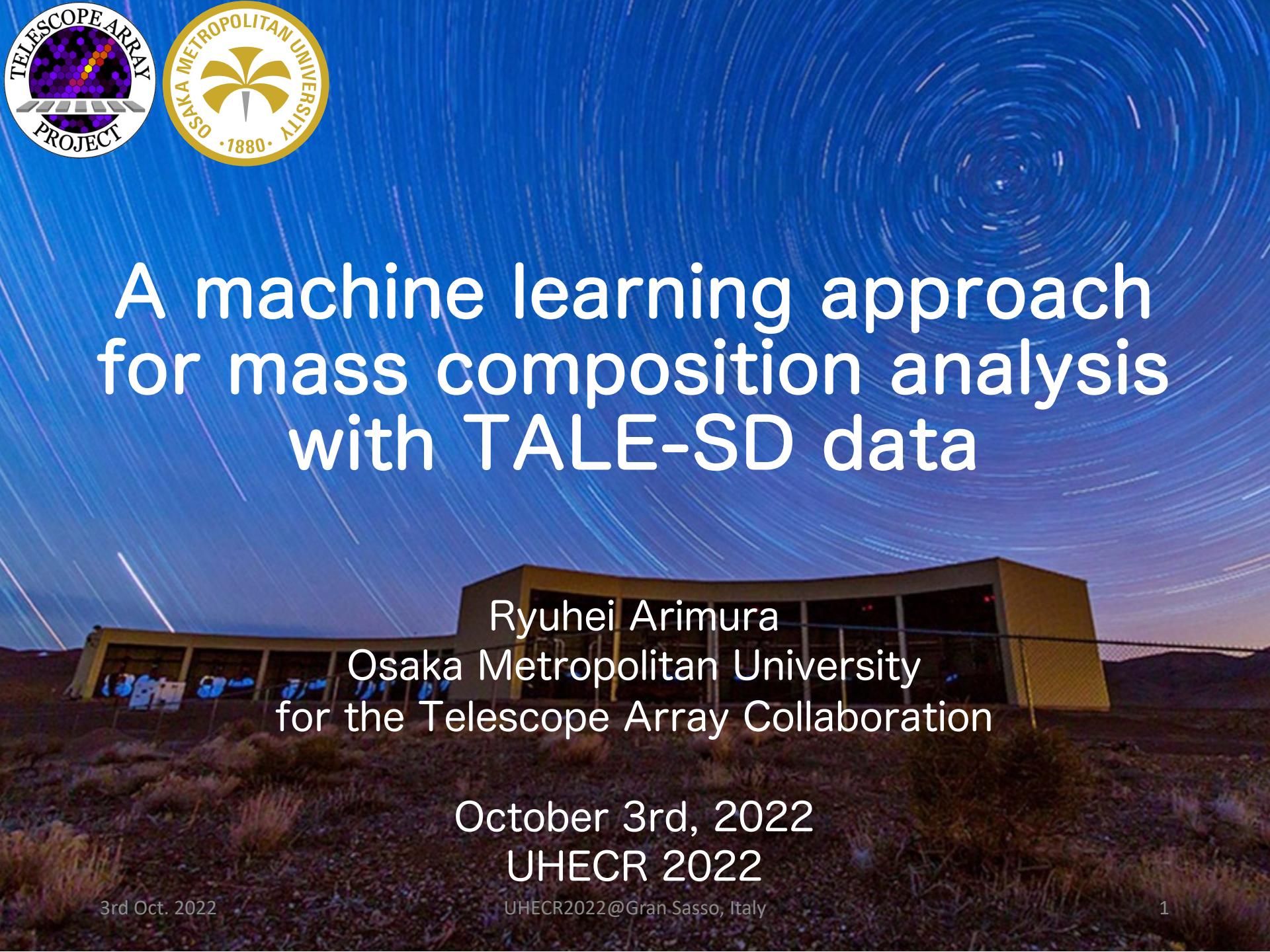




A machine learning approach for mass composition analysis with TALE-SD data

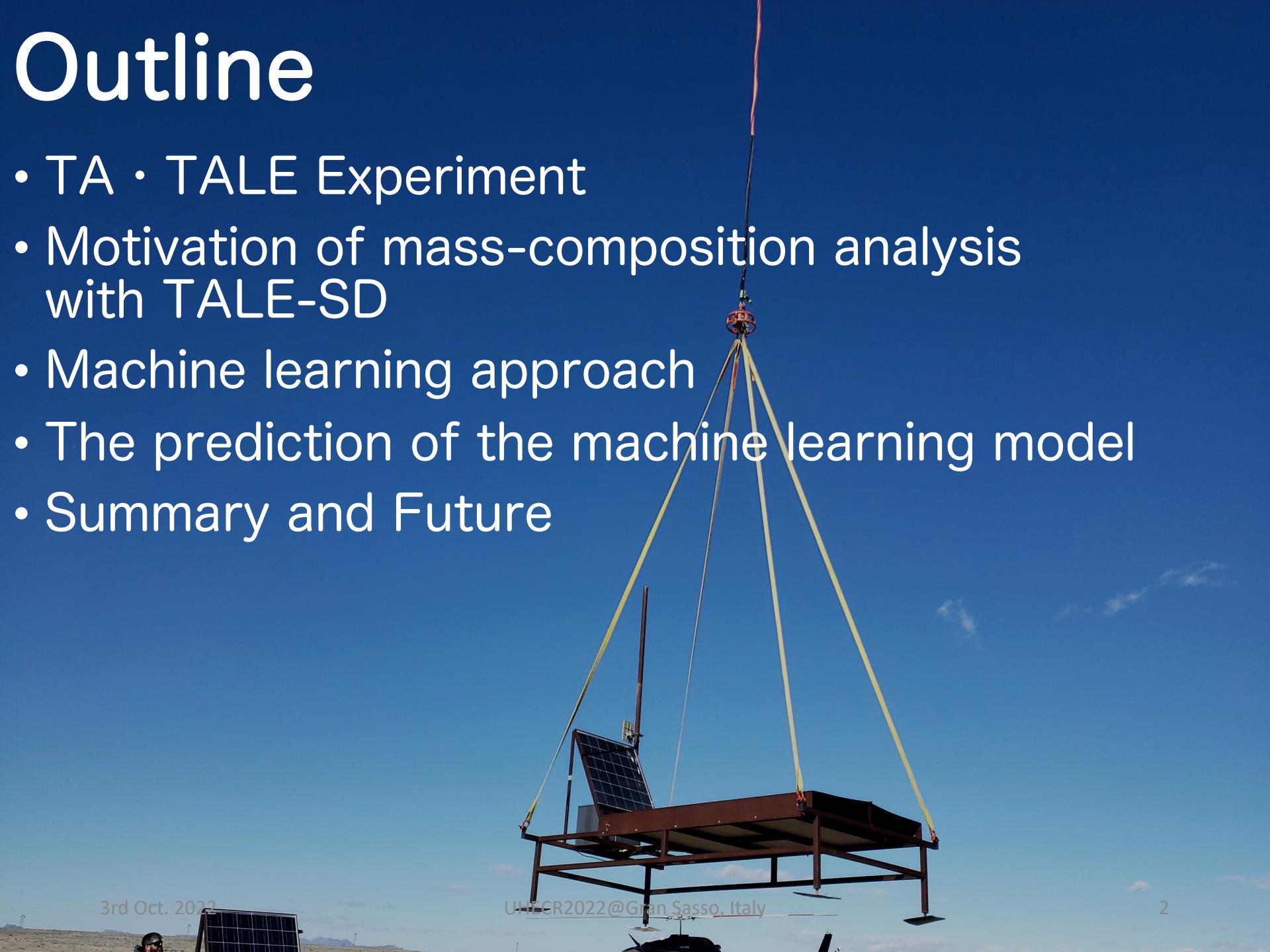
A photograph of a dark, open landscape under a night sky filled with star trails. In the foreground, there's some low-lying vegetation. In the middle ground, there's a chain-link fence. Behind the fence, there's a large, rectangular, modern-looking building with a dark facade and some lights on. The overall atmosphere is mysterious and scientific.

Ryuhei Arimura
Osaka Metropolitan University
for the Telescope Array Collaboration

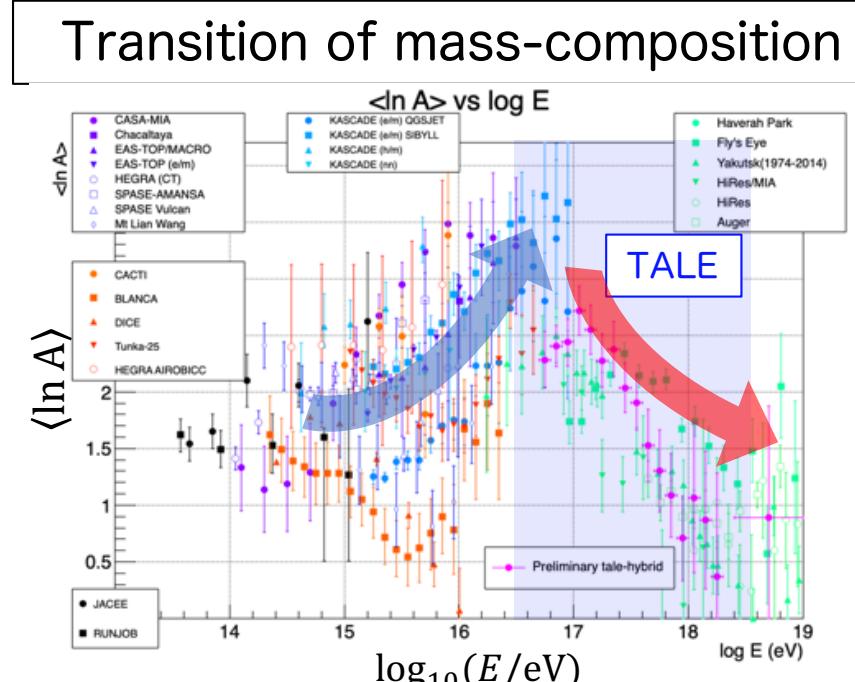
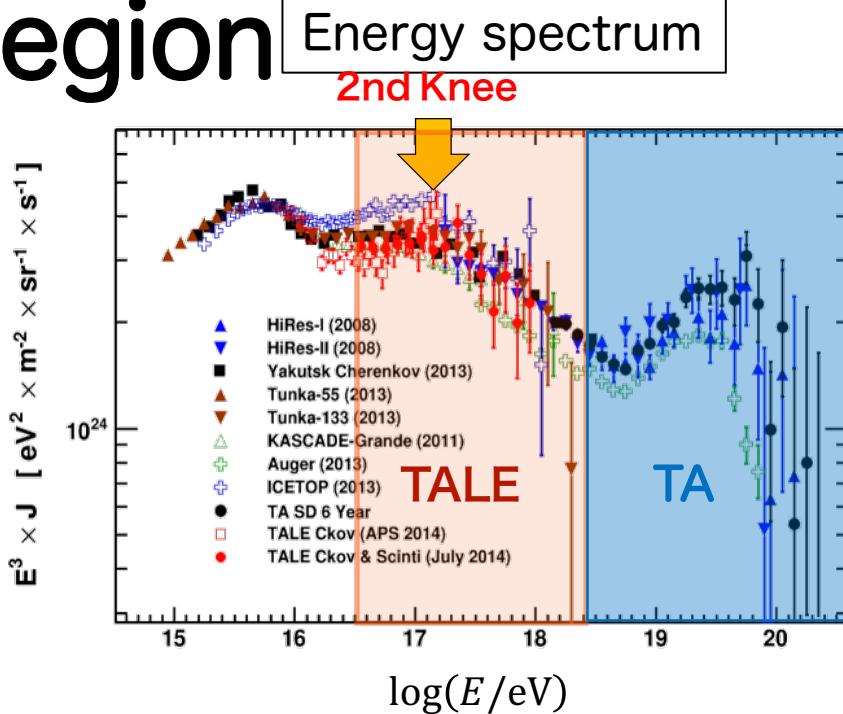
October 3rd, 2022
UHECR 2022

Outline

- TA • TALE Experiment
- Motivation of mass-composition analysis with TALE-SD
- Machine learning approach
- The prediction of the machine learning model
- Summary and Future



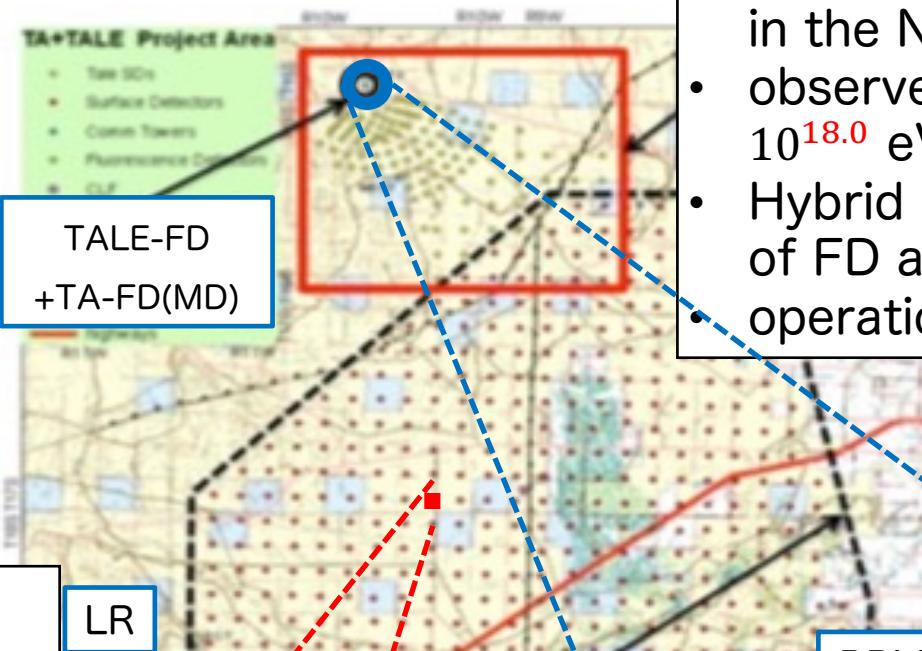
Cosmic ray around the 2nd knee region



Origin of cosmic rays around the 2nd knee region ($\sim 10^{17}$ eV)
→ galactic ↔ extragalactic transition?

Precise measurement of the energy spectrum and mass-composition of cosmic rays in the 2nd knee region by fluorescence detector and surface detector

TA→Telescope Array Experiment



- largest detected area in the Northern Hemisphere
- observe cosmic rays above $10^{18.0}$ eV in Utah (U.S.A.)
- Hybrid measurement of FD and SD operation since May 2008

SD :
Surface Detector
• 1200 m spacing

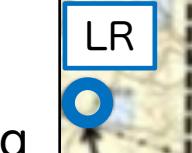
SD

①communication antenna

②solar panel

③Scintillator Box

3rd Oct. 2022



TA-SD array
(507 SDs, 700 km²)

↑ TA general map

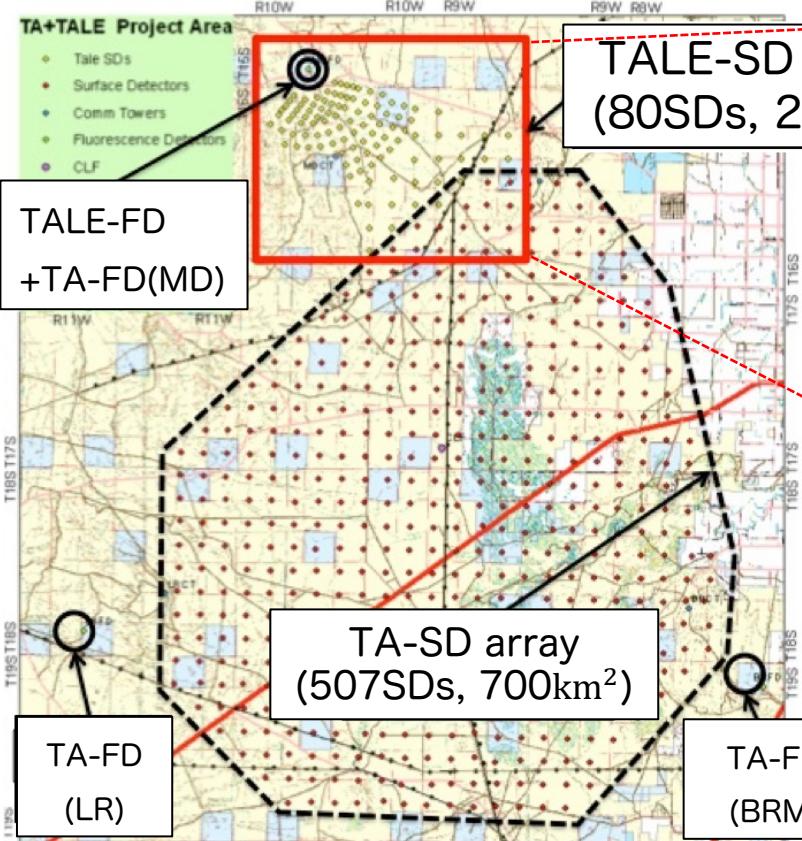
BRM

FD : Fluorescence Detector
• 38 units in 3 stations

TA-FD

TALE-FD

TALE→TA Low energy Extension



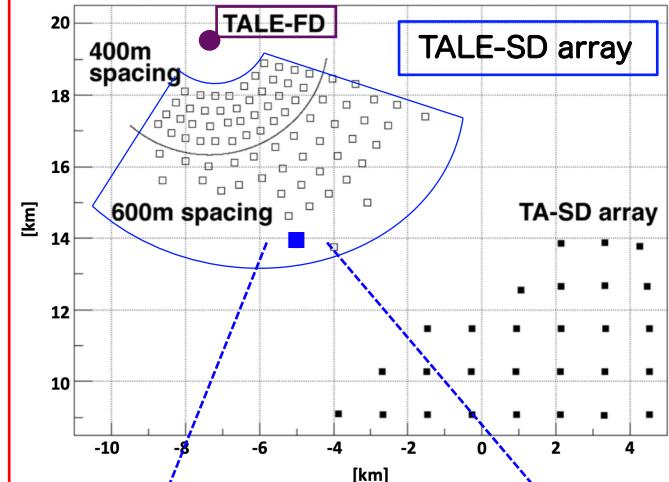
TALE-SD array
(80SDs, 21km²)

TALE-FD
+TA-FD(MD)

TA-SD array
(507SDs, 700km²)

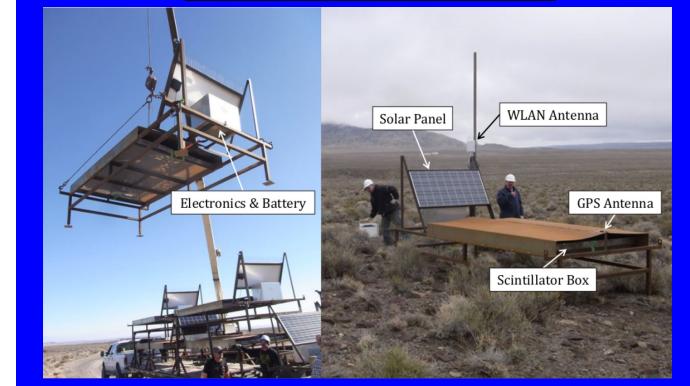
TA-FD
(LR)

TA-FD
(BRM)



80SDs →
40SDs (400m spacing)
+ 40SDs (600m spacing)

Surface Detector

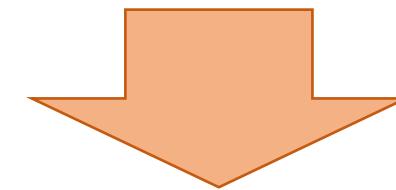
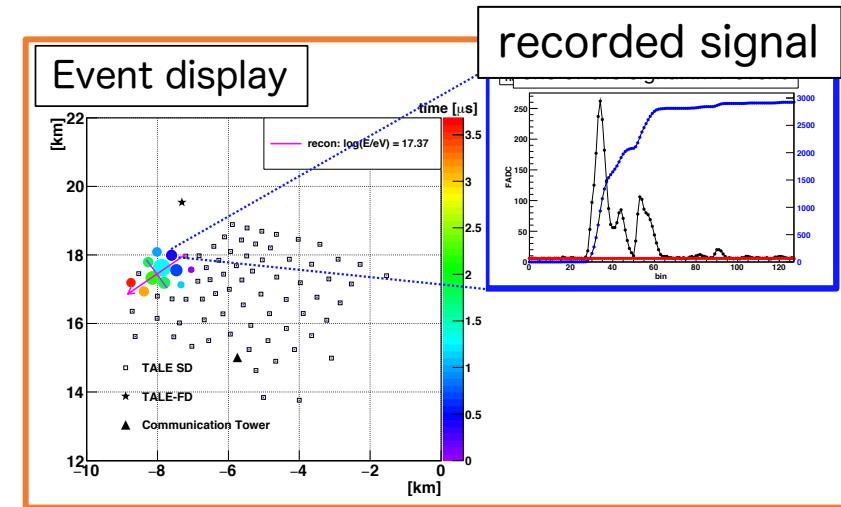


TALE-SD array

- are located in the northwest of TA-SD
- observe cosmic rays above $10^{16.5}$ eV
- have high statistics and uniform sensitivity

Motivation of Mass composition analysis with TALE-SD

- Advantage of SD array
 - over 90% duty cycle,
high statistics
(FD → 10% duty cycle)
 - Application to
cosmic ray anisotropy
per composition
- Approach
 - machine learning



Composition-sensitive parameters
 $(\alpha_1, \alpha_2, \alpha_3, \dots)$

Machine learning approach

Machine learning model Outline

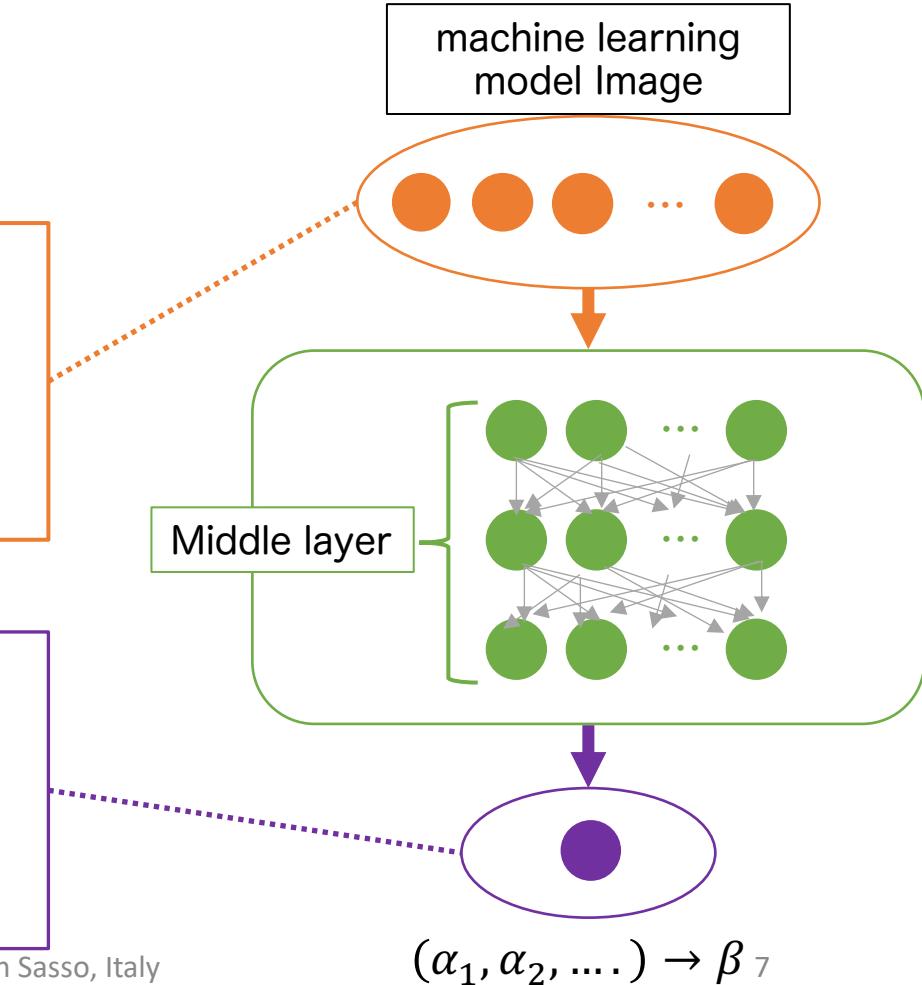
- **proton/iron** cosmic rays by MC simulation discrimination model
(binary classification)

Input : α

- **22** composition-sensitive parameters
 - parameters of geometry fit function
 - parameter of lateral distribution
 - the thickness of an air shower

Output : β

- Real number from **0** to **1**
 - labeled **proton** to **0**, **iron** to **1**
 - $\beta < 0.5 \rightarrow$ consider as **proton**
 - $\beta \geq 0.5 \rightarrow$ consider as **iron**



22 Composition-sensitive parameters with TALE-SD

22 mass-composition sensitive parameters from the characteristics of air showers →

- parameters of geometry fit function

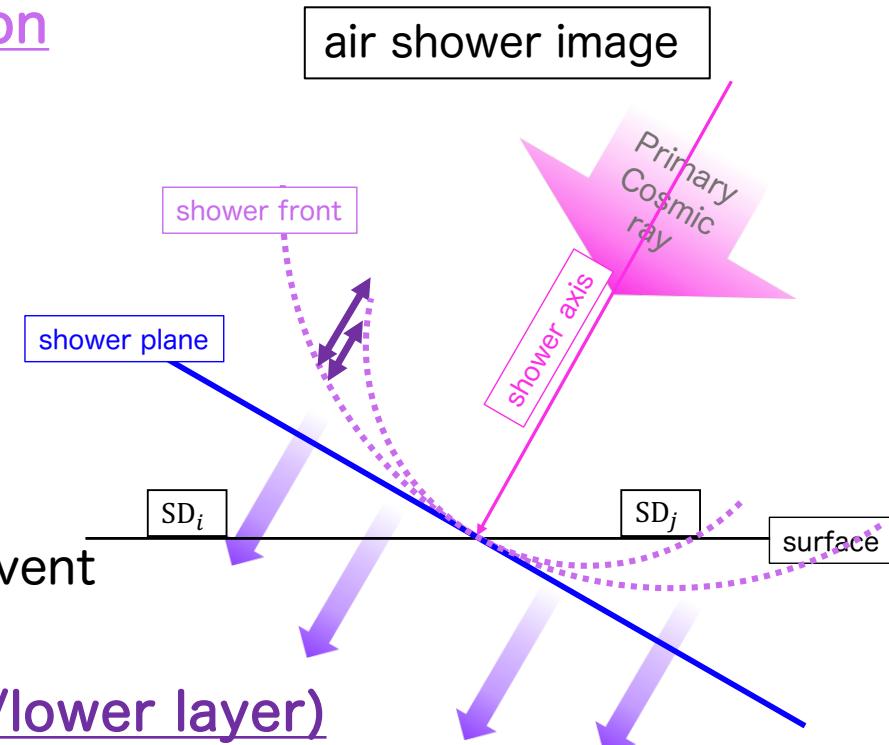
- Linsley curvature parameter
- time delay average

- parameters of lateral distribution

- lateral trigger probability
- energy estimator
- LDF function's parameter
- number of triggered stations in 1 event

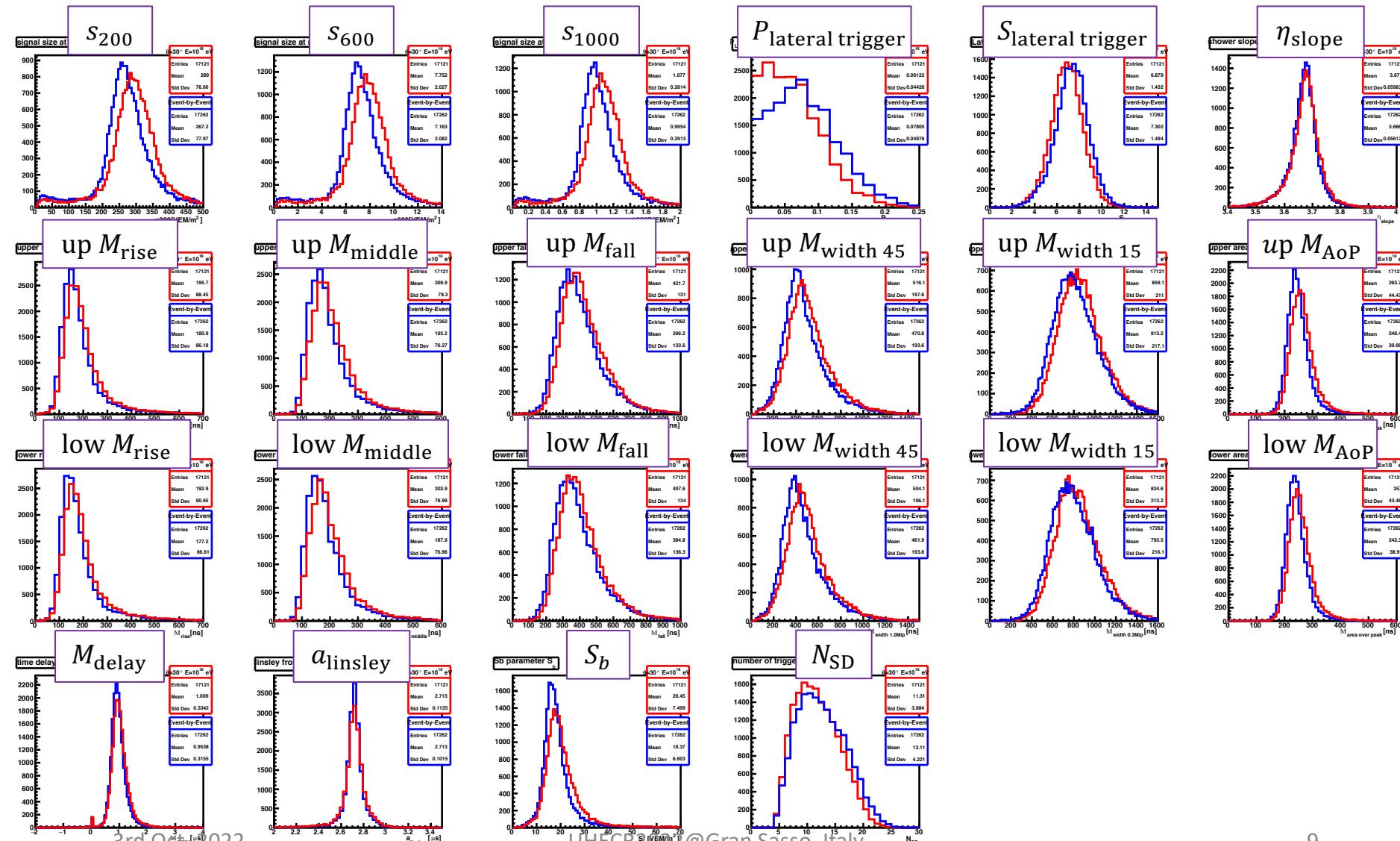
- thickness of an air shower (upper/lower layer)

- rise · middle · fall time stations' average
- time width stations' average
- area over peak stations' average



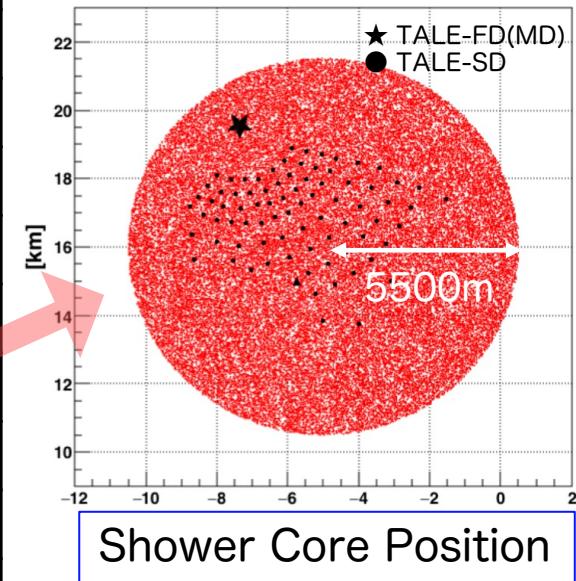
Composition-sensitive parameters with TALE-SD

- proton
- iron



MC simulation Data set and Cut

Data Set	proton	iron
MC Simulation	CORSIKA	
Interaction model	QGSJET II-04	
Energy	10^{18} eV (fixed)	
Zenith Angle	30° (fixed)	
Azimuth Angle	$0^\circ - 360^\circ$	
Core position	Uniform in the red circle	
$N_{\text{TALE-SDs}}$	78 (trigger : any4)	
N_{generate}	100,000	100,000
$N_{\text{reconstructed+after Event Cut}}$	17,121	17,262
N_{used}	<u>17,121</u>	<u>17,121</u>



Event Cut

- $N_{\text{SD}} \geq 5$
- $\chi_G^2/\text{d. o. f.} \leq 4, \chi_L^2/\text{d. o. f.} \leq 2$
- $(\sigma_\theta^2 + \sin^2 \theta \sigma_\phi^2)^{0.5} \leq 2.5 \text{ deg}$
- $\sigma_{S_{600}}/S_{600} \leq 0.25$
- $N_{\text{useSD}} \geq 1$ (in 1 event)

3rd Oct. 2022

N_{useSD} condition (SD selection)

- $t_{\text{record WF}} \leq 2.56 \mu\text{s}$ (128 bin)
- $N_{\text{bin}} (>15 \text{ FADC count}) \geq 2$
- $N_{\text{bin}} (>45 \text{ FADC count}) \geq 1$
- Not saturate
- $400 \text{ m} \leq r \leq 700 \text{ m}$

Machine learning model Details

Hidden layer

Input : $\alpha \rightarrow$
22 cosmic ray event parameters

Dimension : 64
Activation function : “tanh”

Dimension : 32
Activation function : “tanh”

Dimension : 16
Activation function : “tanh”

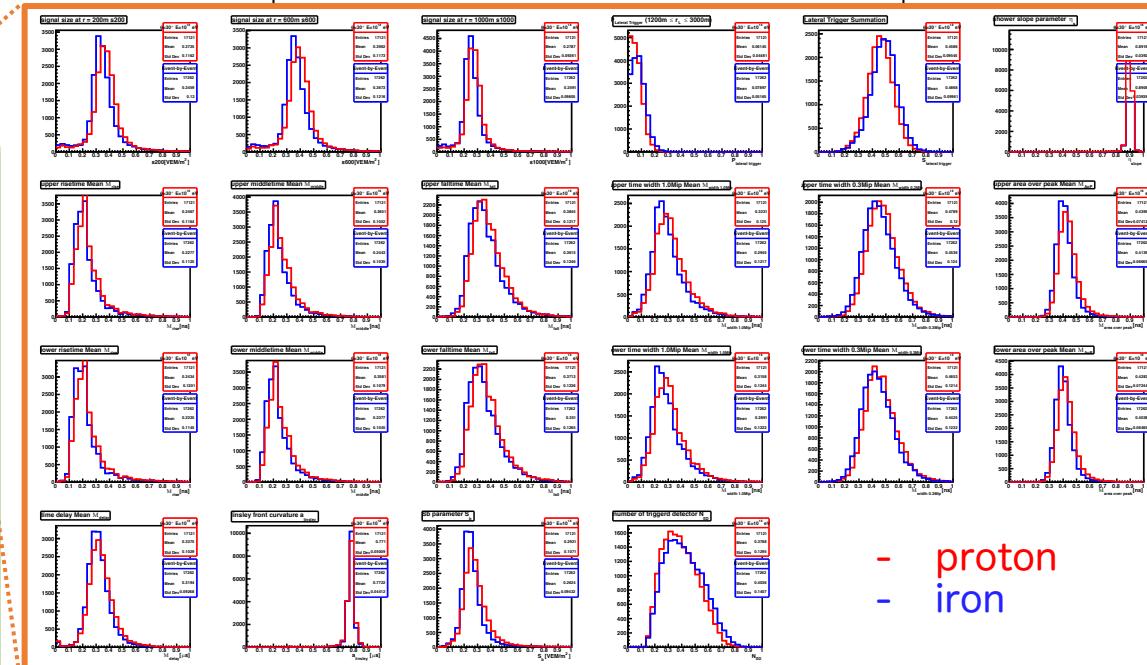
Dimension : 16
Activation function : “tanh”

Dimension : 8
Activation function : “tanh”

Output : $\beta \rightarrow 0 \leq \beta \leq 1$
Activation function : “sigmoid”

- labeled proton to 0, iron to 1
- $\beta < 0.5 \rightarrow$ consider as proton
- $\beta \geq 0.5 \rightarrow$ consider as iron

Input parameters (normalized) α



- proton
- iron

Model information

All data	34242 events
Training data	27392 events (80% of All data)
Validation data	3424 events (10% of All data)
Test data	3424 events (10% of All data)
Optimizer	Adam
Loss function	Cross-Entropy Error function

The prediction of the machine learning model

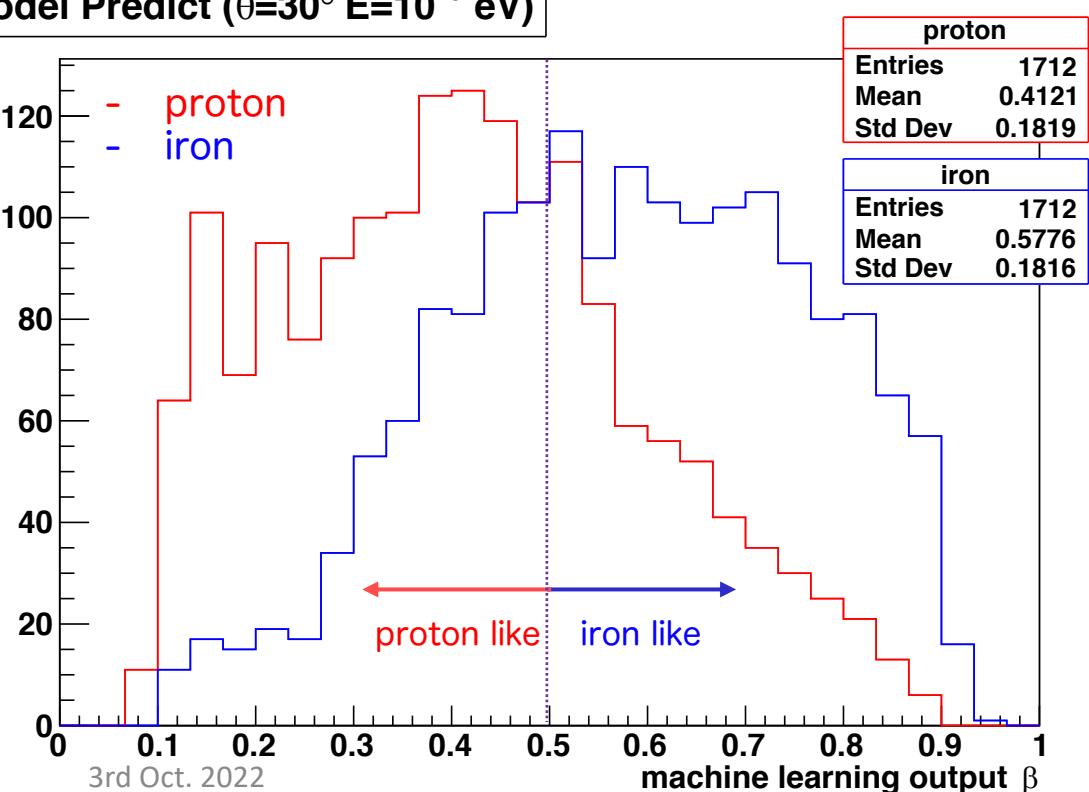
Input α :

- 22 composition-sensitive parameters
 - parameters of geometry fit function
 - parameter of lateral distribution
 - the thickness of an air shower



output $\beta : 0 \leq \beta \leq 1$
➤ proton → 0
➤ iron → 1

Model Predict ($\theta=30^\circ$ E= 10^{18} eV)



predict	answer	
Proton	Proton	Iron
	1180 (68.93%)	532 (31.07%)
Iron	593 (34.64%)	1119 (65.36%)

▣ :
classification results
for all test events
→ accuracy = 67.14%

Summary and Future

- Summary

- Search composition-sensitive parameters
- Input **22** Event-by-Event parameters to machine learning
- Current Accuracy : **67.14%** ($\theta = 30^\circ, E = 10^{18}$ eV)

- Future

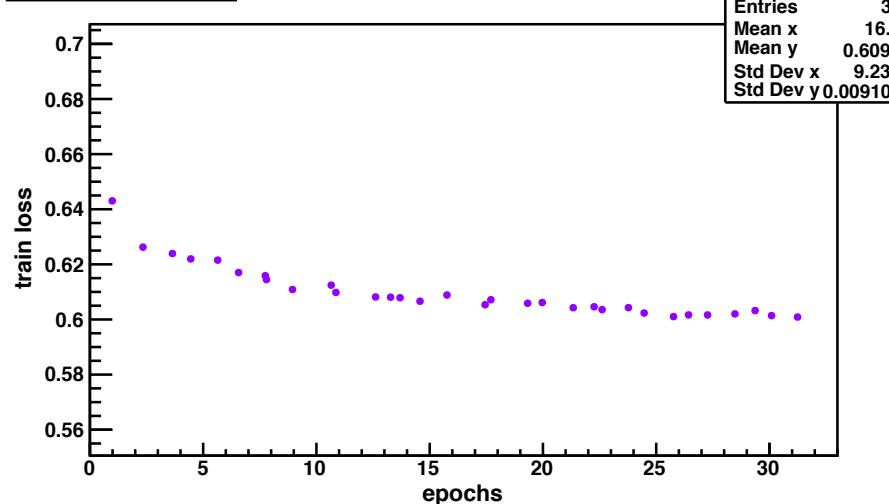
- To improve the accuracy of machine learning model,
 - Addiction of Detectors' parameters as input
 - Search new composition-sensitive parameters
 - Using Graph Neural Network

Backup

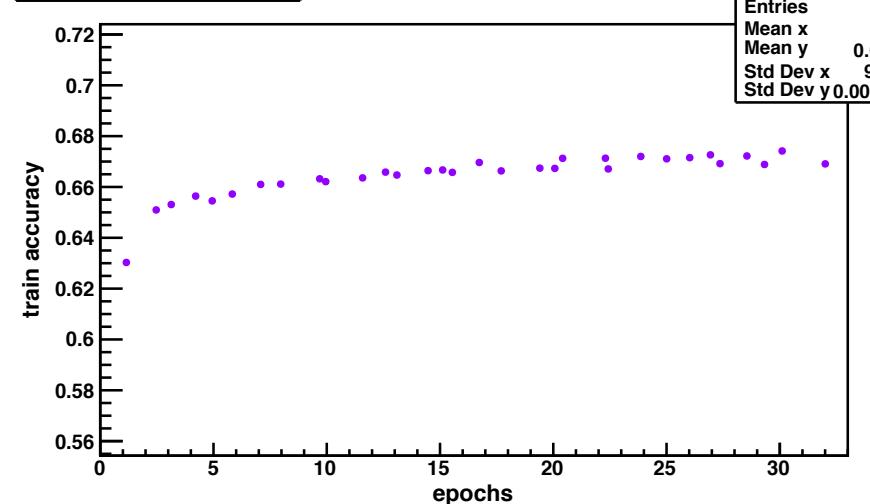
- Accuracy and loss transition
- Zenith angle and energy distribution
- Parameter definition
 - Parameter of geometry fit function
 - Parameter of lateral distribution
 - Thickness of an air shower
(upper/lower layer)

Accuracy and loss transition

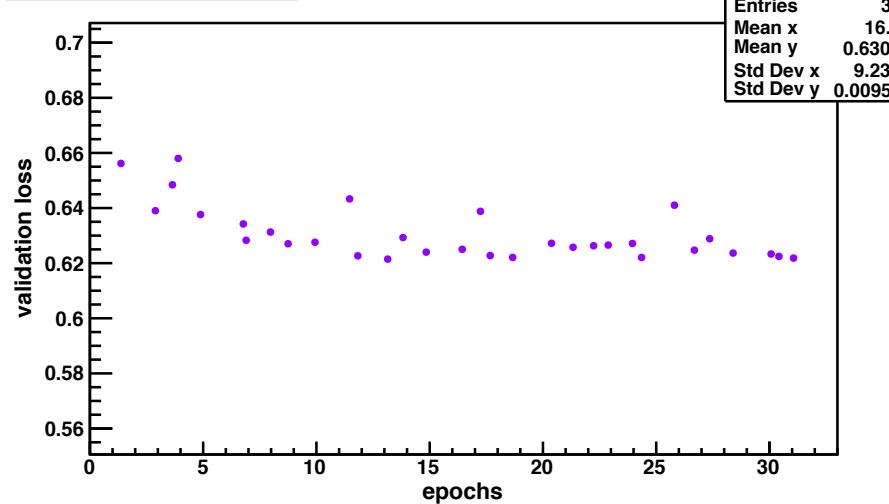
Loss for train data



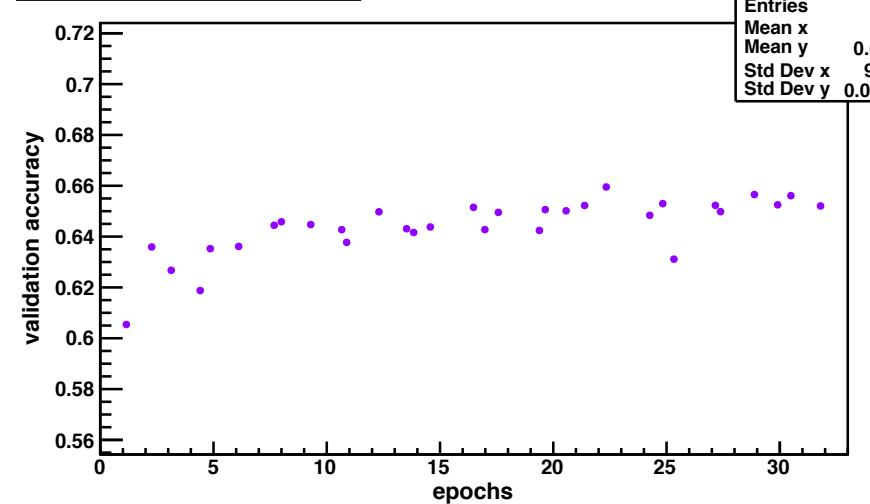
Accuracy for train data



Loss for validation data



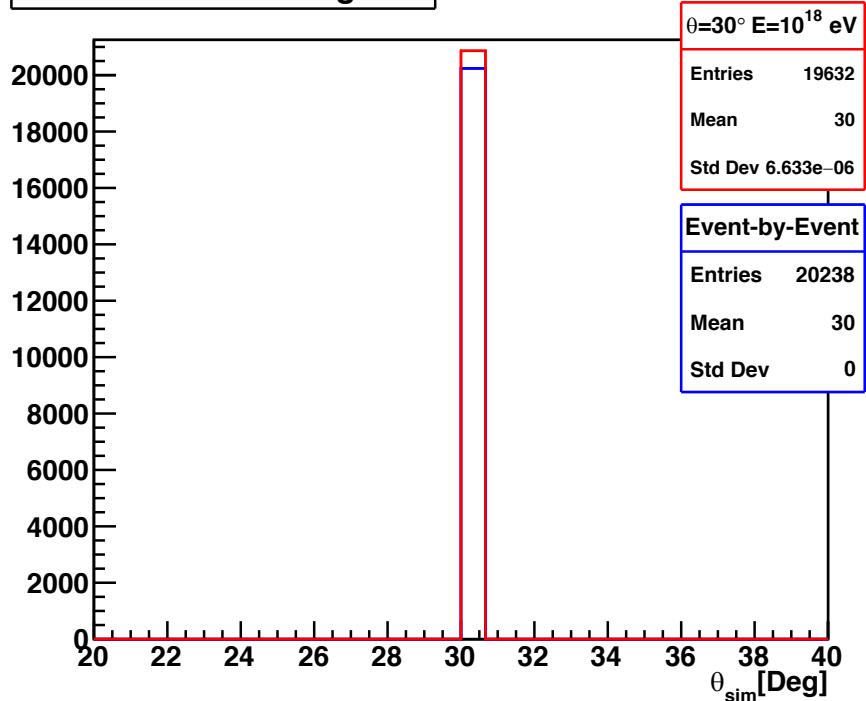
Accuracy for validation data



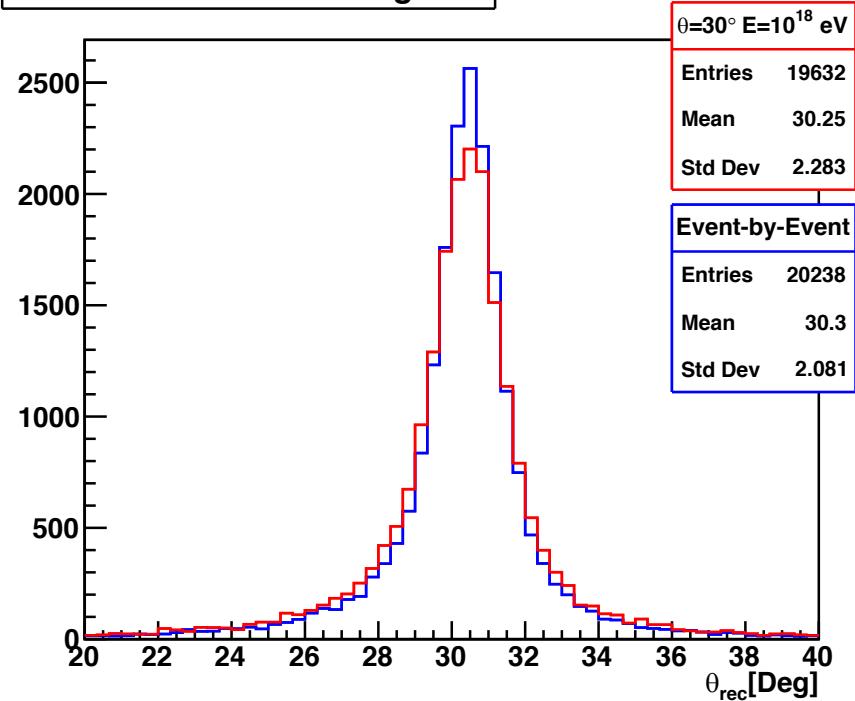
Zenith angle

- proton
- iron

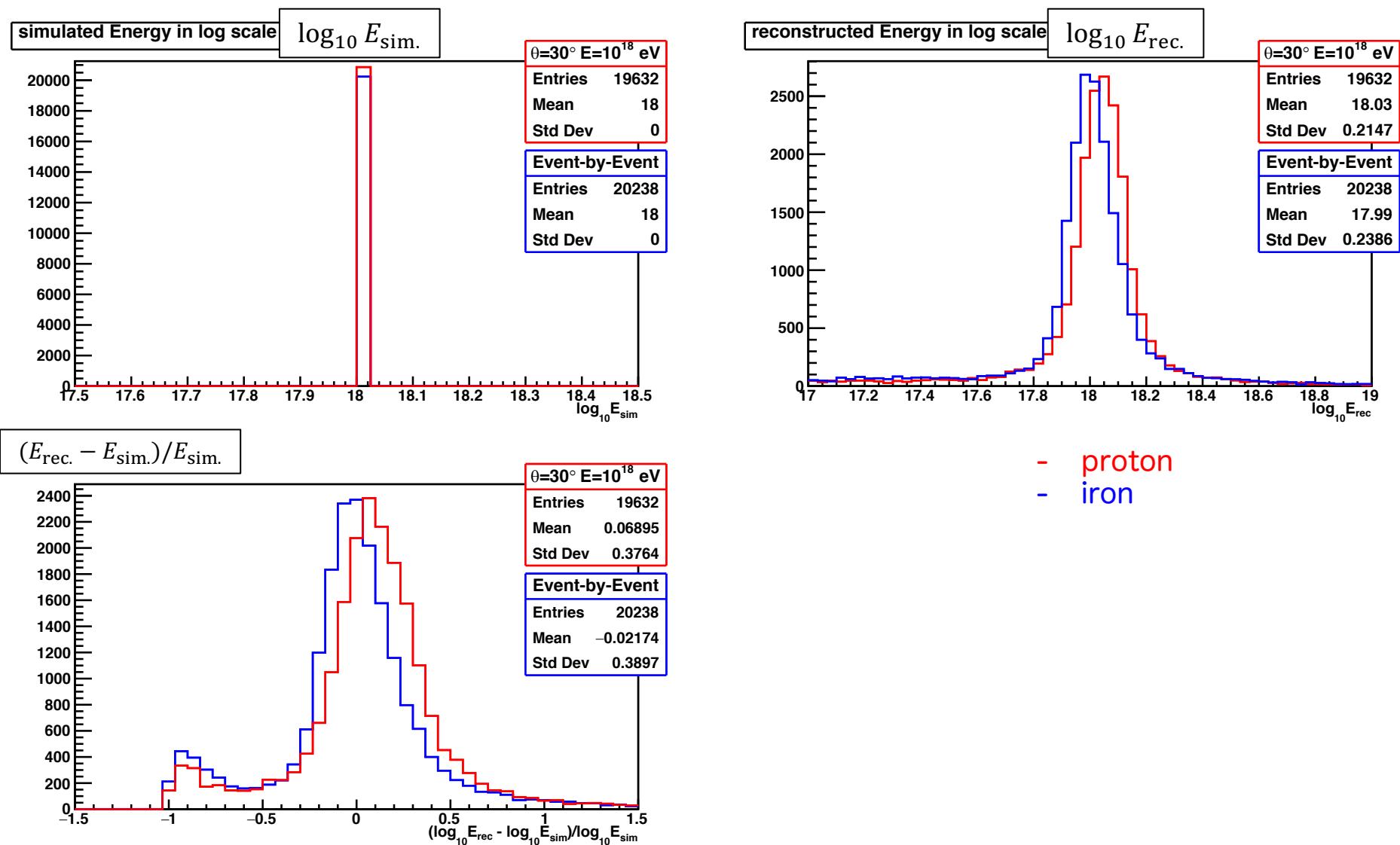
simulated Zenith Angle θ



reconstructed Zenith Angle θ



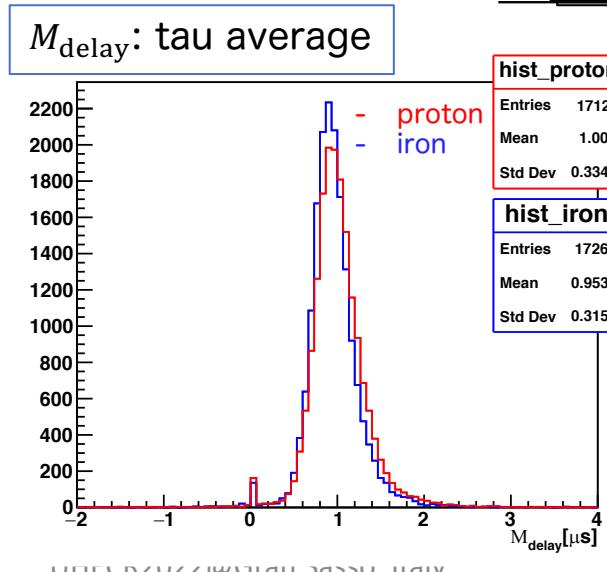
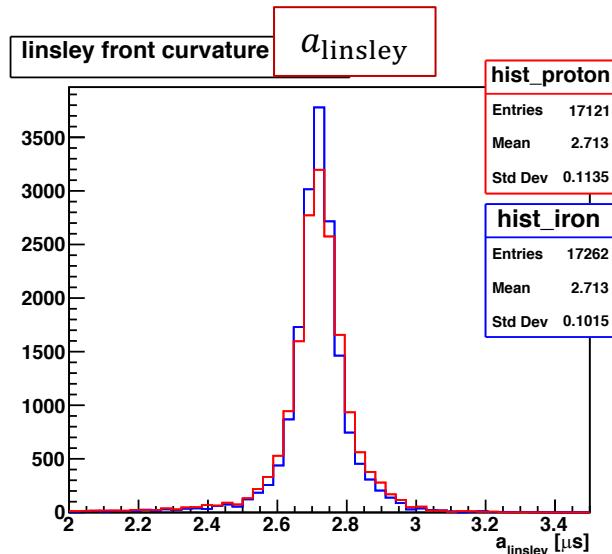
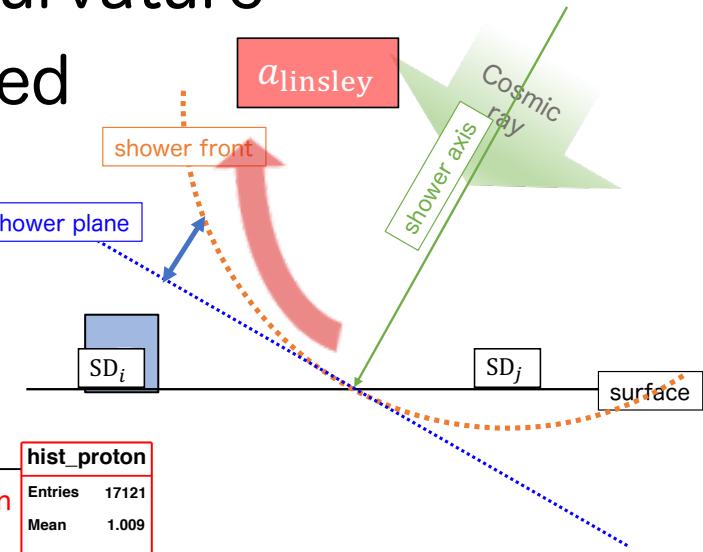
Primary Energy



Parameters of geometry fit function

$$\tau = a_{\text{linsley}} \left(1 - \frac{l}{12000 \text{ m}}\right)^{1.05} \left(1 + \frac{r}{30 \text{ m}}\right)^{1.35} \rho^{-0.3}$$

- the parameter of air shower front curvature
- As a_{linsley} is free parameter, τ is fitted using χ^2 test, r , and ρ



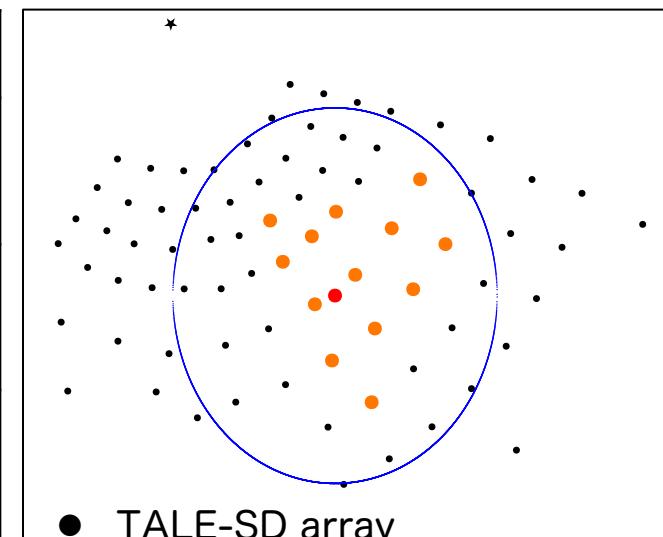
Lateral Trigger Probability

→ $P_{\text{Lateral Trigger}}$

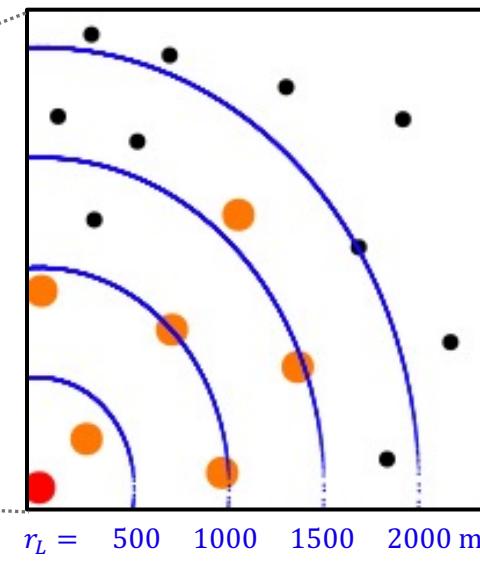
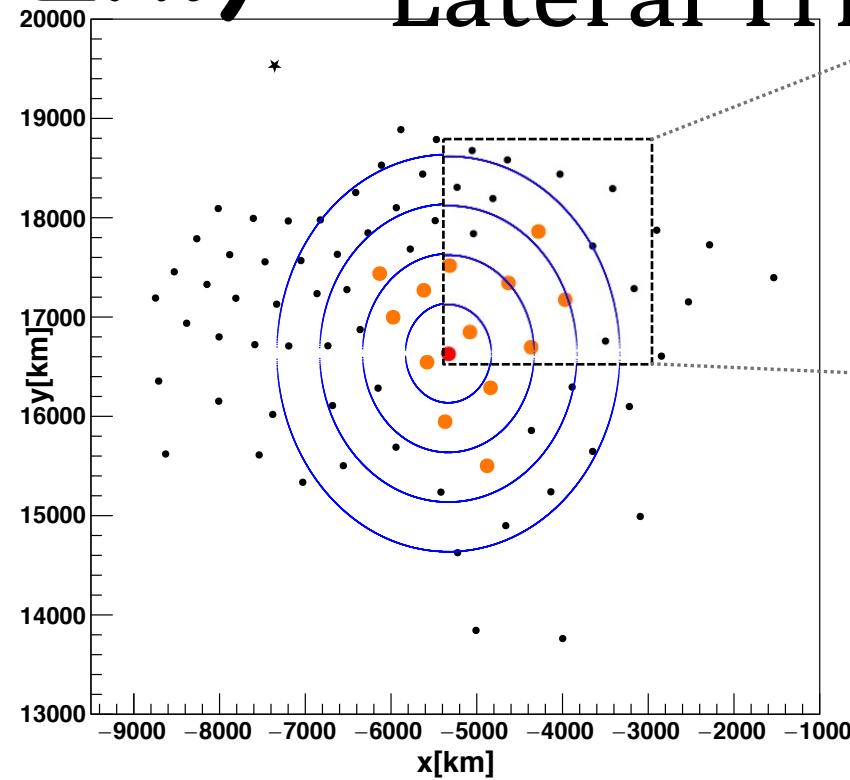
$$P_{\text{Lateral Trigger}}(r_L) = \frac{N_{\text{Trigger}}}{N_{\text{Trigger}} + N_{\text{non-Trigger}}}$$

Ex) $P(2000\text{m}) = 13/40 = 0.325$

parameter	
r_L	Lateral distance between shower core and detector position
N_{trigger}	number of triggered stations contained within radius = r_L
$N_{\text{non-trigger}}$	number of untriggered stations contained within radius = r_L



Ex.) $P_{\text{Lateral Trigger}}$



Left Fig. : Event display of MC(proton)

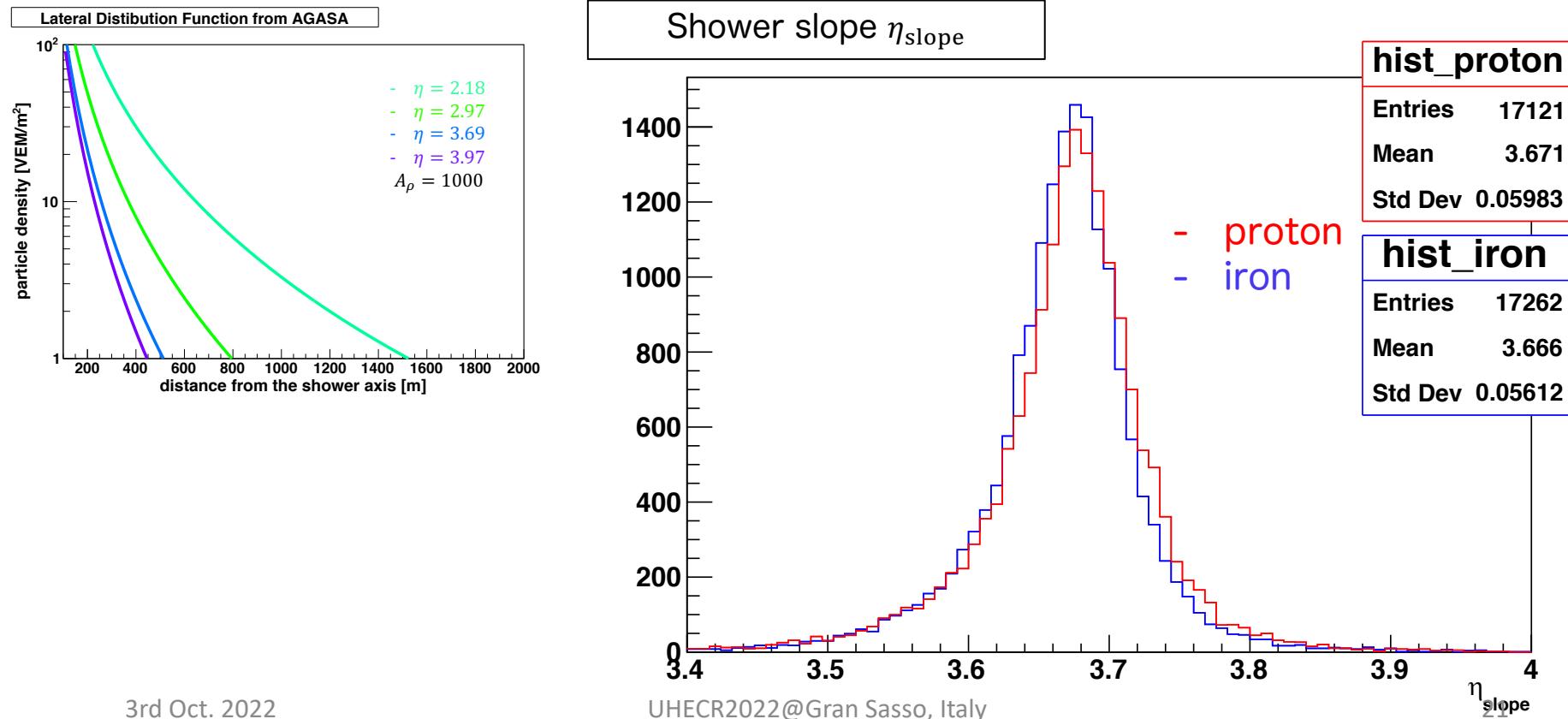
- TALE-SD array
- ★ TALE-FD(MD)
- Triggered SD
- Core Position (Recon)
- Circle of radius “ r_L [m]”

計算例	N_{trigger}	$N_{\text{non-trigger}}$	N_{all}	$P_{\text{Lateral Trigger}} = N_{\text{trigger}}/N_{\text{all}}$
$r_L \leq 500 \text{ m}$	2	0	2	1.000
$r_L \leq 1000 \text{ m}$	8	1	9	0.889
$r_L \leq 1500 \text{ m}$	12	12	24	0.500
$r_L \leq 2000 \text{ m}$	13	27	40	0.325

LDF function slope

$$\rho(r) = A \left(\frac{r}{91.6 \text{ m}} \right)^{-1.2} \left(1 + \frac{r}{91.6 \text{ m}} \right)^{-\eta_s + 1.2} \left(1 + \left[\frac{r}{1000 \text{ m}} \right]^2 \right)^{-0.6}$$

- the parameter of air shower longitudinal development
- As η_s is a free parameter, $\rho(r)$ is fitted using χ^2 test and r



S_b parameter

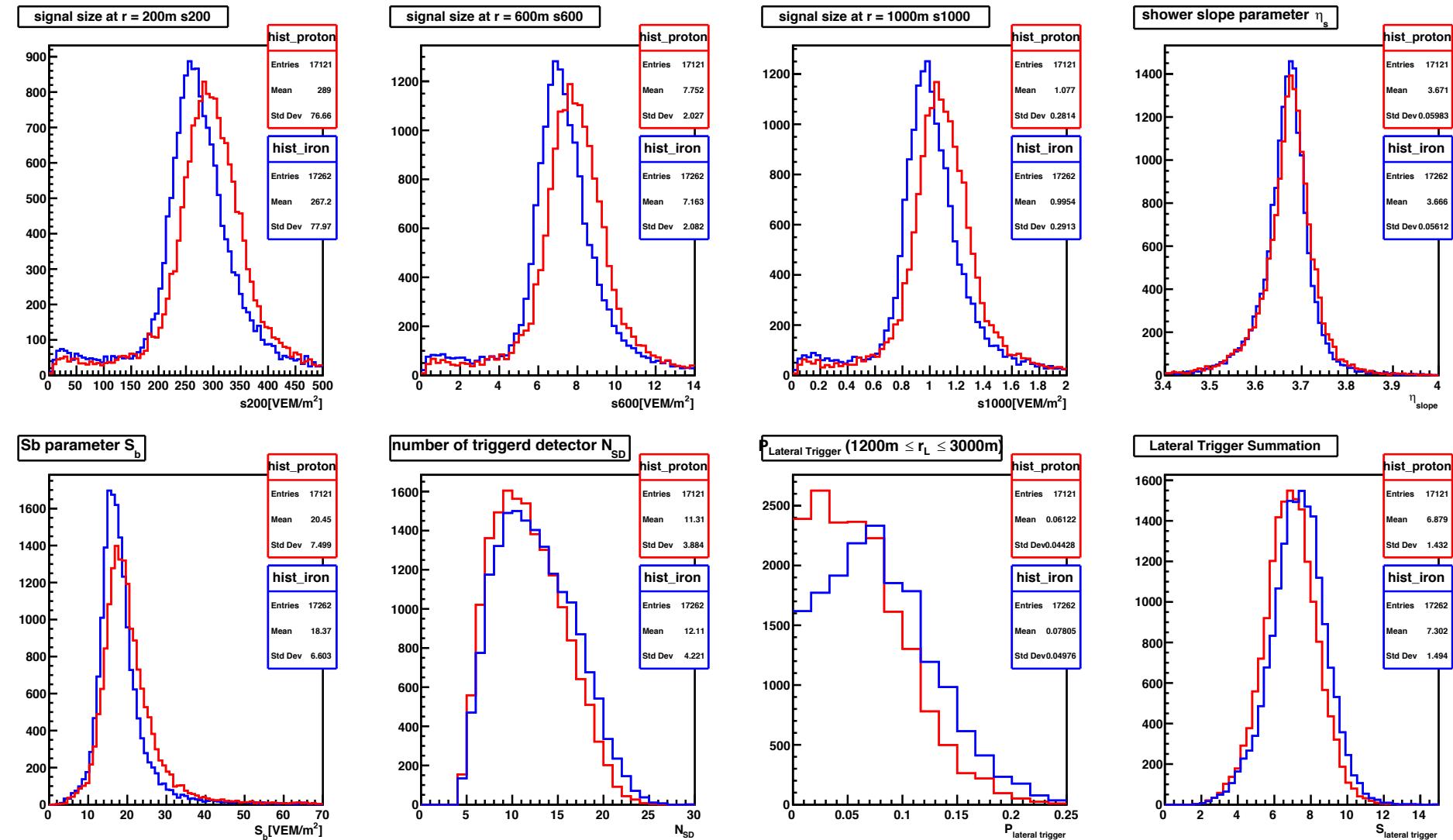
- Definition

$$S_b = \frac{1}{N_{SD}} \sum_i^{N_{SD}} \left[\rho_i \times \left(\frac{r_i}{r_0} \right)^b \right]$$

ρ_i	Particle density of i -th detector
r_i	Distance from the shower axis [m]
r_0	Reference distance [m] $\rightarrow 400$ m
b	Separation parameter $\rightarrow b = 2$

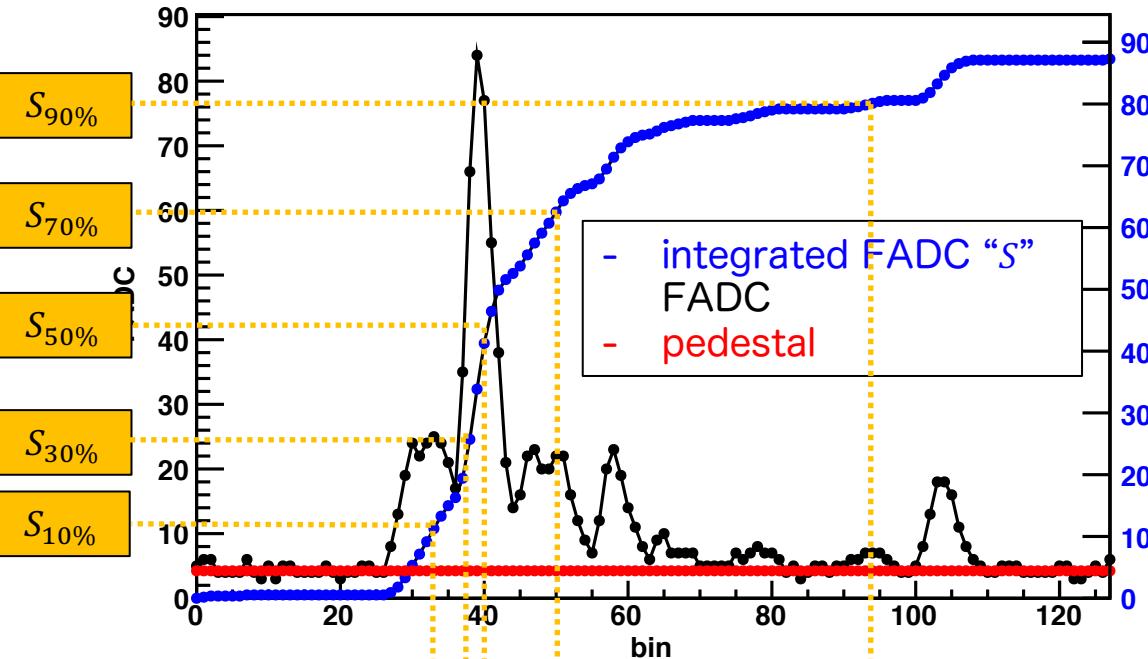
Parameters of lateral distribution

- proton
- iron



The thickness of an air shower

$$\rightarrow t_{\text{rise}} \cdot t_{\text{middle}} \cdot t_{\text{fall}} \cdot t_{\text{width}}^{45} \cdot t_{\text{width}}^{15}$$



Represents the amount of increase in the S

- integrated FADC “ S ”
- FADC
- pedestal

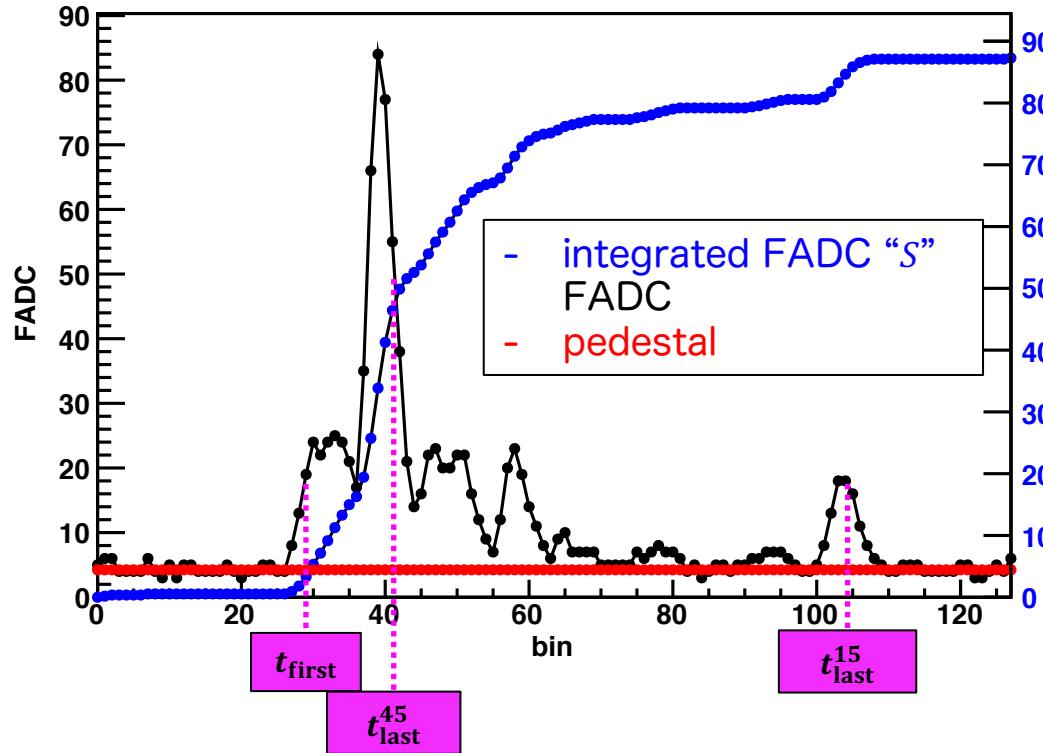
- $t_{\text{rise}} = t_{50\%} - t_{10\%}$
- $t_{\text{middle}} = t_{70\%} - t_{30\%}$
- $t_{\text{fall}} = t_{90\%} - t_{50\%}$

➤ $t_{xx\%}$: time bin when the S is $xx\%$ of the whole

Figure : Signal display
MC iron SD6010

The thickness of an air shower

$$\rightarrow t_{\text{rise}} \cdot t_{\text{middle}} \cdot t_{\text{fall}} \cdot t_{\text{width}}^{45} \cdot t_{\text{width}}^{15}$$



Represents the signal time difference

$$t_{\text{width}}^{45} = t_{\text{last}}^{45} - t_{\text{first}}$$

$$t_{\text{width}}^{15} = t_{\text{last}}^{15} - t_{\text{first}}$$

➤ t_{first} :
time bin when the 0.3mip
(=15fadc) signal is first input

➤ $t_{\text{last}}^{\text{xx}}$:
time bin when the xx fadc signal is last input

Figure : Signal display
MC iron SD6010

Detectors' average

- Definition of Mean for thickness parameters
→ “ $M_{\text{thickness}}$ ”

$$M_{\text{thickness}} = \frac{1}{N_{\text{good SD}}} \sum_i^{N_{\text{good SD}}} t_{\text{thickness}}$$

$t_{\text{thickness}} \rightarrow$	$t_{\text{rise}}, t_{\text{middle}}, t_{\text{fall}}, t_{\text{width 1.0Mip}}, t_{\text{width 0.3Mip}}$	
$N_{\text{good SD}}$	# of SDs that meet the following Signal and r Cut	
Signal Cut	r Cut	
➤ WF recording time is within 128 bin	t_{rise}	$400 \text{ m} \leq r \leq 1000 \text{ m}$
➤ 2 or more bins higher than 15FADC	t_{middle}	$0 \text{ m} \leq r \leq 800 \text{ m}$
➤ 1 or more bins higher than 45FADC	t_{fall}	$0 \text{ m} \leq r \leq 800 \text{ m}$
➤ Not saturate	$t_{\text{width}}^{1.0\text{Mip}}$	$100 \text{ m} \leq r \leq 700 \text{ m}$
	$t_{\text{width}}^{0.3\text{Mip}}$	$100 \text{ m} \leq r \leq 800 \text{ m}$

Thickness of an air shower (upper/lower layer)

- proton
- iron

