



INITIUM

Multivariate techniques for track reconstruction and particle identification

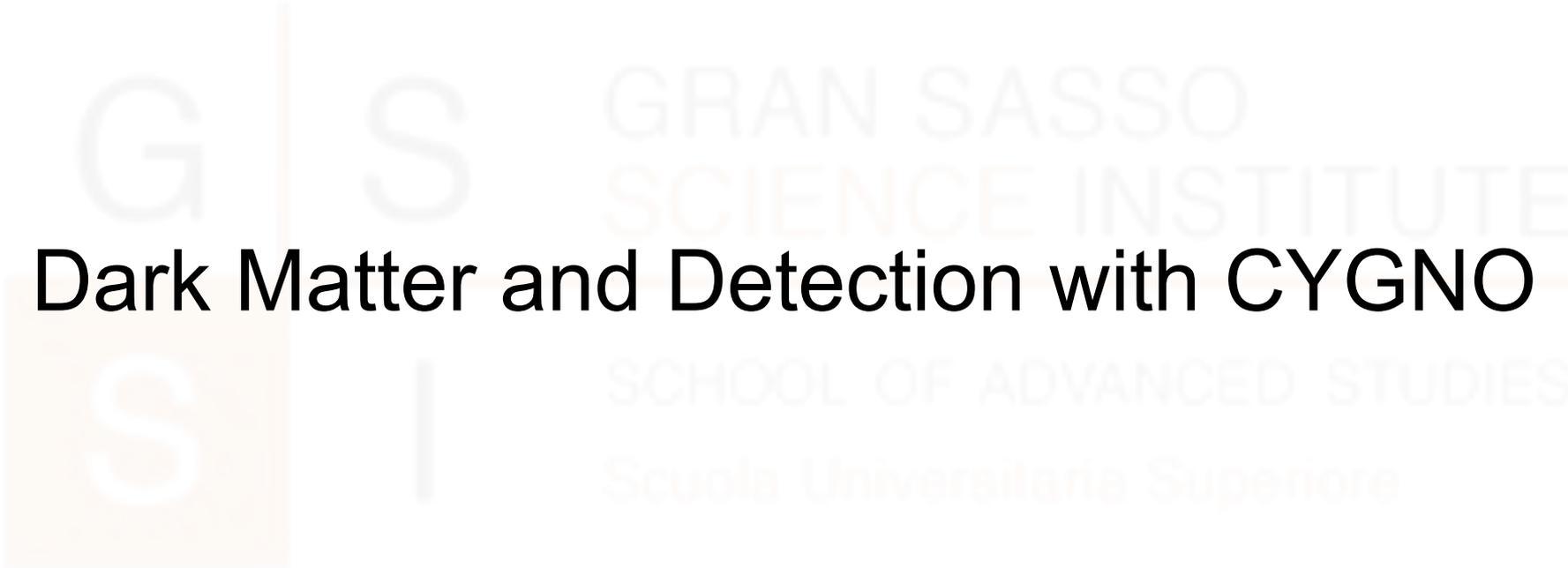
INITIUM: Innovative Negative Ion Time projection chamber for Underground dark Matter searches

Candidate: Atul Prajapati

Supervisor: Prof. E. Baracchini

Outline

- ❖ Dark Matter and It's detection
- ❖ Low Energy X-Rays data with LIME detector
- ❖ Background Rejection with Deep Learning Models
- ❖ Starting With Convolutional Neural Networks (CNN)
- ❖ Future Work



Dark Matter and Detection with CYGNO

CYGNO/INITIUM Detector

Dark Matter Signal

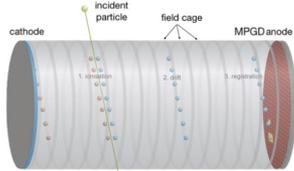
Wind of DM particles



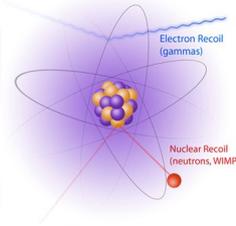
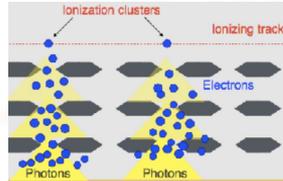
Elastic scattering with ordinary matter



Recoiling Nuclei



Amplification Region + Readout (sCMOS + PMT)

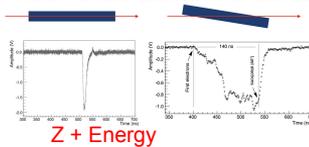


Important Sources of Background:

- ❖ Neutrinos from Sun and Atmosphere
- ❖ Cosmic Rays and Cosmogenic activation of detector materials
- ❖ Natural Radioactivity



- ❖ CYGNO uses He:CF₄ gas mixture
- ❖ 3 GEM stack is used for charge amplification
- ❖ INITIUM is a part of CYGNO project which focuses on the development of TPCs with negative ion drift using SF₆ gas



sCOMS:

- Single photon Sensitivity
- High Granularity
- Large area for detection

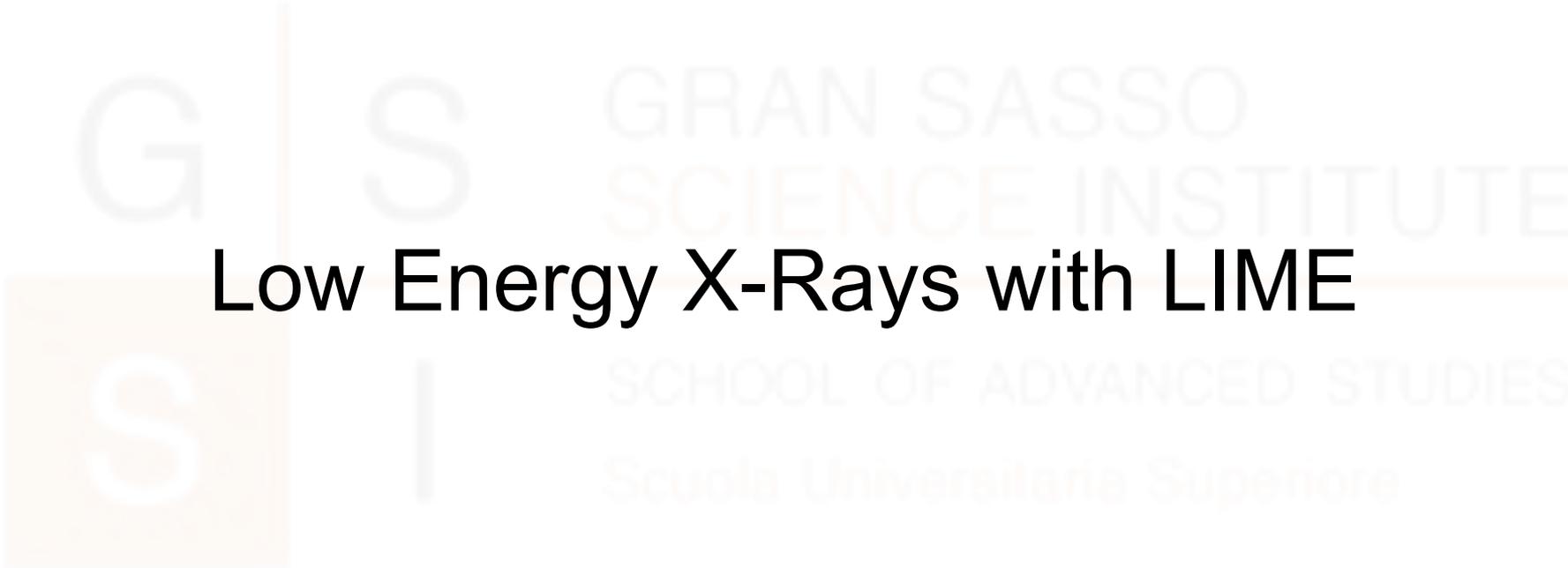
PMT:

- Fast
- Integrated energy measurement

Background can be reduced by:

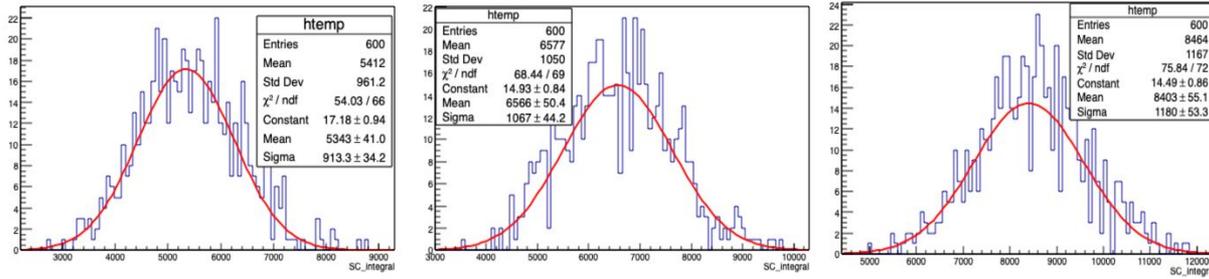
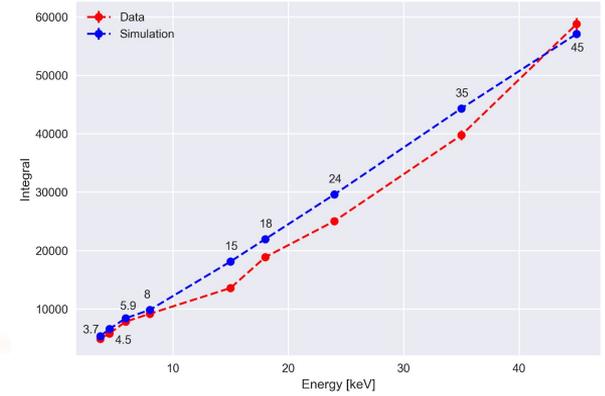
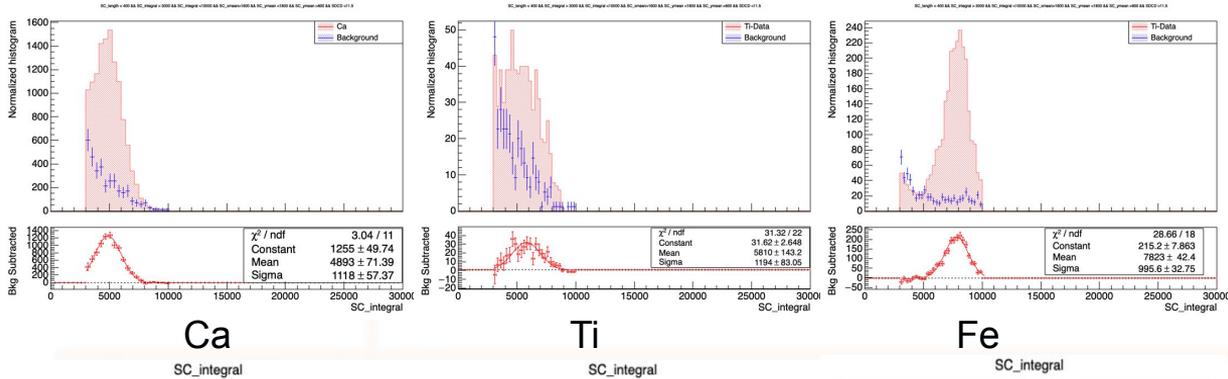
- Background Control (Cleaning, Shielding,..)
- Underground Operation (Reduces muon induced neutron)
- Electron and Nuclear recoil discrimination**

Interaction rate is extremely low and background rate is very high.

The background features a large, faint watermark of the Gran Sasso Science Institute logo. It consists of a vertical line on the left, with a large 'G' above and a large 'S' below it. To the right of this line, there is another large 'S' above and a large 'I' below it. Further to the right, the text 'GRAN SASSO SCIENCE INSTITUTE' is written in a large, light font, with 'SCHOOL OF ADVANCED STUDIES' and 'Scuola Universitaria Superiore' in smaller fonts below it.

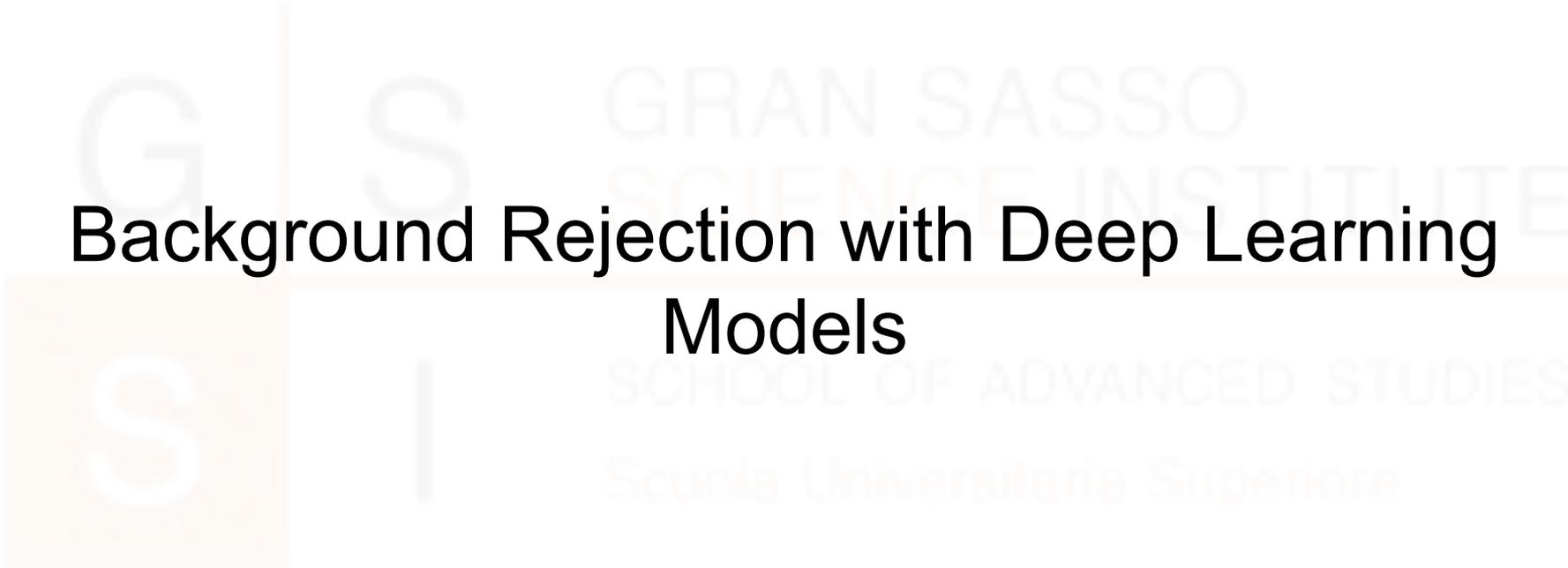
Low Energy X-Rays with LIME

Data/MC comparison of Ca, Ti and Fe data



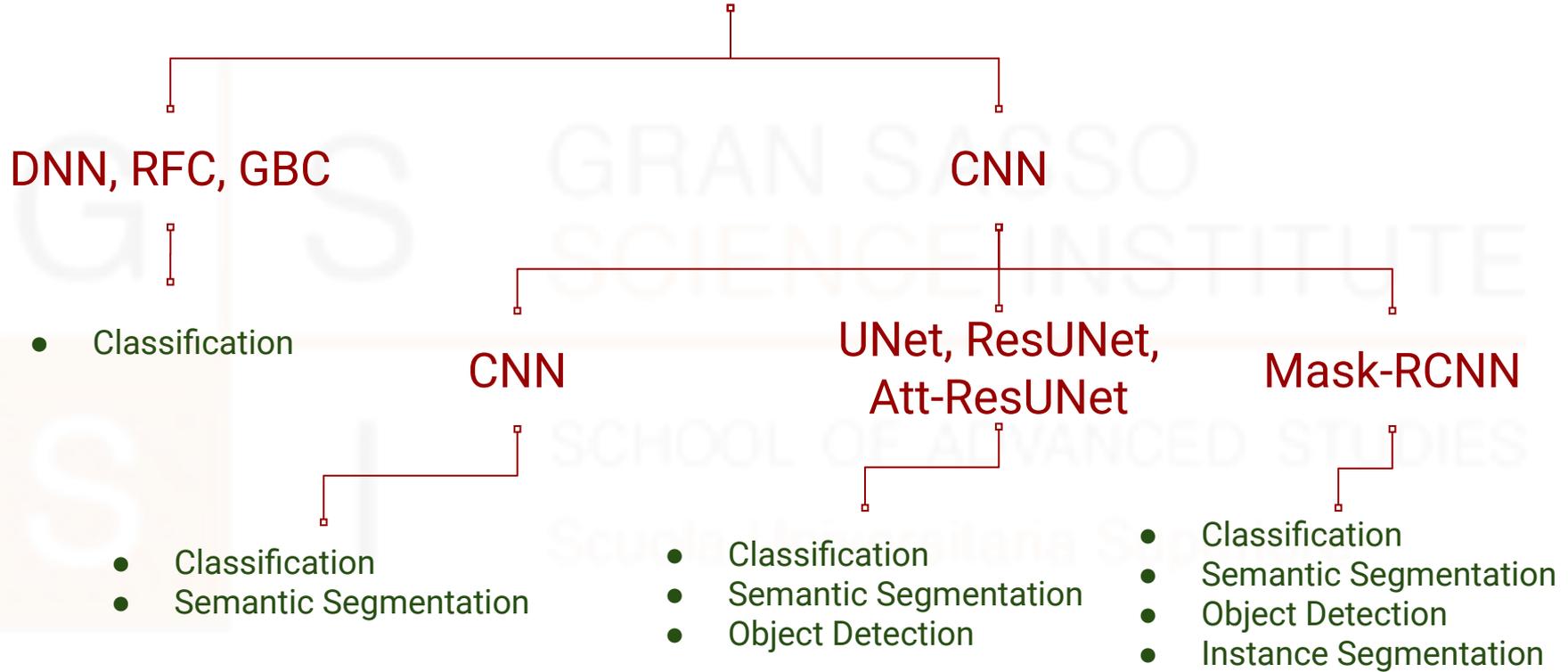
- ❖ I managed to select the events down to 3.5 keV energy.
- ❖ This work was very important to validate the data/MC agreement down to very low energy.

- ❖ I applied multiple cuts on variables to select the events.
- ❖ Integral of the data and simulated events matches well.
- ❖ Other variables were also used for data/MC comparison.



Background Rejection with Deep Learning Models

Models

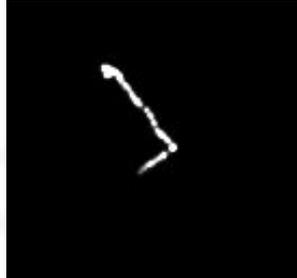


Classification



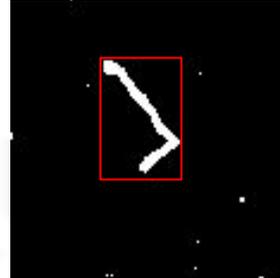
- ❖ Models: DNN, RFC, GBC
- ❖ Discriminating variables are computed
- ❖ Classification (Classify into ER and NR)

Semantic Segmentation



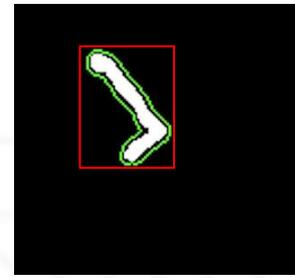
- ❖ Output of CNN, UNets, ResNet, Att-ResNet
- ❖ Classification
- ❖ Semantic Segmentation (Each pixel is classified as noise or track)

Object Detection



- ❖ Output of UNets, ResNet, Att-ResNet
- ❖ Classification
- ❖ Semantic Segmentation
- ❖ Object Detection (Finds a bounding box around the track and specifies if it is a ER or NR)

Instance Segmentation



- ❖ Output of Mask-RCNN
- ❖ Classification
- ❖ Semantic Segmentation
- ❖ Object Detection
- ❖ Instance segmentation (Finds the cluster around the tracks for each object (track) detected.)

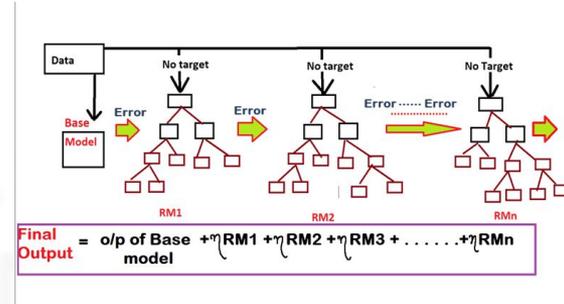
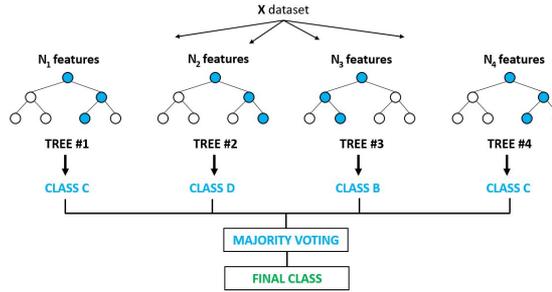
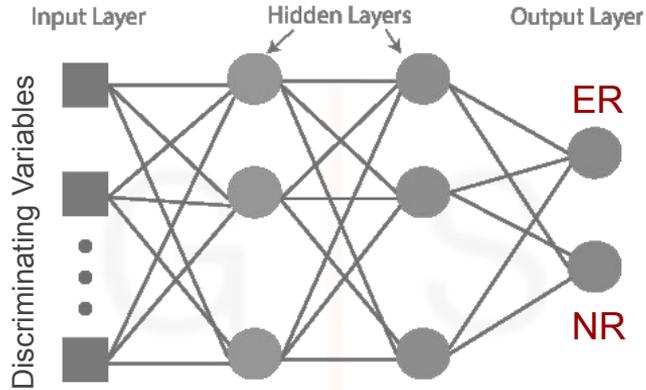
Models

DNN, RFC, GBC

- Classification

- ❖ Deep Neural Network
- ❖ Random Forest Classifier
- ❖ Gradient Boosted Classifier

Deep Learning Models



1) Deep Neural Network

- ❖ Weights of the network is optimised iteratively
- ❖ Result is the output of the last layer.
- ❖ 3 hidden layers, 10 neurons in each layer

2) Random Forest Classifier

- ❖ It can build each tree independently.
- ❖ Results are combined at the end of the process.
- ❖ 400 trees

3) Gradient Boosted Classifier

- ❖ It builds one tree at a time.
- ❖ It combines results along the way.
- ❖ 400 trees

Preparing the dataset for training

ER & NR simulation

Digitization

Reconstruction

Discriminating Variables



Interaction of the particles with gas is simulated using either GEANT4 (for ER) or SRIM (for NR).

These tracks are then projected to a 2D plane and detector effects are added like diffusion, camera noise, effective ionisation, gain fluctuation and geometrical acceptance etc.

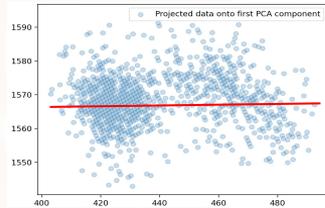
Digitized tracks are reconstructed for the tracks using an iterative density based scanning algorithm called IDBSCAN.

Reconstructed tracks are used to build several discriminating variables like skeleton, Length along principal axis, Charge uniformity, Maximum density, Slimness, Integral etc.

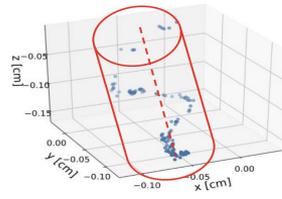
E Baracchini et. al., "Identification of low energy nuclear recoils in a gas TPC with optical readout", arXiv:2007.12508v1

Training the Models

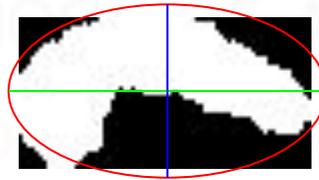
- ❖ Energy range used for training is 1-36 keV for both ER and NR.
- ❖ 5000 events for each energy for **ER**.
- ❖ 3000 events for each energy for **NR**.
- ❖ **Variables:** thin_track, SDCD, CylThick, ChargeUnif, LAPA, MaxDen, eta, curlyness, SC_nhits, SC_integral, SC_length, SC_width, delta, slimness



LAPA



CylThick



slimness



thin_track

Observables for recoil identification in gas TPCs: arXiv:2012.13649v1
GEM-based TPC with CCD Imaging for Directional Dark Matter Detection: arXiv:1510.02170v3

I presented the development of all the discriminating variables in previous year's presentation.

Classical Approach for Background Rejection

- ❖ Applying cuts on all the variables that I used for training.

Signal Events (N_{signal}) = No. of NR events from the variable passing the cut

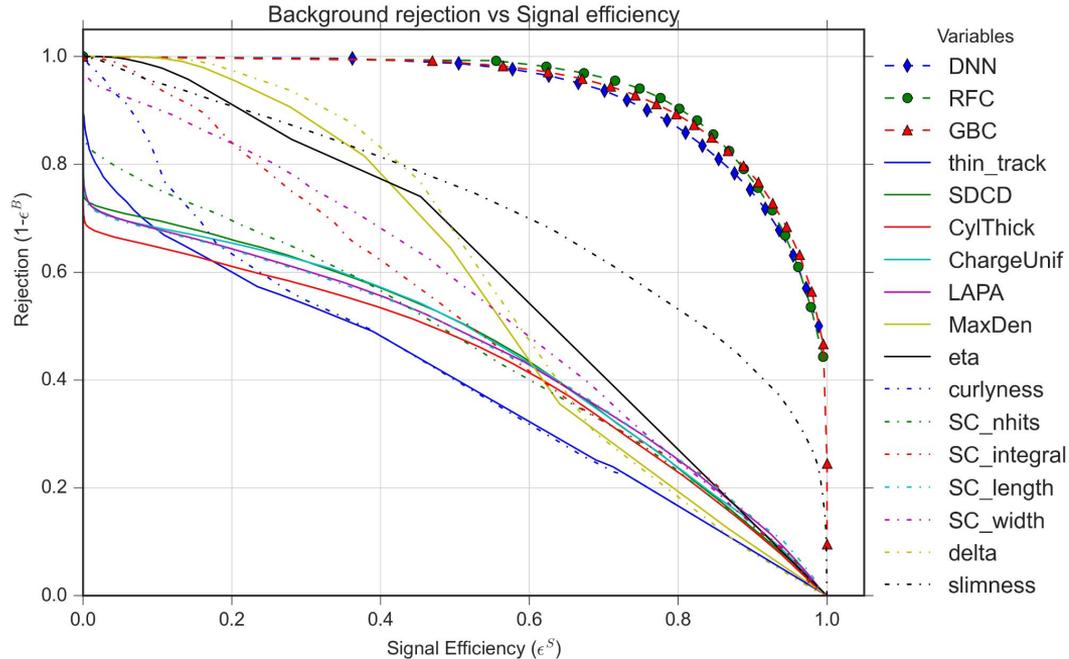
Bkg Events (N_{bkg}) = No. of ER events from the variable passing the cut

Signal efficiency (S_{eff}) = $N_{\text{signal}}/N_{\text{total,sig}}$

Bkg. Efficiency (B_{eff}) = $N_{\text{bkg}}/N_{\text{total,bkg}}$

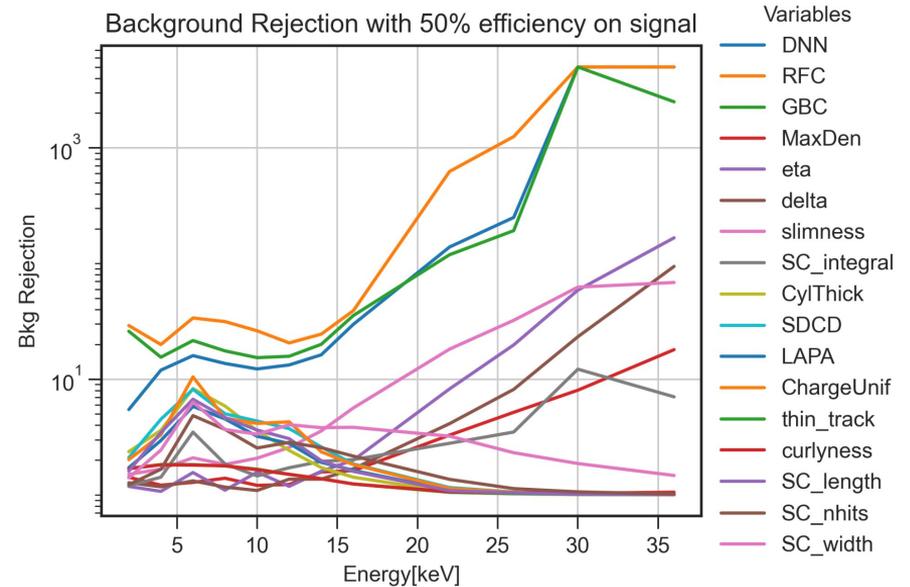
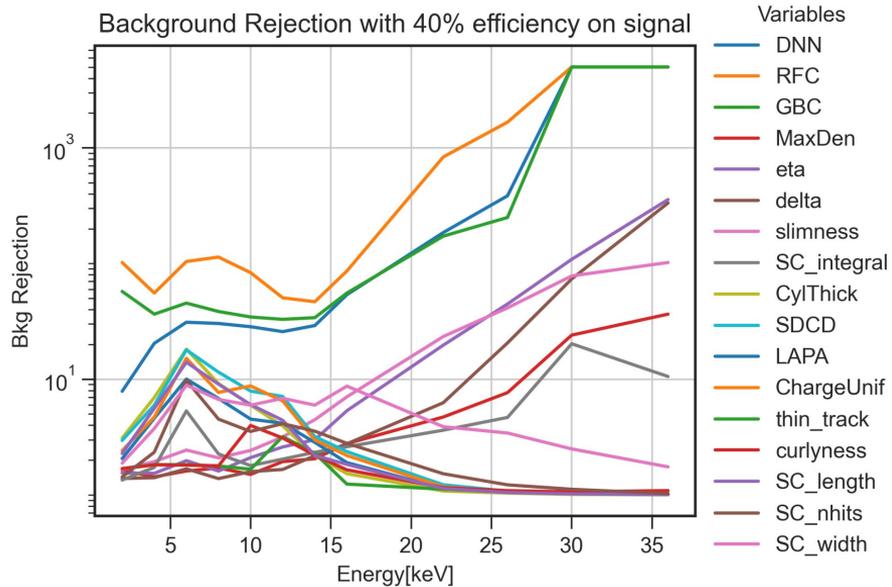
Bkg. Rejection = $N_{\text{total,bkg}}/N_{\text{bkg}}$

Rejection efficiency vs Signal Efficiency



- ❖ All the variables shown in the plot show the rejection efficiency with classical approach.
- ❖ Rejection of background events was then computed at 40% and 50% signal efficiency.

Background Rejection



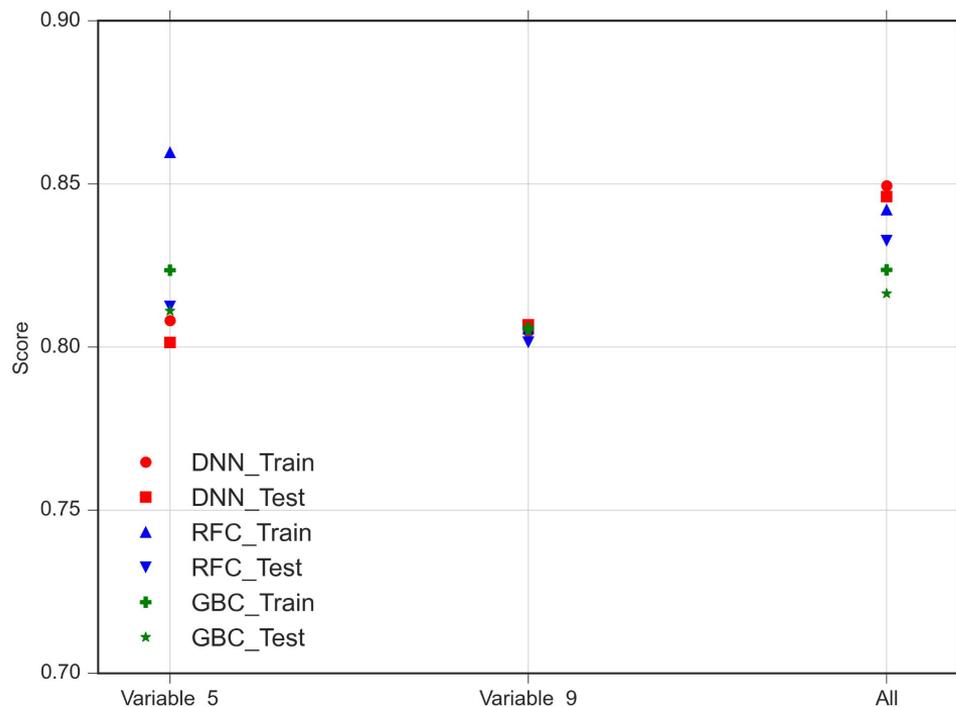
- ❖ Background Rejection is plotted with 40% and 50% signal efficiency in each energy bin.
- ❖ All the variables shown in the plot show the background rejection with classical approach.

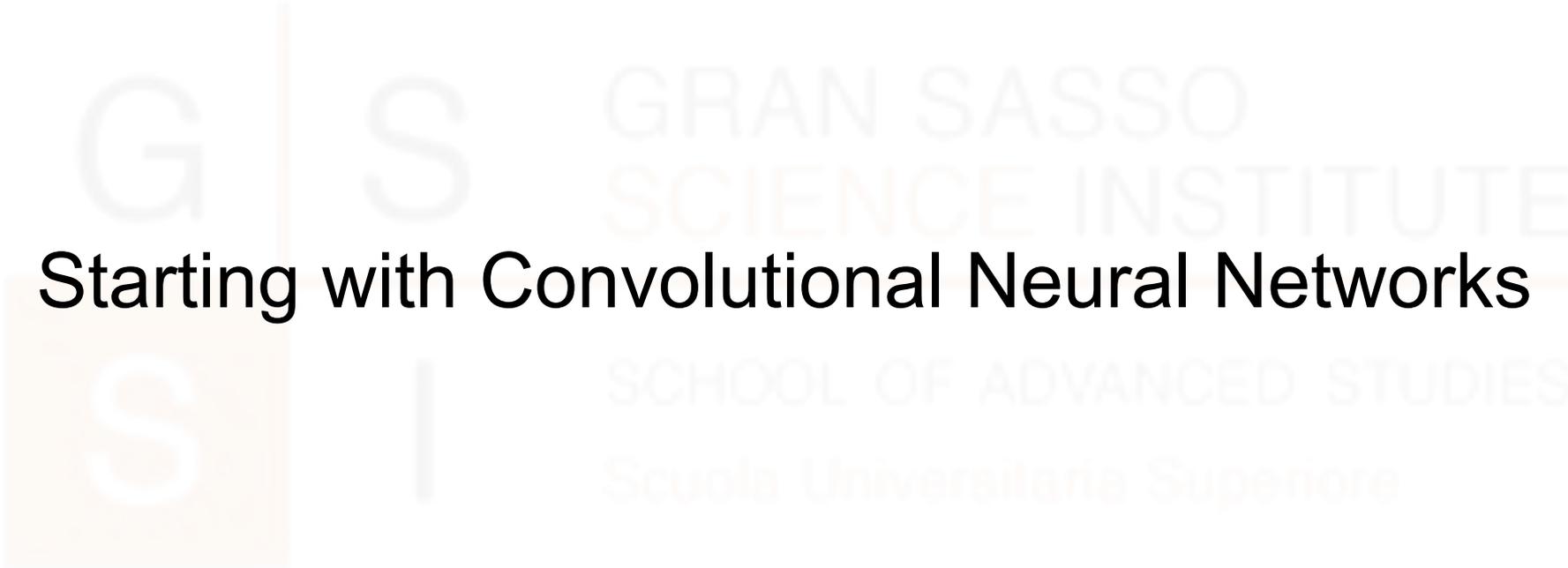
Training and Testing Scores for all 3 models

- ❖ All the 3 models were trained on all 3 different datasets namely: **All, Variable_5**: MaxDen, eta, delta, slimness, integral.

Variable_9: thin_track, SDCD, CylThick, ChargeUnif, LAPA, curlyness, SC_nhits, SC_length, SC_width

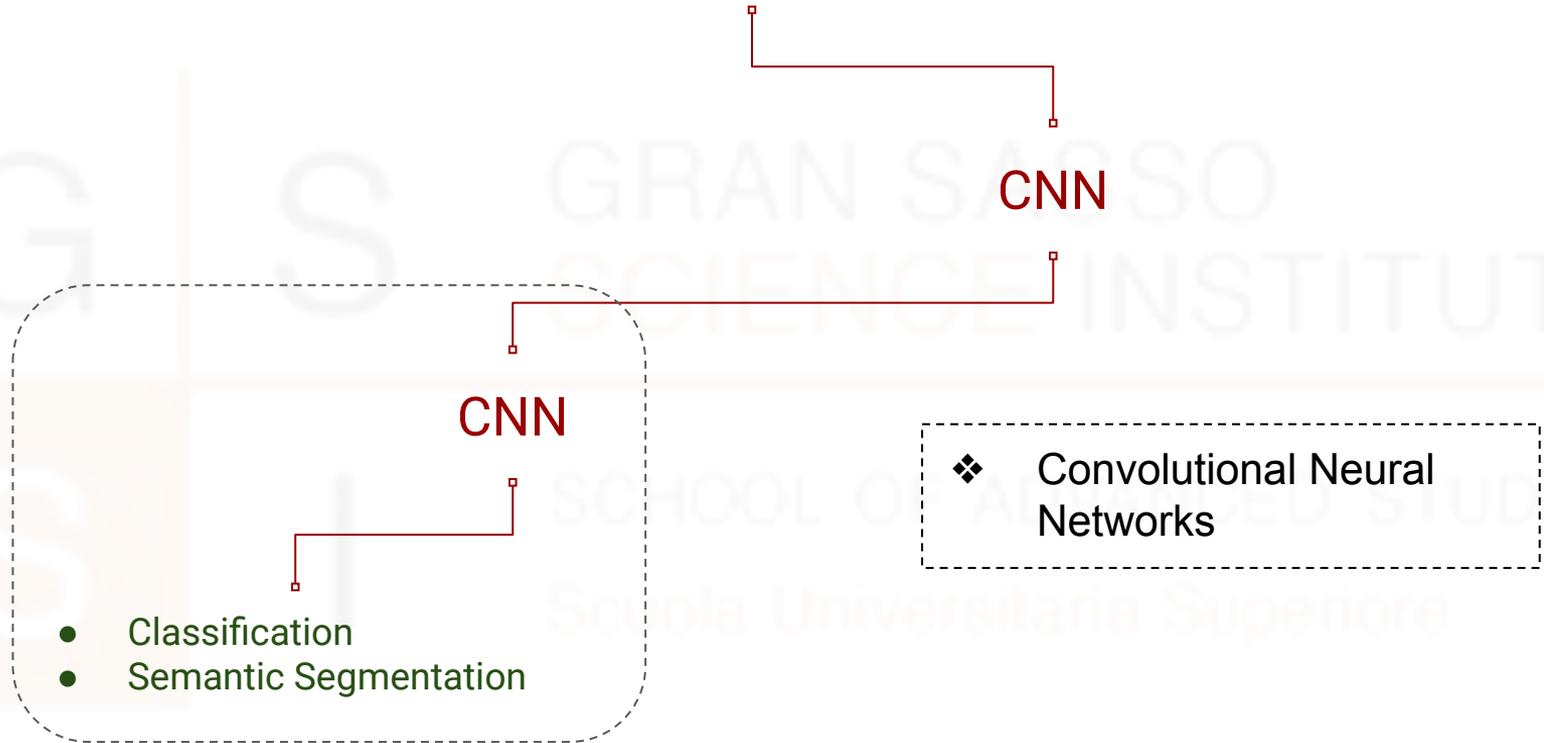
- ❖ Score of training and testing of all the models is plotted for the different datasets.



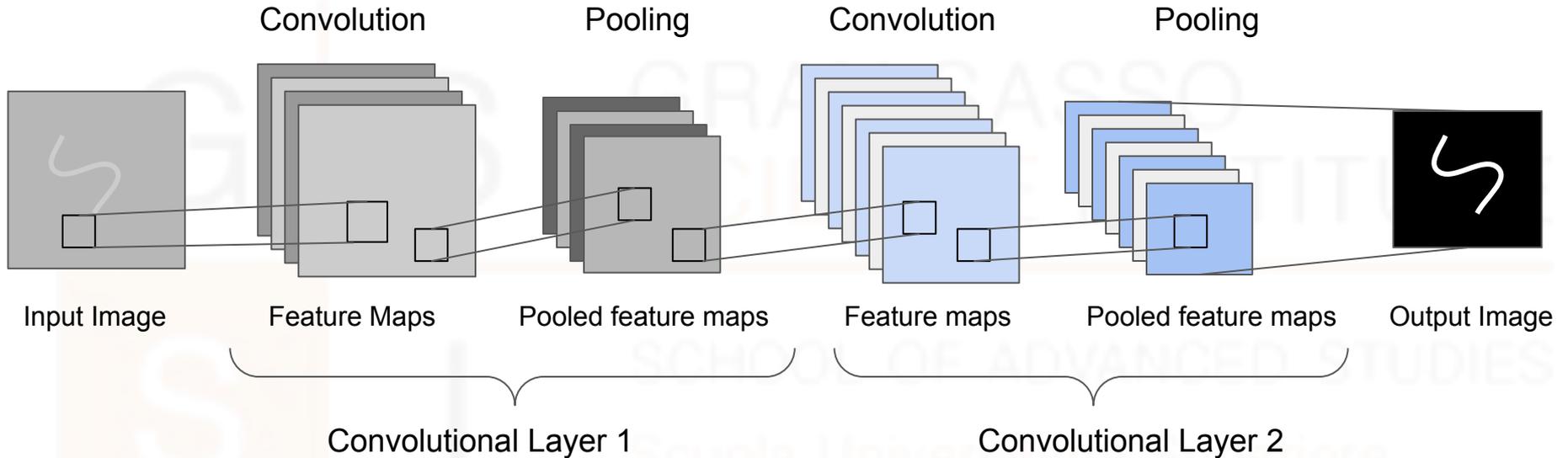
The background features a large, faint watermark of the Gran Sasso Science Institute logo. It consists of a vertical orange bar on the left with a white 'G' above and a white 'S' below. To the right of this bar is a white 'S' above a white 'I'. Further to the right, the text 'GRAN SASSO SCIENCE INSTITUTE' is written in a light orange color, with 'SCIENCE INSTITUTE' on a second line. Below that, 'SCHOOL OF ADVANCED STUDIES' and 'Scuola Universitaria Superiore' are written in a light grey color.

Starting with Convolutional Neural Networks

Models



Convolutional Neural Networks



- ❖ A CNN (ConvNet) is a Deep Learning algorithm which can take in an input image, assign importance to various aspects/objects in the image and be able to differentiate from other.
- ❖ The pre-processing required in CNN is much lower as compared to other classification algorithms.

Models

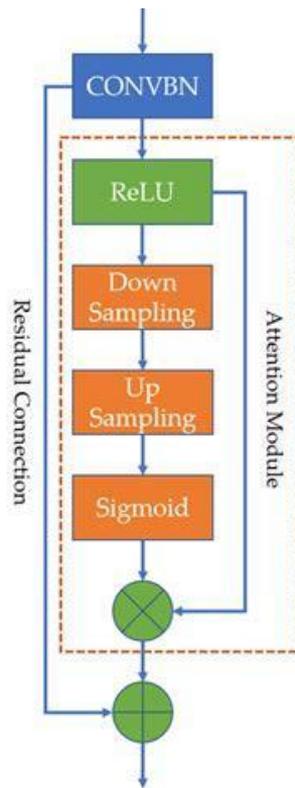
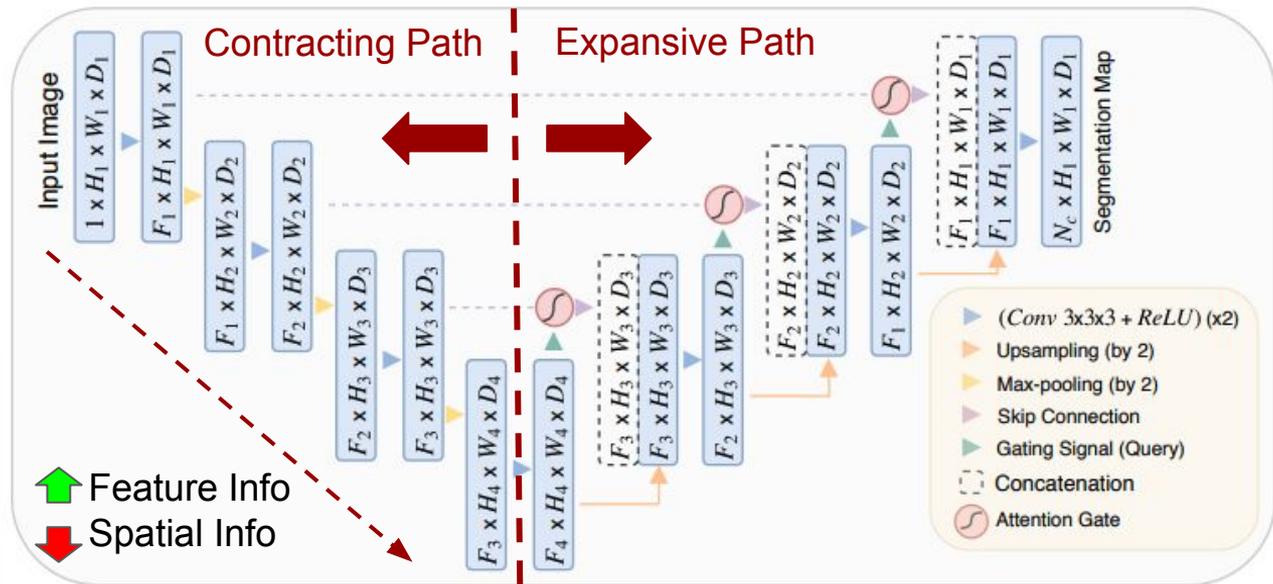
CNN

UNet, ResNet,
Att-ResNet

- ❖ UNets
- ❖ Residual UNets
- ❖ Attention Residual UNets

- Classification
- Semantic Segmentation
- Object Detection

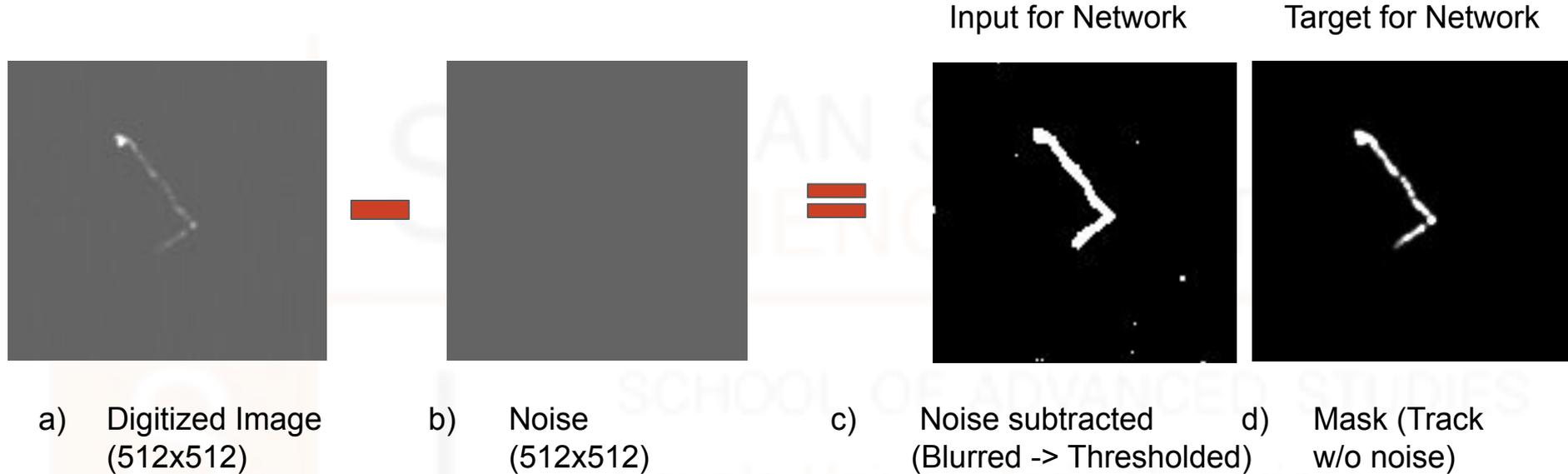
Attention Residual UNets



Architecture is very similar to ResNet, except there is an extra block called attention block. Attention in U-Nets is a method to highlight only the relevant activations during the training.

It reduces computation resources wasted on irrelevant activations and provides better generalization of the network.

Preparing Data for training CNNs



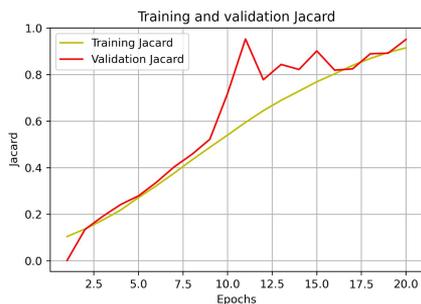
Noise map is subtracted from the digitized image. Noise subtracted image is then passed through a median filter with a kernel size of 3. Blurred image is thresholded with a threshold of 1 (pixels with intensity more than 1 becomes 255 and rest 0). These images are input for the network.

Masks are produced by digitizing the tracks without noise. Network is trained to produce images similar to masks.

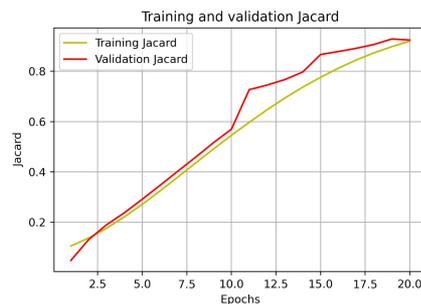
Training and Validation accuracy



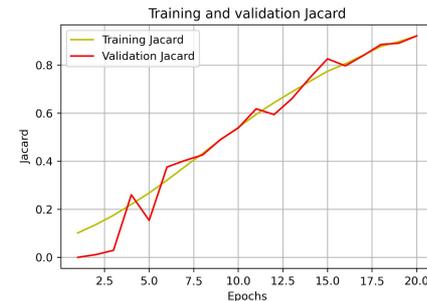
(a) CNN



(b) UNet



(c) Att-UNet

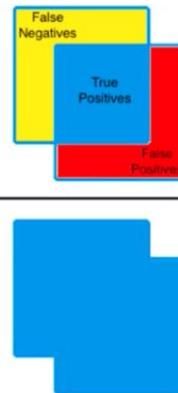


(d) Att-ResUNet

Jaccard Coefficient = Intersection over Union

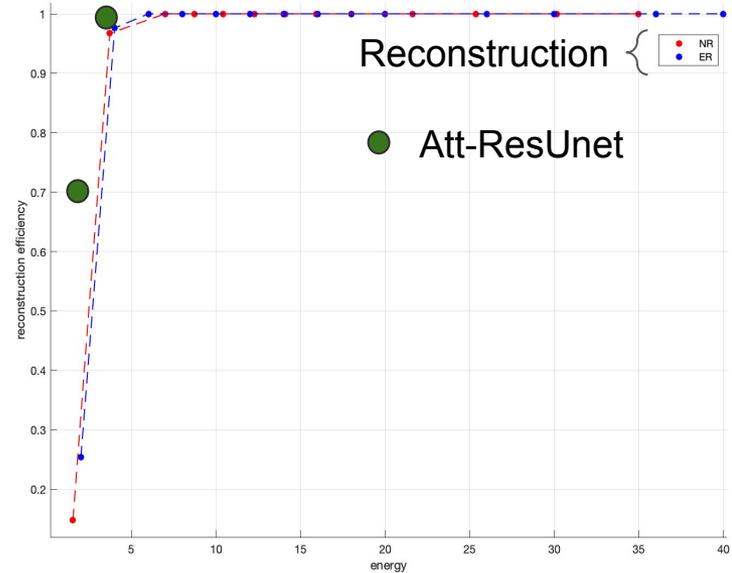
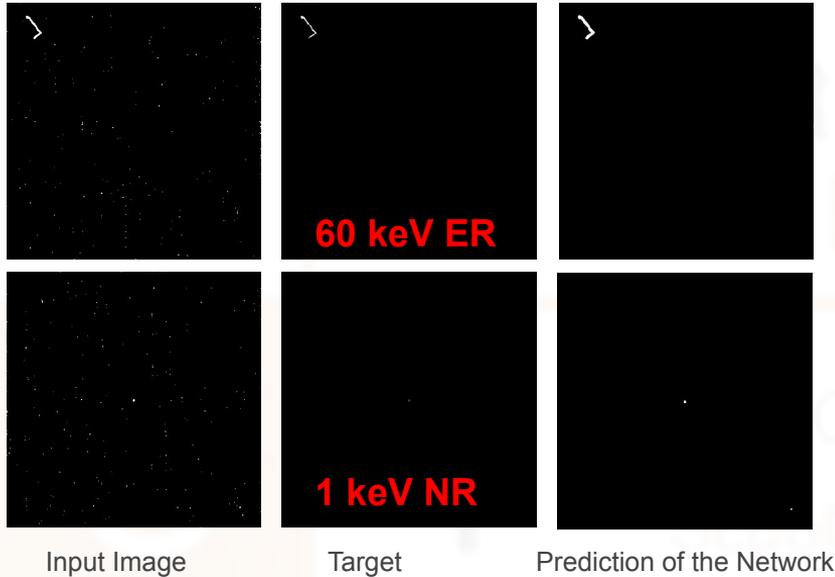
Jaccard Coefficient measures the similarity between between 2 sets of data. The closer to 1 means more similar data.

$$IoU = \frac{\text{Area of Overlap}}{\text{Area of Union}}$$



All the models were trained with 600 images and for 20 epochs.

Prediction from Att-ResUNet



- ❖ Predicted images were used to find the cluster around the track using OpenCV.
- ❖ Reconstruction efficiency at 1 keV of NR is ~ 70% which is ~10% with usual reconstruction algorithm and at 3 keV NR is 100% and with usual reconstruction algorithm it is ~97%.

Models

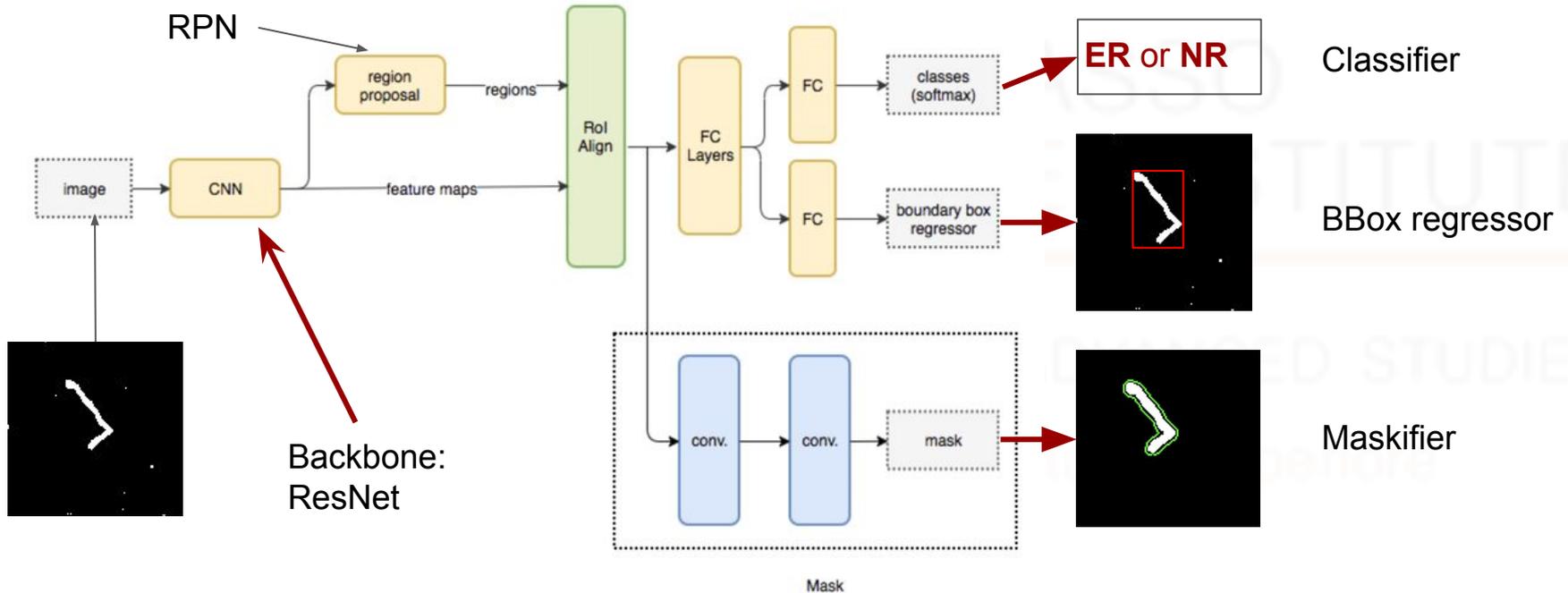
CNN

- ❖ Mask - Region based Convolutional Neural Networks (Mask-RCNN)

Mask-RCNN

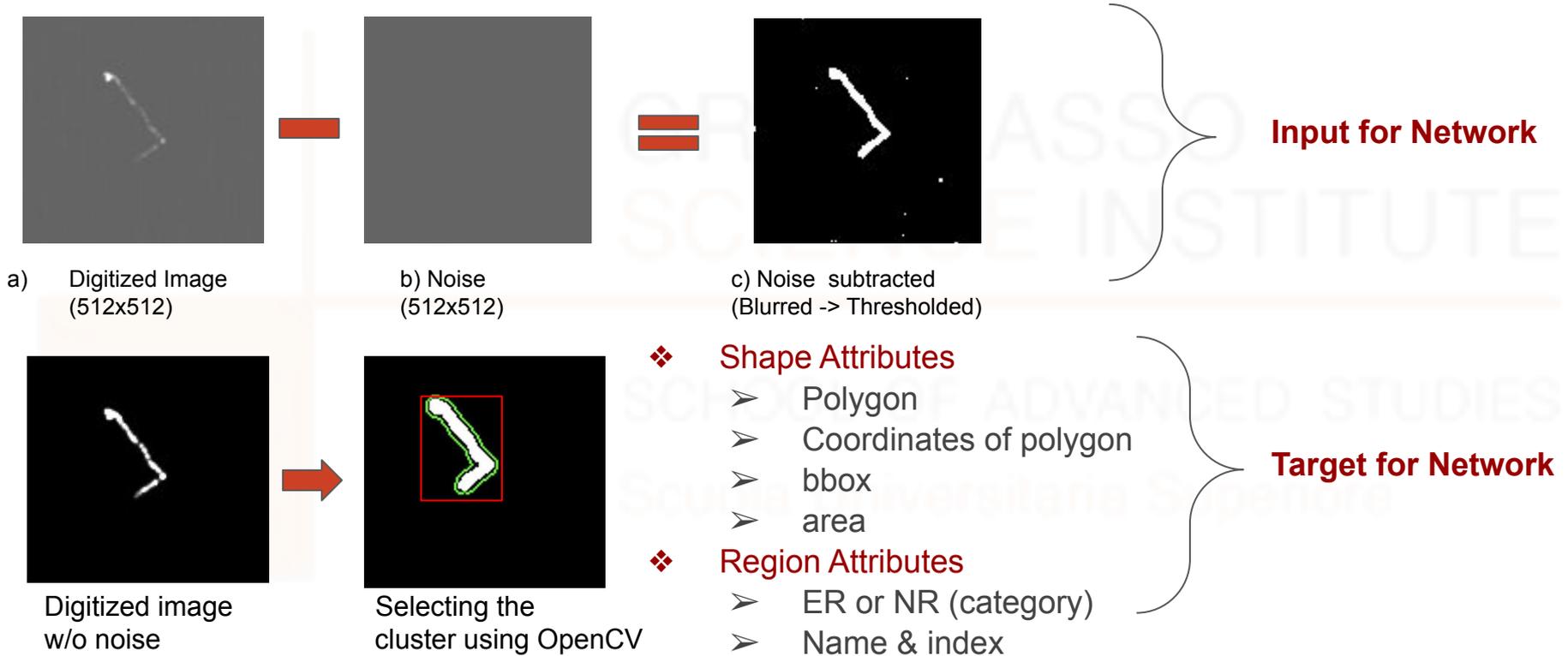
- Classification
- Semantic Segmentation
- Object Detection
- Instance Segmentation

Architecture of Mask-RCNN



Mask-RCNN Paper: <https://doi.org/10.48550/arXiv.1703.06870>

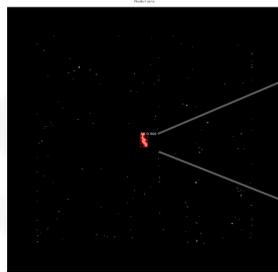
Data for Mask-RCNN



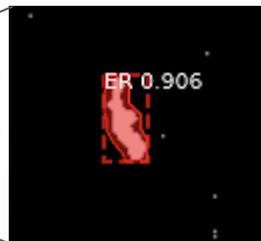
First results from Mask-RCNN



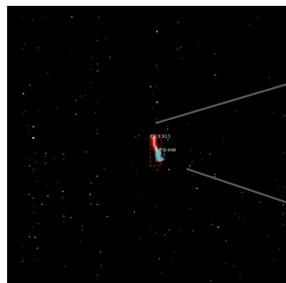
Input



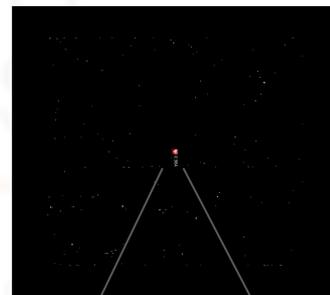
Prediction



Zoomed Track



Input



Prediction



Model was trained just with 4 input images and for 20 epochs.

Future Work

- ❖ We are working on simulating larger samples for training the networks to test the rejection capabilities of the network
- ❖ To train Mask-RCNN with larger sample of simulated data
- ❖ Working with Gustavo V. Lopez from CS department at GSSI to build a network in which we can integrate the sCOMS data and PMT data
- ❖ Test all these models on data
- ❖ Will work with the negative ions data from MANGO detector

References

1. G. Bertone et. al., “Particle Dark Matter: Evidence, Candidates and constraints”, [arXiv:hep-ph/0404175](https://arxiv.org/abs/hep-ph/0404175)
2. V.C. Antochi et. al., “Combined readout of a triple-GEM detector”, JINST13 (2018) 05, P05001
3. Physics Reports Volume 627, 20 April 2016, Pages 1–49
4. E. Baracchini et al.: Measur.Sci.Tech.32(2021)2,025902
5. C. J. Martoff et. al., “Suppressing drift chamber diffusion without magnetic field”, Nucl. Instr. and Meth. in Phys. Res. A, 440 (2000), p. 355.
6. E Baracchini et. al., “Identification of low energy nuclear recoils in a gas TPC with optical readout”, [arXiv:2007.12508v1](https://arxiv.org/abs/2007.12508v1)
7. M. Ghrear et. al., “Observables for Recoil Identification in High-Definition Gas Time Projection Chambers”, [arXiv:2012.13649v1](https://arxiv.org/abs/2012.13649v1)
8. N. S. Phan et. al., “GEM-based TPC with CCD Imaging for Directional Dark Matter Detection”, [arXiv:1510.02170v3](https://arxiv.org/abs/1510.02170v3)

Schools/Conferences

- [1] School on Underground Physics, SOUP 2021, 28 June - 2 July (Online)
- [2] 4th International School on Deep Learning, DeepLearn 2021 Summer, Las Palmas de Gran Canaria, Spain, 26 July - 30 July (In person)
- [3] Quantum Sensors for Fundamental Physics, QSFP 2021, 6 September -17 September (Online)
- [4] Advanced Computing and Analysis Techniques in Physics Research (ACAT) 2021, 29 November - 3 December 2021, (Online)
 - **Presented a Poster and won the best poster award**
- [5] CYGNO Collaboration meeting, 21-22 Dec 2021 at GSSI. (Online)
 - **Oral Presentation**
- [6] SiPM Radiation: Quantifying Light for Nuclear, Space and Medical Instruments under Harsh Radiation Conditions, 25-29 April 2022 (Online)
- [7] INFN School of Statistics 2022, Paestum, 15-20 May 2022. (In person)
- [8] International conference on Machine Learning for Astrophysics - ML4Astro, Catania, 30 May - 1 June 2022. (In person)
 - **Presented a poster**
- [9] PyHEP 2022 Workshop, 12- 16 Sept 2022. (Online)
- [10] Micro Pattern Gaseous Detectors 2022, 11 -16 December 2022, Submitted an abstract for oral presentation.

Publications

- [1] F. D. Amaro, E. Baracchini et. al., “**Directional Dark Matter Searches with CYGNO**”, MDPI, July 2021, DOI: [10.3390/particles4030029](https://doi.org/10.3390/particles4030029)
- [2] F. D. Amaro, E. Baracchini et. al., “**The CYGNO Experiment**”, MDPI, February 2022, [arXiv:2202.05480](https://arxiv.org/abs/2202.05480)
- [3] F. D. Amaro et al., “**Study of Performance of Different Photodetectors and Electrical Signal for Fast Detection in Fifty Litres CYGNO Prototype (LIME)**,” IEEE, 2021, pp. 1-5, [doi:10.1109/NSS/MIC44867.2021.9875803](https://doi.org/10.1109/NSS/MIC44867.2021.9875803)
- [4] F. D. Amaro et. al., “**Performances of a 3D optical readout TPC for the CYGNO experiment**”, 2022, DOI: <https://doi.org/10.22323/1.398.0799>
- [5] C.A.J. O’Hare, D. Loomba et. al., **Recoil imaging for directional detection of dark matter, neutrinos, and physics beyond the Standard Model**, July 2022, [arXiv:2203.05914v3](https://arxiv.org/abs/2203.05914v3) (Submitted)
- [6] A. Prajapati et. al., **Reconstruction and Particle Identification with CYGNO Experiment**, Springer Nature, Astrophysics and Space Science Proceedings, August 2022 (Submitted)



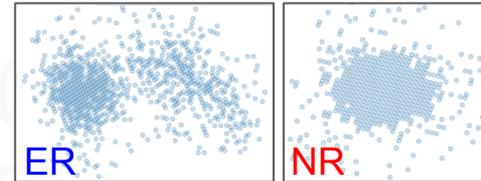
Backup

Variables

Observables for recoil identification in gas
TPCs
arXiv:2012.13649v1

❖ Standard Deviation of Charge Distribution 2D(SDCD_2D):

$$SDCD = \sqrt{\frac{\sum_{i=1}^N (\mathbf{r}_i - \bar{\mathbf{r}})^2}{N}}$$



- Electron recoils (ER) are longer, so the spread of charge is higher for ER when compared to Nuclear recoils (NR).

❖ Charge Uniformity 2D (ChargeUnif_2D):

- For each point within the charge distribution, find the average distance to all other points.
- ChargeUnif_2D is standard deviation of values computed in step 1.
- Electron recoils tend to have charge distribution which is dense in some areas and sparse in other areas, while nuclear recoils are generally uniform.

❖ Maximum Density 2D (MaxDen_2D):

- MaxDen is the value of most intense pixel from the image after rebinning it by a factor 2.
- Electrons lose their energy at a slower rate than nuclei, this suggests that electron recoils are travel greater distance between interactions resulting in more sparse energy distribution.

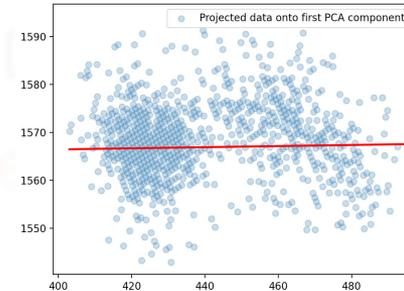
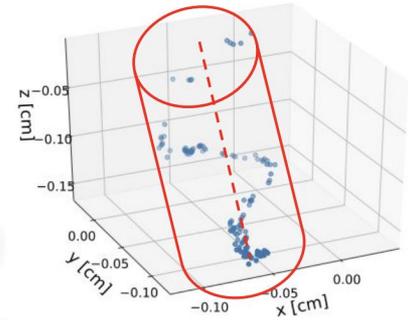
Variables

❖ Cylindrical Thickness 2D (CylThick_2D):

- For each charge, calculate the squared distance from the principal axis.
- CylThick is the sum of all squared distances.
- It is a measure of how much a recoil track deviates from the trajectory approximated by the principal axis.
- Electrons experience far more scattering compared to nuclei, so principal axis approximates NR's trajectory much more accurately than it does for ER.

❖ Length Along Principal Axis 2D (LAPA_2D):

- Project all the points in the charge distribution on to the principal axis.
- LAPA is the difference between maximum and minimum projected value.
- ER are longer compared to NRs, therefore projection is also longer.

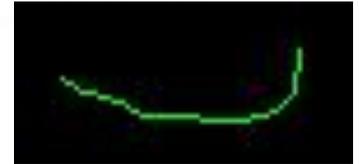


Variables

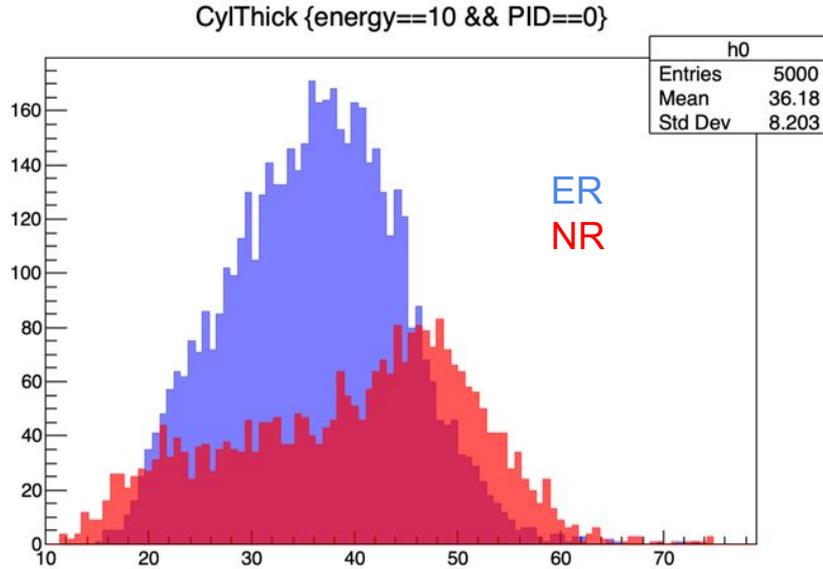
- ❖ eta:
 - MaxDen_2D divided by length (found by skeletonization)
- ❖ Light Density (delta):
 - Integral of the track divided by number of pixels in the track.
 - NR deposit higher energy over a short distance, therefore Light Density is higher for NR.
- ❖ Slimness:
 - Ratio of minor over major axis of the ellipse which bounds the track.
 - Electrons recoils suffer more scattering, so minor axis of the bounding ellipse is bigger when compared to NR which are generally straight.
- ❖ Skeleton length (thin_track):
 - Length in pixels found by thinning procedure.

GEM-based TPC with CCD Imaging for Directional Dark Matter Detection
arXiv:1510.02170v3

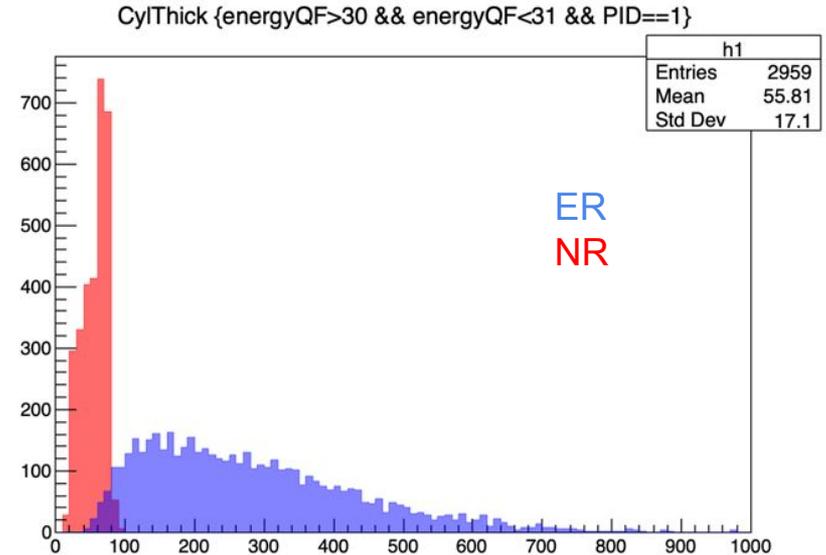
E Baracchini et. al., "Identification of low energy nuclear recoils in a gas TPC with optical readout",
arXiv:2007.12508v1



Variables with decreasing rejection at higher energy

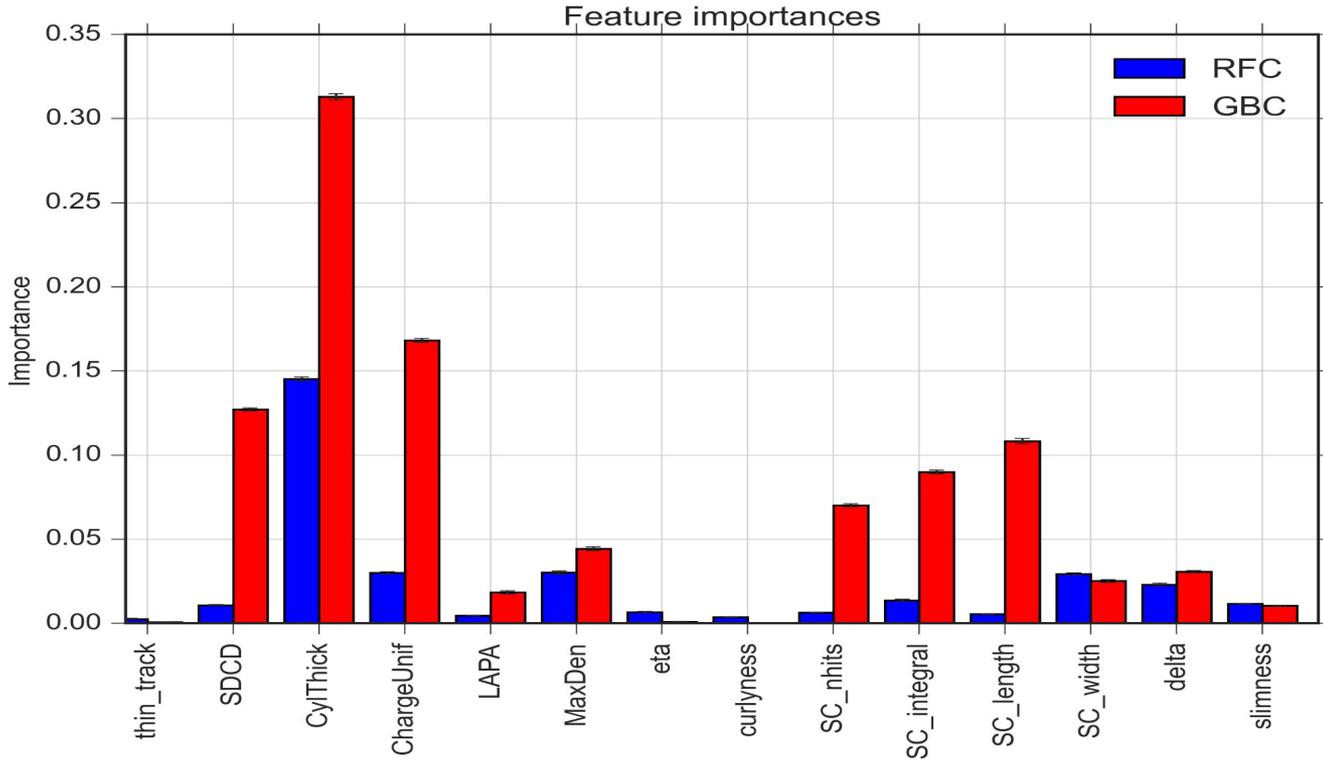


CylThick @ 10 keV

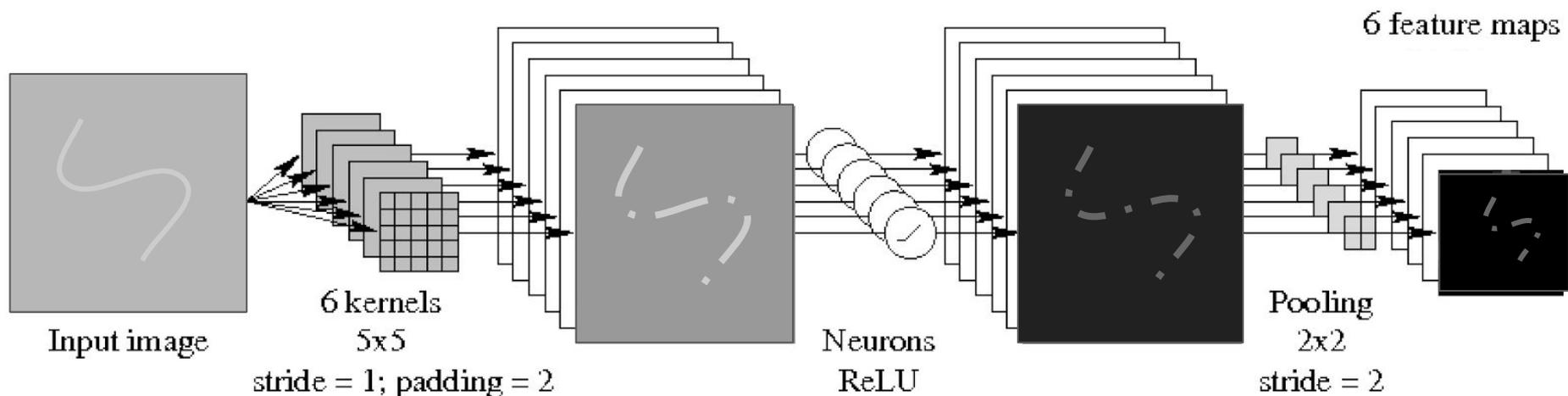


CylThick @ 30 keV

Feature Importance



Convolutional Layer

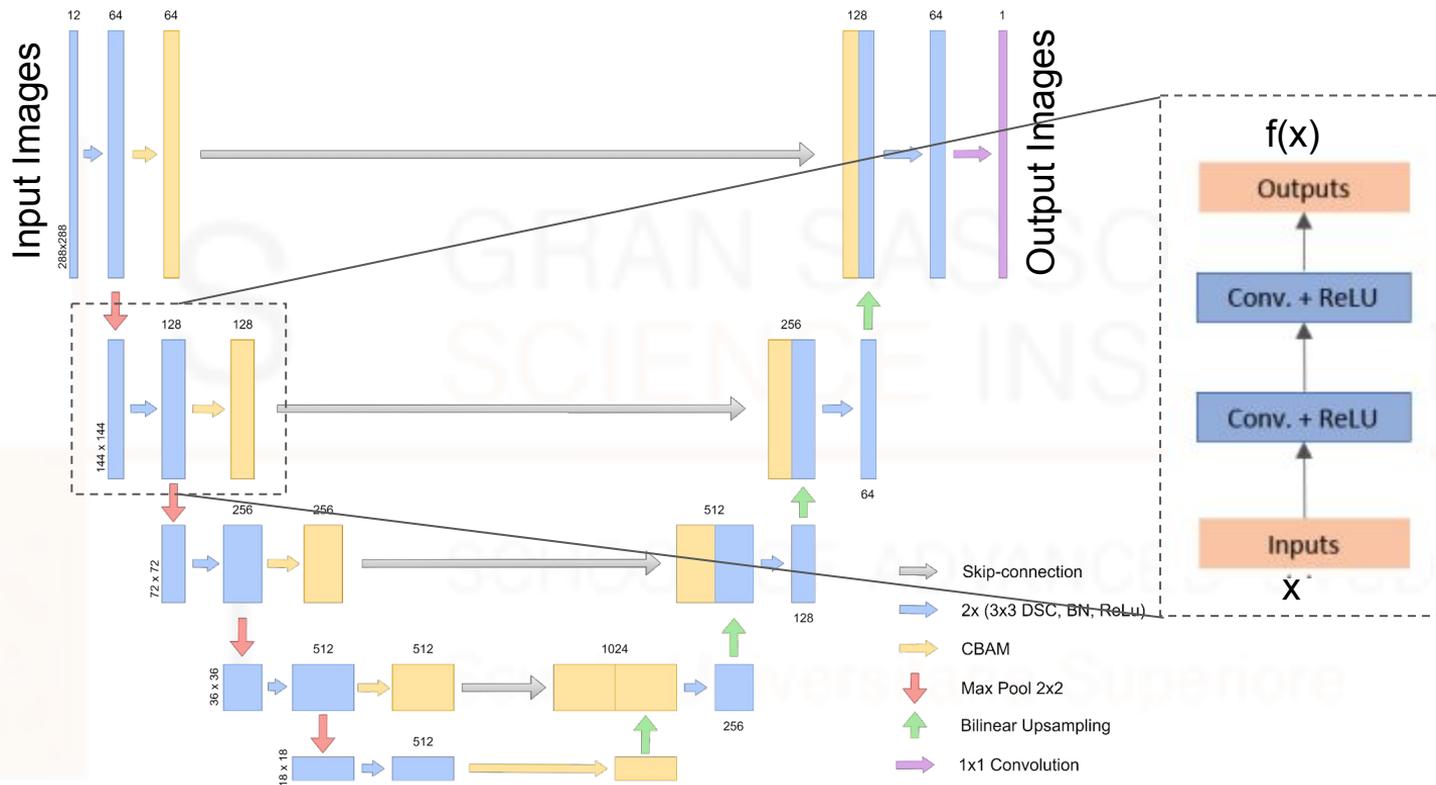


Feature Maps: Feature maps are generated by applying filters (kernels) to the input image. Filters try to gain some understanding of what features our CNN detects.

Activation Function: Activation functions decide if the neuron would fire or not.

Pooling: Pooling reduces the number of parameters and computation in the network, controlling overfitting by progressively reducing the spatial size of the network.

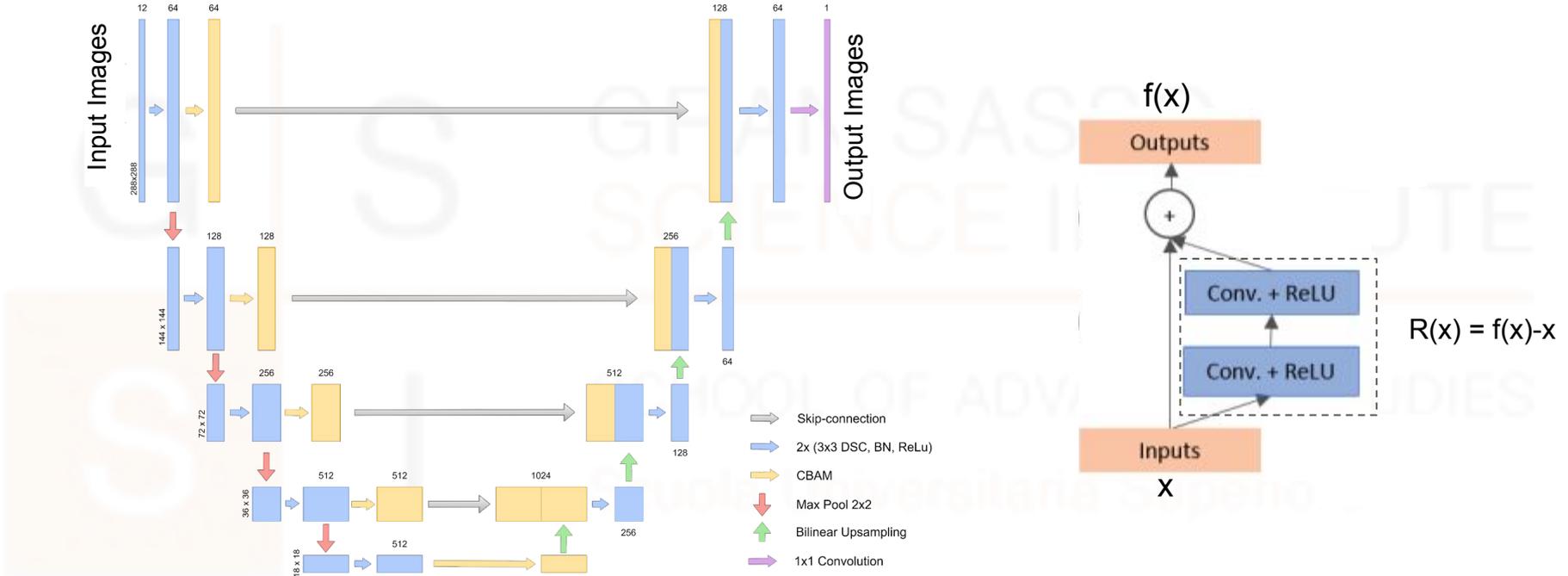
UNet



Layers are trying to learn $f(x)$ for the given input x .

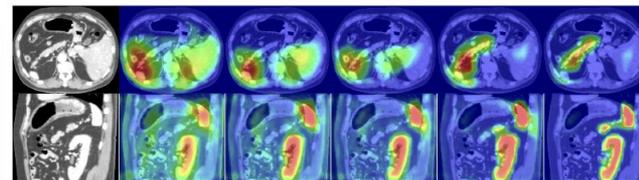
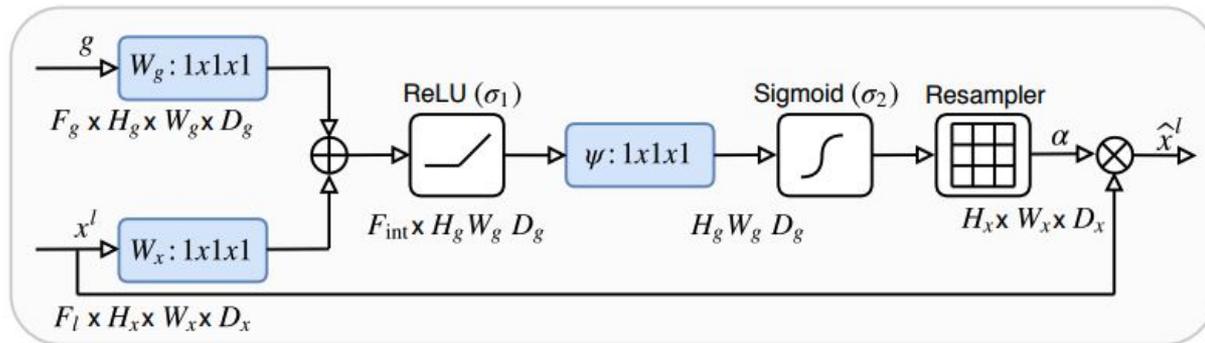
Recurrent Residual Convolutional Neural Network based on U-Net (R2U-Net) for Medical Image Segmentation
<https://arxiv.org/pdf/1802.06955.pdf>

ResUNet

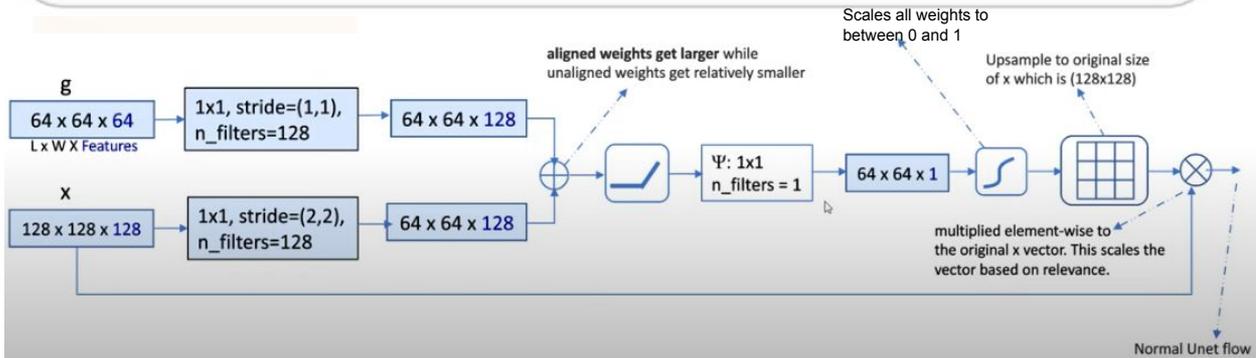


Here, Layers are trying to learn the residual unlike UNet where they try to learn $f(x)$. ResNet helps in solving the problem of vanishing gradients and also of overfitting to an extent.

Attention Block



Attention U-Net: Learning Where to Look for the Pancreas
<https://arxiv.org/pdf/1804.03999.pdf>



It reduces computation resources wasted on irrelevant activations and provides better generalization of the network.