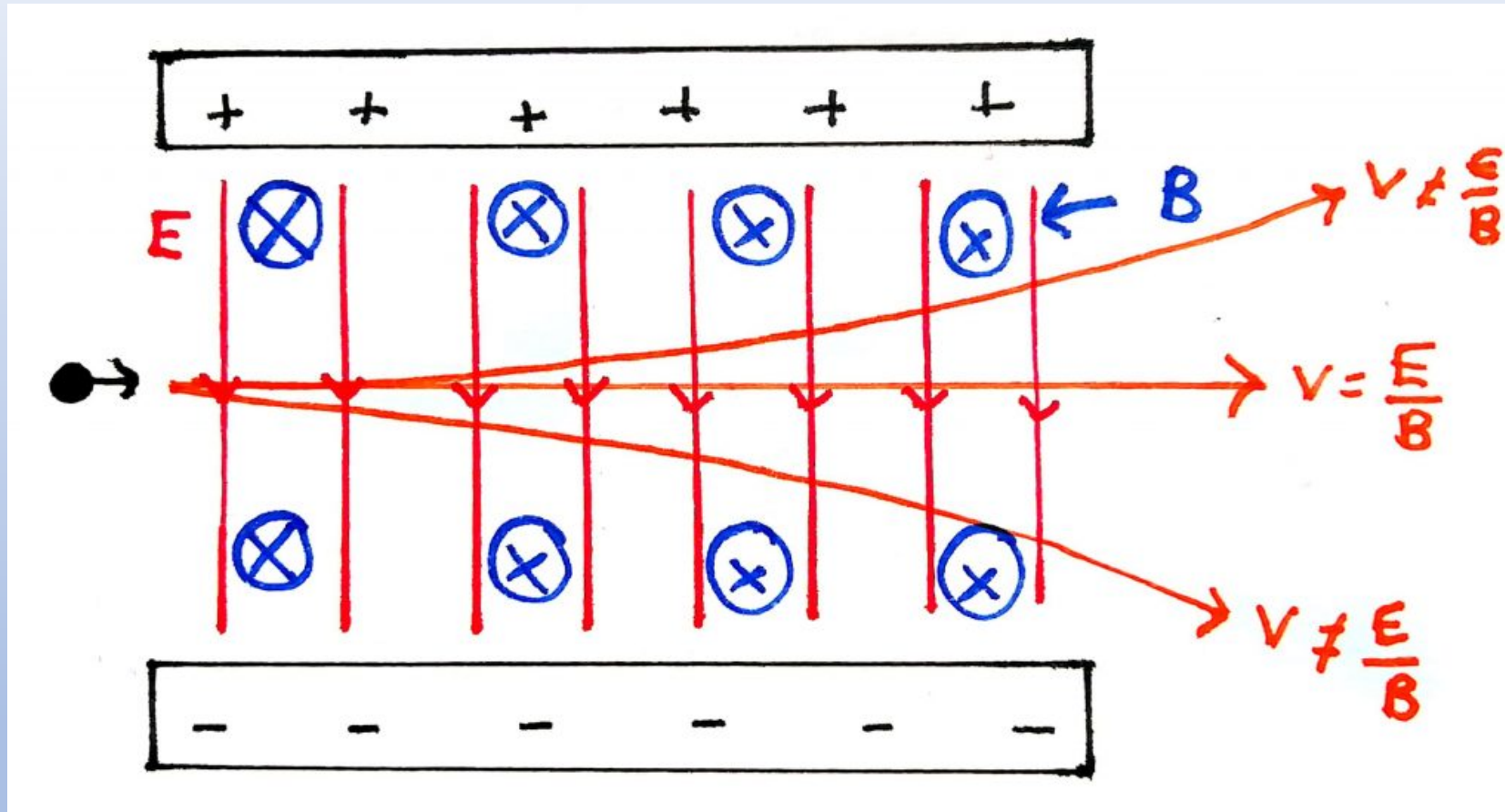
few $\times 10^{-6}$ energy resolution set by m_ν
 KATRIN $\sim 10^{-4}$ (current limitation)

PTOLEMY: $10^{-4} \times 10^{-2}$
 (compact filter) x (microcalorimeter)

Motion of a Charged Particle in a B Field



Classic Velocity Selector

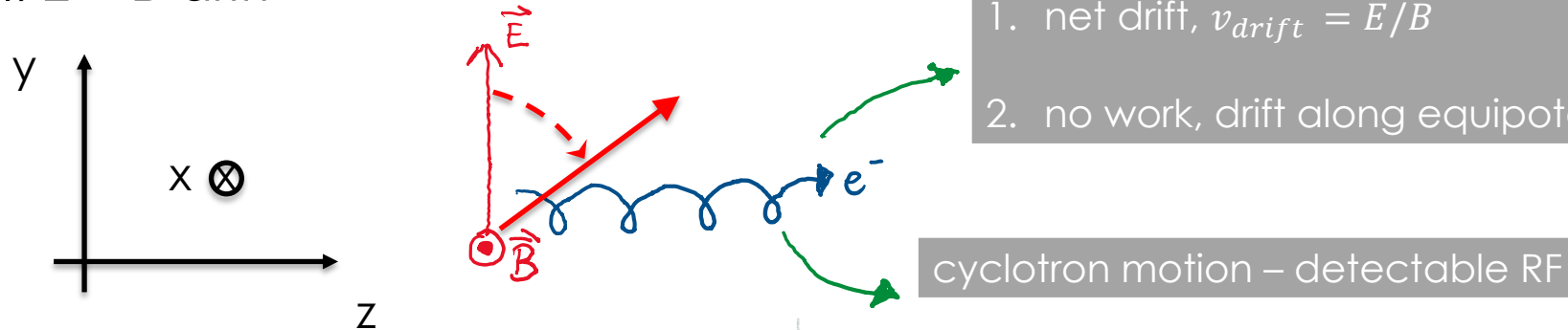


Is this the only way to select velocity?

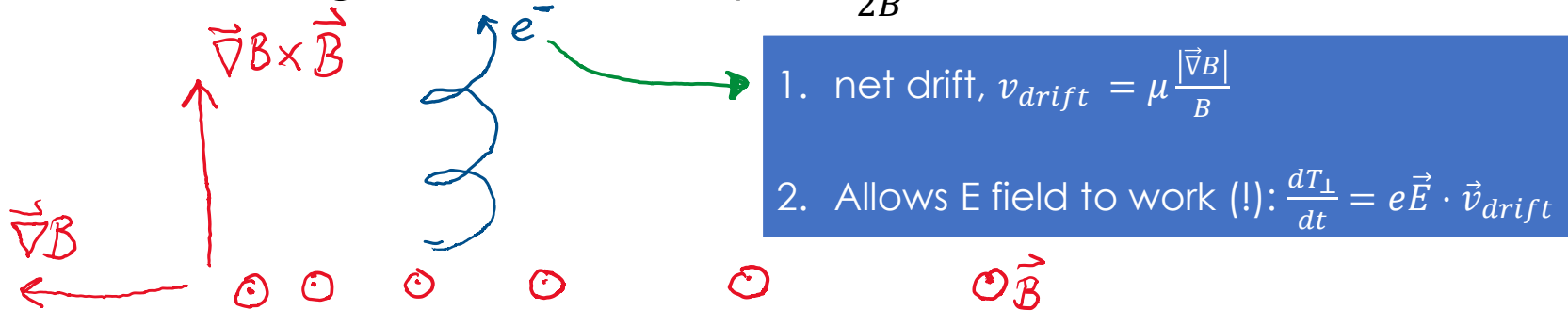
PTOLEMY Filter Concept

Auke Pieter Colijn (PATRAS 2019)

I: $\vec{E} \times \vec{B}$ drift



II: $\frac{\mu}{B^2} \vec{\nabla} B \times \vec{B}$ drift, with magnetic moment $\mu = \frac{m_e v_{\perp}^2}{2B}$



$$\mathbf{V}_{E \times B}^y(z)|_{x,y=0} = \frac{\mathbf{E} \times \mathbf{B}}{B_x^2} = \frac{E_z B_x \hat{y}}{B_x^2} = \frac{E_z}{B_x} \hat{y}$$

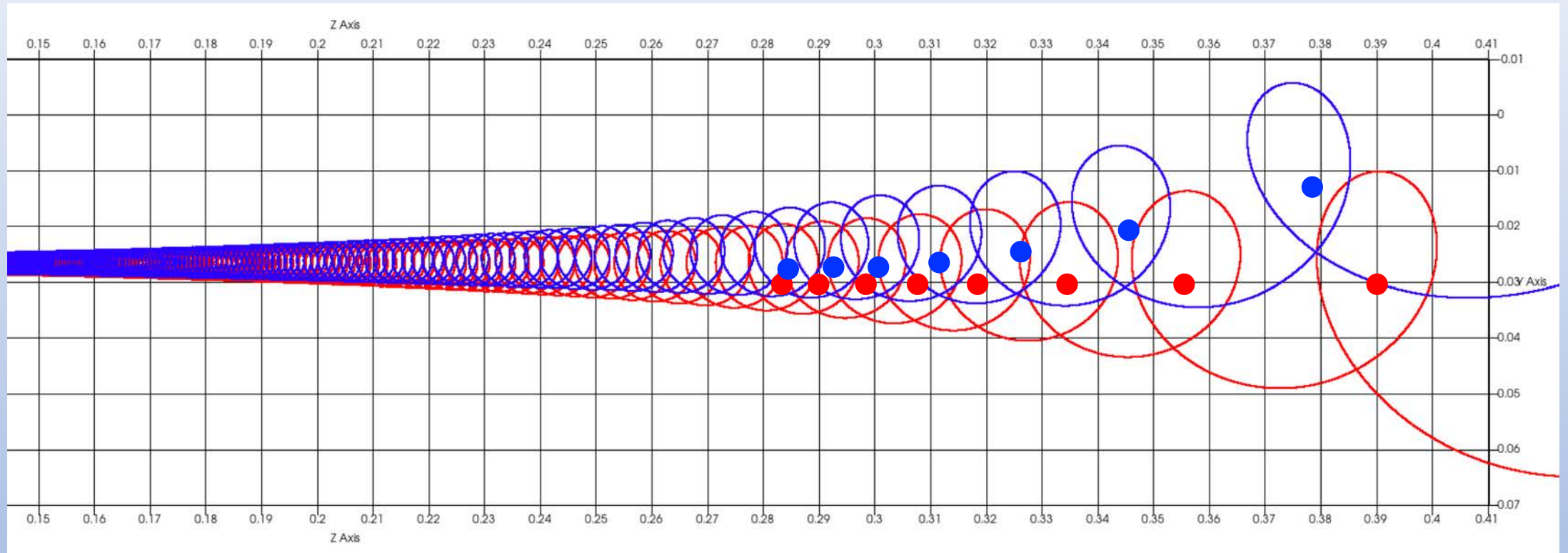
$$\mathbf{V}_{\nabla B}(z)|_{x,y=0} = -\frac{\mu \times \nabla_{\perp} \mathbf{B}(z)}{qB(z)} = -\frac{\mu}{qB_x} \frac{dB_x}{dz} \hat{y}$$

Enforce zero drift in y (rotate E):

yields

$$E_z(z)|_{y=0} = -\frac{\mu}{q} \frac{dB_x(z)}{dz}$$

Bingo!



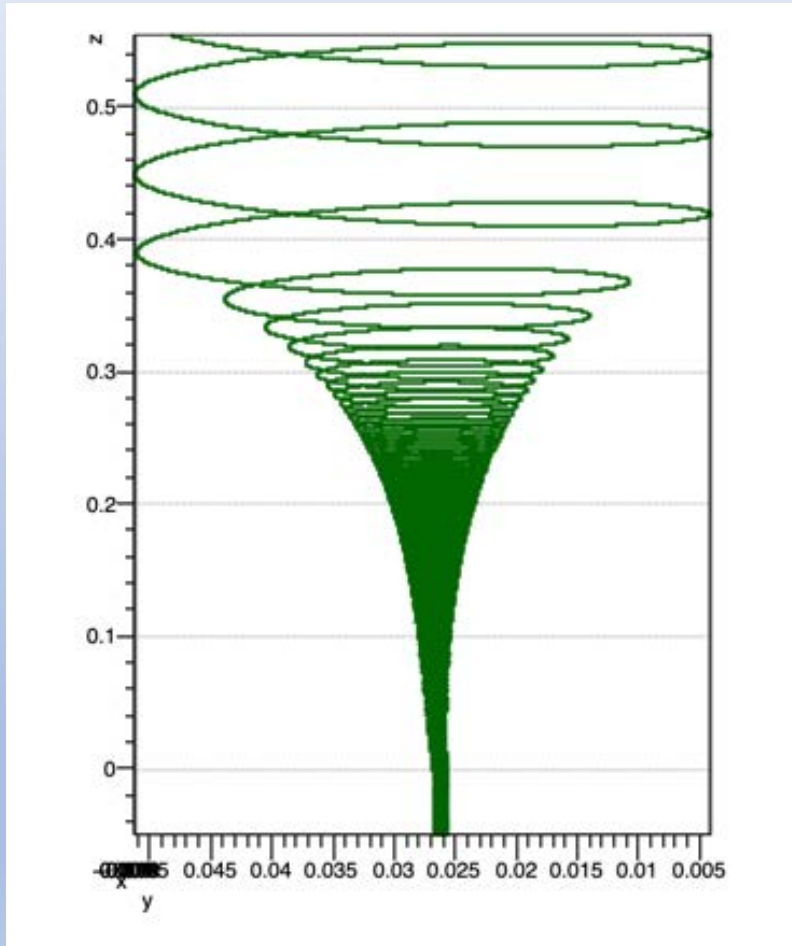
Selected velocity based on cyclotron drift

New type of particle accelerator (useful for fusion reactor heating) <https://www.intechopen.com/chapters/82927>

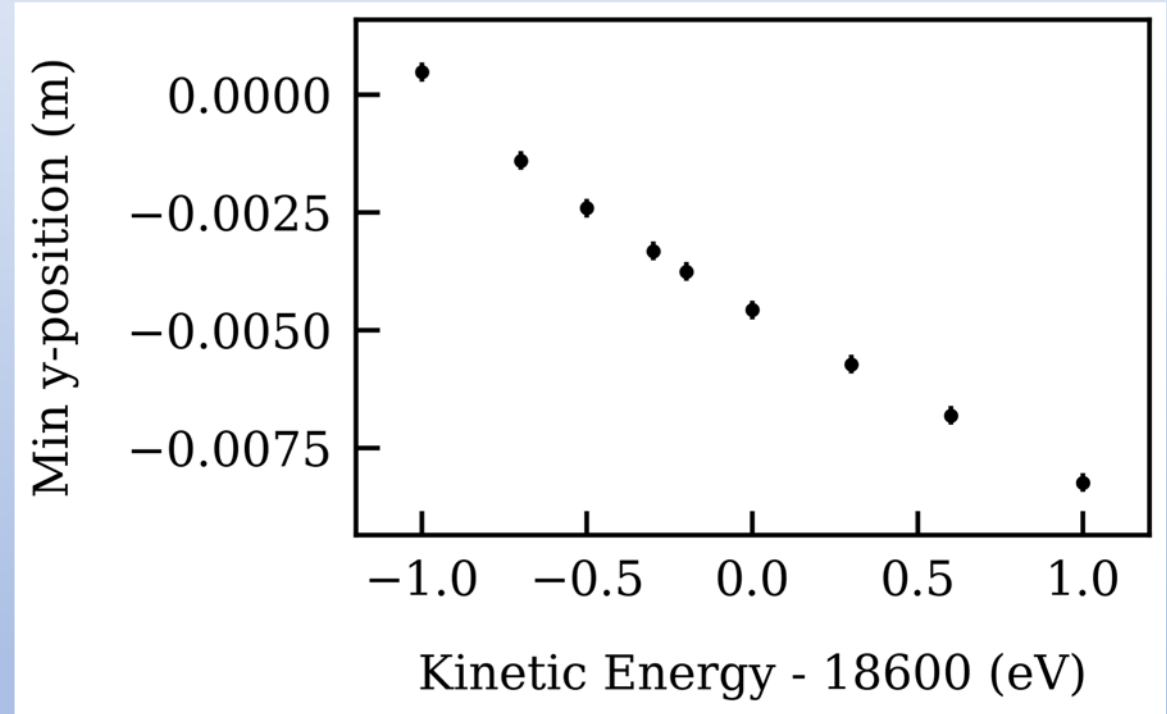
How well can one select velocities?



400mm



45mm

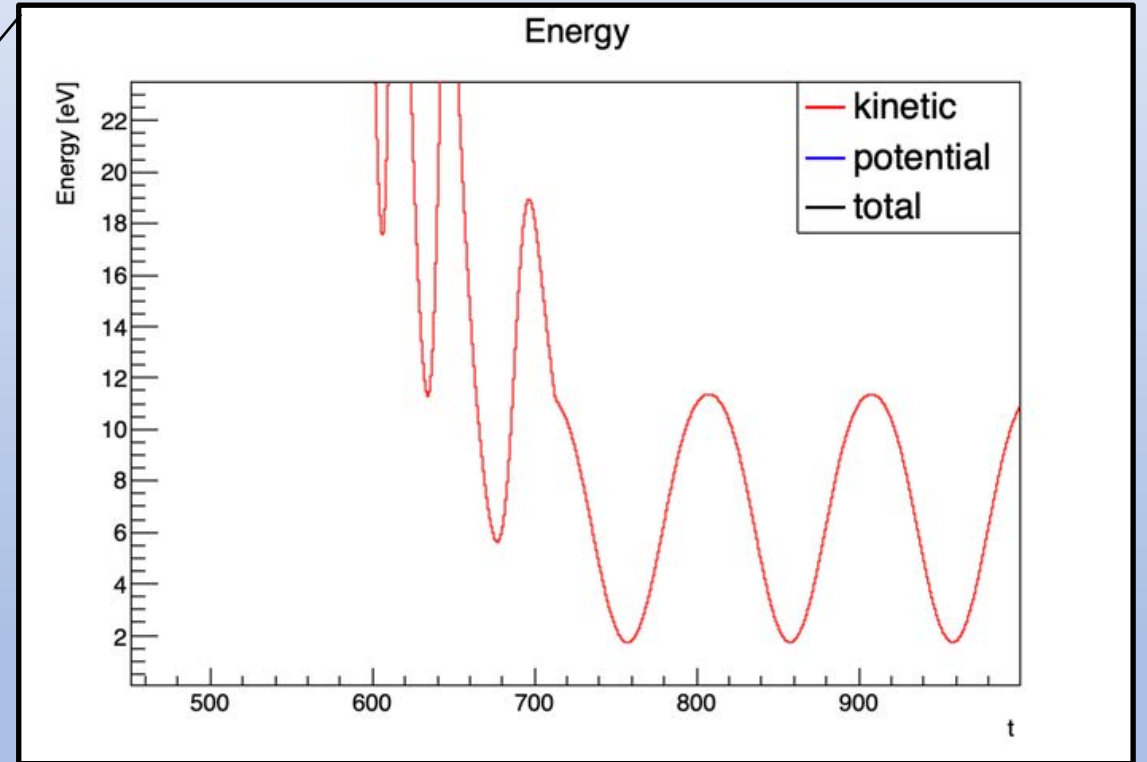
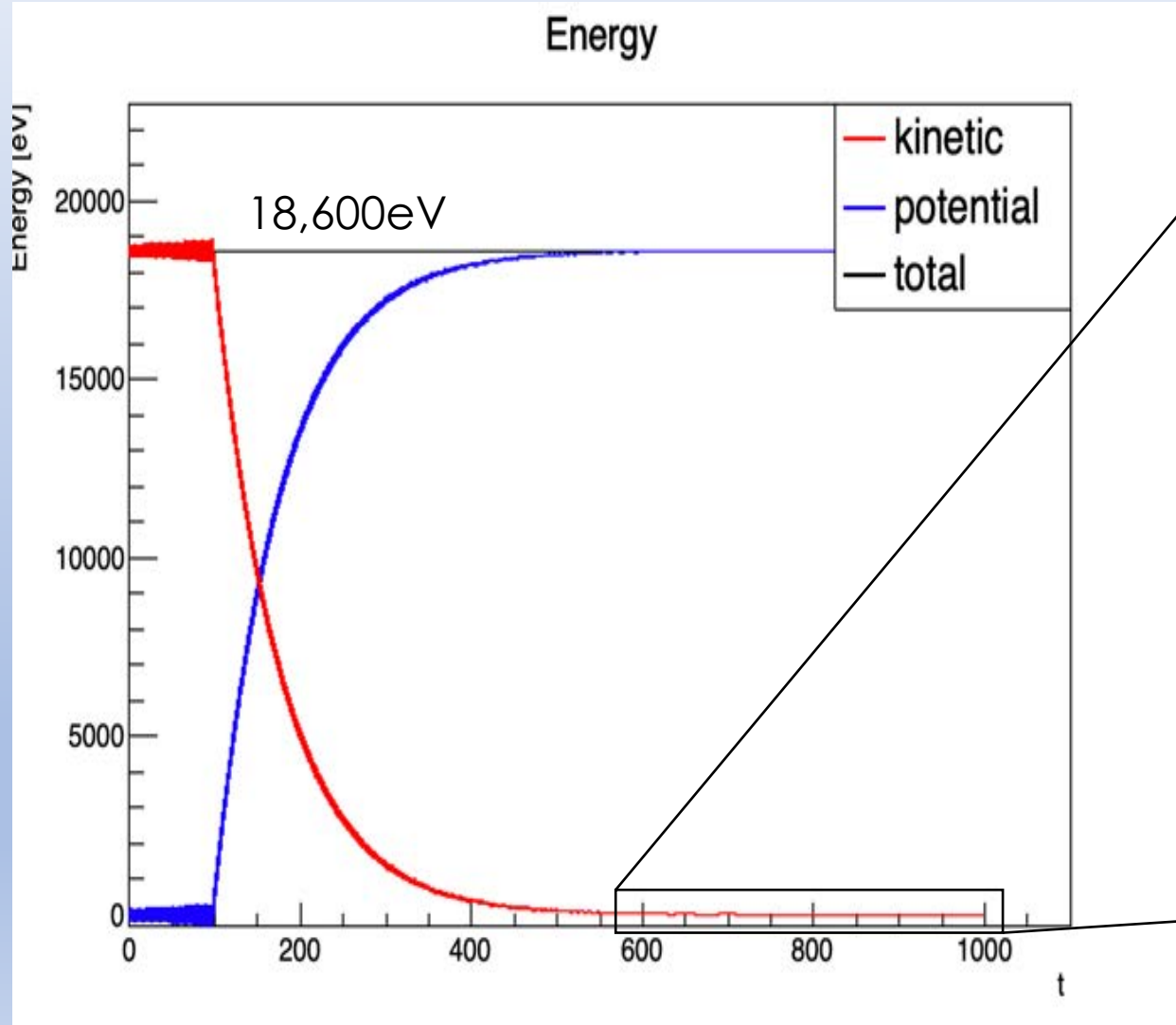


Deflection is $\sim 4\text{mm/eV}$
(over a drift distance of 400mm)

Lorentz4 Code

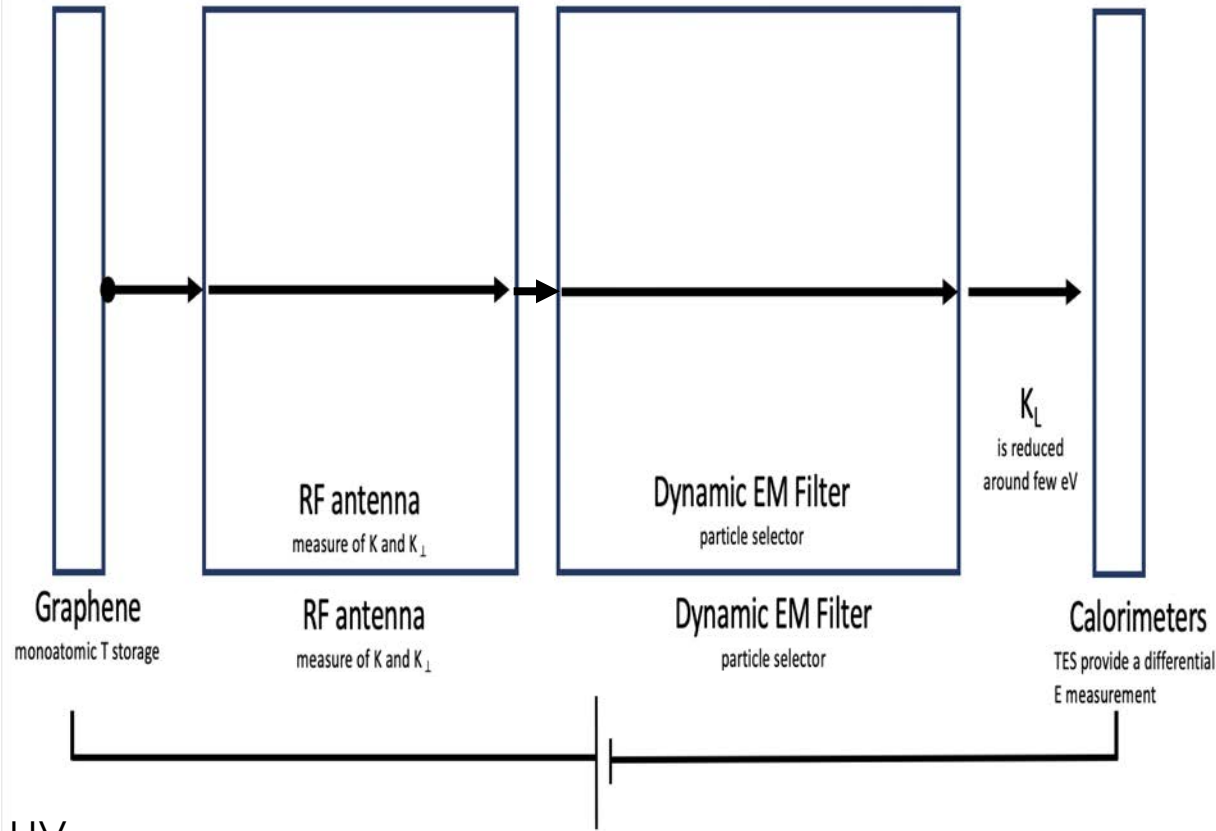
<https://github.com/gkrossi/lorentz4>

LNGS software
Nicola Rossi
Marcello Messina

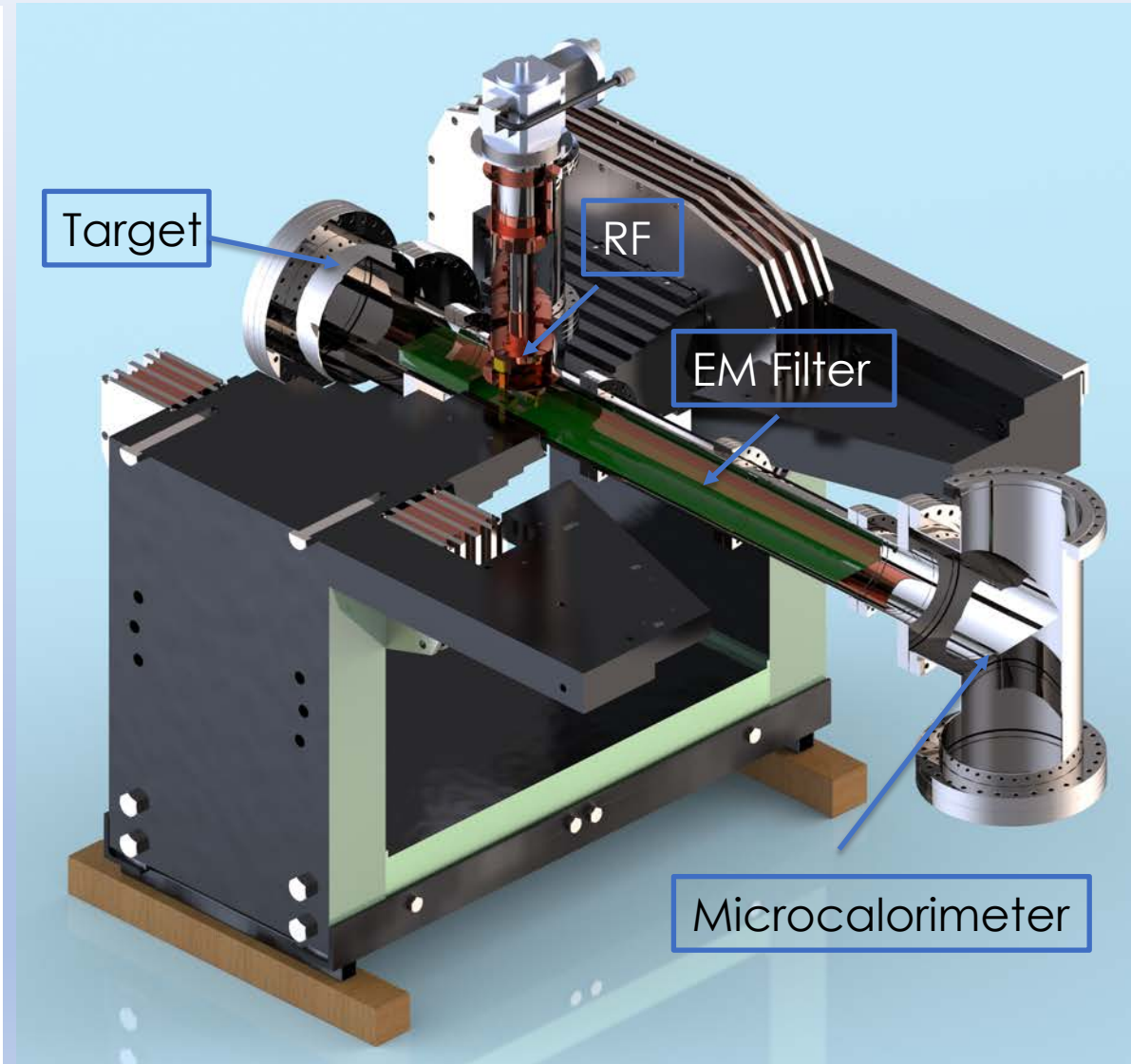


PTOLEMY R&D Development Setup

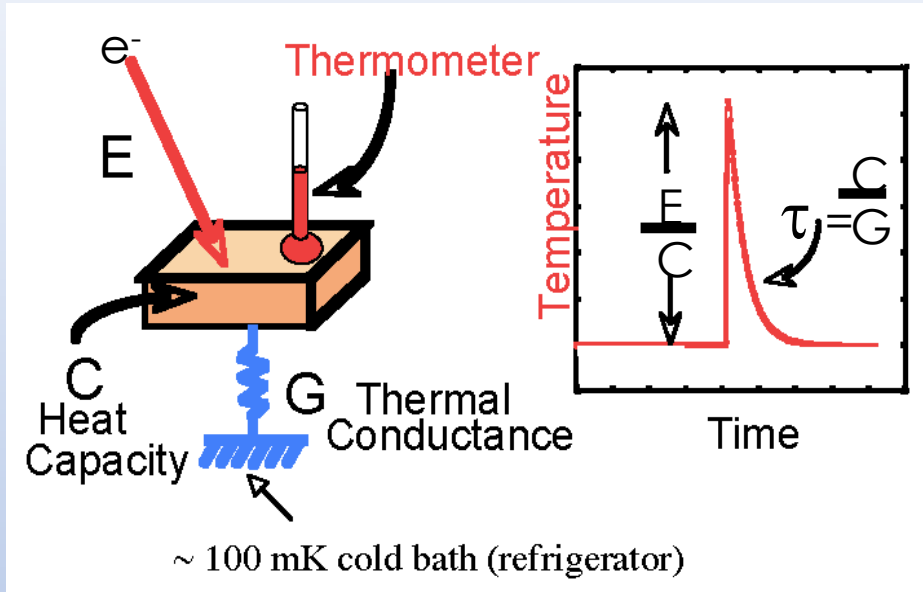
Andi Tan (Princeton)



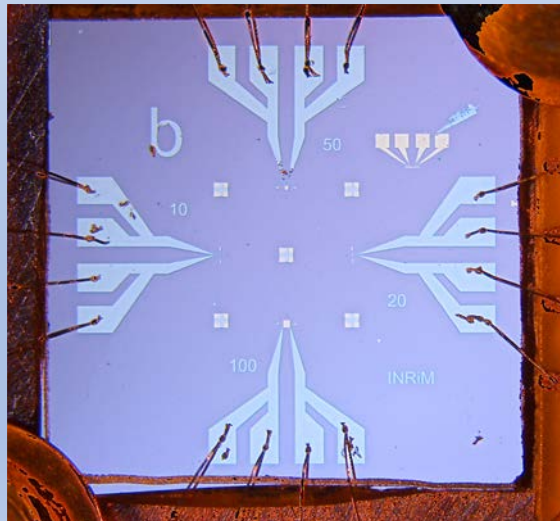
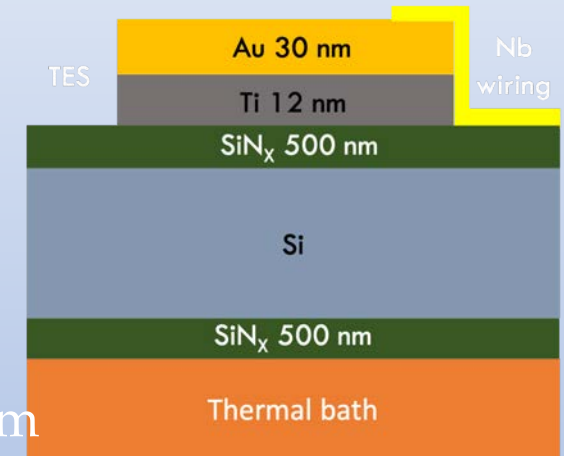
$$E_{Total} = q(V_{TES} - V_{Target}) + E_{RFcorr} + E_{cal}$$



Measurement Arm: μCal



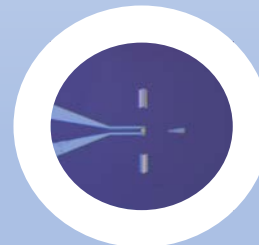
Thin sensors:
~1 eV electron can be stopped with very small C



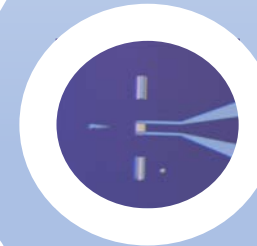
10x10 μm



50x50 μm



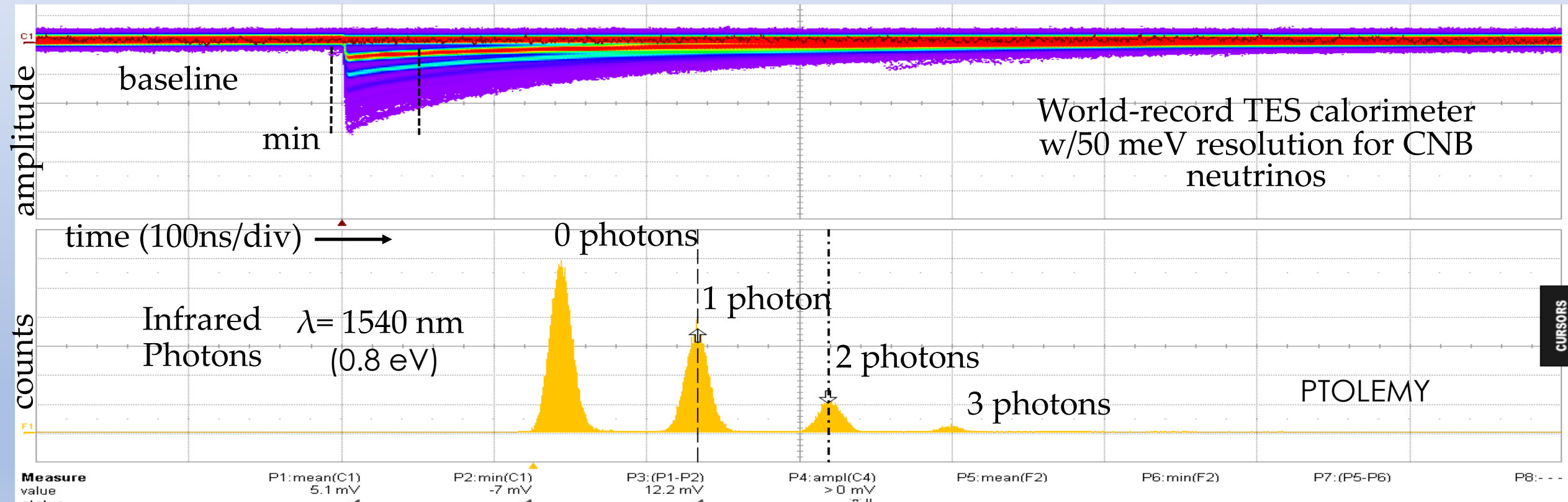
100x100 μm

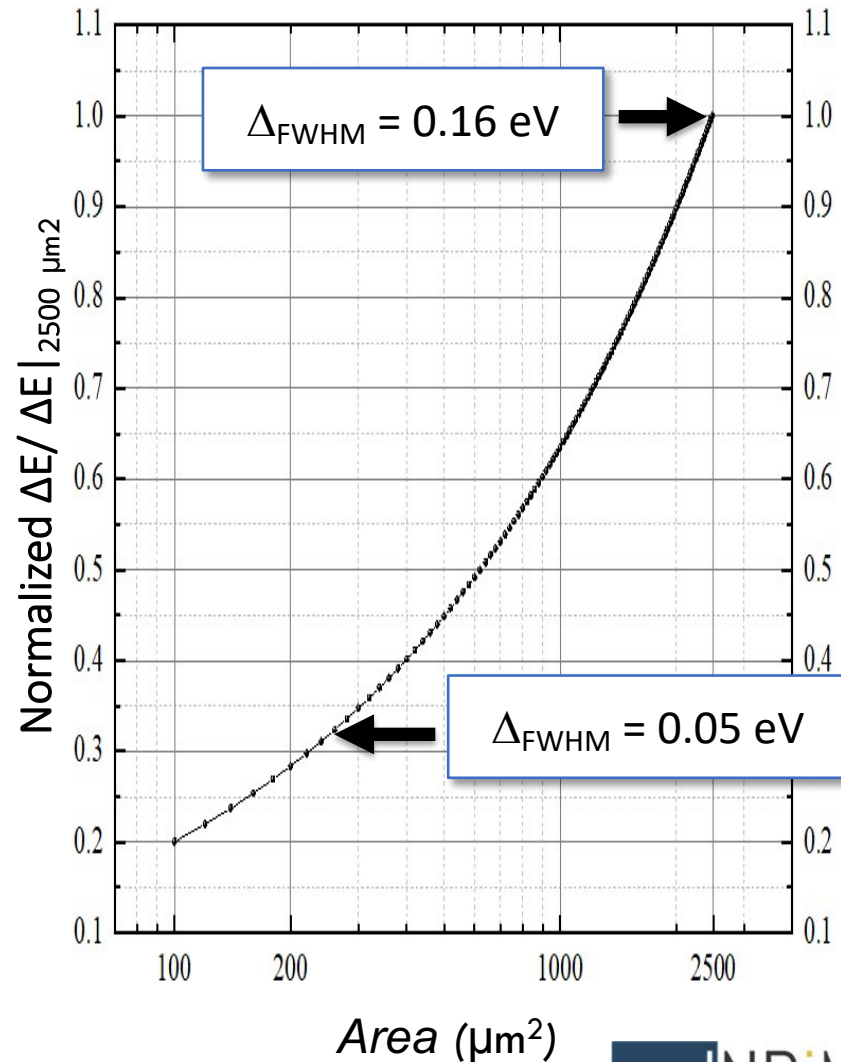
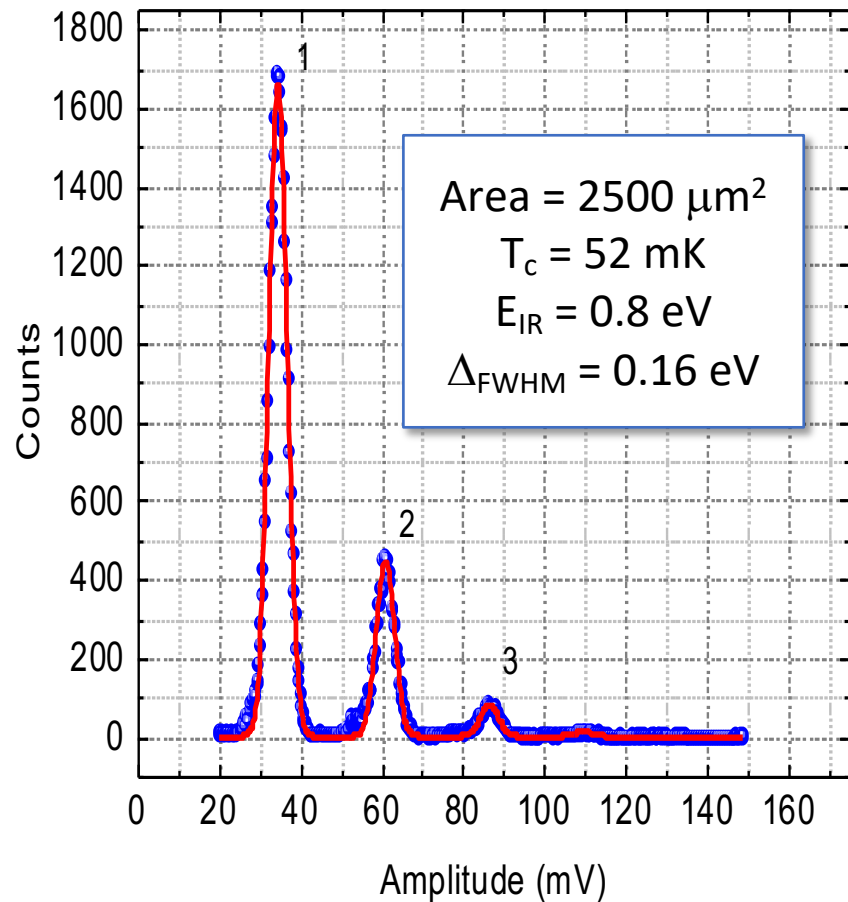


20x20 μm

C. Pepe, E. Monticone, M. Rajteri

1% energy resolution at optical photon energies, i.e. measures the wavelength of a 500nm photon to a few nm





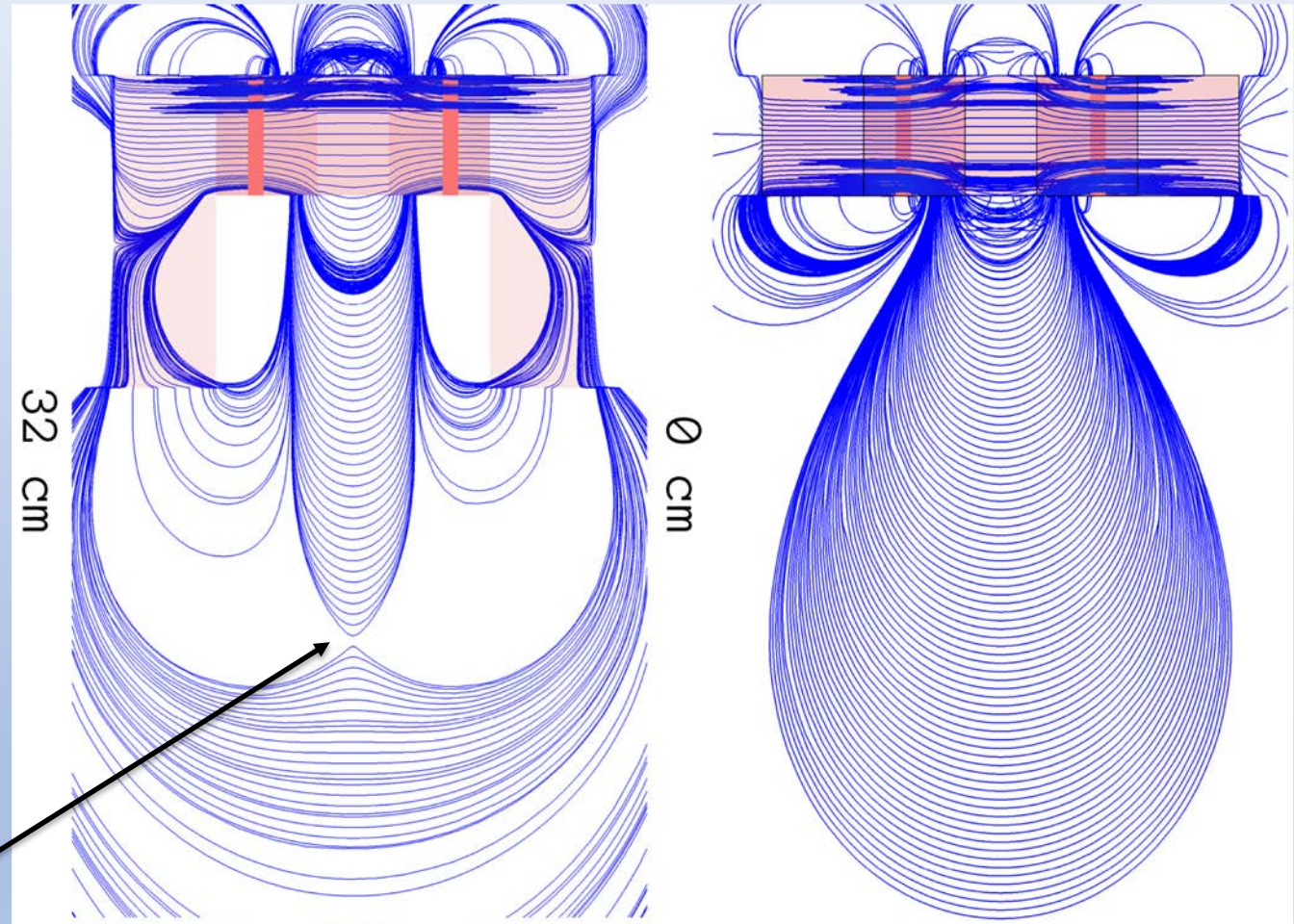
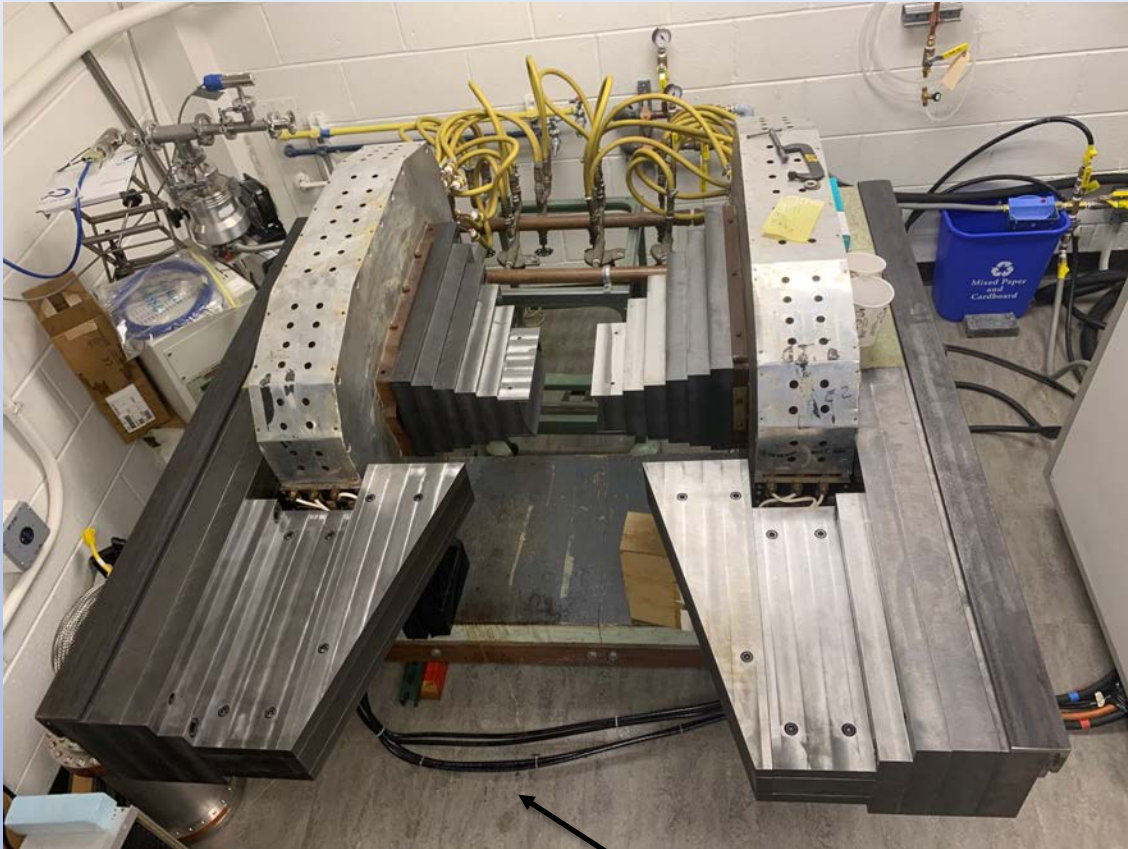
Resolution of $\sim m_v$:
 Area $\sim 15 \mu\text{m} \times 15 \mu\text{m}$

→ Demonstrate with electrons

First Version of the PTOLEMY filter

PTOLEMY
filter@Princeton

Wonyong
Chung

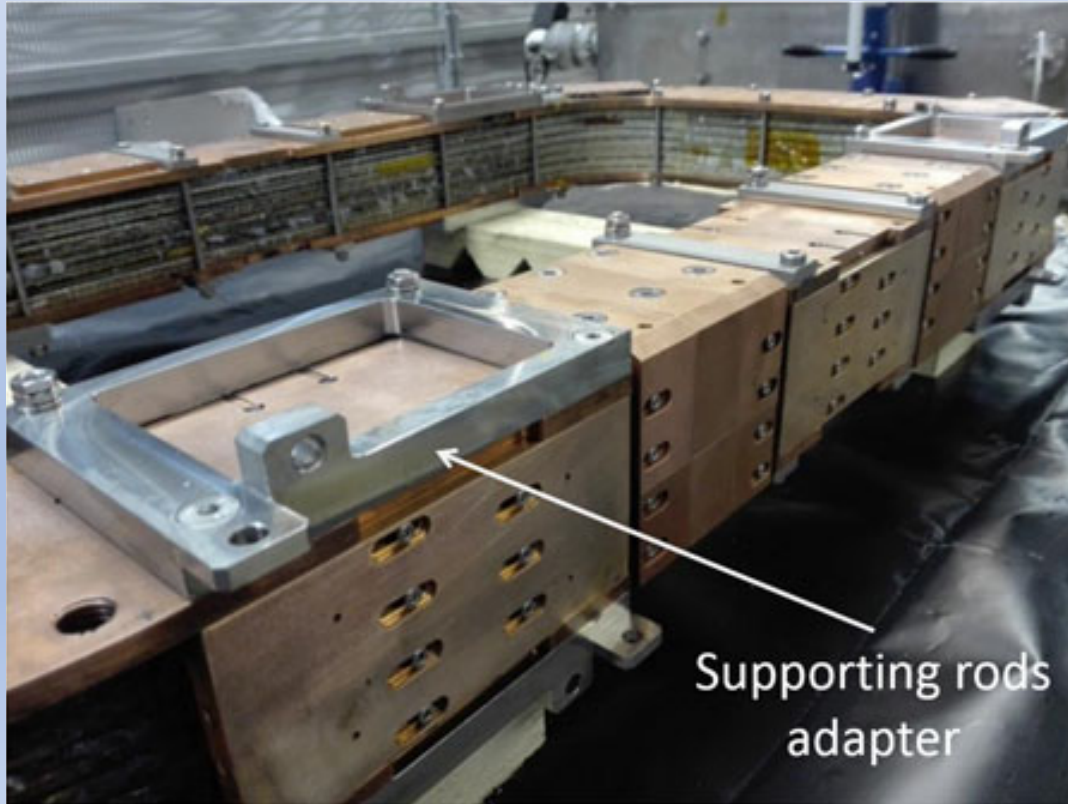


Andi
Tan

Zero field (location for TES microcalorimeter)

Conduction-Cooled Superconducting Coils

- LNGS magnet specifications within ~20% of a Wind Generator system made by ANSALDO designed in Spain (~10 Open MRI similar commercial systems sold per year)
 - This is the preferred option and reduces 70kW → 10kW power



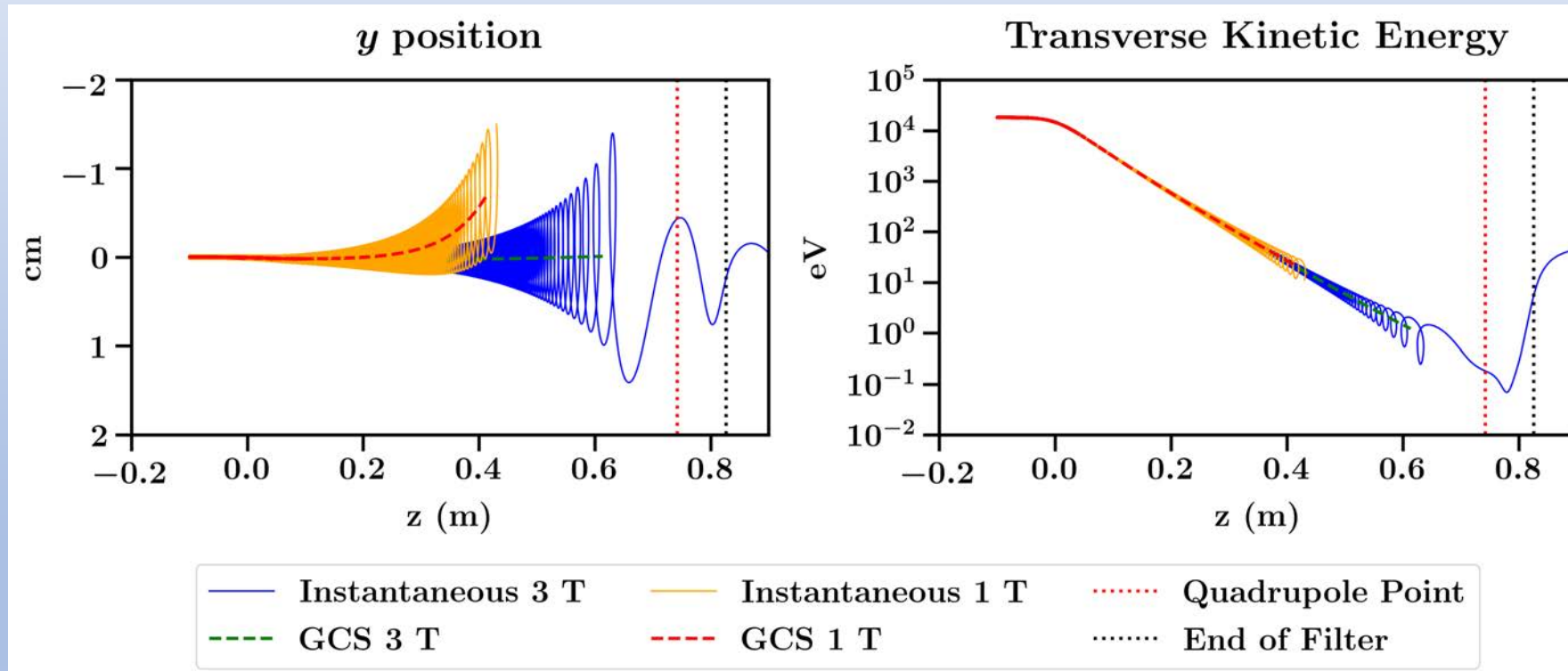
Packed with multi-layer thermal insulation

Filter Performance

Improves as B^2 for a fixed filter dimension

18.6 keV @ 1T \rightarrow \sim 10eV (in 0.4m)

18.6 keV @ 3T \rightarrow \sim 1eV (in 0.6m)



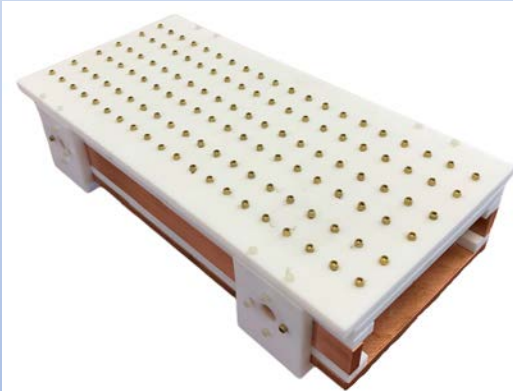
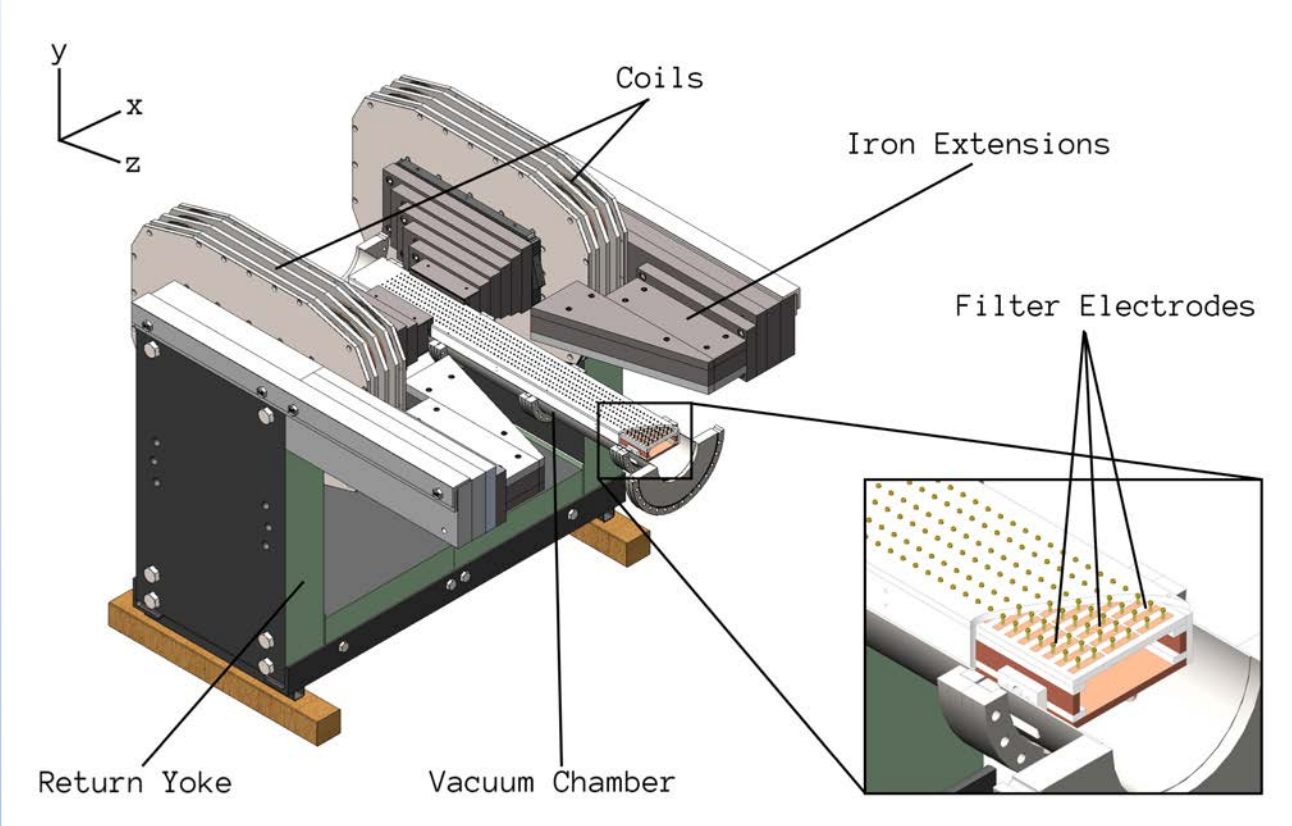
PTOLEMY Collaboration, <https://arxiv.org/abs/2108.10388>

"Implementation and Optimization of the PTOLEMY Electromagnetic Filter"

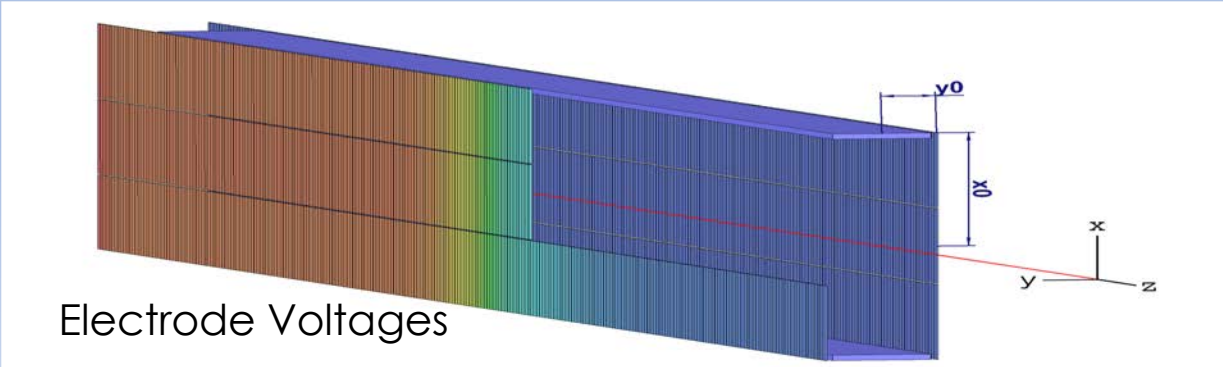
<https://iopscience.iop.org/article/10.1088/1748-0221/17/05/P05021>

Electrode Prototype

Andi Tan (Princeton)

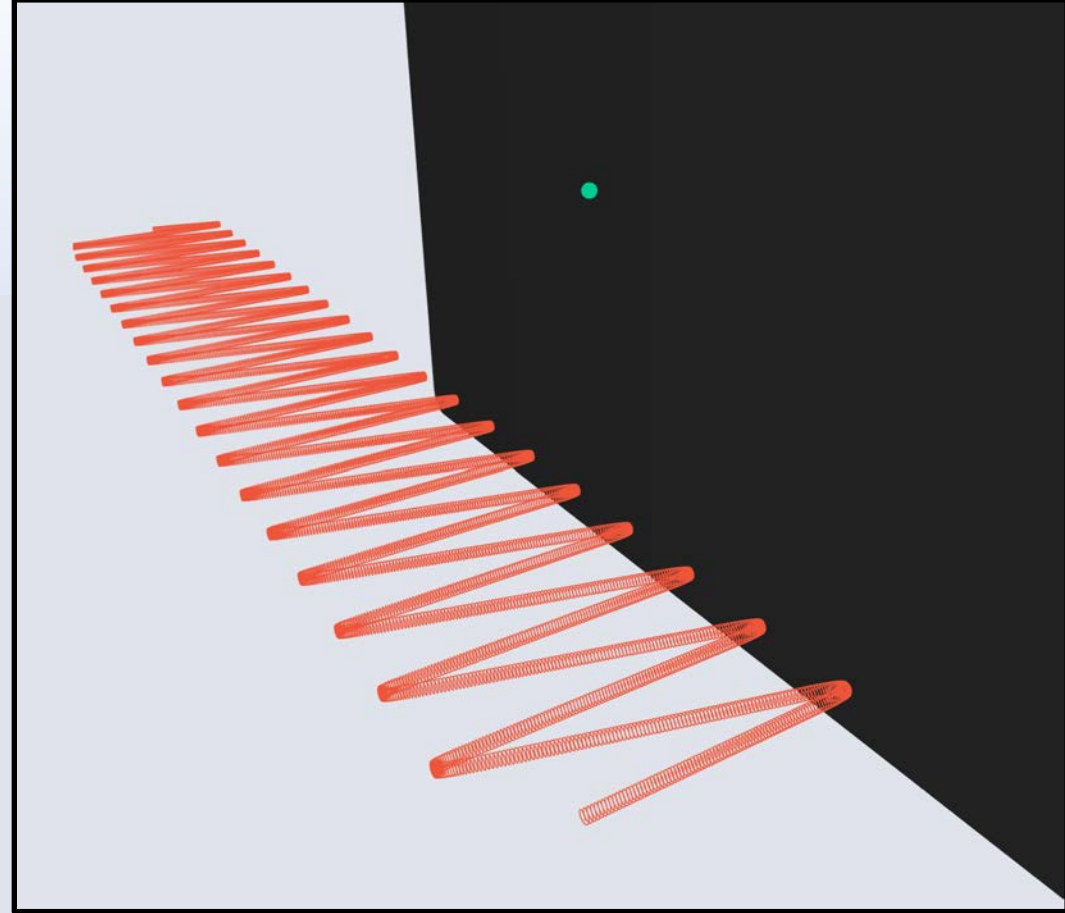
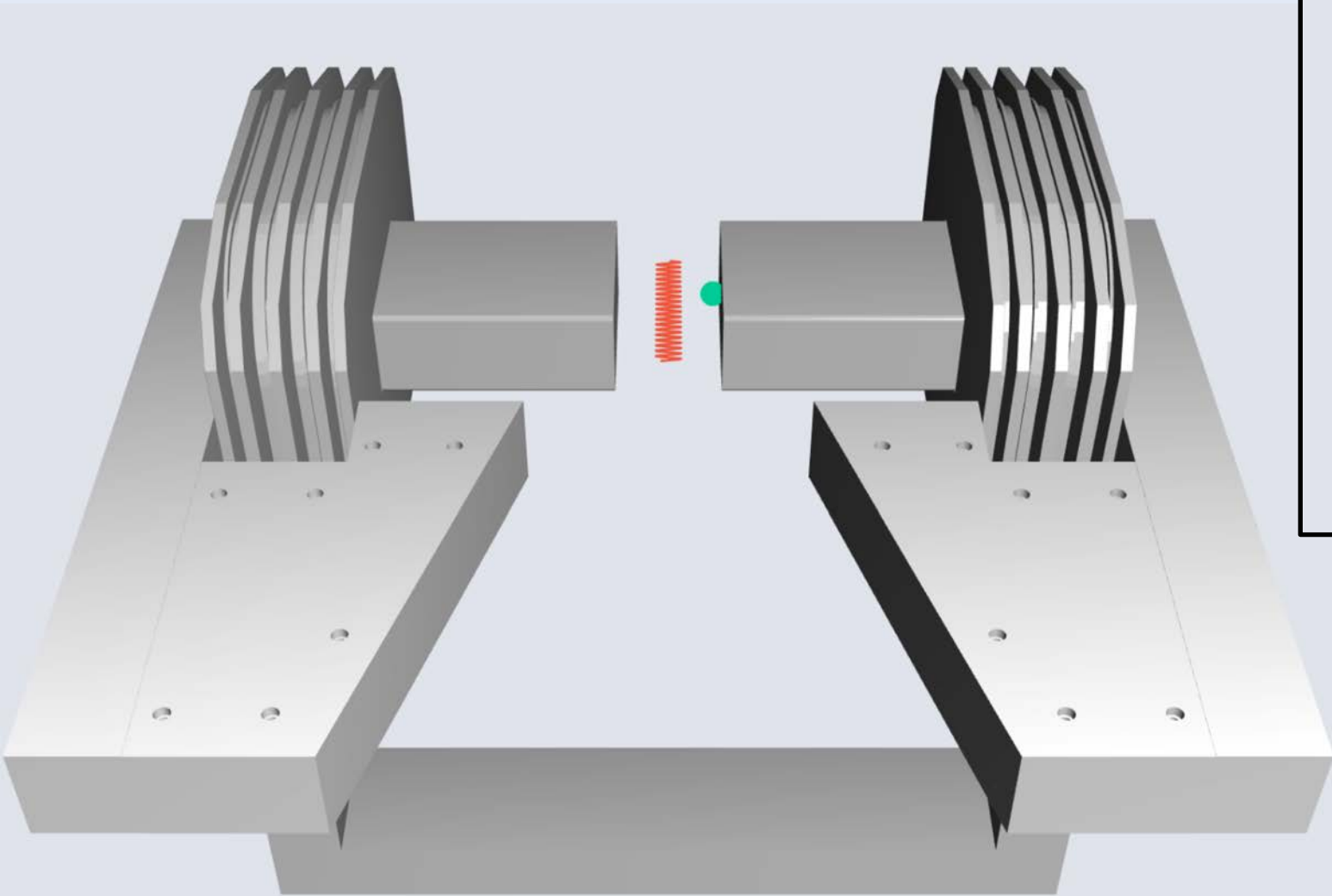


Wonyong
Chung
(Princeton)

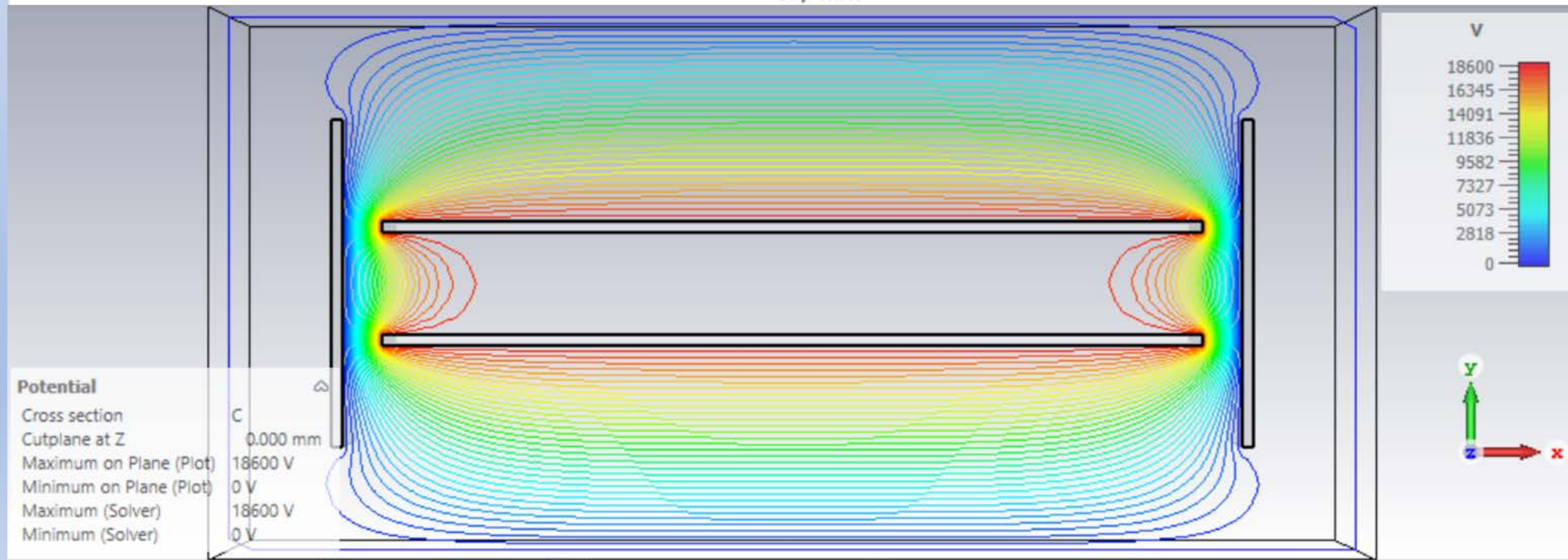
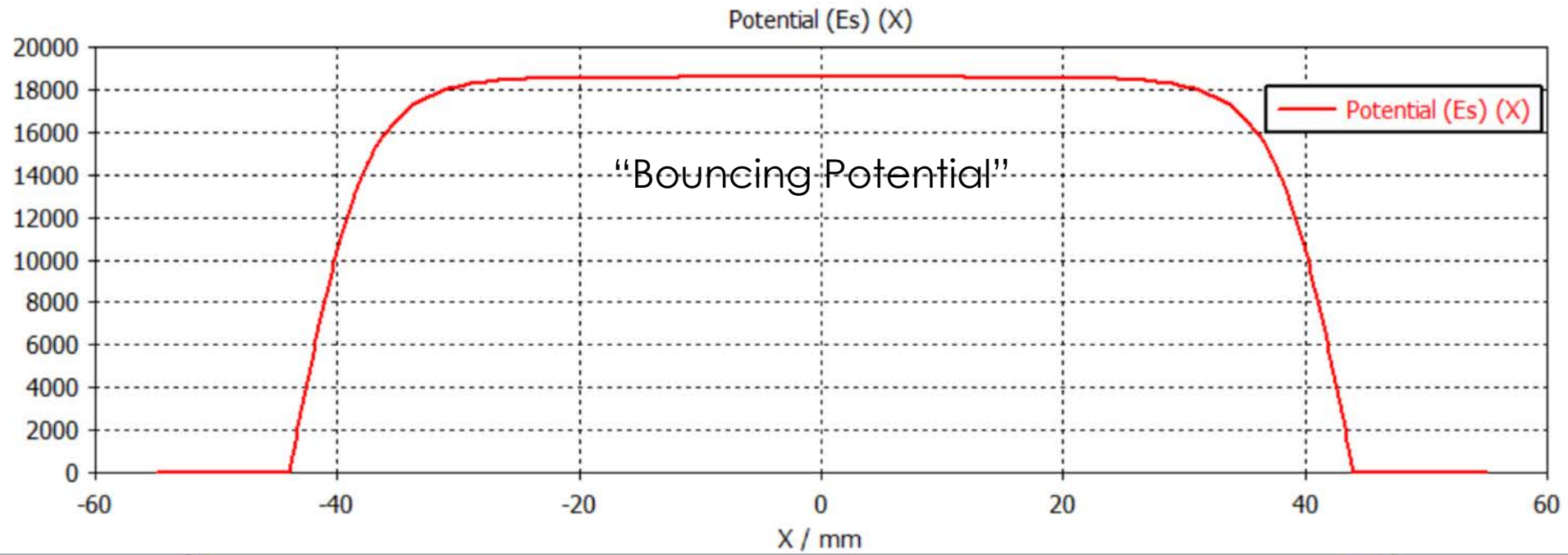


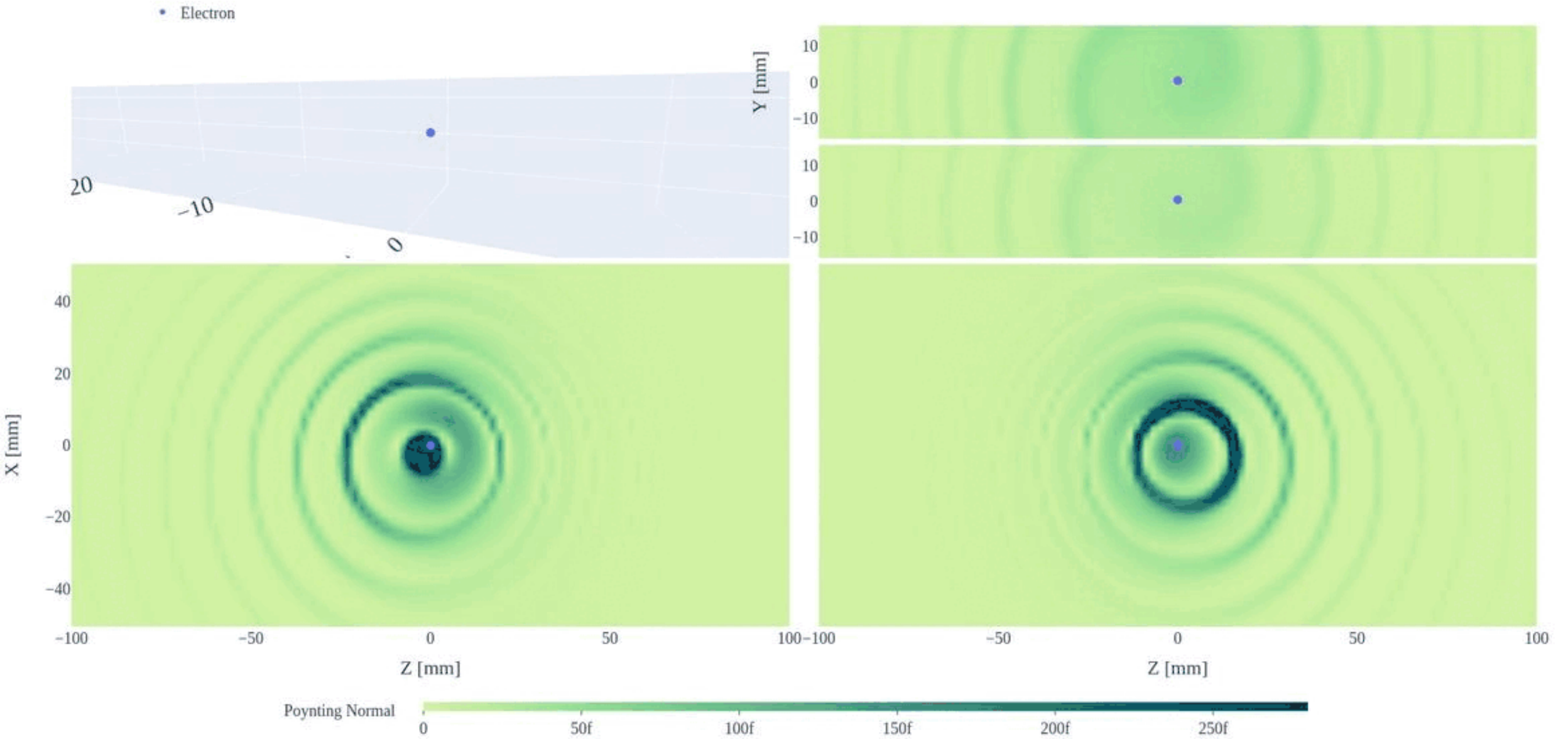
Electrode Voltages

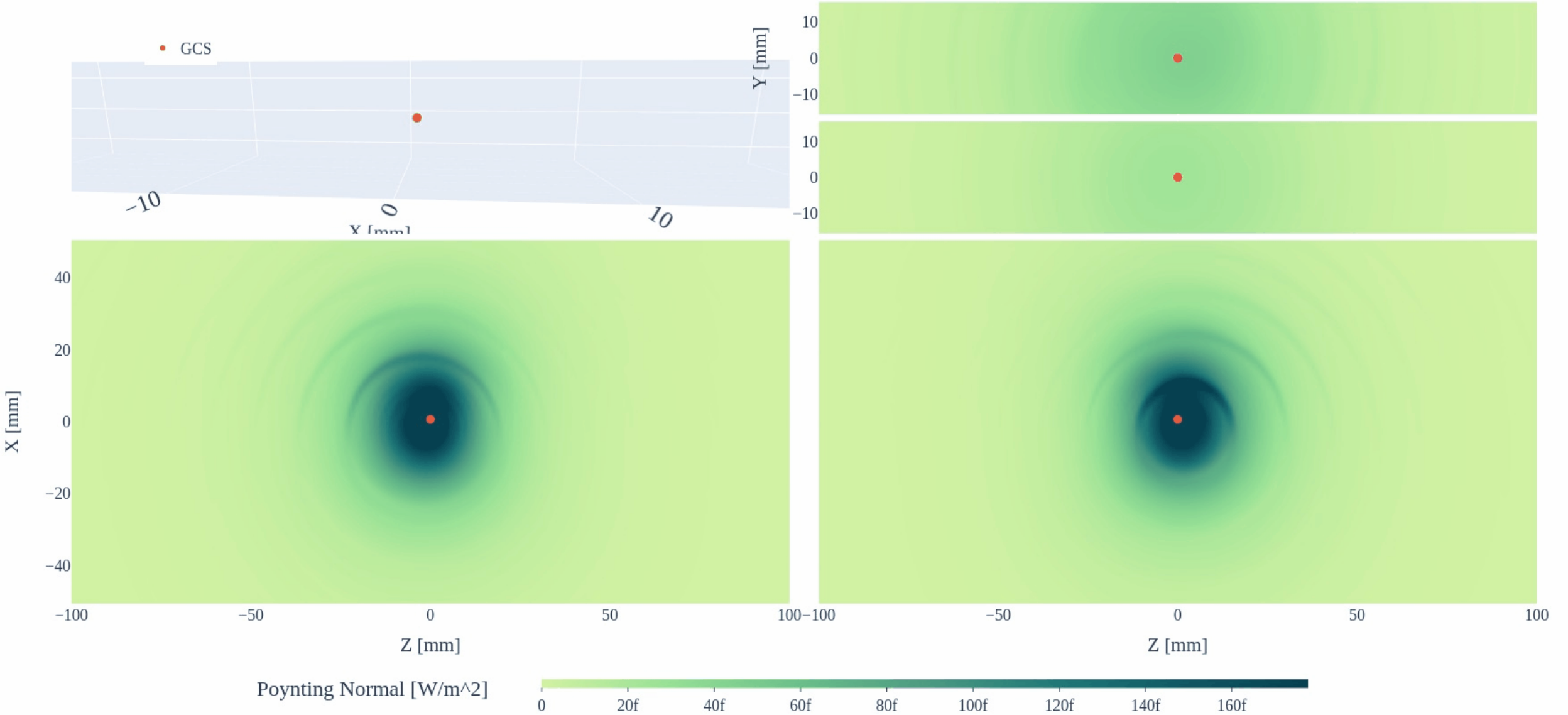
Pitch 85 Long Trajectory



Andi Tan (Princeton)

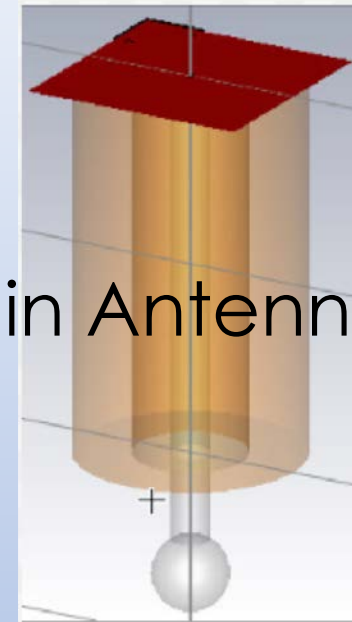
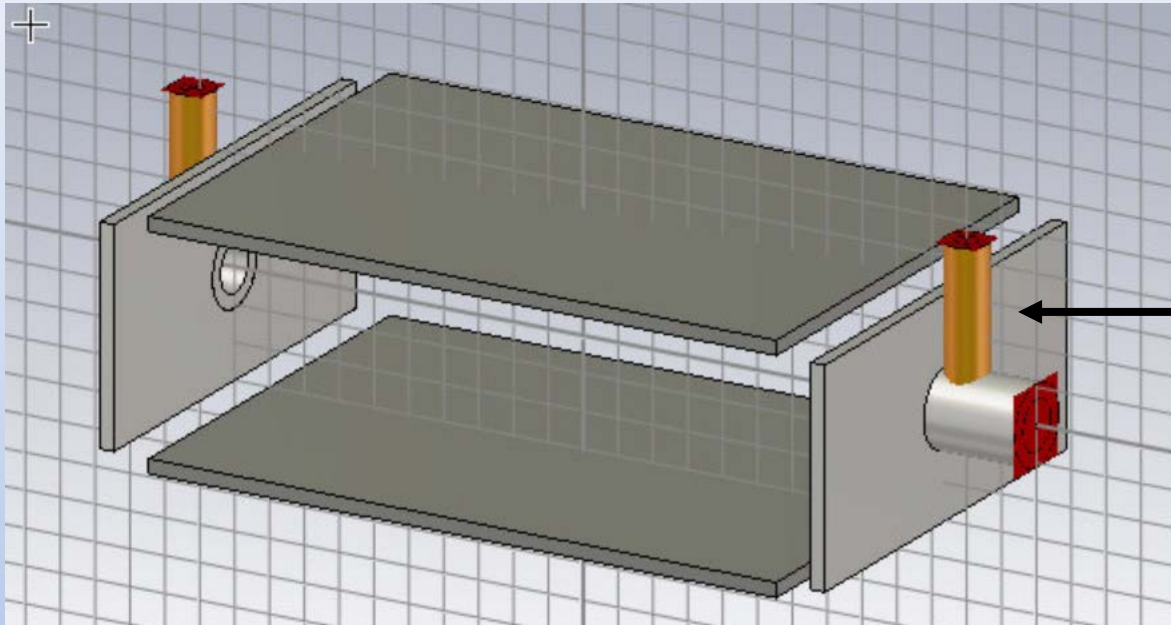






Antenna Design Studies

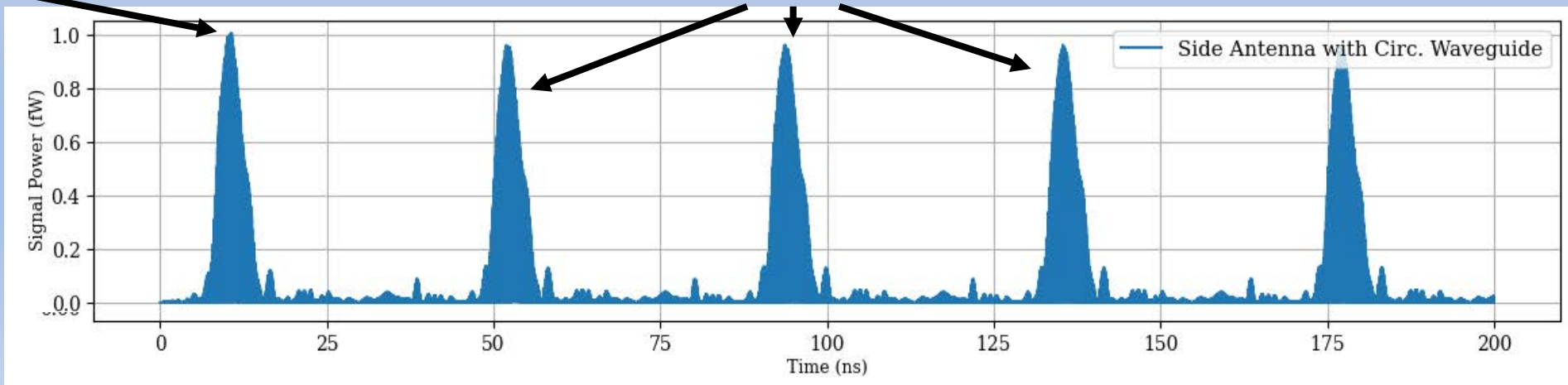
Yuno Iwasaki (Princeton)

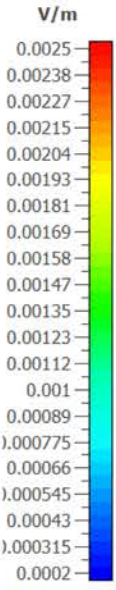
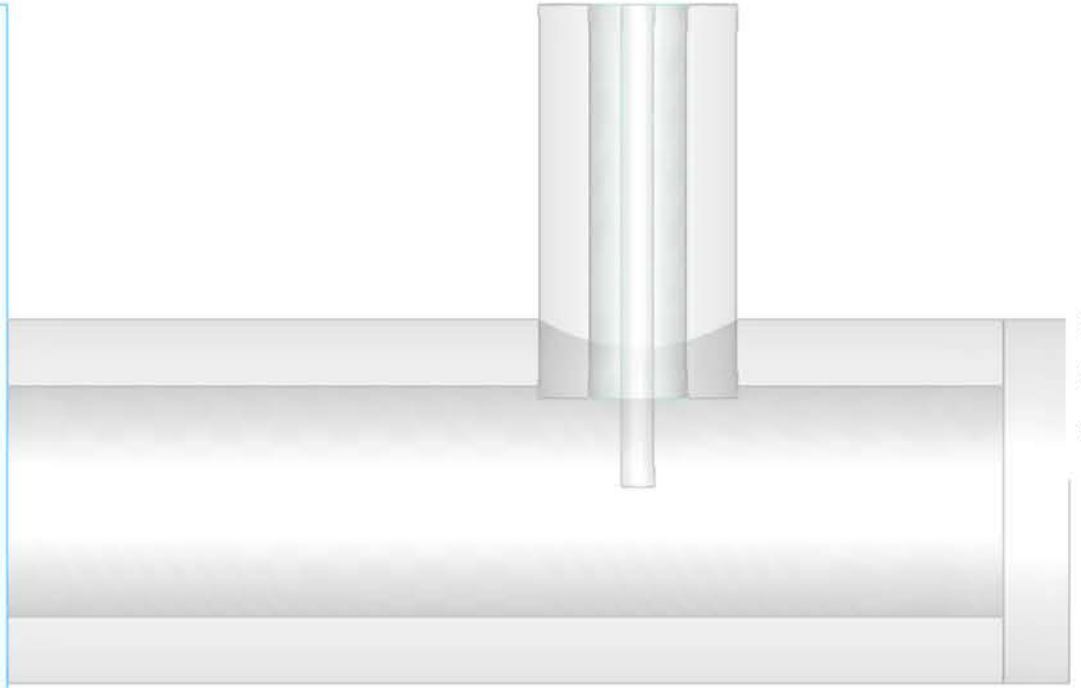
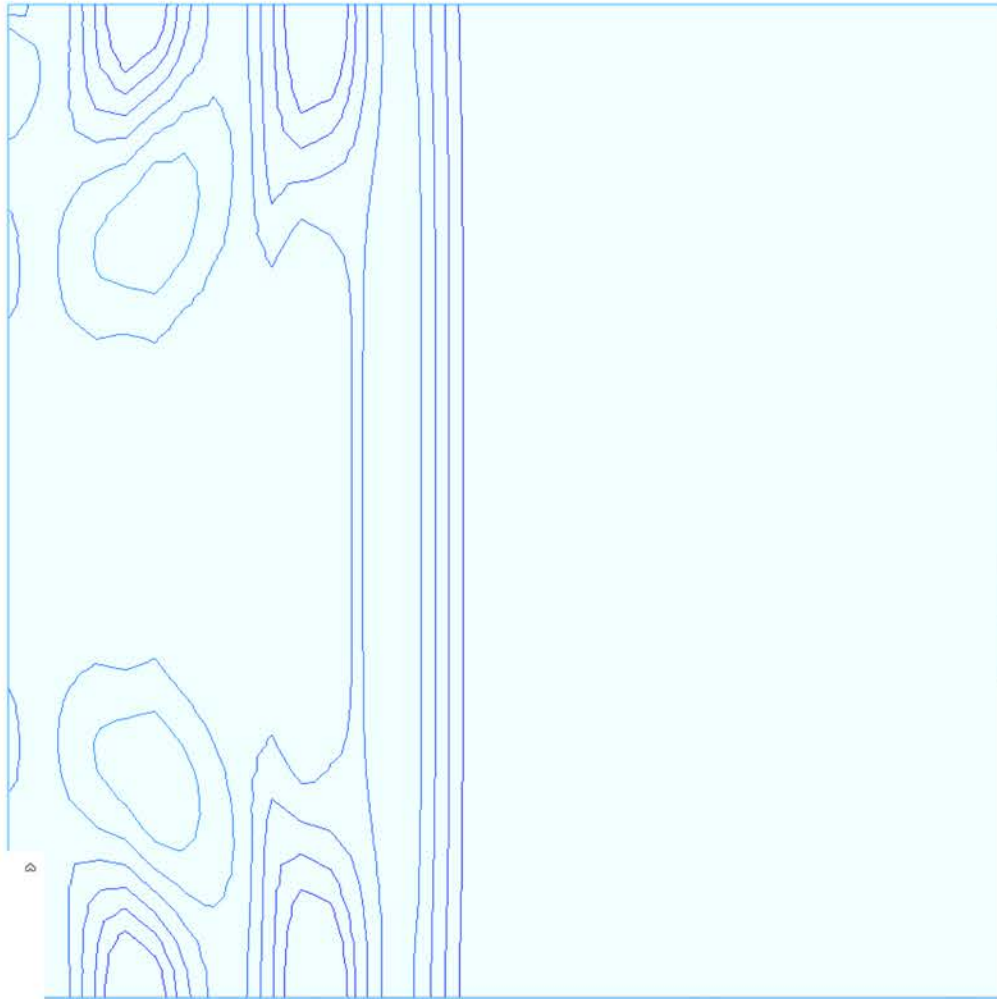


Pin Antenna

1fW peak (87.3% of total, Avg. 5%)
@27 GHz

Bouncing (every ~40ns)





e-field (t=0..end(0.5)) [pw]

Component	Abs
Sample	100/1458
Time	49.5 ps
Cross section	A
Cutplane at Z	0.000 mm
Maximum on Plane (Sample)	0.000540068 V/m
Maximum (Sample)	0.000592264 V/m
Maximum (Global)	0.00764982 V/m



RF Antenna and Readout

Dutch-led Consortium: *started 9/1/21 (5-year)

<https://www.nwo.nl/en/projects/nwa129219231>



Find funding Research policy NWO Research & results

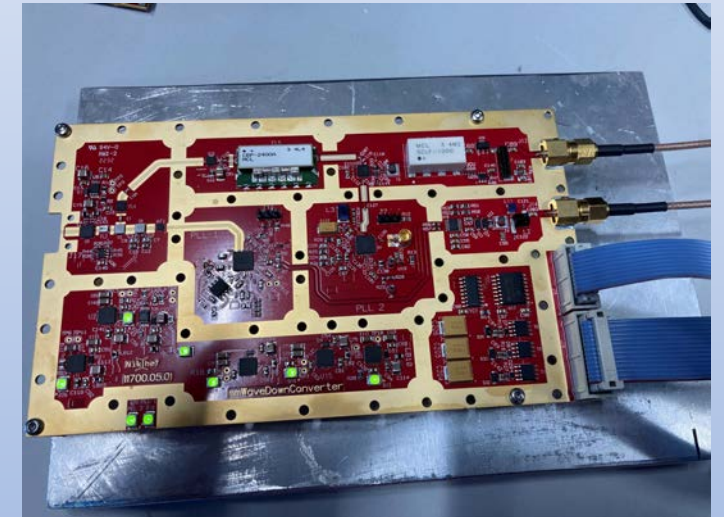
One second after the Big Bang

Every second, Earth is bombarded with an enormous number of neutrinos from the cosmos. These neutrinos were created in the primordial soup one second after the Big Bang, but they have never been observed. The researchers will develop an experiment to observe “relic neutrinos” by investigating the decay of heavy-hydrogen tritium.

Official secretary on behalf of the consortium: Prof. Auke Colijn – University of Amsterdam

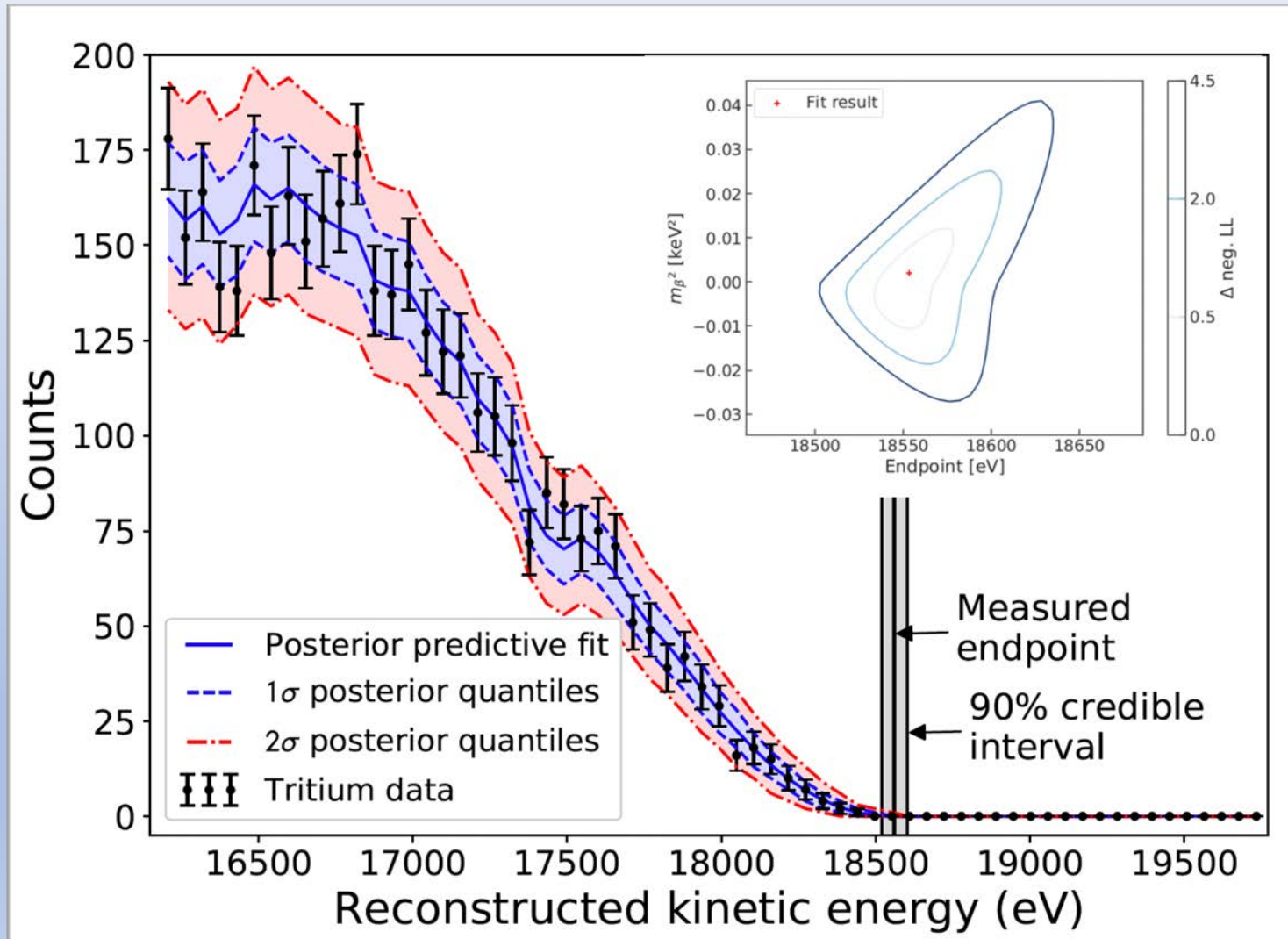
Consortium: University of Amsterdam, Nikhef, Radboud University, The Hague University of Applied Sciences, TNO, Princeton Physics Department, Gran Sasso National Laboratory (LNGS), Netherlands' Physical Society, Ampulz, Karlsruhe Institute of Technology

Amount awarded: 1.1 million euros



Further Antenna design work by TNO and digital processing at Univ. of Amsterdam

Recent Project 8 Tritium Measurement



RF measurement background levels extremely low.

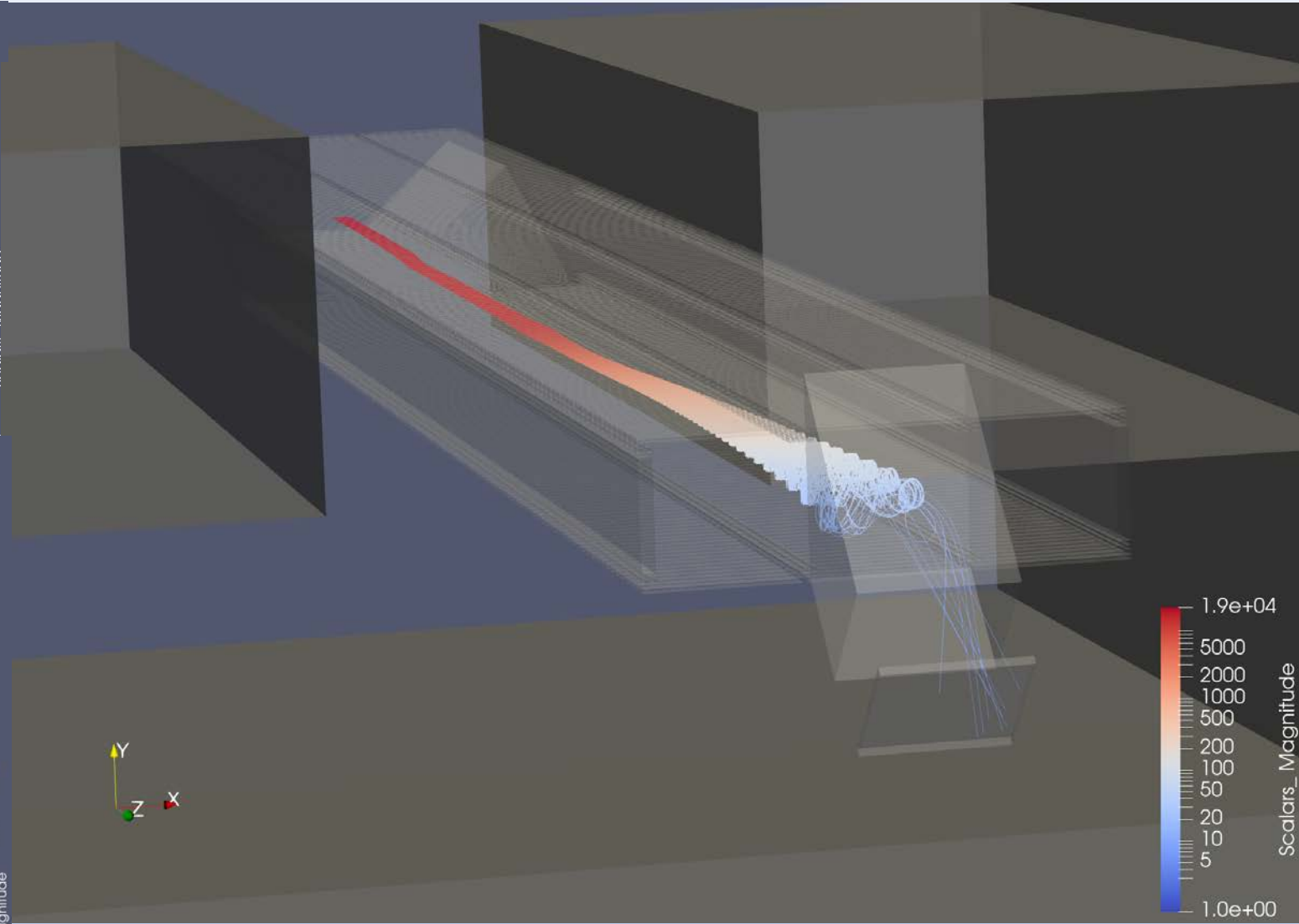
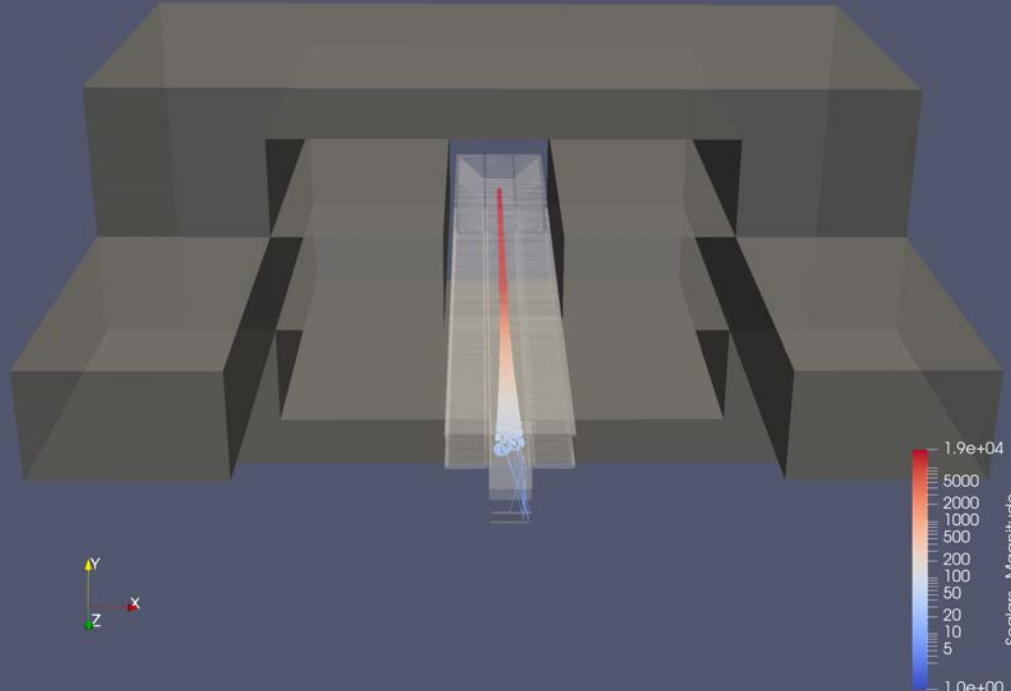
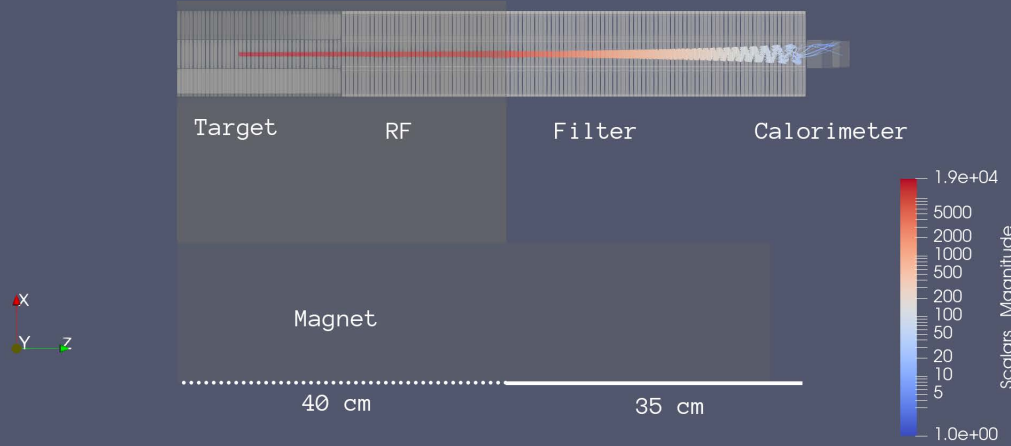
No events observed above endpoint,
Setting upper limit on background rate

$< 3 \times 10^{-10}$ /eV/s (90% CL)

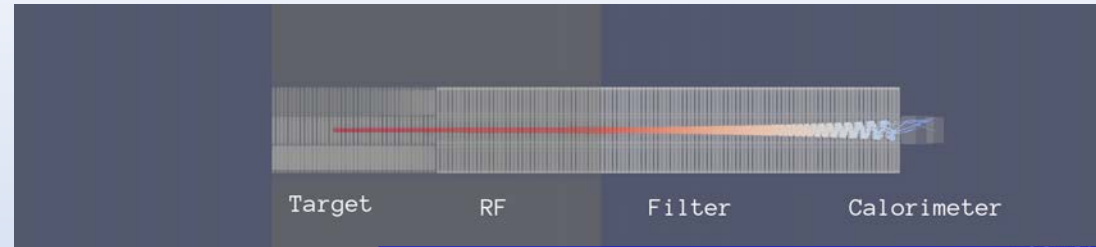
$\rightarrow < 1$ event per eV in 100 years!

End-to-end Transport w/Kassiopeia

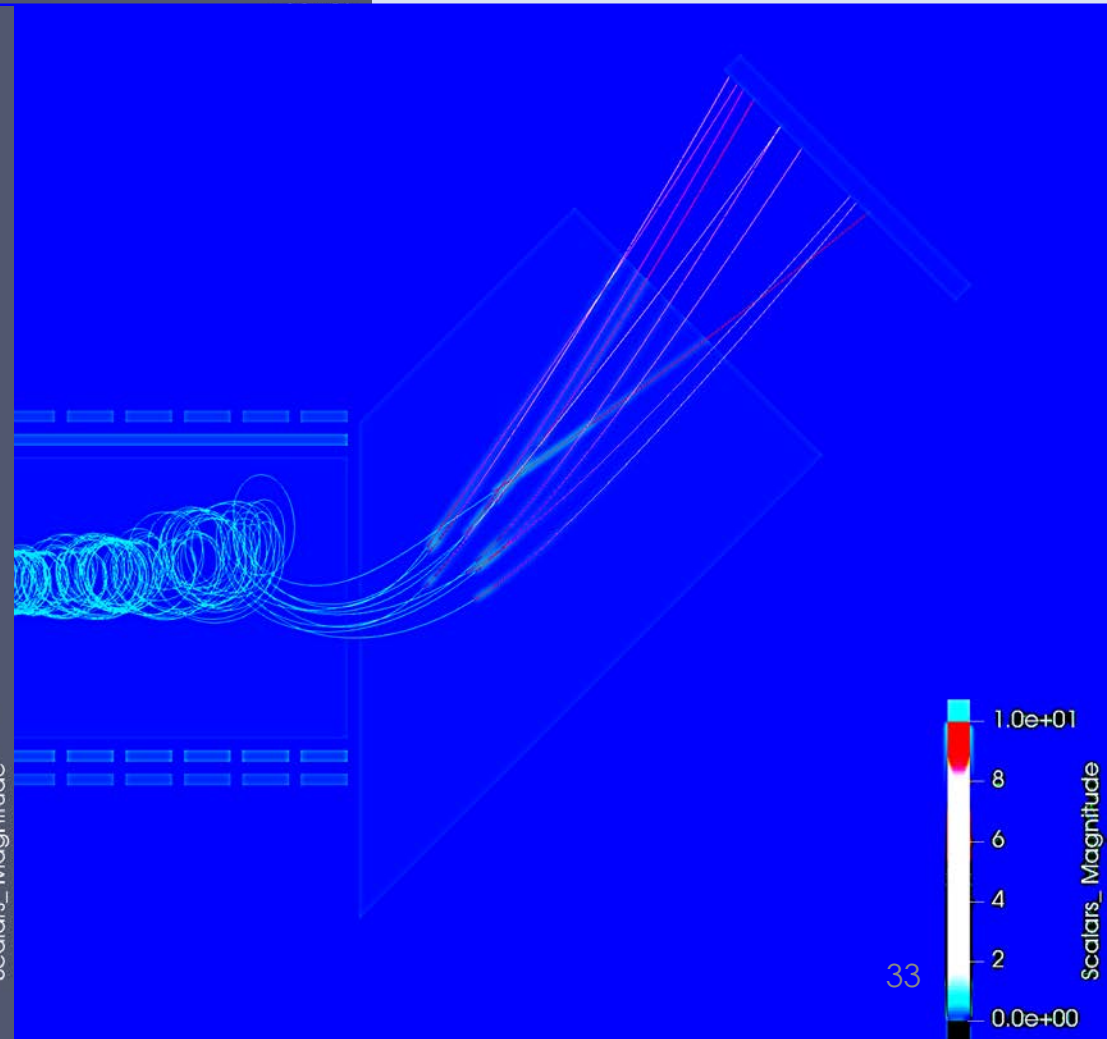
pitch 85 electrons shown



Zero-Field Calorimeter Transition



Wonyong Chung (Princeton)



Gap Opening in Double-Sided Highly Hydrogenated Free-Standing Graphene

Maria Grazia Betti,* Ernesto Placidi, Chiara Izzo, Elena Blundo, Antonio Polimeni, Marco Sbroscia, José Avila, Pavel Dudin, Kailong Hu, Yoshikazu Ito, Deborah Prezzi,* Miki Bonacci, Elisa Molinari, and Carlo Mariani



Cite This: *Nano Lett.* 2022, 22, 2971–2977



Read Online

ACCESS |



Metrics & More

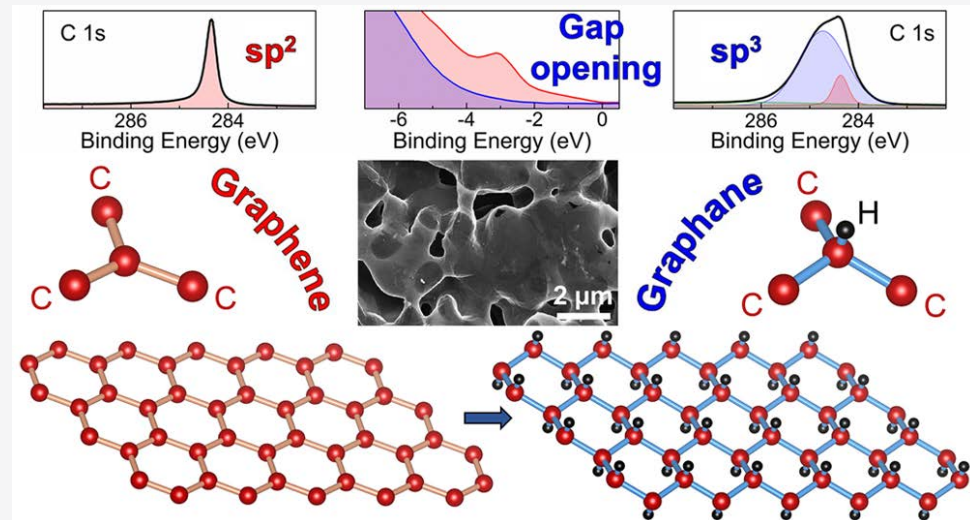


Article Recommendations



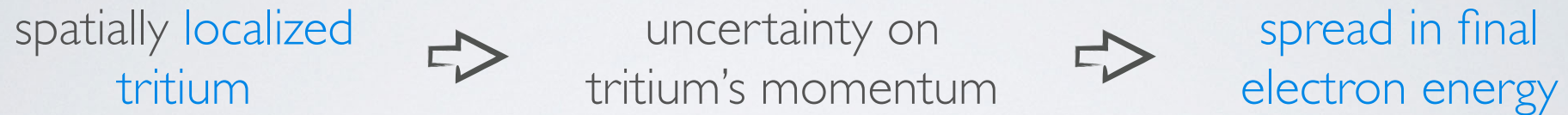
Supporting Information

ABSTRACT: Conversion of free-standing graphene into pure graphane—where each C atom is sp^3 bound to a hydrogen atom—has not been achieved so far, in spite of numerous experimental attempts. Here, we obtain an unprecedented level of hydrogenation ($\approx 90\%$ of sp^3 bonds) by exposing fully free-standing nanoporous samples—constituted by a single to a few veils of smoothly rippled graphene—to atomic hydrogen in ultrahigh vacuum. Such a controlled hydrogenation of high-quality and high-specific-area samples converts the original conductive graphene into a wide gap semiconductor, with the valence band maximum (VBM) ~ 3.5 eV below the Fermi level, as monitored by photoemission spectromicroscopy and confirmed by theoretical predictions. In fact, the calculated band structure unequivocally identifies the achievement of a stable, double-sided fully hydrogenated configuration, with gap opening and no trace of π states, in excellent agreement with the experimental results.



QUANTUM SPREAD

- Distributing tritium on flat graphene has one drawback



[Cheipesh, Cheianov, Boyarsky – PRD 2021, 2101.10069]

- A simple semi-classical estimate:

fluctuating momenta

$$\begin{aligned} \mathbf{p}_T &= \Delta\mathbf{p}_T \\ \mathbf{p}_{He} &= \bar{\mathbf{p}}_{He} + \Delta\mathbf{p}_{He} \\ \mathbf{p}_e &= \bar{\mathbf{p}}_e + \Delta\mathbf{p}_e \end{aligned}$$

energy and momentum conservation returns

$$\Delta E_e \simeq \left| \frac{\mathbf{p}_e \cdot \Delta\mathbf{p}_T}{E_{He}} \right| \sim \frac{p_e}{m_{He}} \frac{1}{\Delta x_T}$$

spread of initial tritium wave function ($\Delta x_T \sim 0.1 \text{ \AA}$)

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spread of initial tritium wave function ($\Delta x_T \sim 0.1 \text{ \AA}$)

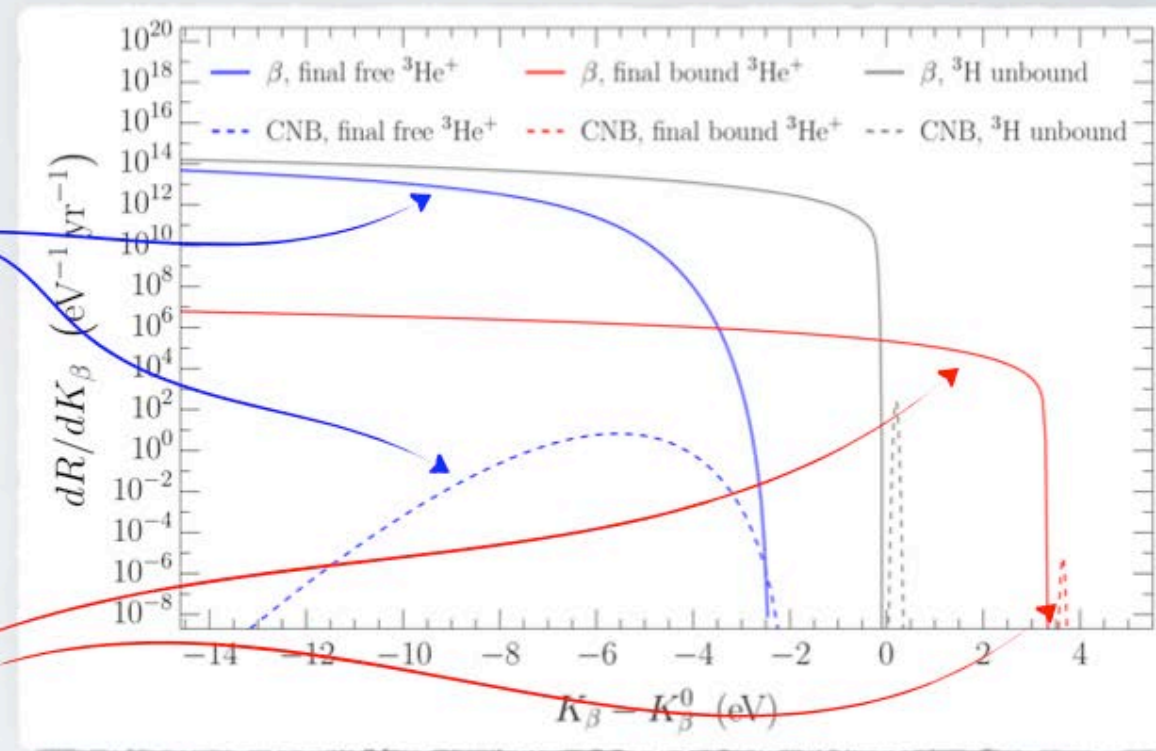
an order of magnitude larger than the wanted energy accuracy

QUANTUM SPREAD

- The resulting rate is

${}^3\text{He}^+$ is mostly freed from the graphene \rightarrow the cosmic neutrino peak disappears under the decay spectrum

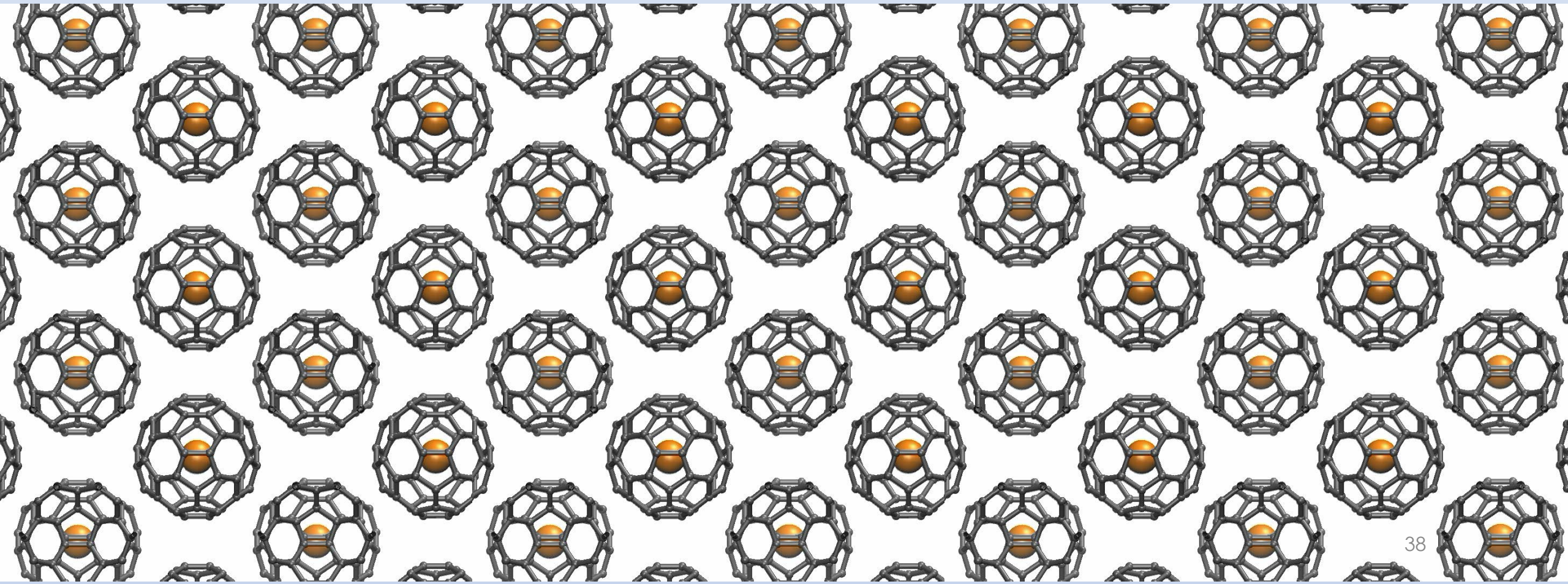
When the ${}^3\text{He}^+$ remains bound in the ground state the peak is well separated \rightarrow it is however exponentially unlikely



[PTOLEMY - 2203.11228]

Collaboration with Savannah River National Laboratory for Tritium Loading

CNT, NPG, CVD-G, and De-localized Atomic T Geometries
~2Å flat potential – not chemically active



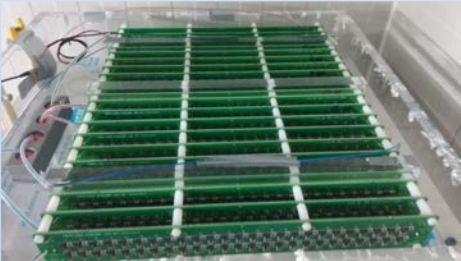
HV Stability and Monitoring

@LNGS

Field mill box



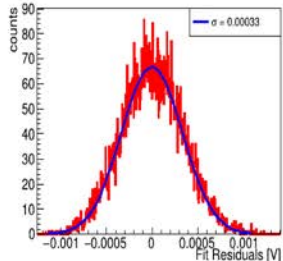
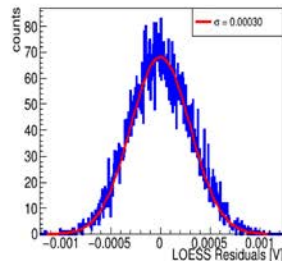
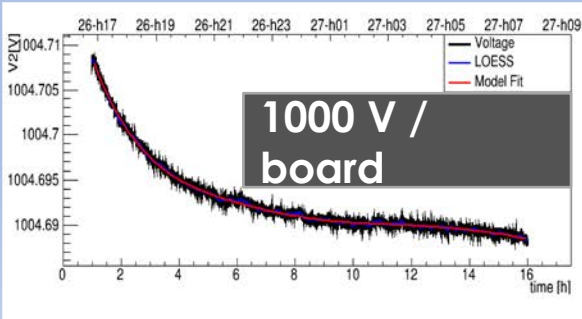
Diodes box



High precision voltmeter
Keysight®

Power supply
Bertan®

environmental parameter stabilization ($dT \sim 0.1^\circ\text{C}$, Pressure < 1mBar, humidity 0%)



Single board

$$\sigma = 0.3 \text{ mV}$$

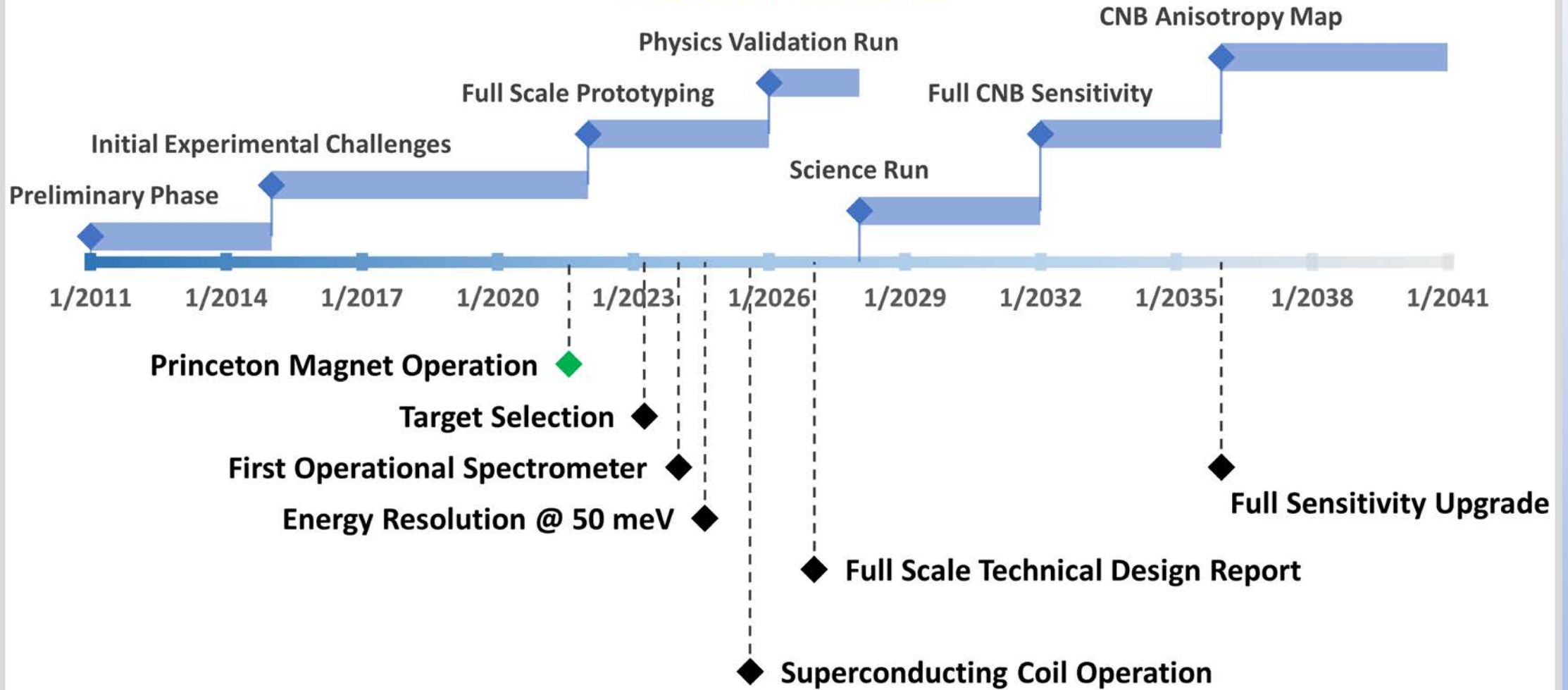
Expect \sqrt{N} boards :
~1.4mV@20kV

Field Mill ~50mV

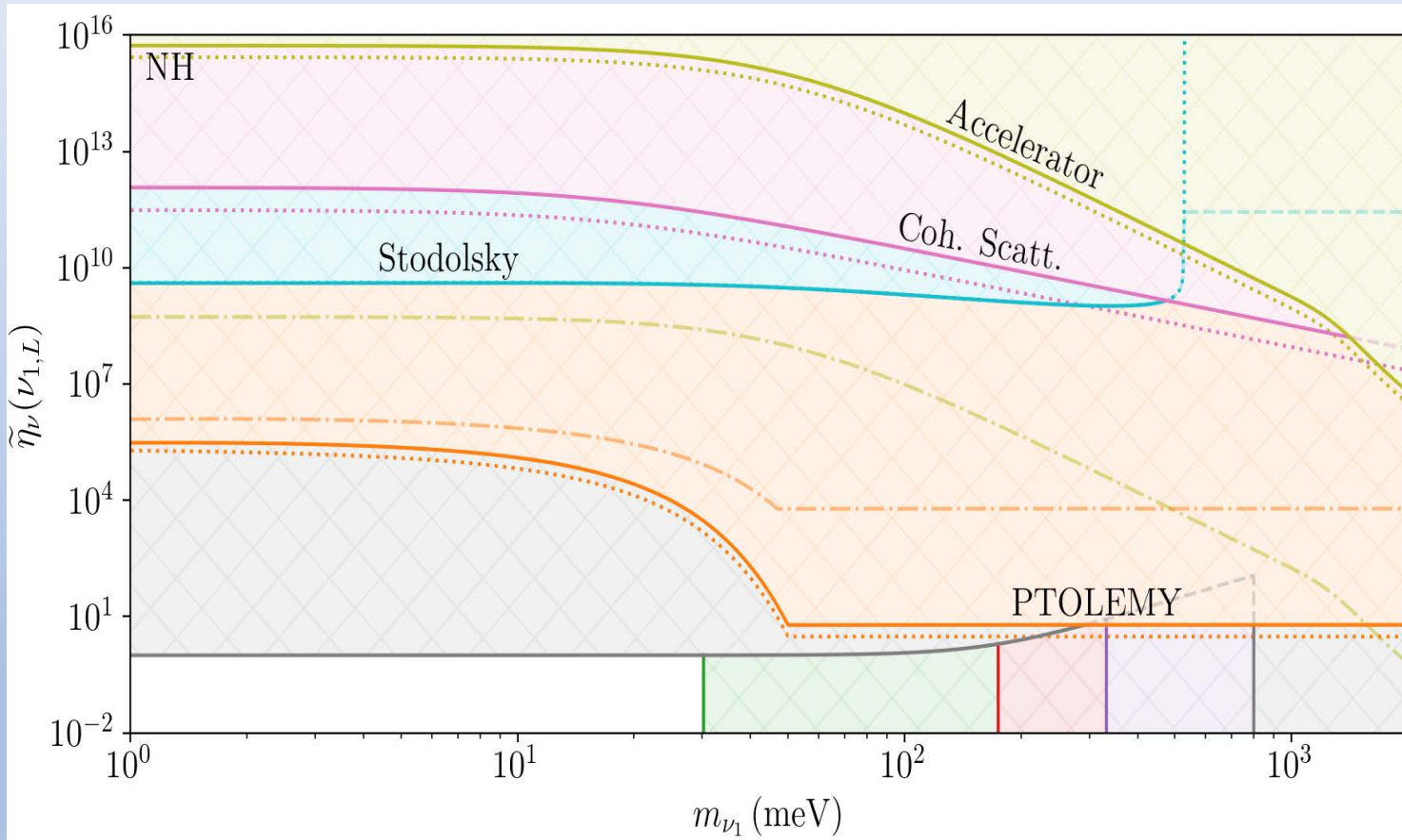
LNGS Full-Scale Prototype

- Features of prototype:
 - Iron-return flux magnet @1T w/ conduction-cooled SC coils
 - small (few cm²) tritium-loaded graphene target from SRNL w/Rome hydrogen loading system
 - RF antenna @26.5GHz from Univ.of Amsterdam/TNO
 - PTOLEMY filter with high precision HV reference
 - Vacuum cryostat interface to TES microcalorimeter fridge
- Fabrication in progress on most elements (SC coil approval soon)
- Operate through 2024 for first tritium data release

PTOLEMY Timeline



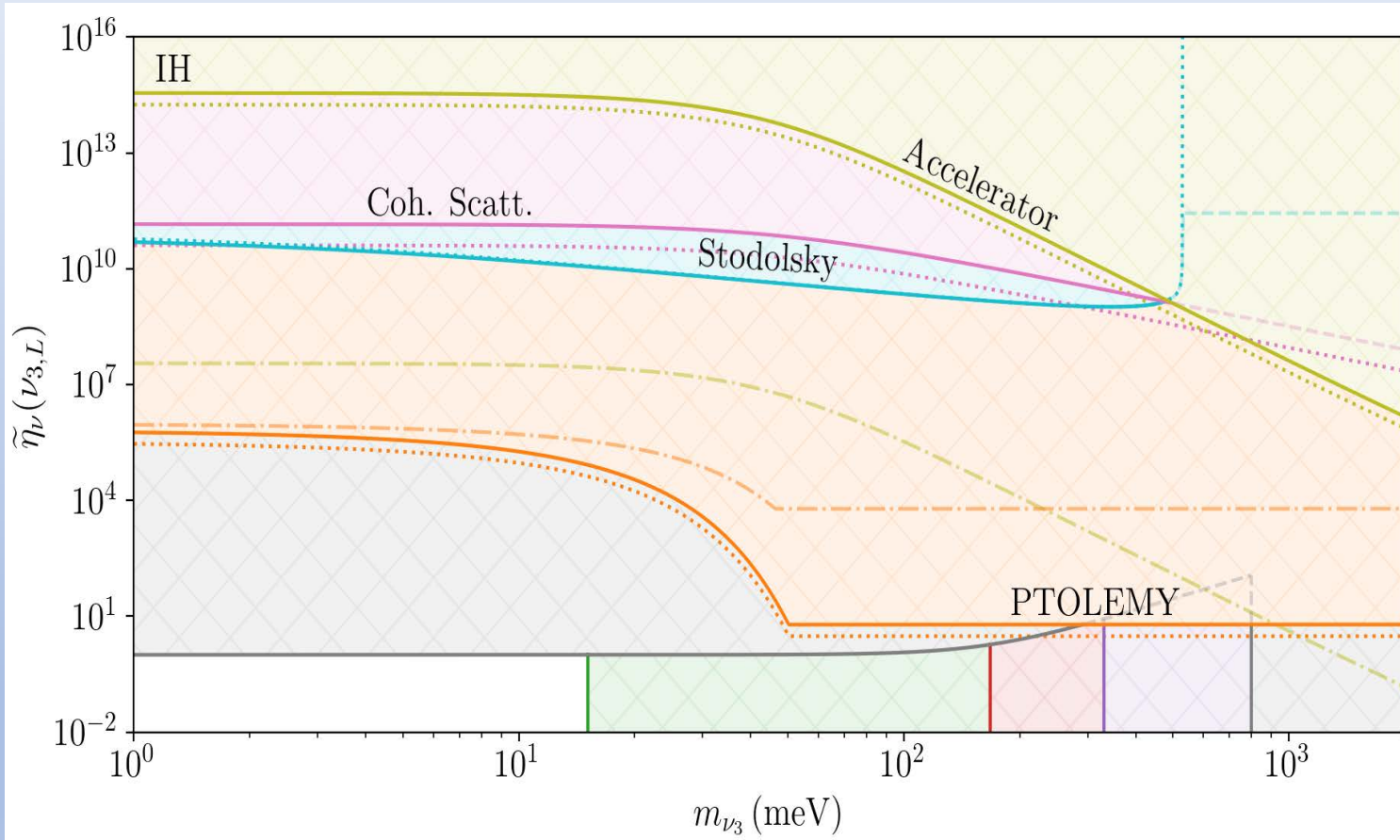
Relic Neutrino Exp Search Sensitivities (NH)



Curve	Description
Orange	PTOLEMY sensitivity to Dirac (solid) and Majorana (dotted) neutrinos, standard method. Time dependent method for Dirac neutrinos (dot-dashed).
Cyan	Stodolsky effect sensitivity to Dirac (solid) and Majorana (dotted) neutrinos.
Pink	Coherent scattering sensitivity to Dirac (solid) and Majorana (dotted) neutrinos.
Light green	Accelerator sensitivity to Dirac (solid) and Majorana (dotted) neutrinos. Using an optimistic setup for Dirac neutrinos (dot-dashed).
Grey	Excluded by theory and experiment for $T_{\nu_i} = T_{\nu,0}$ (solid, Figures 9 and 10). Excluded by KATRIN (dashed, Figures 11 and 12).
Blue	Excluded by Pauli exclusion principle for $T_{\nu_i} \neq T_{\nu,0}$.
Purple	Strongest mass bound on unstable Dirac neutrinos, from cosmology.
Red	Strongest mass bound on unstable Majorana neutrinos, from KamLAND-Zen.
Green	Strongest mass bound on stable neutrinos, from cosmology.

<https://arxiv.org/abs/2207.12413>

Relic Neutrino Exp Search Sensitivities (IH)



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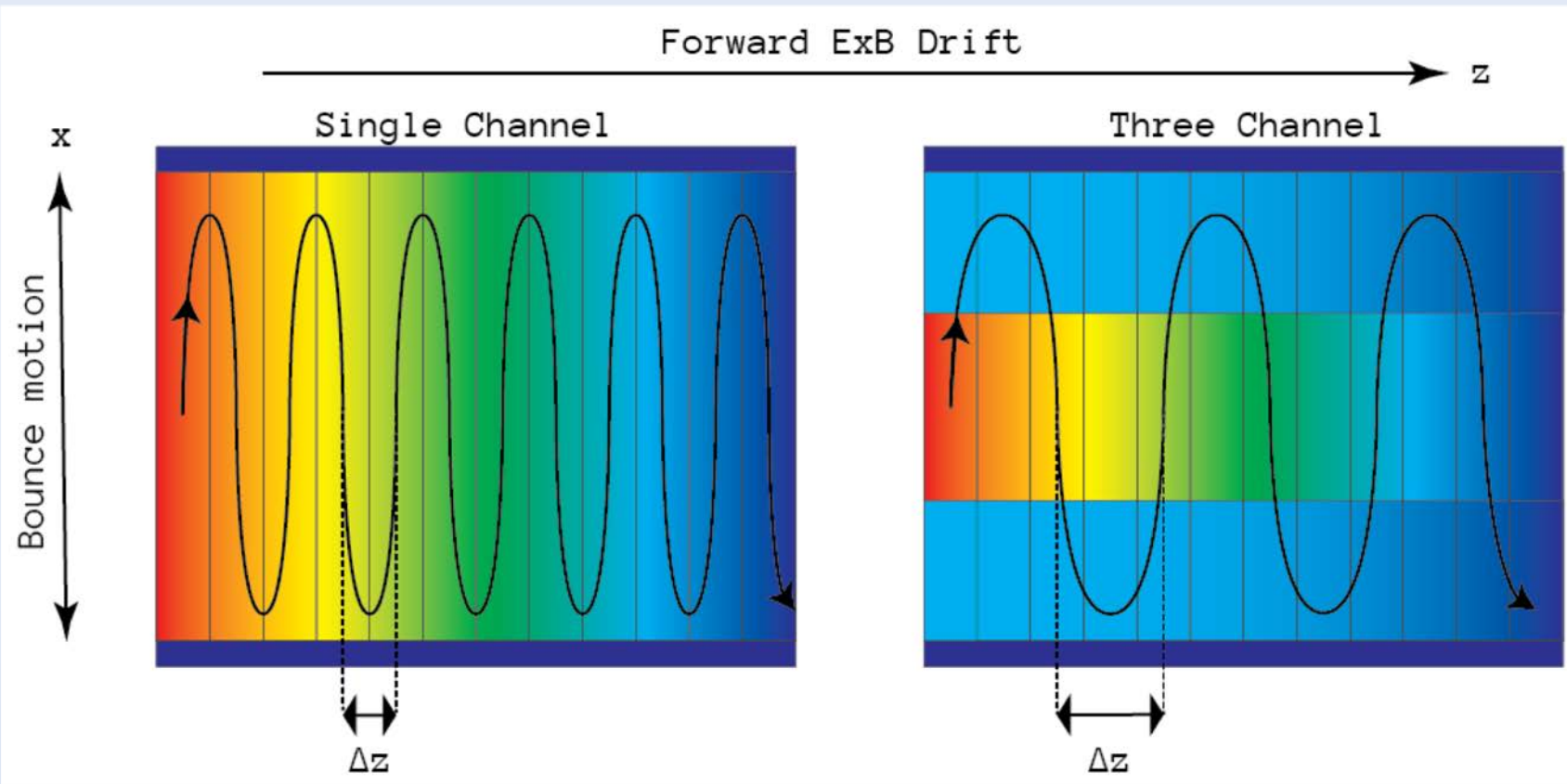
<https://arxiv.org/abs/2207.12413>

Outlook

- The sustained effort in PTOLEMY R&D is a testament to the importance our dedicated group of collaborators place on pushing the frontier of early Universe neutrino cosmology – clever ingenuity is behind many advances
- PTOLEMY expects to become the leader in tritium endpoint energy measurement resolution within the next 2 years
 - Next stop (w/ more target mass): absolute neutrino mass

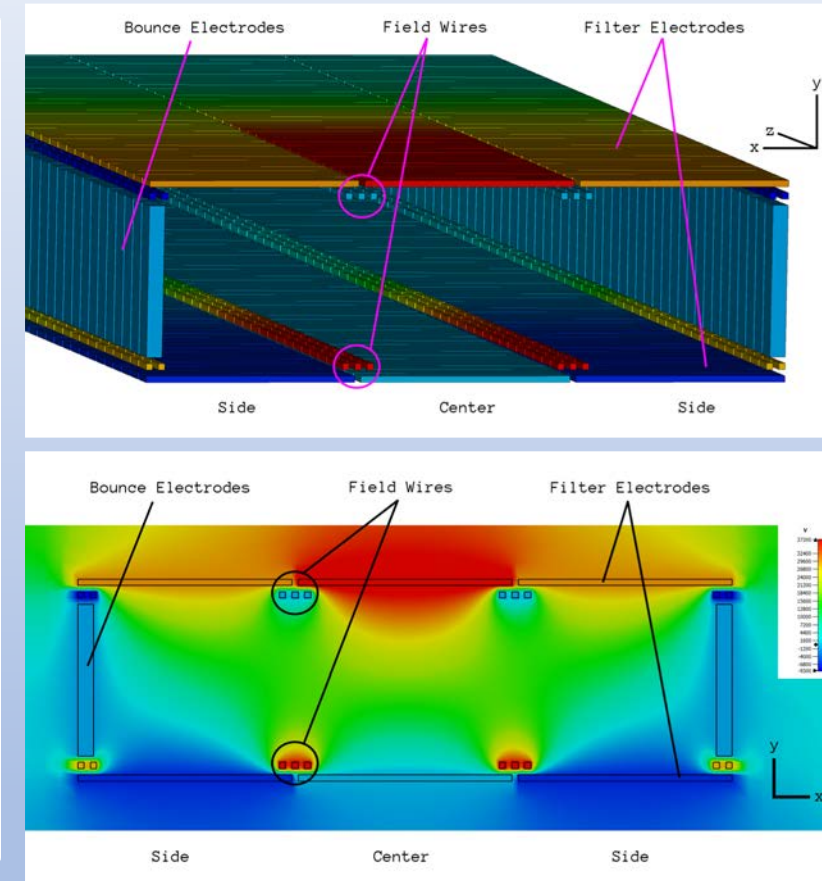
Backup

Bobsledding (pushing electron up potential)

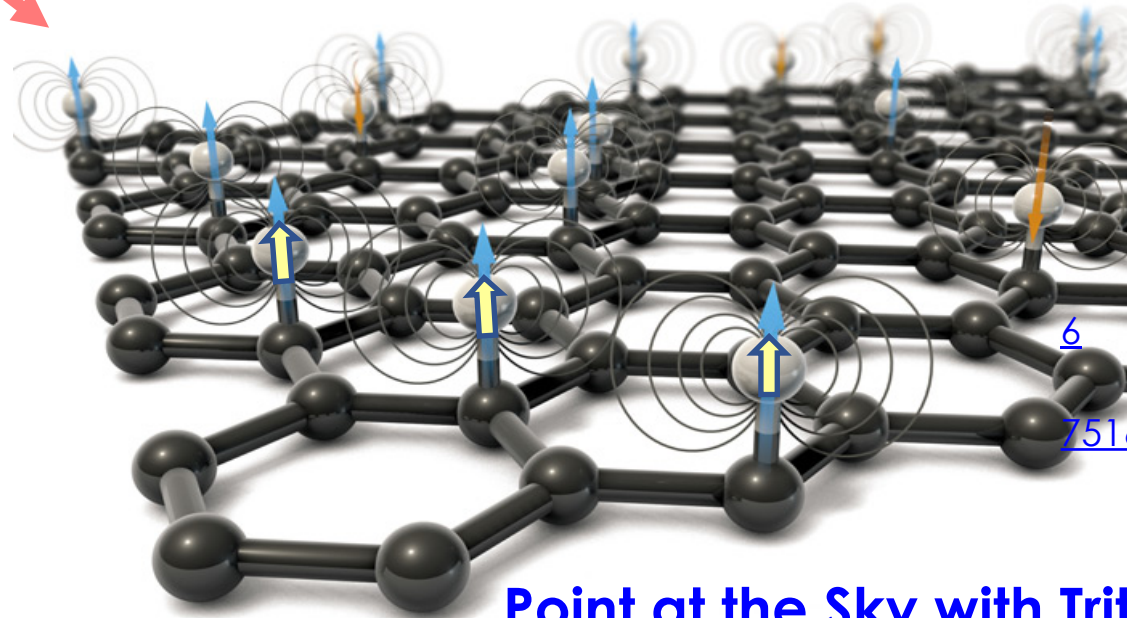
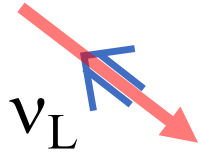


Transverse “Selector”
(one channel)

Dynamically Adjusted
(side channels)
to **Total Energy** “Selector”



Polarized Tritium Target



Lisanti, Safdi, CGT, 2014.
[10.1103/PhysRevD.90.07300](https://arxiv.org/abs/10.1103/PhysRevD.90.07300)

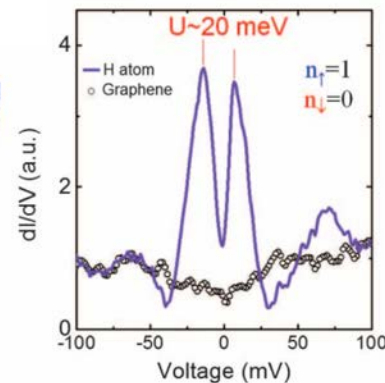
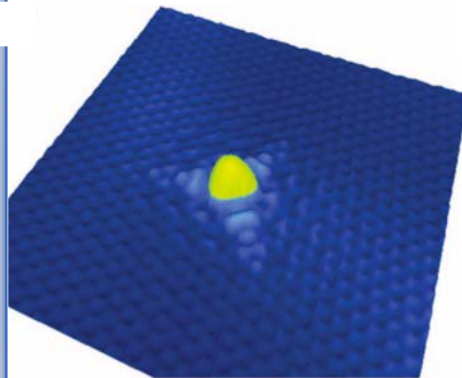
[6](#)
Akhmedov, 2019.
[10.1088/1475-7516/2019/09/031](https://arxiv.org/abs/10.1088/1475-7516/2019/09/031)

Point at the Sky with Tritium Nuclear Spin ↑

Detection (capture) of cold neutrinos:

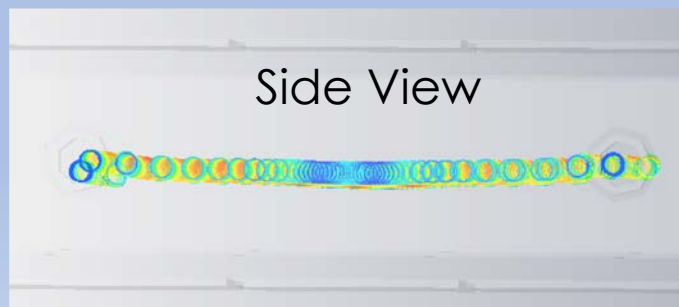
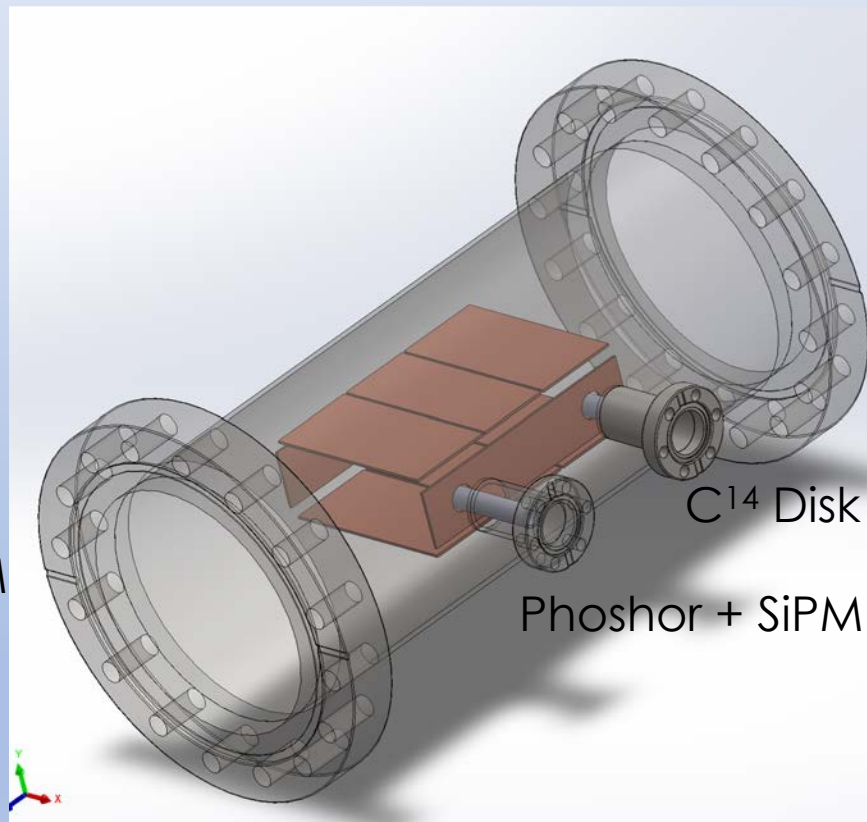
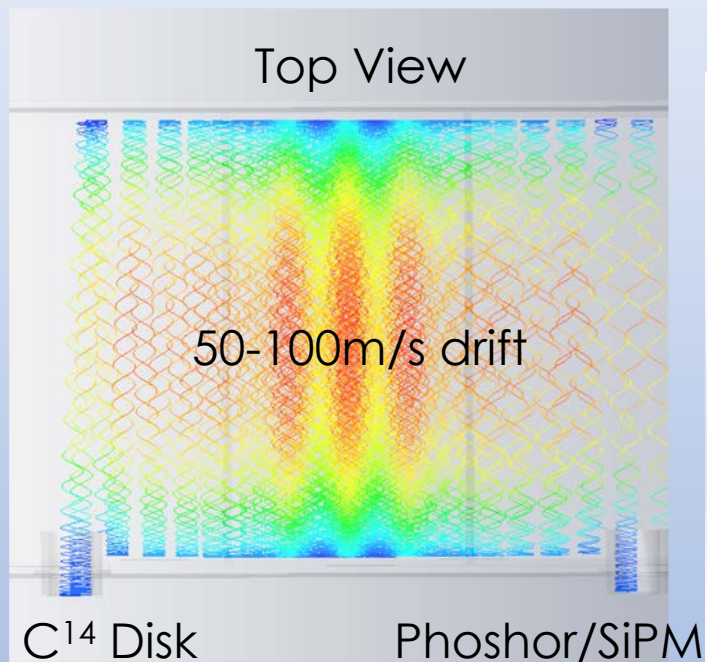
$$d\sigma/d\cos\theta (v/c) \sim (1+\cos\theta)$$

Hydrogen doping on graphene reveals magnetism

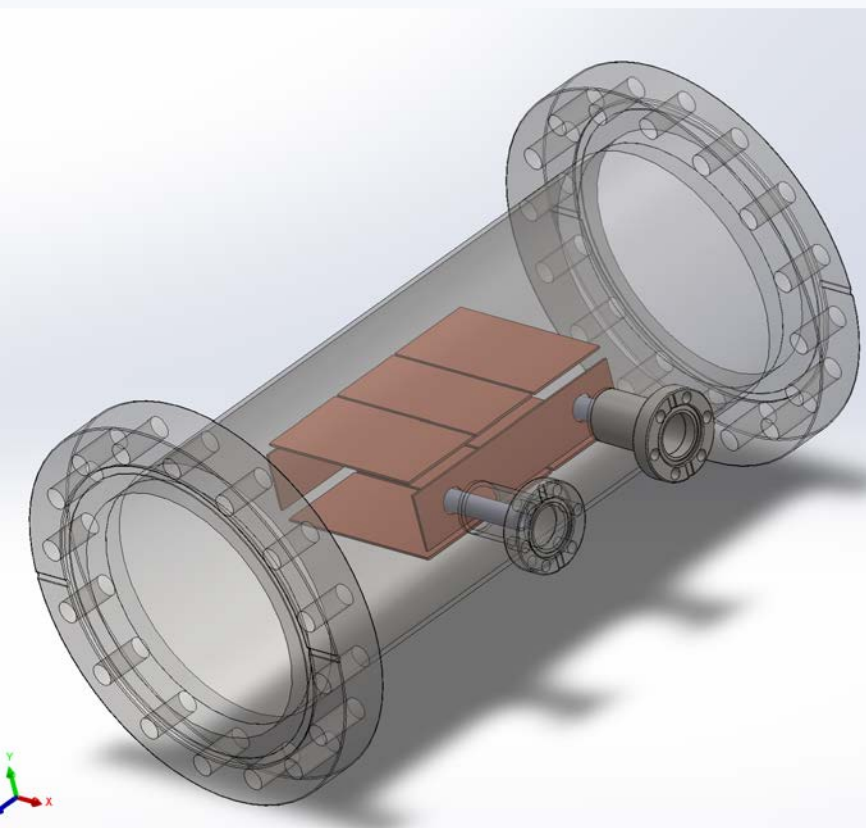


Gonzalez-Herrero, H. *et al.* Atomic-scale control of graphene magnetism by using hydrogen atoms. *Science* (80). **352**, 437–441 (2016).

Ultra-slow electron drift region

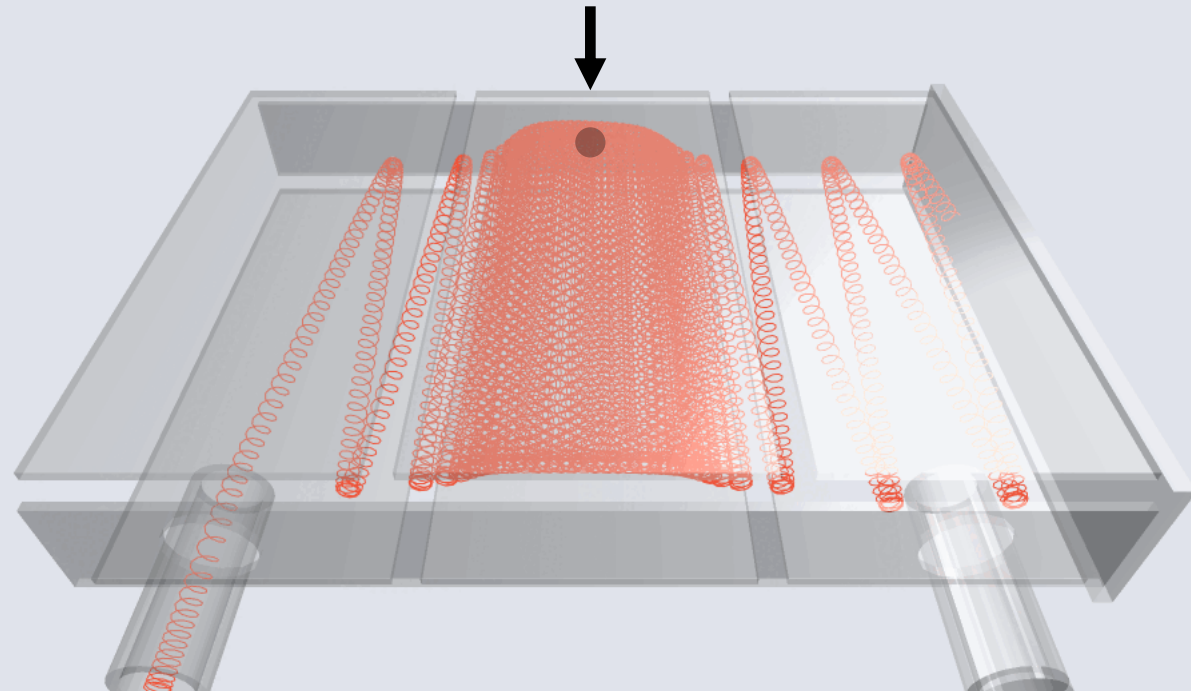


Scanning ExB Voltages to Maximize Duration of Antenna Signal



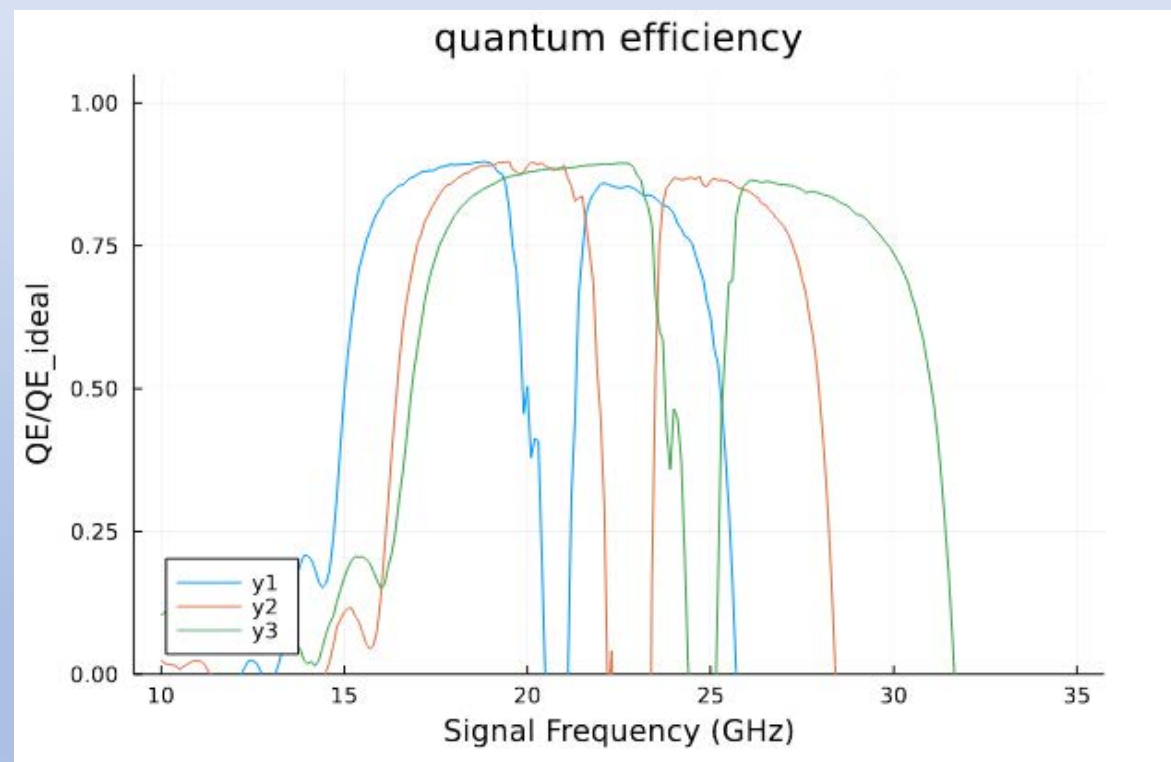
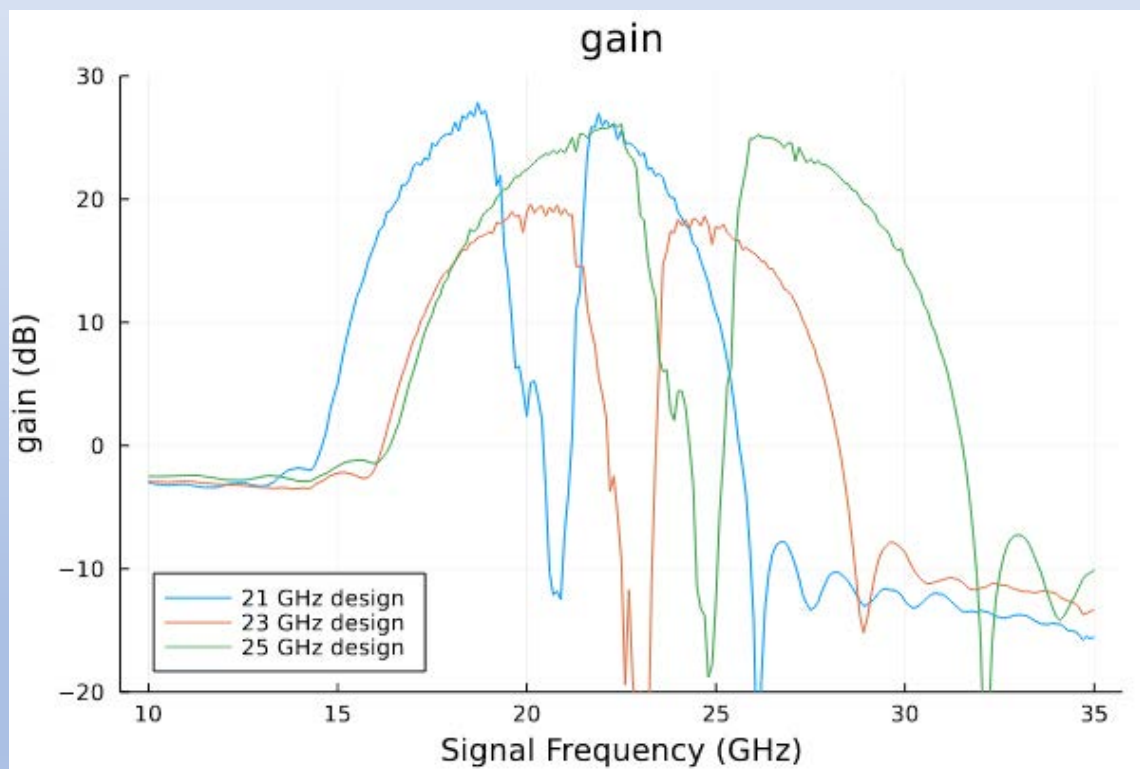
Middle electrodes at ± 2 V

Circular Cavity w/ Pin Antenna



Quantum-Limited Parametric Amp

High Frequency Josephson Traveling Wave Parametric Amp (TWPA)



Joint Project w/ MIT and Project 8

Electromagnetic Filters

MAC-E filter

Magnetic Adiabatic Invariance

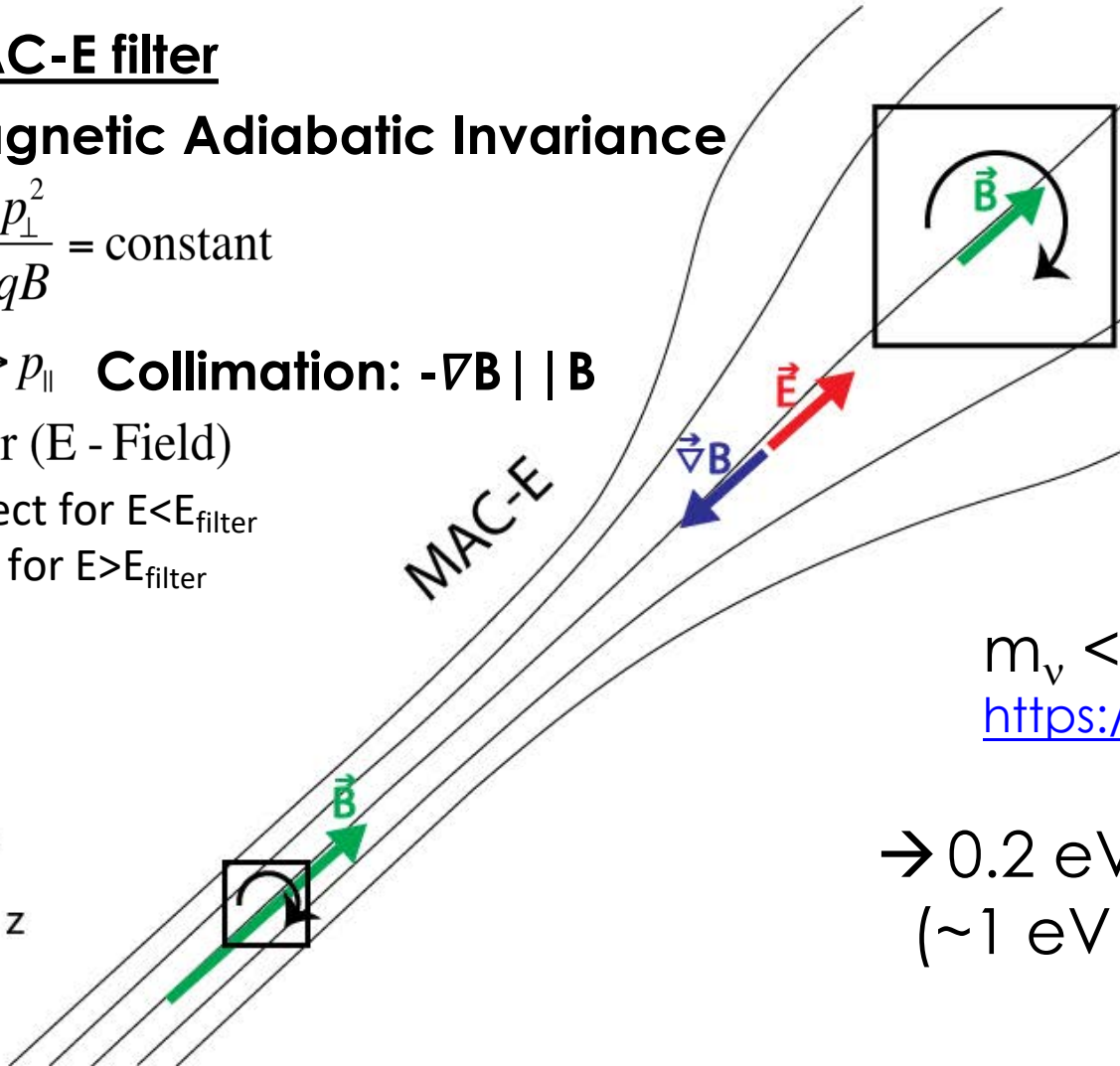
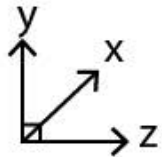
$$\mu = \frac{p_{\perp}^2}{qB} = \text{constant}$$

$p_{\perp} \rightarrow p_{\parallel}$ **Collimation: $-\nabla B \parallel B$**

Filter (E - Field)

Reflect for $E < E_{\text{filter}}$

Pass for $E > E_{\text{filter}}$



KATRIN

~1200m³

$m_{\nu} < 0.8 \text{ eV}/c^2$ (90% CL)
<https://arxiv.org/abs/2105.08533>

→ 0.2 eV/c² Sensitivity Goal
 (~1 eV energy resolution)

Electromagnetic Filters

Transverse Drift filter

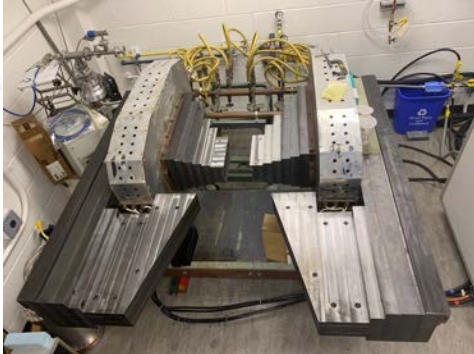
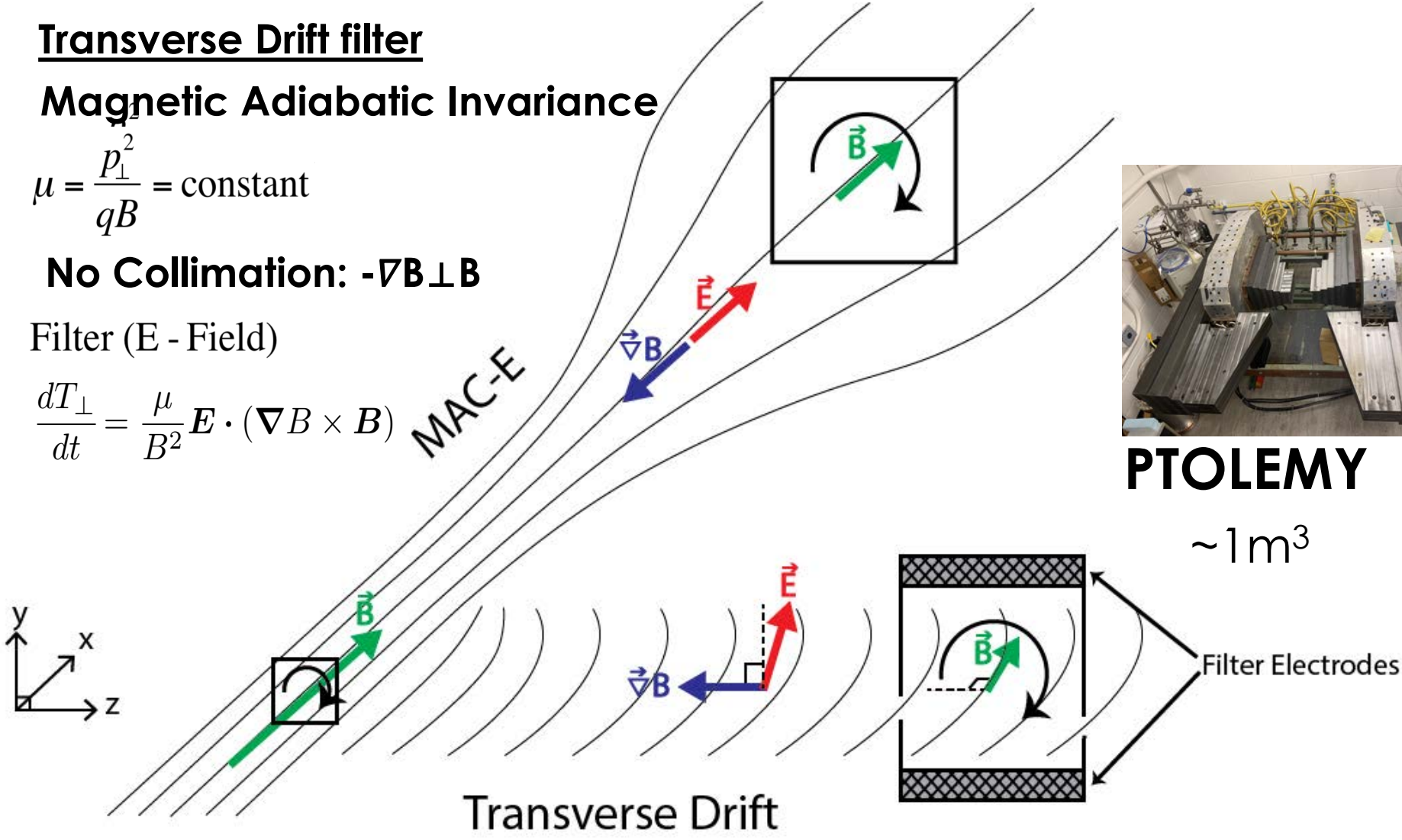
Magnetic Adiabatic Invariance

$$\mu = \frac{p_{\perp}^2}{qB} = \text{constant}$$

No Collimation: $-\nabla B \perp B$

Filter (E - Field)

$$\frac{dT_{\perp}}{dt} = \frac{\mu}{B^2} \mathbf{E} \cdot (\nabla B \times B)$$

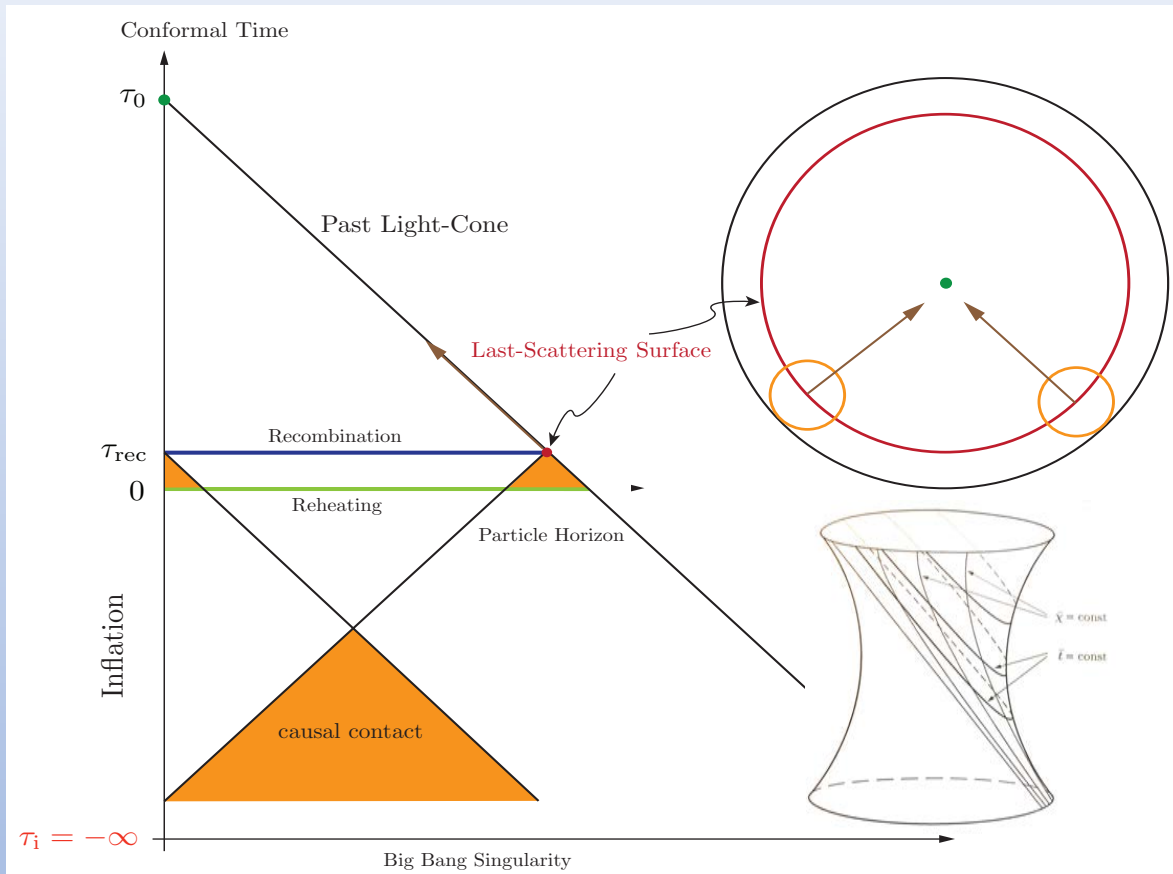


PTOLEMY

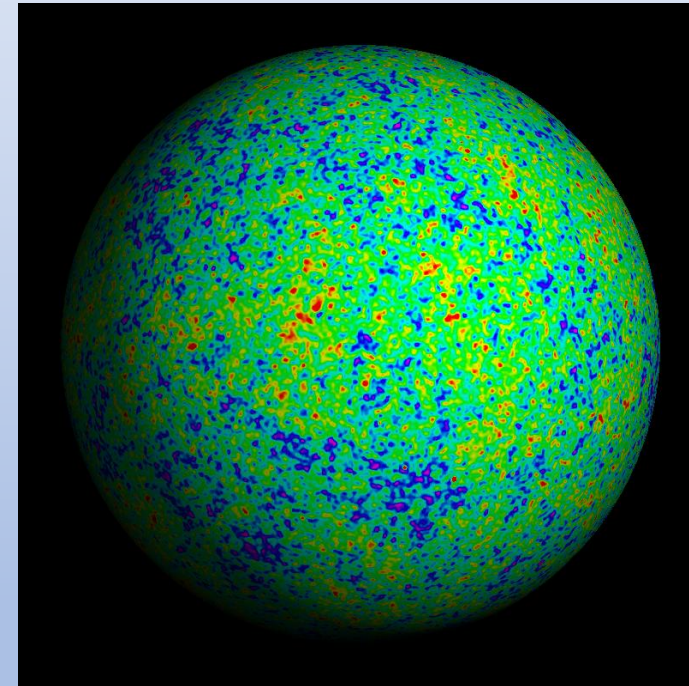
~1m³

Filter Electrodes

Big Bang Cosmology



Adiabatic Density Anisotropies
 $\delta \sim 10^{-5}$ at $z \sim 1100$



Where we think there is an initial $\tau_i=0$ Big Bang Singularity is believed to be the “end” of an inflation period that slowly pulled out (>60 e-folds $a(\tau) \sim e^{H\tau}$) of a “de Sitter”-like spacetime

Axions and Relic Neutrinos

- Early Universe models trend toward having axion and neutrino sector implications:
 - Both involve relatively low masses
 - Both have unique coupling terms including non-SM
 - Both see some stopping blocks at BBN and CMB
 - Both have possibilities for late Universe generation in warm-inflation-like models
 - Experiments are distinct in terms of wave-like or target-based particle detection, but interesting parameter spaces can be correlated depending on the model

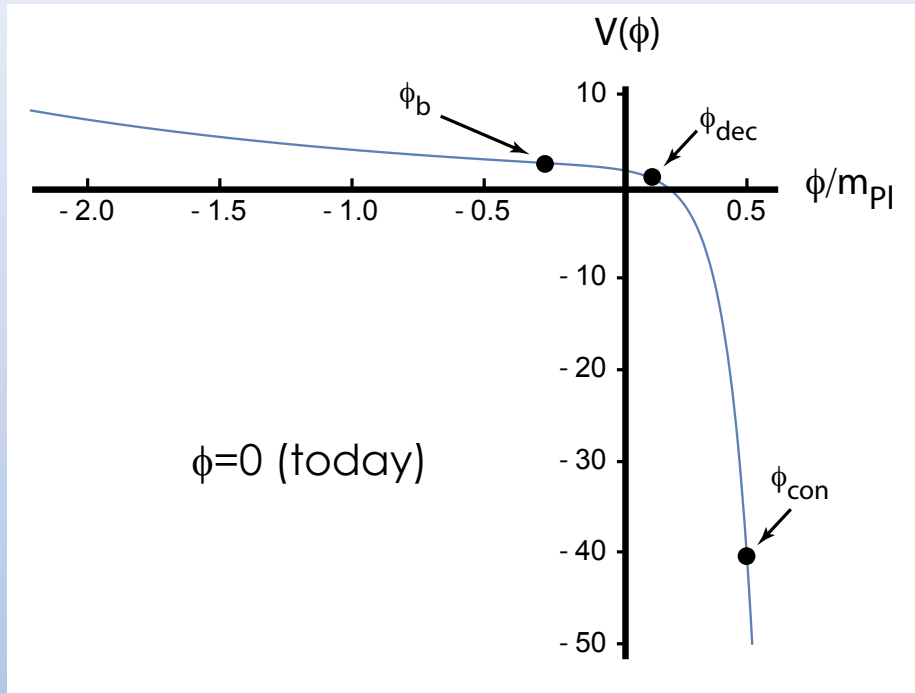
Where neutrinos come in

1 MeV \sim 70x10³⁹ km/s per Mpc

- Example: F. Takahashi, W. Yin, and A. Guth, “QCD axion window and low-scale inflation”, 10.1103/PhysRevD.98.015042 (2018).
<https://arxiv.org/abs/1805.08763>
 - $H_{\text{inf}} < \mathcal{O}(1)$ MeV no fine-tuning of misalignment angle needed
 - Upper bound on axion scale relaxed \rightarrow wave-like particles with dark matter abundance
 - *Reheating?* Right-handed neutrinos coupled to the inflaton (B-L Higgs?)
 - *Inflaton-Radiation equality, followed by perturbative right-handed neutrino decays to Higgs and leptons (resonant leptogenesis?)*
 - *Heavier right-handed neutrino radiative correction increases spectral index*

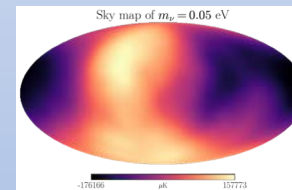
Rich era of axion experiments – but with neutrinos a close buddy

End of Expansion



Frictional term $\propto \dot{\phi}$
may couple to gauge singlets:
- Right-handed neutrinos??

→ CNB w/ high local density
and much more uniform



Dipole $\sim 8\% \rightarrow <1\%$
Quad $\sim 4\% \rightarrow <0.5\%$

C. Andrei, A. Ijjas and P.J. Steinhardt, The End of Expansion, <https://arxiv.org/abs/2201.07704>

A. Ijjas and P.J. Steinhardt, The End of Expansion and Dark Radiation (tentative title), in preparation

Also: K. Berghaus, P.W. Graham, D.E. Kaplan, G.D. Moore and S. Rajendran, Dark Energy Radiation, <https://arxiv.org/abs/2012.10549>

D. Green, D.E. Kaplan and S. Rajendran, Neutrino Interactions in the late universe, <https://arxiv.org/abs/2108.06928>