



The **CYGNO** Experiment: A Directional Dark Matter Detector with Optical Readout

INITIUM: Innovative Negative Ion Time projection chamber for Underground dark Matter searches

*A. Prajapati** on behalf of *CYGNO* collaboration

*Gran Sasso Science Institute, L'Aquila, Italy / Ph.D. student / Email: atul.prajapati@gssi.it

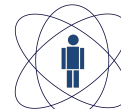
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Part of this project has been funded by the European Union's Horizon 2020 research and innovation programme under the ERC Consolidator Grant Agreement No 818744



UNIVERSIDADE D COIMBRA



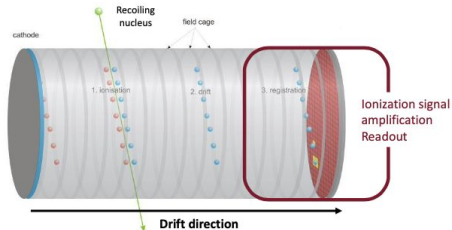
CBPF



UNIVERSIDADE FEDERAL DE JUIZ DE FORA

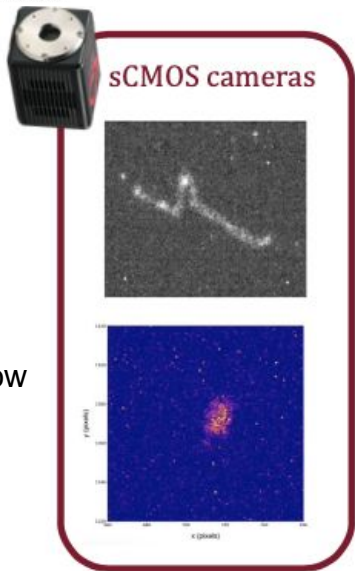


The University Of Sheffield.

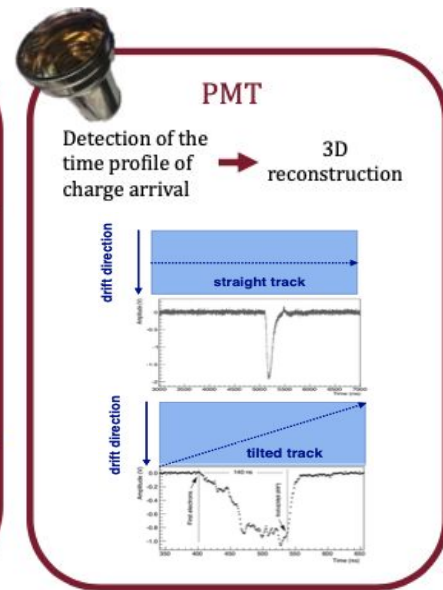
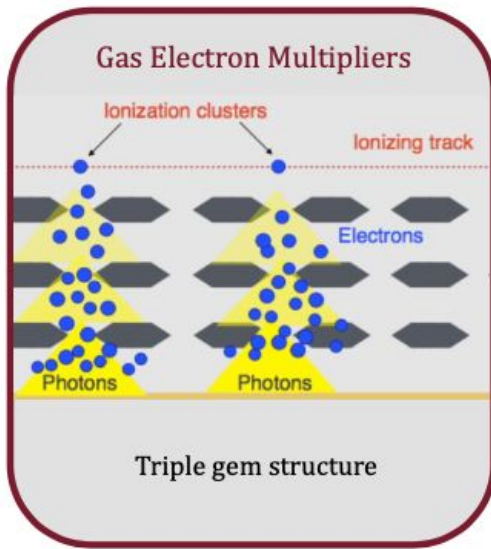


Triple GEM
Charge amplification & light production

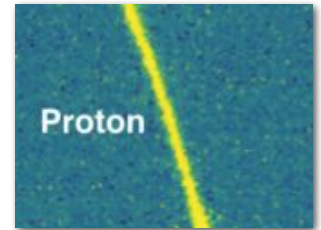
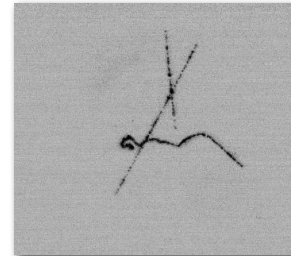
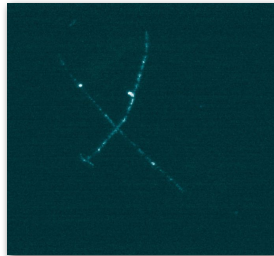
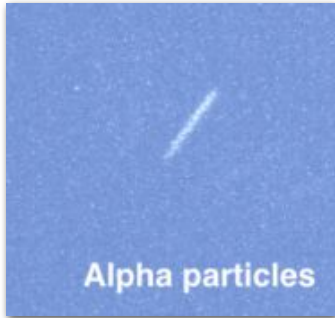
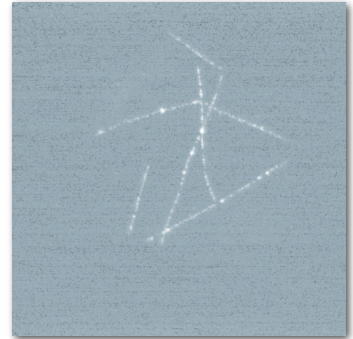
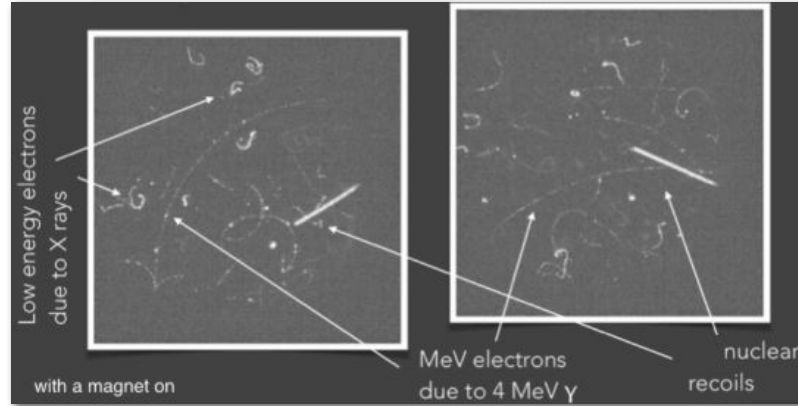
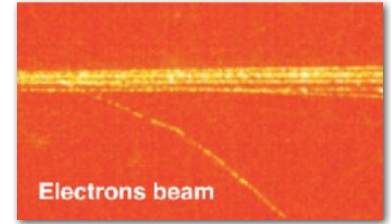
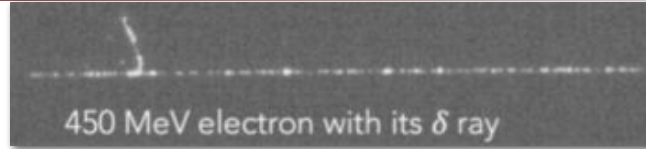
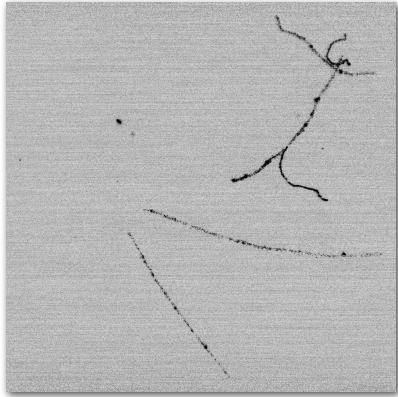
Camera & PMT
Light produced by the de-excitation of the gas molecules during electron multiplication is optically read by sCMOS and PMT

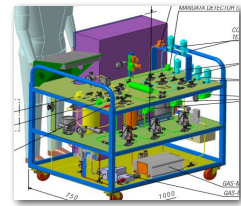
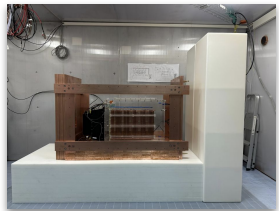
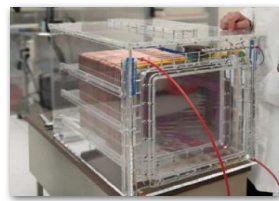
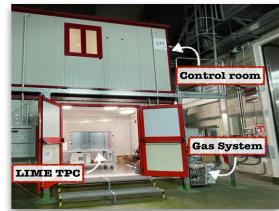
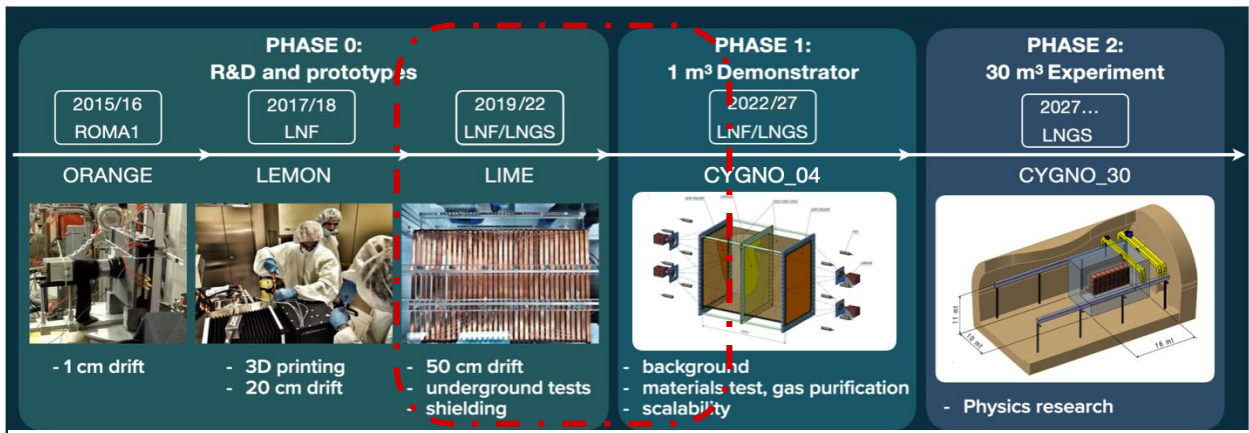


We can measure **energy** and **X-Y coordinate** using sCMOS's high granularity and low readout noise.



PMT measures the integrated **energy** and **time of arrival (dZ)** of charge carriers with high sampling rates.





Ongoing studies:

- ❖ Performance and stability test
- ❖ 3D reconstruction
- ❖ Directionality
- ❖ ER vs. NR discrimination
- ❖ Shielding materials
- ❖ Data/MC comparison

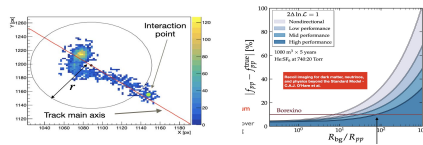
❖ Parallel research with MANGO detector for studying different GEM configuration, gas mixtures and Negative Ion Drift.

E. Baracchini et. al, JINST 13(2018) no.04, P04022



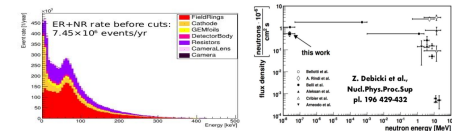
E. Baracchini

- Directionality
- Neutrino feasibility study



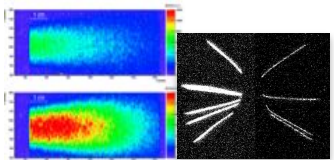
S. Torelli

- Neutron Simulation
- Background Simulation and Shielding



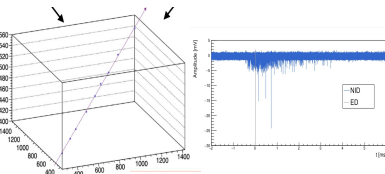
F. Di Giambattista

- CYGNO Sensitivity
- EL and NID studies with MANGO



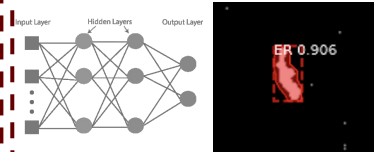
G. Dho

- PMT Signal
- 3D Reconstruction



D. Marques

- Track Reconstruction
- PID with ML

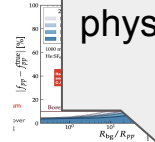
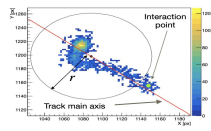


A. Prajapati



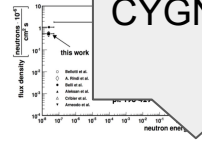
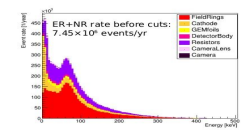
E. Baracchini

- Directionality
- Neutrino feasibility study



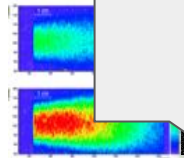
If CYGNO_30 is funded then it becomes a physics case.

- Neutron Simulation
- Background Simulation and Shielding



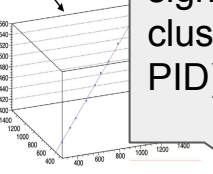
Measurement of annual variation of neutron with CYGNO_04.

- CYGNO
- EL
- with



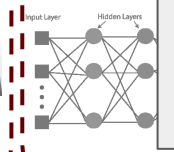
- 1) Use of NID to reduce diffusion
- 2) Study of EL and Gas mixture to amplify light output.

- PMT
- 3D

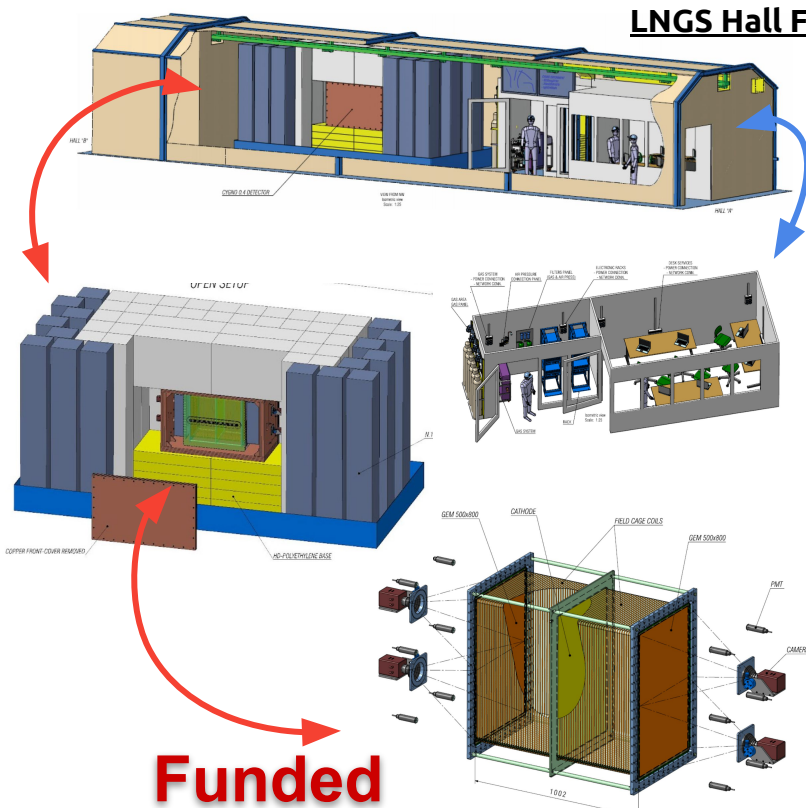


Study of NID with PMT, which is showing the signs of single ionization cluster (best way to do PID).

- Track
- PID

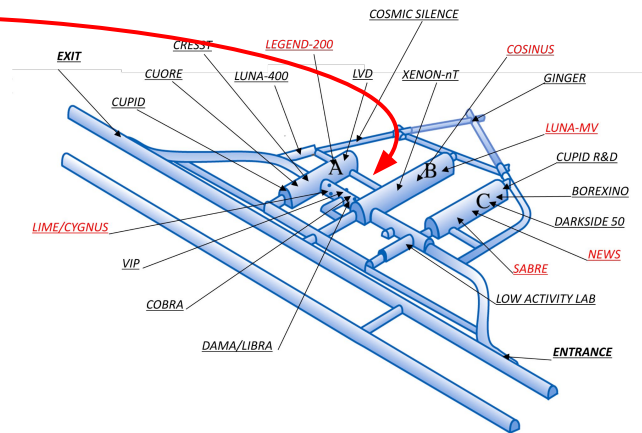


Improvement in simulation and use of different algorithm to improve Reconstruction and PID.

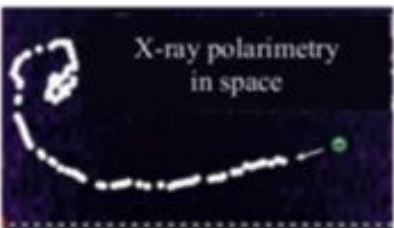


Funded

LNGS Hall F



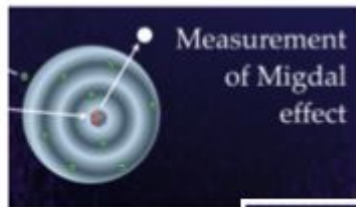
- Setting limits with **CYGNO_04**
- **CYGNO_04** data analysis
- Minimisation of radioactivity for **CYGNO_04** and **CYGNO_30**



Funded!
"HypeX: High Yield Polarimetry Experiment in X-rays"
 (PRIN 2020 Prot. 2020MZ884C)

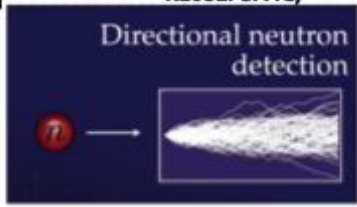
HypeX (X-Ray Polarimetry) -> Funded
 - Process of hiring a PostDoc
 (There's a dedicated poster)

FINEM (Migdal measurement) -> Funded

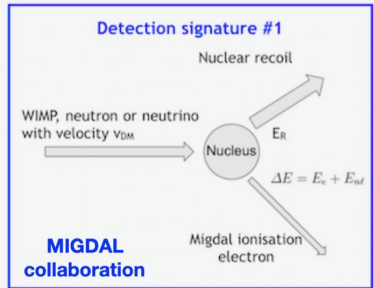


Funded!
"FINEM: Full Imaging of Nuclear recoils for Experimental Migdal measurement"
 (FARE 2020 Prot. R208LP3A4C)

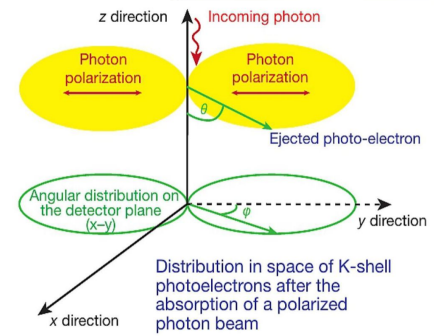
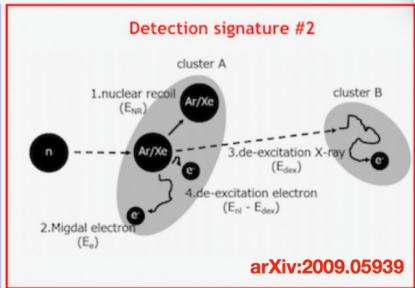
Funded!
"Zero Radioactivity for Future Experiments"
 (PRIN 2017 Prot. 2017T54J9J)



PLEASE NOTE: signature #1 is ALWAYS present (no X-ray needed) but might be difficult to distinguish ER + NR from same vertex



PLEASE NOTE: signature #2 is required for high density mixtures/low granularity readout BUT need an atom that makes X-ray



$$\frac{d\sigma_{ph}}{d\Omega} = \frac{\sigma_{ph}^{tot}}{4\pi} \left[1 + \frac{b}{2} \left(\frac{3 \sin^2 \theta \cos^2 \phi}{(1 + \beta \cos \theta)^4} - 1 \right) \right]$$

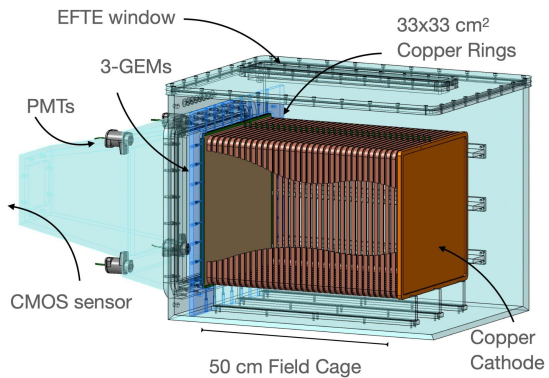


...for more info:
**CYGN0 - Directional
Dark Matter Search**

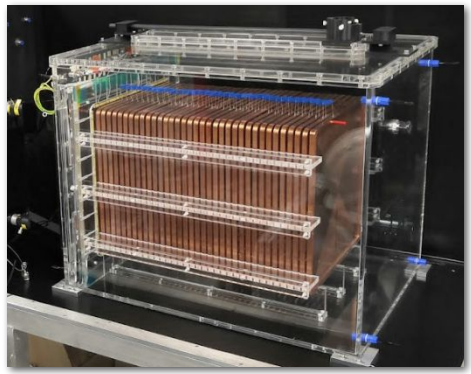
[https://www.facebook.com/
cygno.experiment](https://www.facebook.com/cygno.experiment)

THANKS!

Backup



- ❖ **50 L** gaseous TPC with **50 cm drift**
- ❖ **He:CF₄ (60:40)** gas mixture at room temperature and atm pressure
- ❖ **Triple 33x33 cm² GEM** stack for amplification
- ❖ **Optical readout**
 - 4 PMTs
 - 1 sCMOS camera (Orca Fusion)



ORCA-Fusion CAMERA SPECS

LOW NOISE AND EXCEPTIONAL
READOUT NOISE UNIFORMITY

READOUT NOISE
0.7 electrons rms
Ultra-quiet Scan

DSNU
0.3 electrons rms

PRNU
0.06 % rms
At 7500 electrons

HIGH SPEED
100 frames/s
At 2304 × 2304

PIXEL SIZE
6.5 μm × 6.5 μm

DYNAMIC RANGE
21 400:1

HIGH RESOLUTION
2304 × 2304
5.3 Megapixels

PEAK QE
80 %



Carbon tetrafluoride (CF₄)

- Significant light yield at the camera's QE peak

