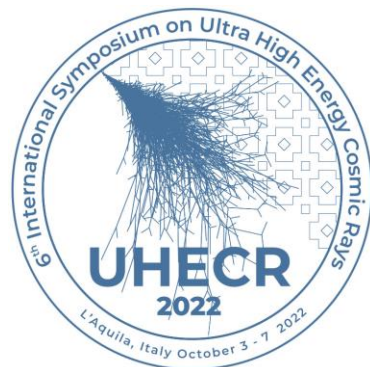


# Indication of a Local Source of Ultra-High-Energy Cosmic Rays in the Northern Hemisphere



HELMHOLTZ WEIZMANN  
RESEARCH SCHOOL  
MULTIMESSENGER ASTRONOMY

Pavlo Plotko  
UHECR 2022, L'Aquila,  
06.10.2022



pavlo.plotko@desy.de

**HELMHOLTZ**



# WHICH ONE



VS

# IS BETTER ?



# WHICH ONE IS BETTER ?





# WHICH ONE IS BETTER?

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# WHICH ONE IS BETTER?

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*“We remain with the dilemma: protons versus heavy nuclei. A clear cut decision cannot be reached yet.”*

**G Cocconi: Fifth International Cosmic Ray Conference, Guanajuato, Mexico, 1955**

Prof. Alan Watson UHECR2022



Pierre Auger Observatory

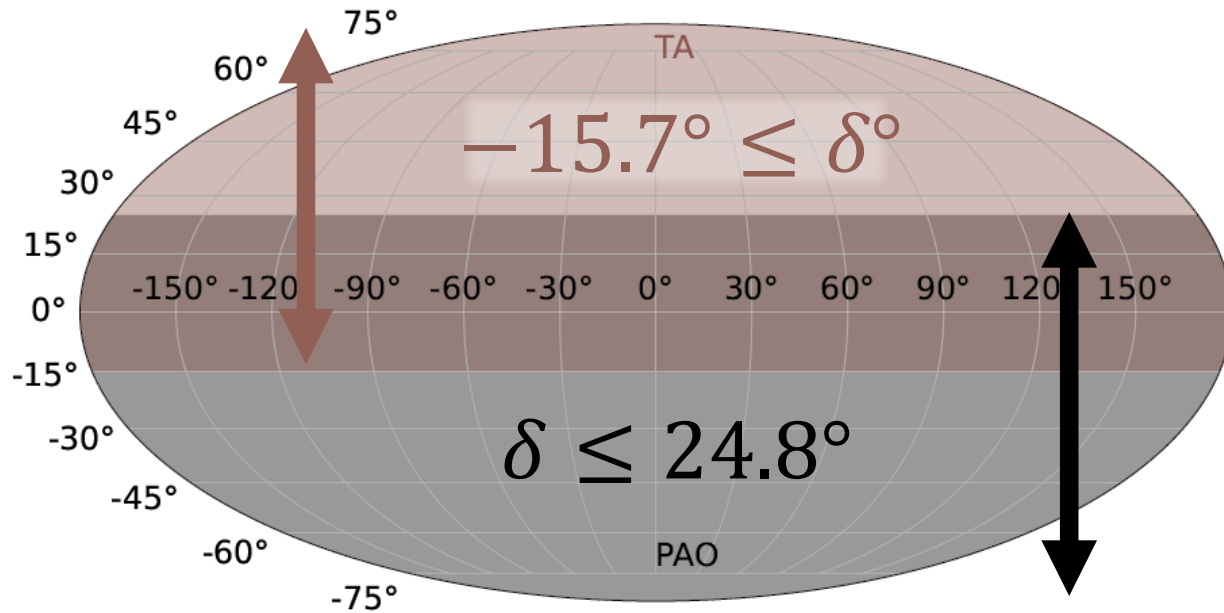
Telescope Array

vs



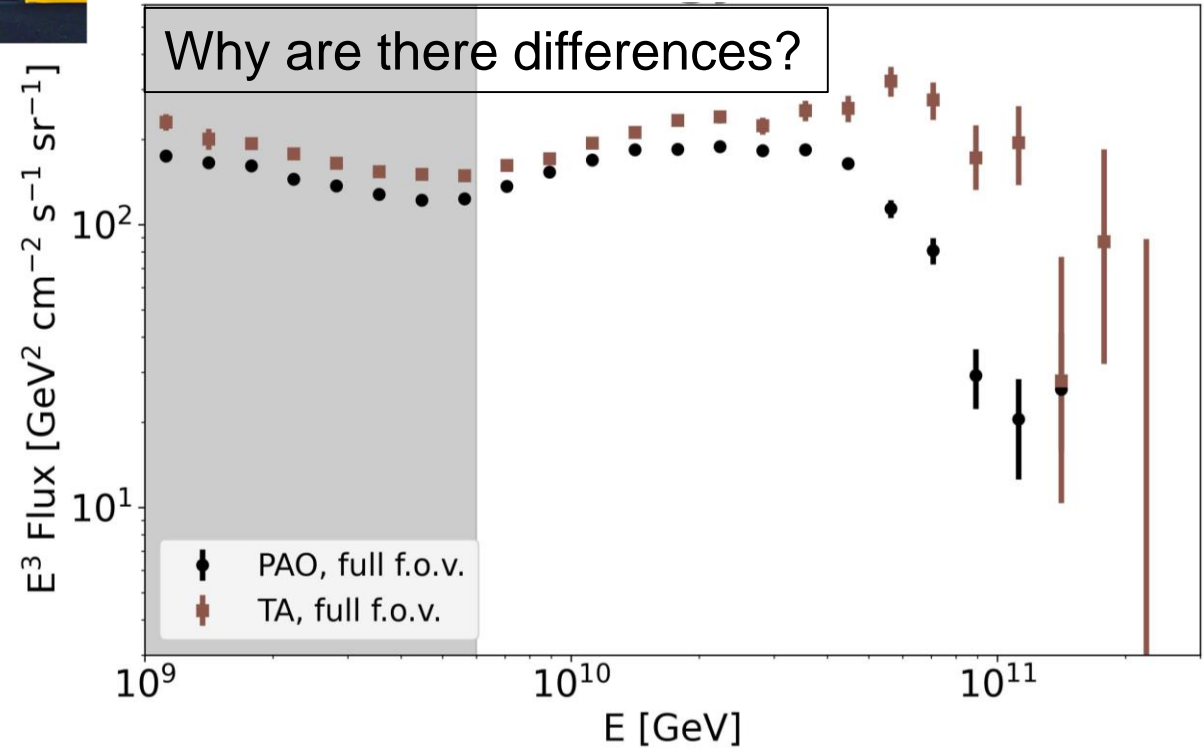
# UHECR detectors: TA vs PAO

Telescope Array (TA), USA



Energy systematic uncertainties:

↗ 14% ↘ 21%



Pierre Auger Observatory (PAO), Argentina



Check presentations:

*Dr Valerio Verzi, Dr Jihyun Kim, Douglas Bergman  
 Quentin Luce, Prof. Shoichi Ogio, Federico Urban*



# UHECR spectrum

## Motivation

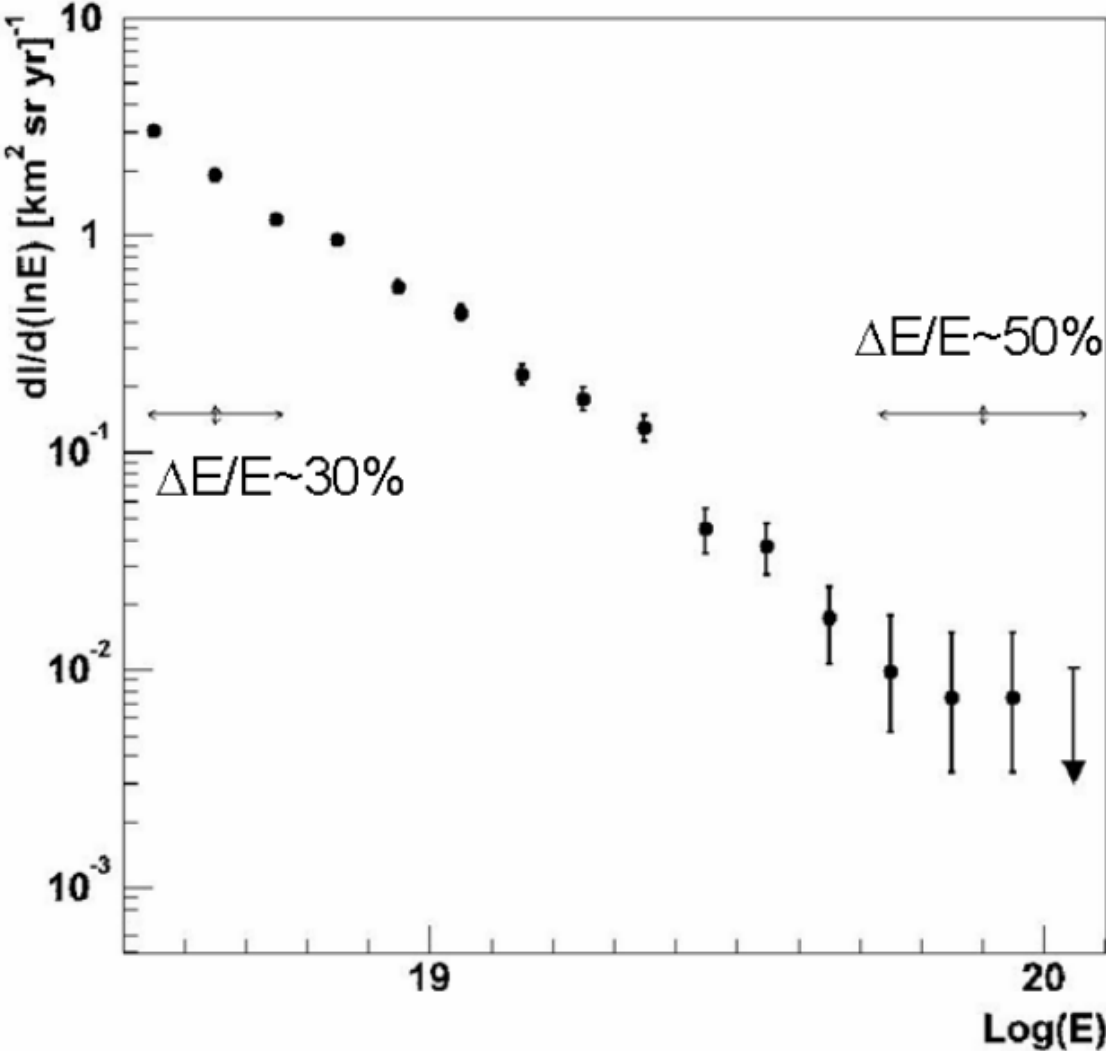
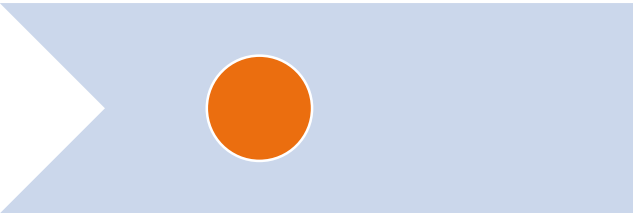




# UHECR time line

PAO

first results  
2004

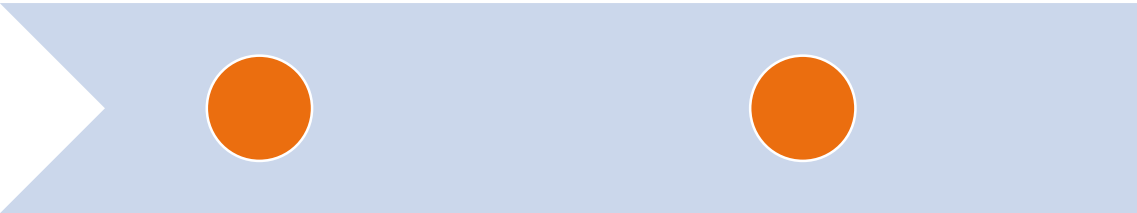


Mantsch et al. (2006)

# UHECR time line

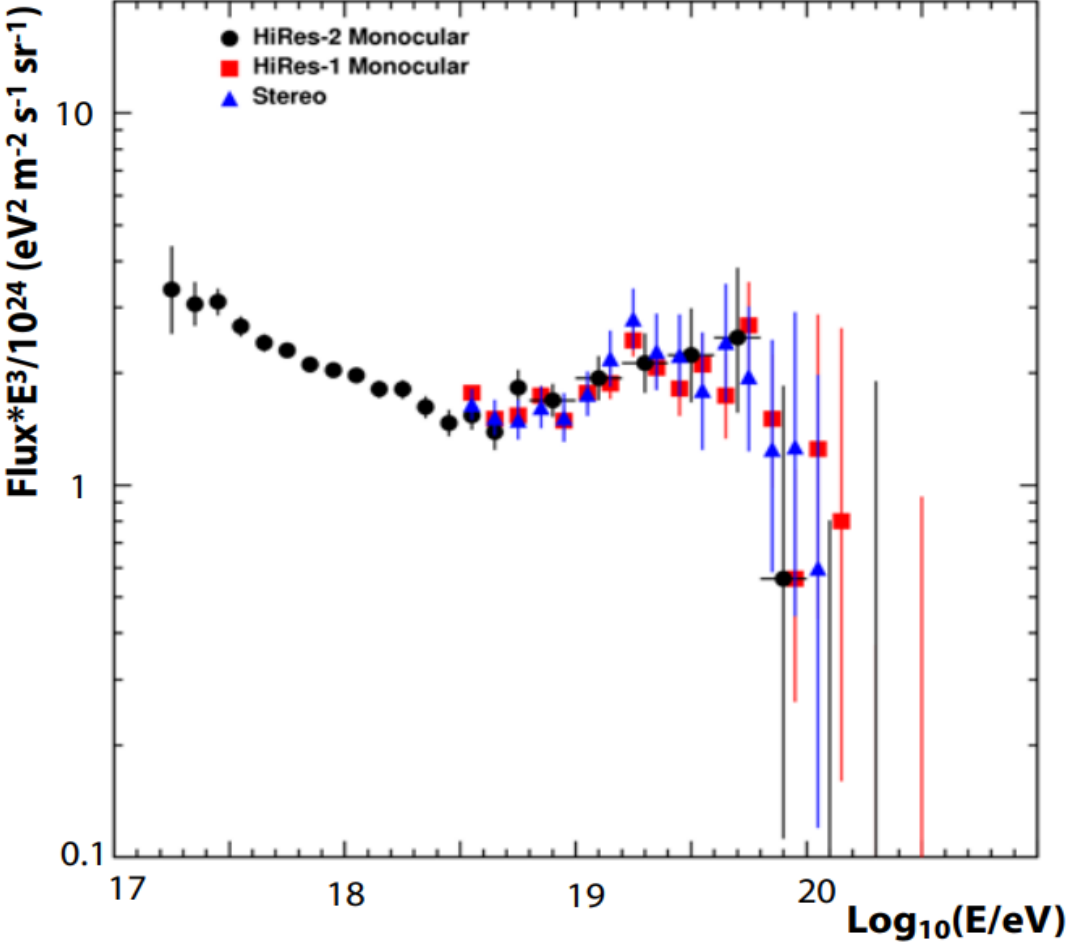
PAO

first results  
2004



TA

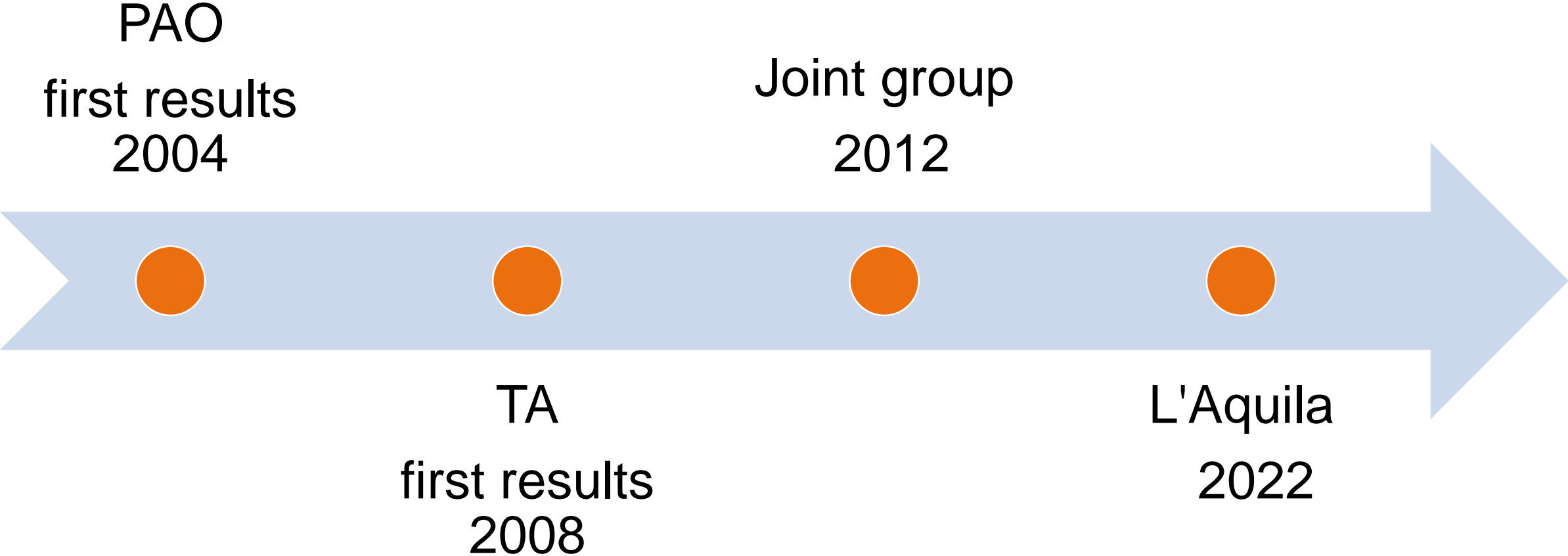
first results  
2008



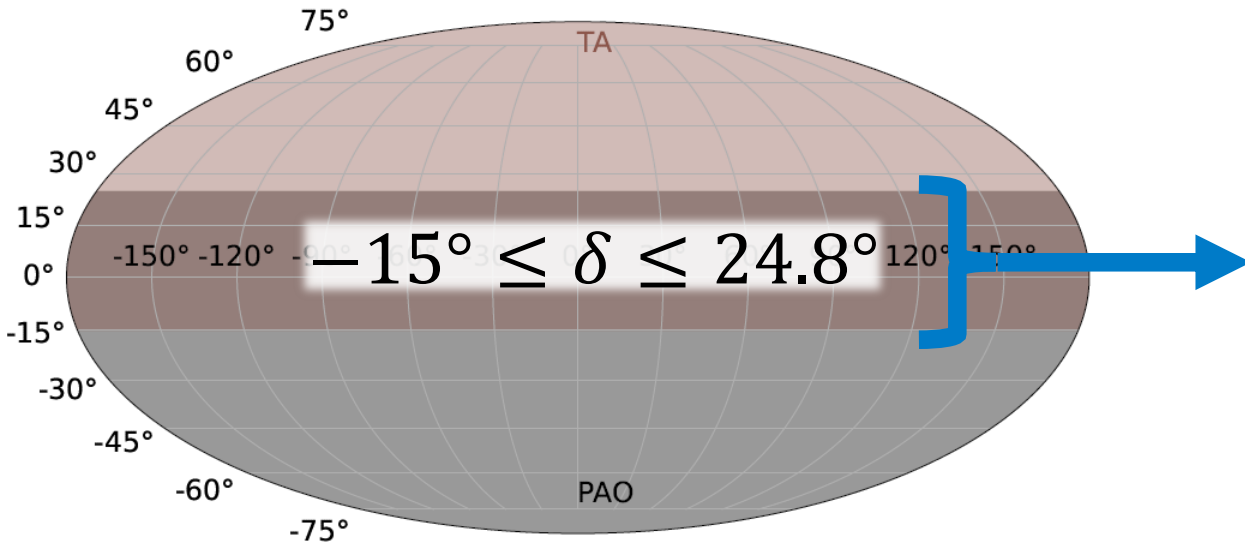
Sokolsky et al. (2008)



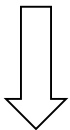
# UHECR time line



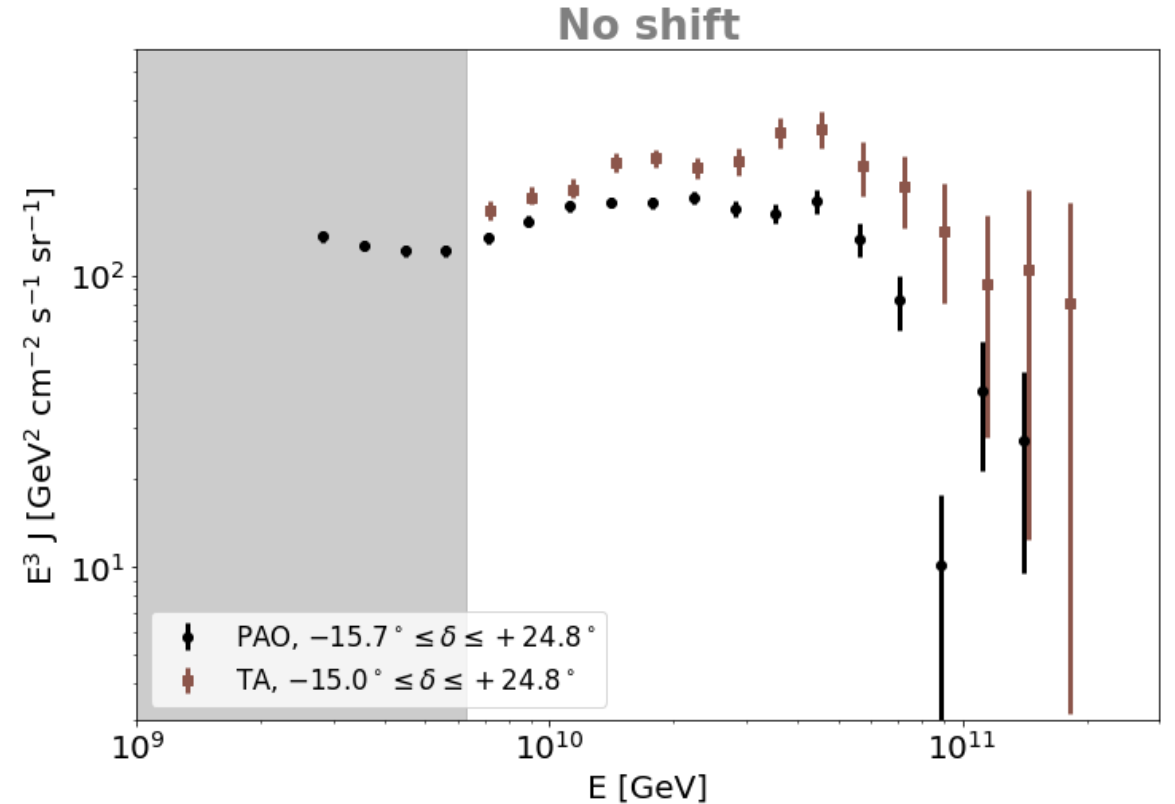
# Joint group for UHECR spectrum



**PAO and TA see the same part of the sky**



**PAO and TA should observe the same spectrum?**

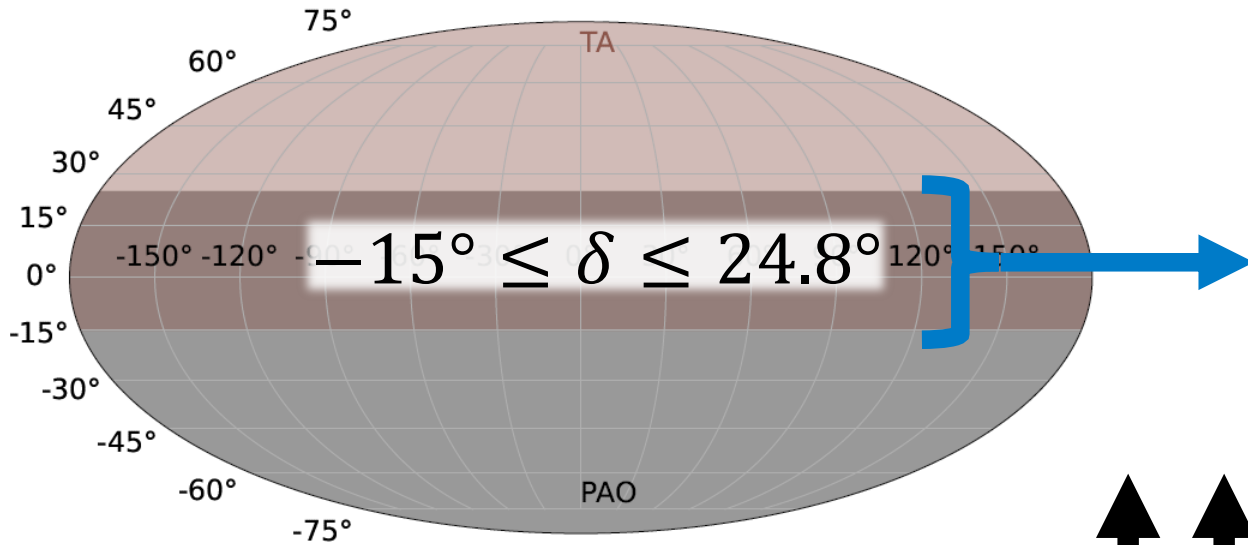


Joint spectrum group  
( ICRC 2017, 2019 and 2021)



# Joint group for UHECR spectrum

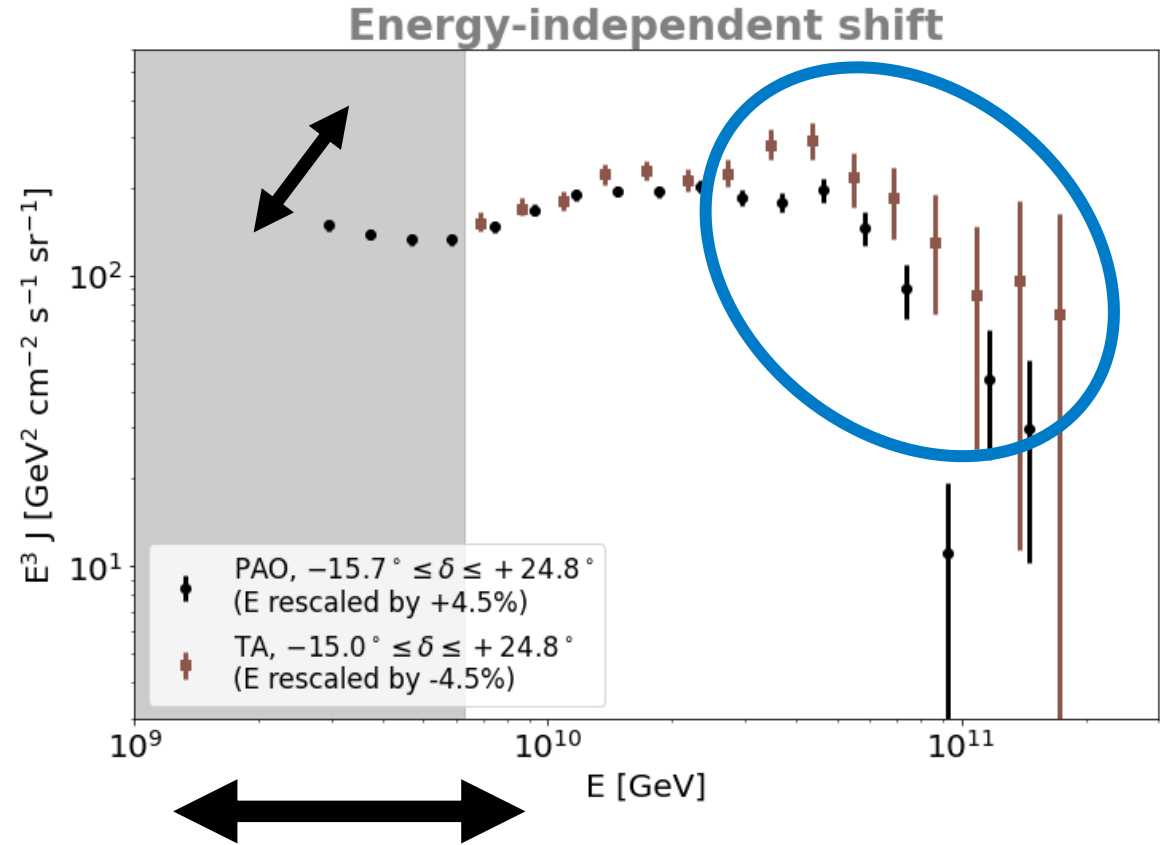
## Simple shift



Energy-independent shift:

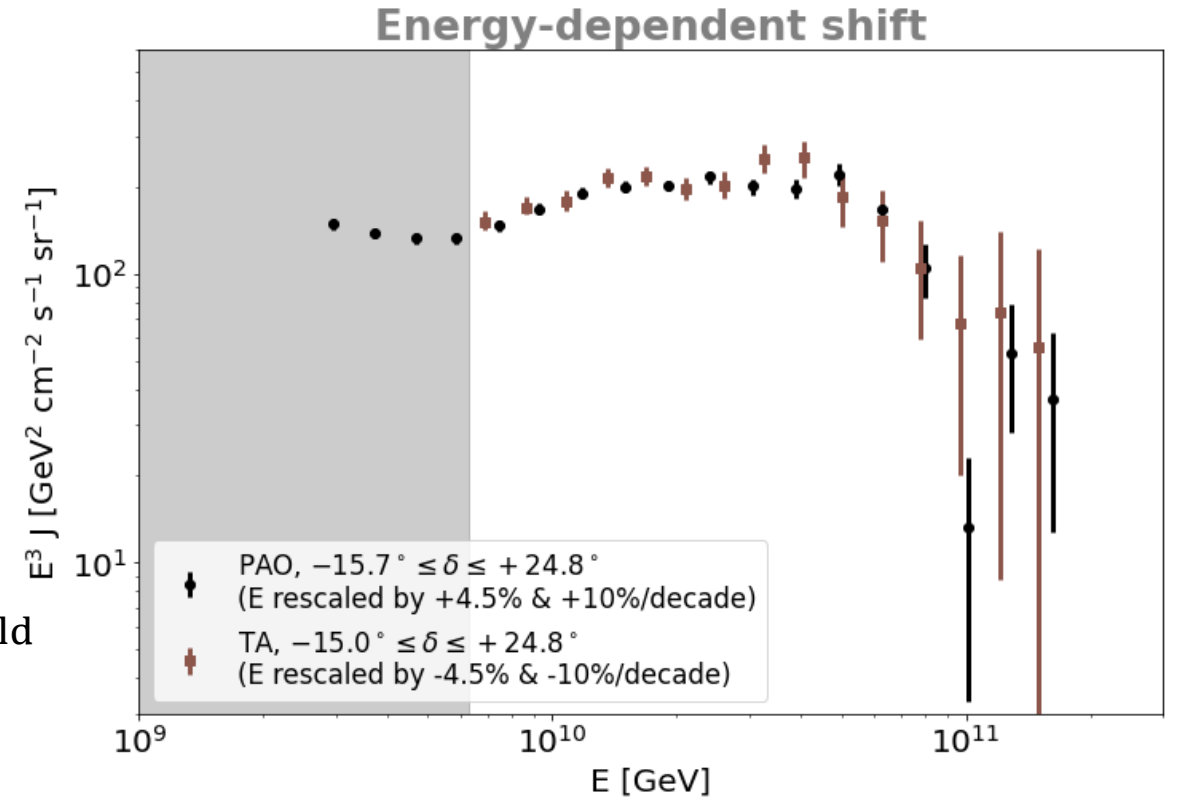
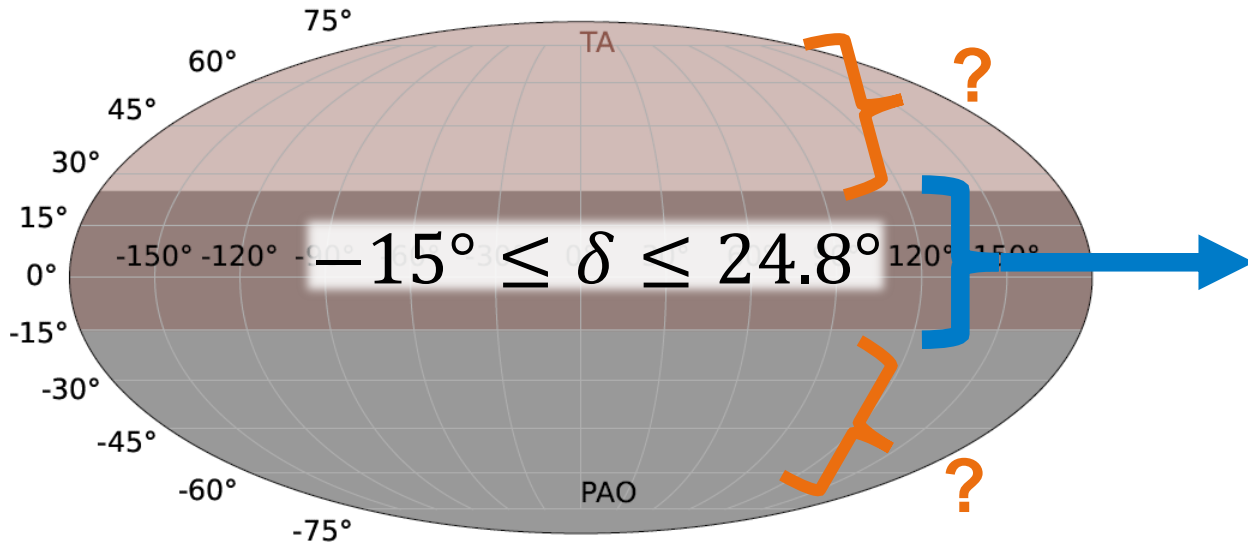
$$E_{\text{new}} = (1 \pm 4.5\%)E_{\text{old}}$$

$$E_{\text{new}}^3 \cdot J_{\text{new}} = (1 \pm 4.5\%)^2 E_{\text{old}}^3 \cdot J_{\text{old}}$$



# Joint group for UHECR spectrum

## Complex shift

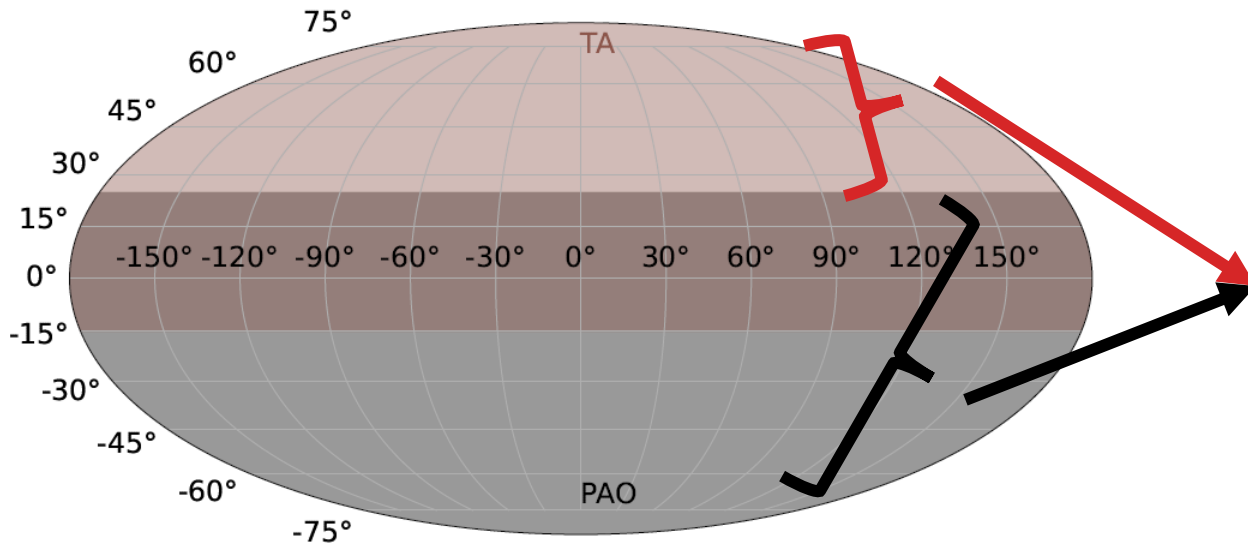


$$E_{\text{new}} = \left( 1 \pm 4.5\% \pm 10\% \cdot \log_{10} \left( \frac{E_{\text{old}}}{10^{10} \text{ GeV}} \right) \right) E_{\text{old}}$$

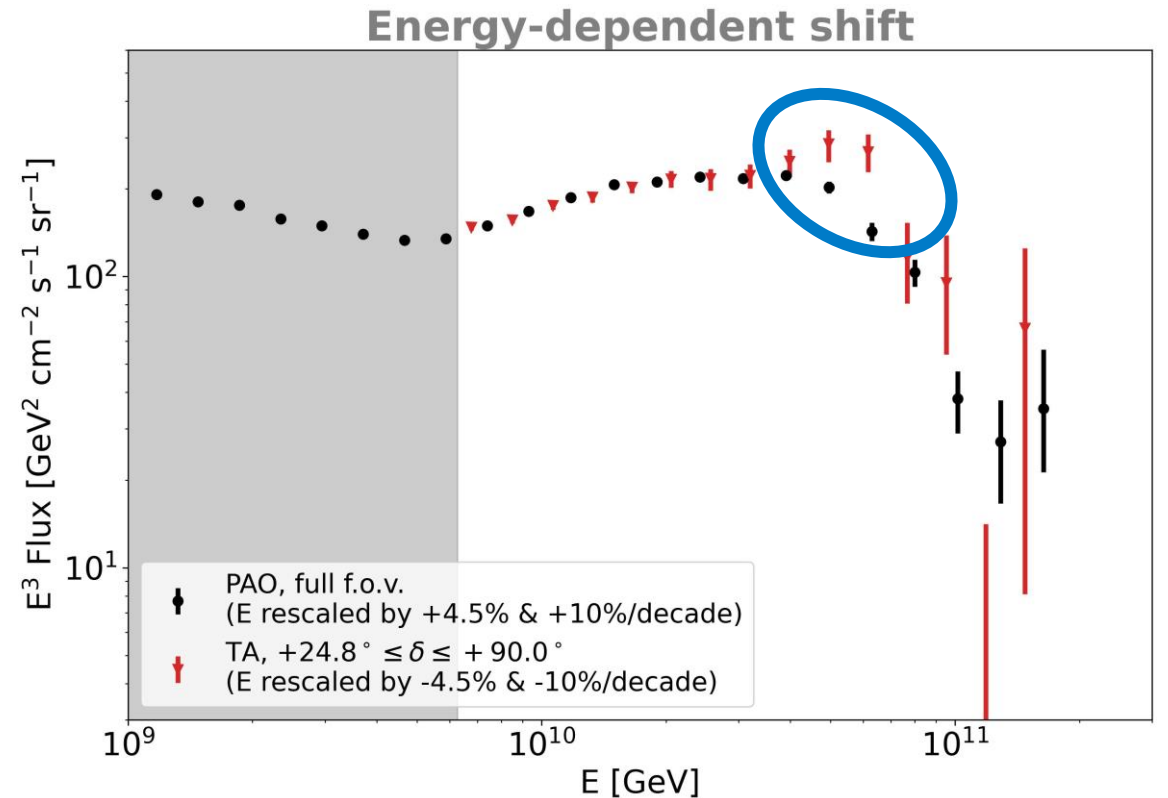
$$E_{\text{new}}^3 \cdot J_{\text{new}} = \left( 1 \pm 4.5\% \pm 10\% \cdot \log_{10} \left( \frac{E_{\text{old}}}{10^{10} \text{ GeV}} \right) \right)^2 E_{\text{old}}^3 \cdot J_{\text{old}}$$



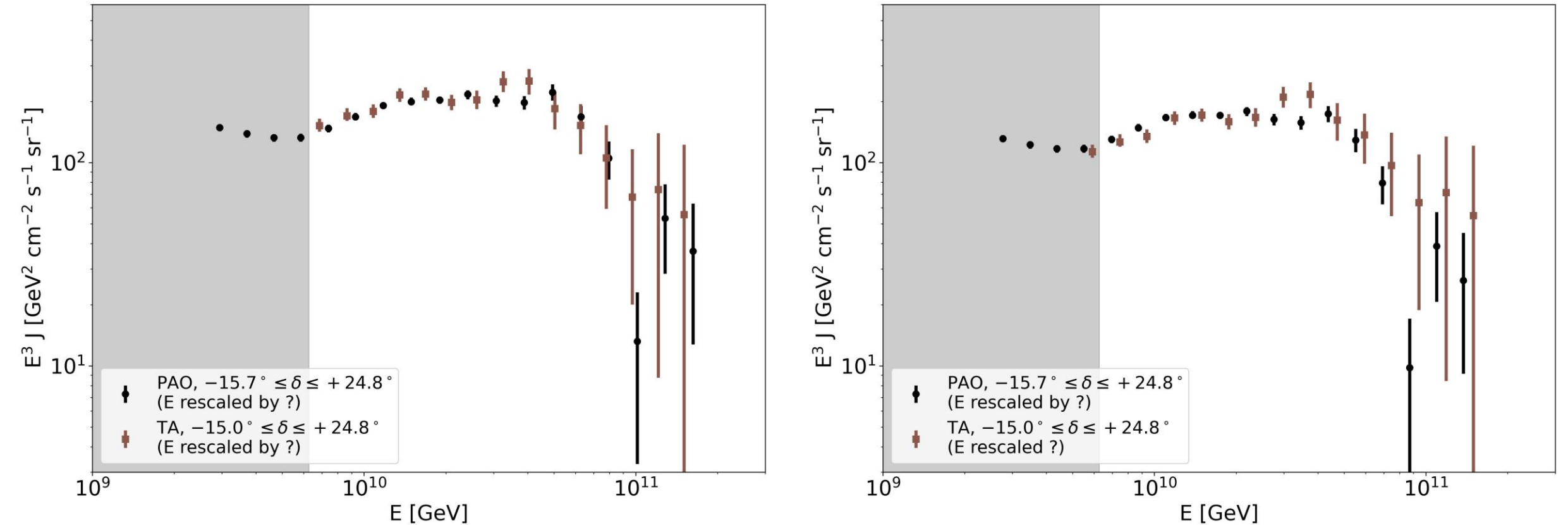
# Joint group for UHECR spectrum



Declination-dependent shift?  
→ Physics behind the shift?

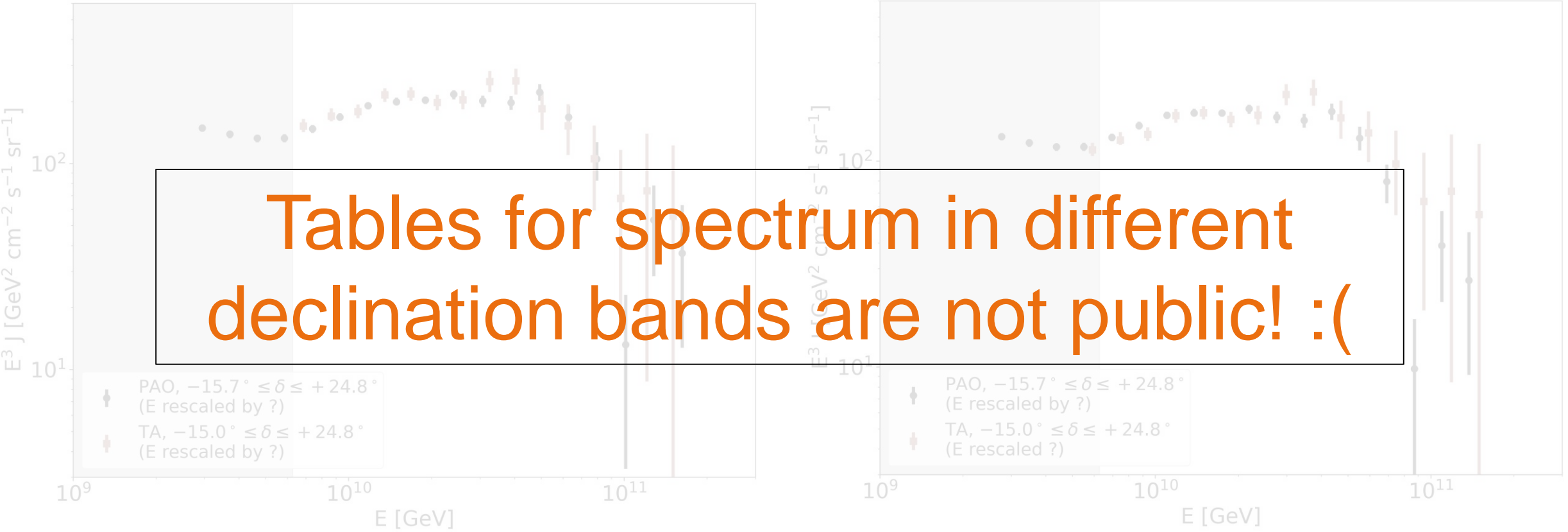


# Energy shifts in common declination band

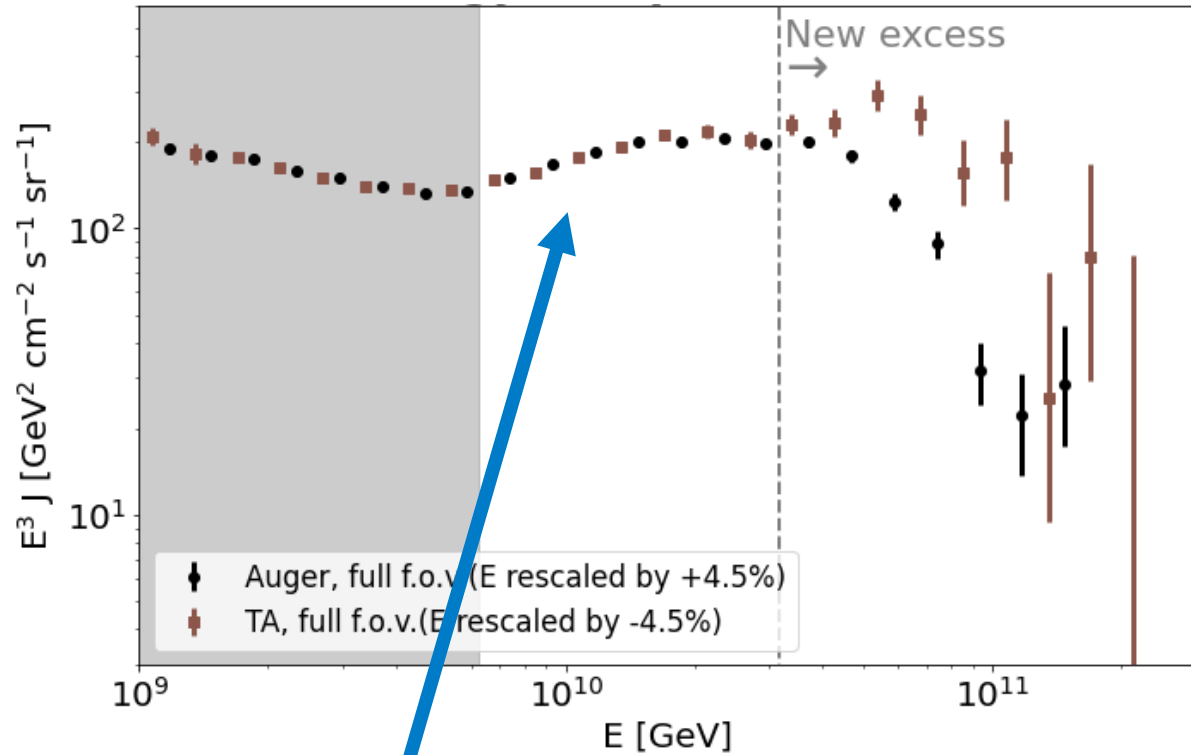


The simple and complex shifts are compatible within the uncertainties

# Energy shifts in common declination band



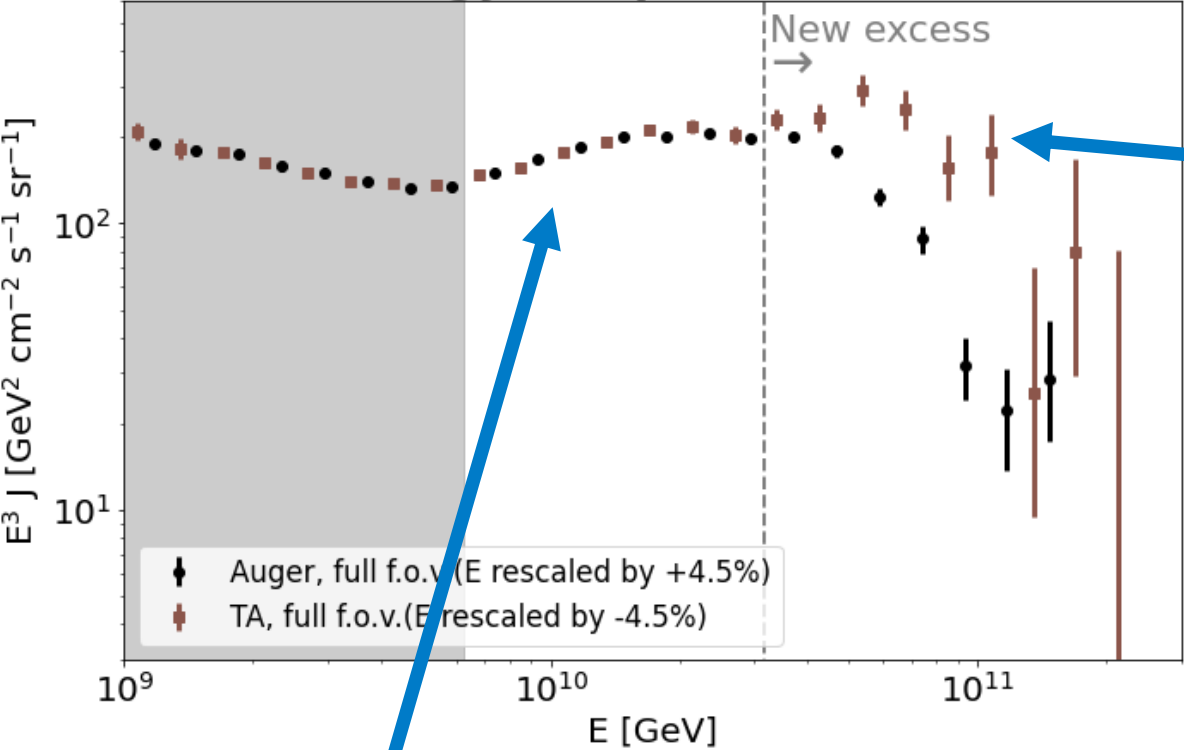
# Our proposal



Can be explained by simple shift



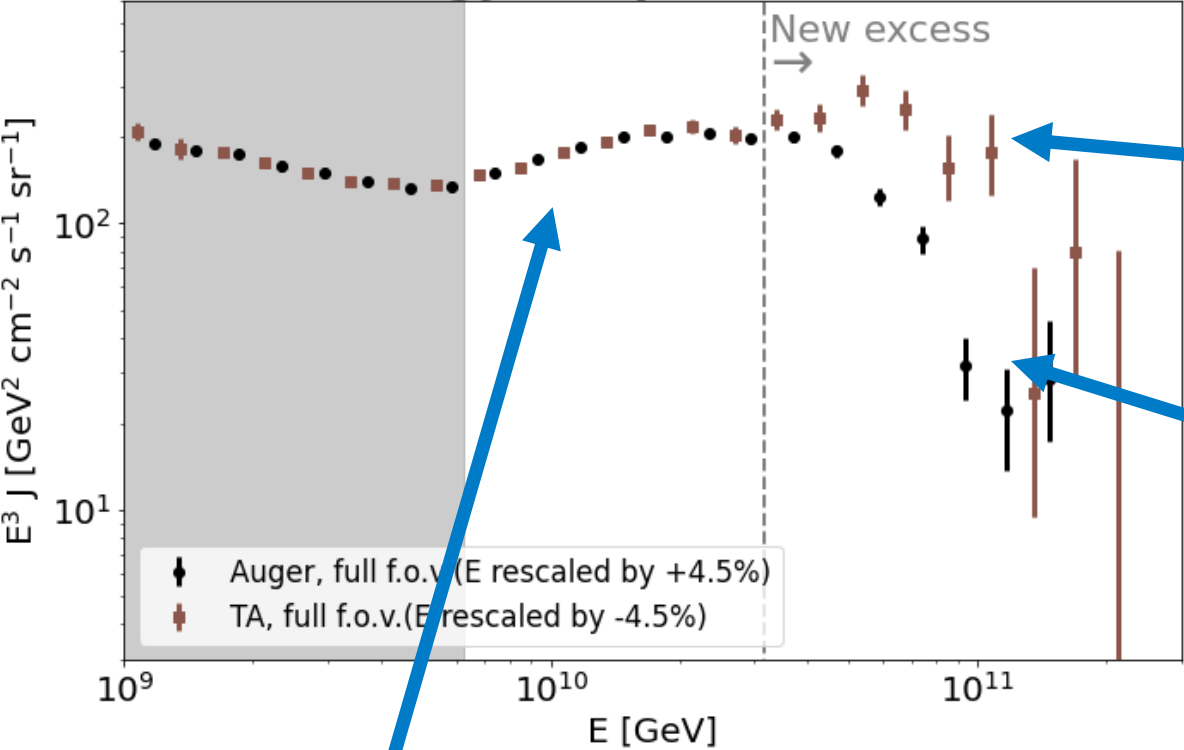
# Our proposal



Systematics?

Can be explained by simple shift

# Our proposal

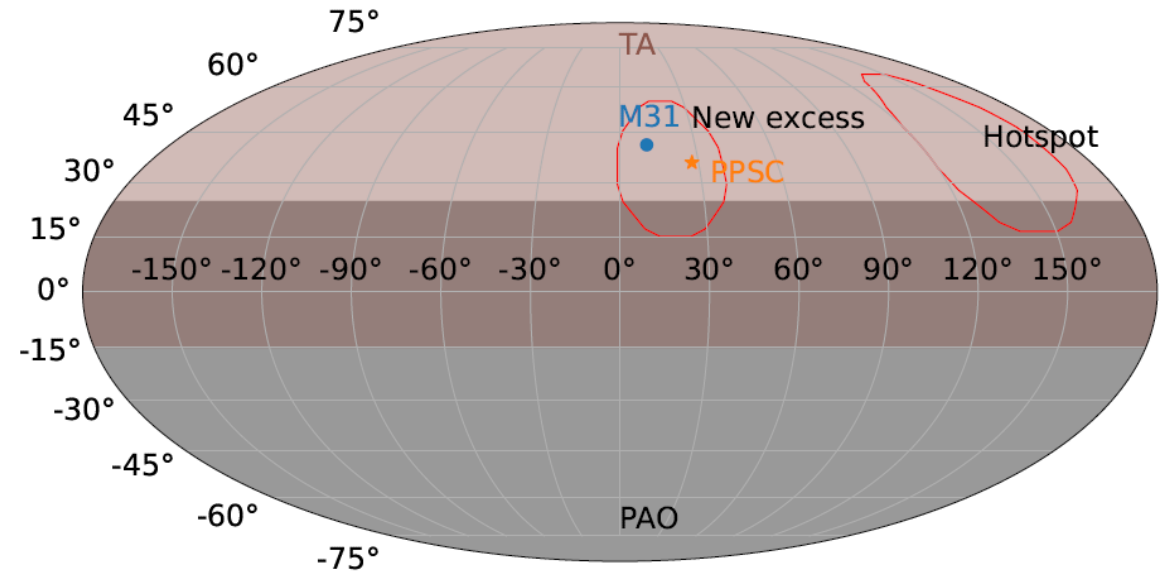
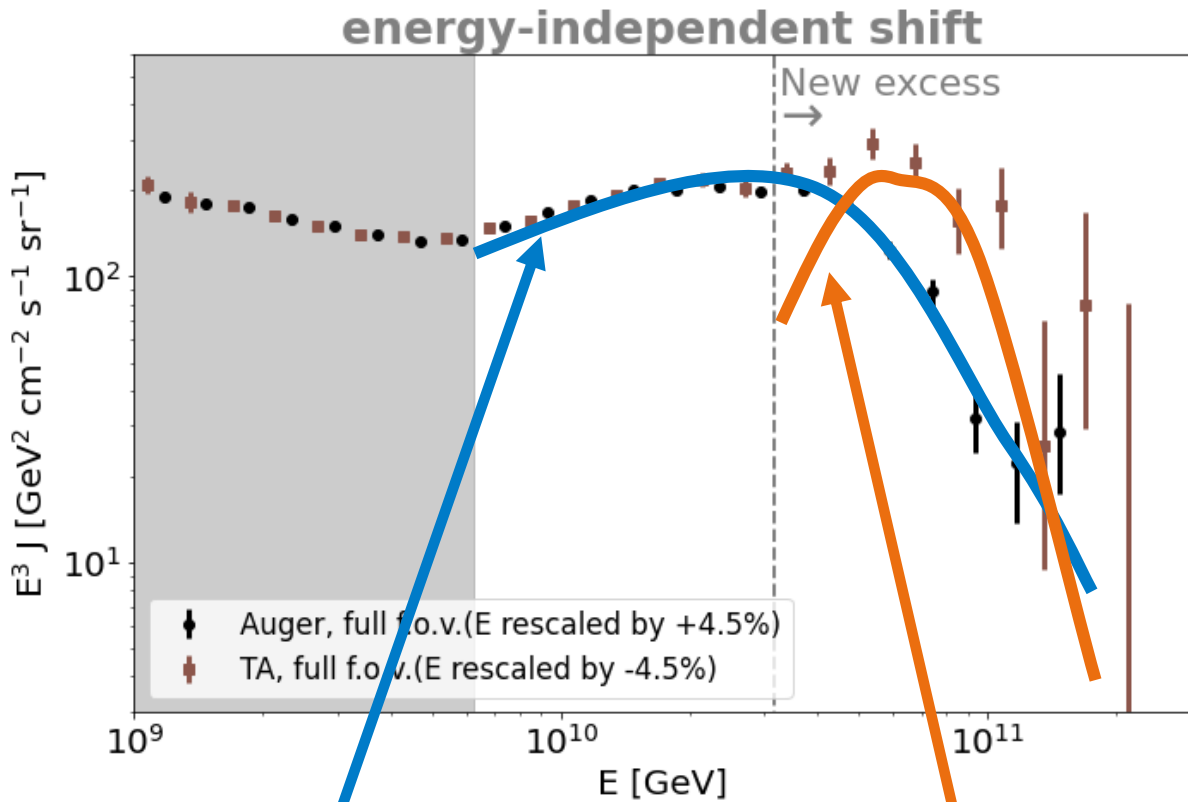


Systematics?

Astrophysical origin?

Can be explained by simple shift

# Our proposal

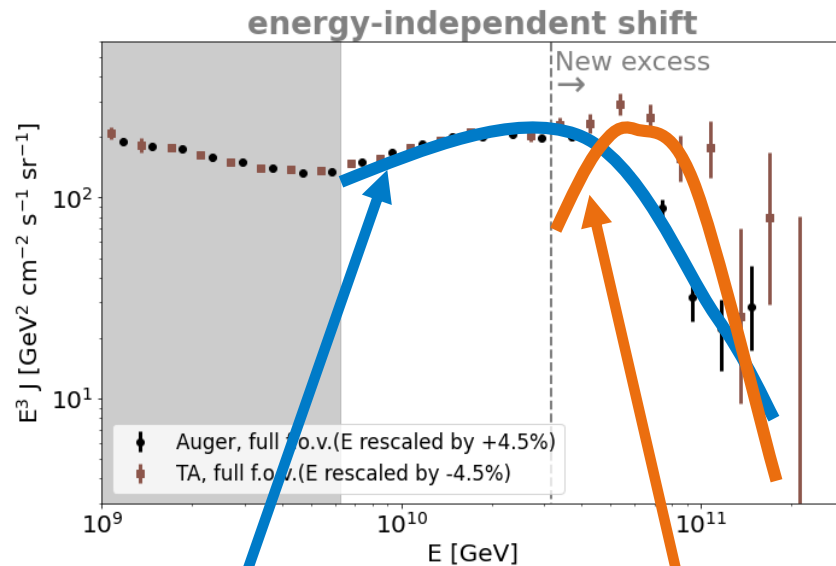


- 1. Cosmological source distribution – TA and PAO
- 2. Local source in Northern hemisphere – TA only

Kim et al. (2021)

Parameters:

- $\lambda_{\text{cosmo}} = (\gamma^{\text{cosmo}}, R_{\text{max}}^{\text{cosmo}}, m^{\text{cosmo}}, \mathcal{L}_{\text{CR}}^{\text{cosmo}}, f_{\text{A}}^{\text{cosmo}})$
- $\lambda_{\text{local}} = (\gamma^{\text{local}}, R_{\text{max}}^{\text{local}}, D^{\text{local}}, L^{\text{local}}, A^{\text{local}})$



1. Cosmological source distribution – TA and PAO
2. Local source in Northern hemisphere – TA only

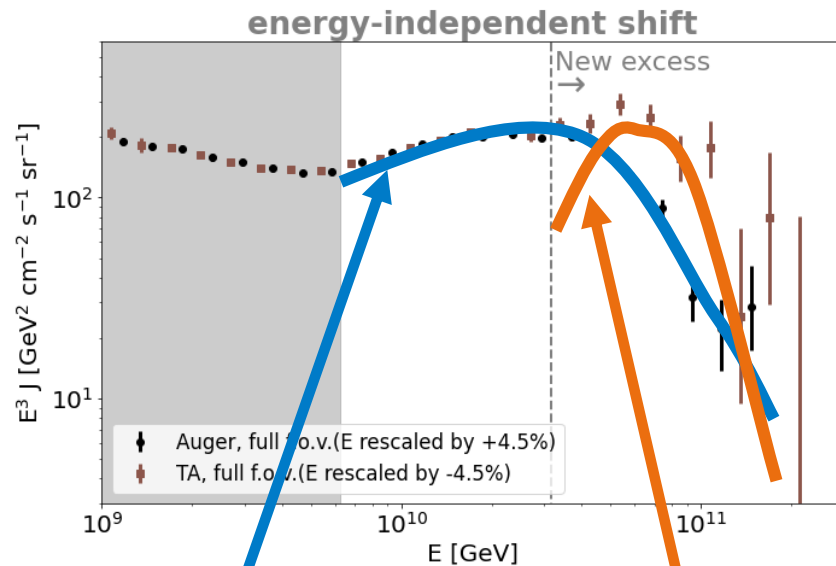


Parameters:

- $\lambda_{\text{cosmo}} = (\gamma^{\text{cosmo}}, R_{\text{max}}^{\text{cosmo}}, m^{\text{cosmo}}, \mathcal{L}_{\text{CR}}^{\text{cosmo}}, f_{\text{A}}^{\text{cosmo}})$
- $\lambda_{\text{local}} = (\gamma^{\text{local}}, R_{\text{max}}^{\text{local}}, D^{\text{local}}, L^{\text{local}}, A^{\text{local}})$

$\chi^2$ test (Combined fit):

- $\chi_{\text{PAO}}^2 = \chi_{\text{PAO}}^2(\lambda_{\text{cosmo}}, \delta_{\text{E}}^{\text{PAO}}, \delta_{\langle X_{\text{max}} \rangle}^{\text{PAO}}, \delta_{\sigma(X_{\text{max}})}^{\text{PAO}})$
- $\chi_{\text{TA}}^2 = \chi_{\text{TA}}^2(\lambda_{\text{cosmo}}, \lambda_{\text{local}}, \delta_{\text{E}}^{\text{TA}}, \delta_{\langle X_{\text{max}} \rangle}^{\text{TA}}, \delta_{\sigma(X_{\text{max}})}^{\text{TA}})$



1. **Cosmological source distribution** – TA and PAO
2. **Local source in Northern hemisphere** – TA only

$\delta$  represent the systematic uncertainties

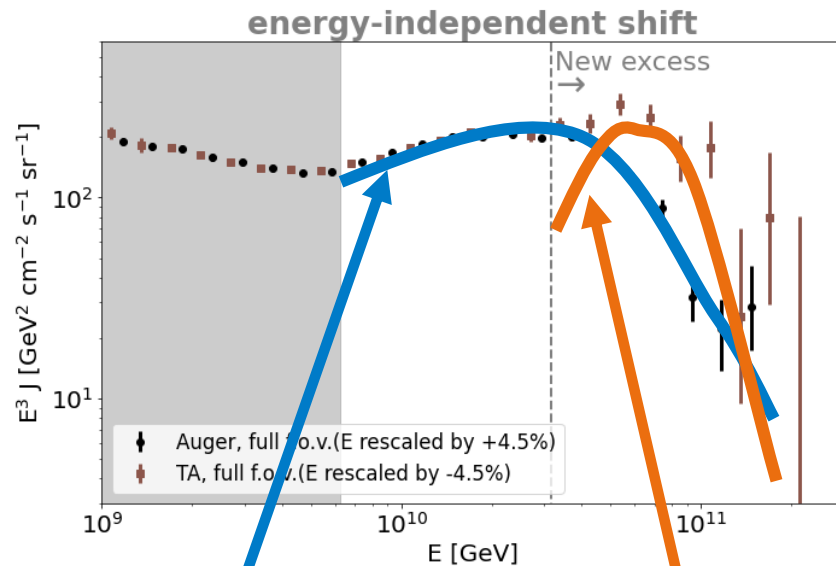
Parameters:

- $\lambda_{\text{cosmo}} = (\gamma^{\text{cosmo}}, R_{\text{max}}^{\text{cosmo}}, m^{\text{cosmo}}, \mathcal{L}_{\text{CR}}^{\text{cosmo}}, f_A^{\text{cosmo}})$
- $\lambda_{\text{local}} = (\gamma^{\text{local}}, R_{\text{max}}^{\text{local}}, D^{\text{local}}, L^{\text{local}}, A^{\text{local}})$

$\chi^2$ test (Combined fit):

$$\chi_{\text{PAO}}^2 = \chi_{\text{PAO}}^2(\lambda_{\text{cosmo}}, \delta_E^{\text{PAO}}, \delta_{\langle X_{\text{max}} \rangle}^{\text{PAO}}, \delta_{\sigma(X_{\text{max}})}^{\text{PAO}})$$

$$\chi_{\text{TA}}^2 = \chi_{\text{TA}}^2(\lambda_{\text{cosmo}}, \lambda_{\text{local}}, \delta_E^{\text{TA}}, \delta_{\langle X_{\text{max}} \rangle}^{\text{TA}}, \delta_{\sigma(X_{\text{max}})}^{\text{TA}})$$



Joint fit:

$$\chi_{\text{global}}^2(\lambda^{\text{cosmo}}, \lambda^{\text{local}}, \delta) = \chi_{\text{PAO}}^2 + \chi_{\text{TA}}^2 +$$

$$+ \left(\frac{\delta_E^{\text{PAO}}}{\sigma_E^{\text{PAO}}}\right)^2 + \left(\frac{\delta_{\langle X_{\text{max}} \rangle}^{\text{PAO}}}{\sigma_{\langle X_{\text{max}} \rangle}^{\text{PAO}}}\right)^2 + \left(\frac{\delta_{\sigma(X_{\text{max}})}^{\text{PAO}}}{\sigma_{\sigma(X_{\text{max}})}^{\text{PAO}}}\right)^2$$

$$+ \left(\frac{\delta_E^{\text{TA}}}{\sigma_E^{\text{TA}}}\right)^2 + \left(\frac{\delta_{\langle X_{\text{max}} \rangle}^{\text{TA}}}{\sigma_{\langle X_{\text{max}} \rangle}^{\text{TA}}}\right)^2 + \left(\frac{\delta_{\sigma(X_{\text{max}})}^{\text{TA}}}{\sigma_{\sigma(X_{\text{max}})}^{\text{TA}}}\right)^2$$

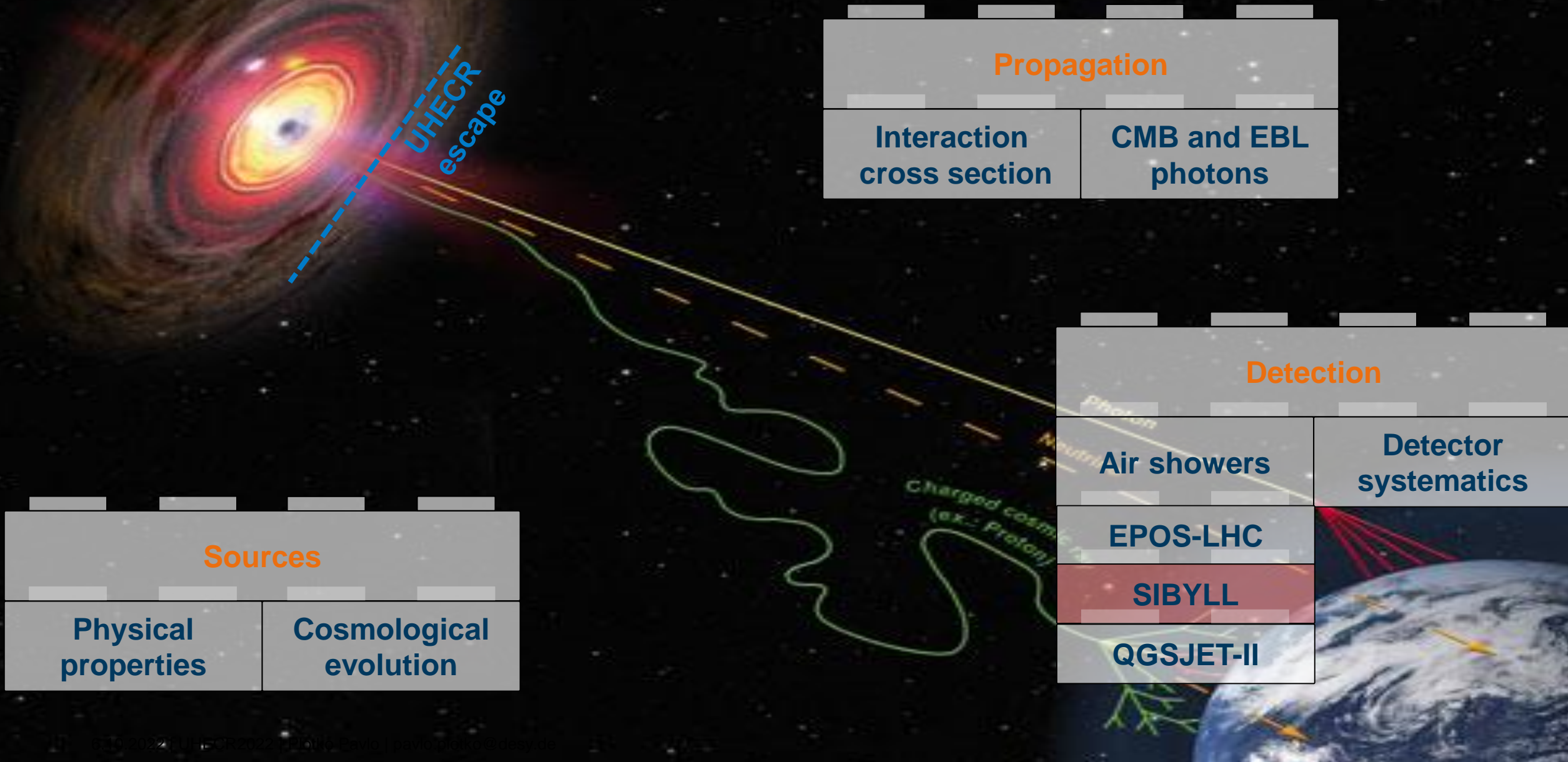
1. Cosmological source distribution – TA and PAO
2. Local source in Northern hemisphere – TA only

$\delta$  represent the systematic uncertainties

~ 20 CPU years to reproduce results

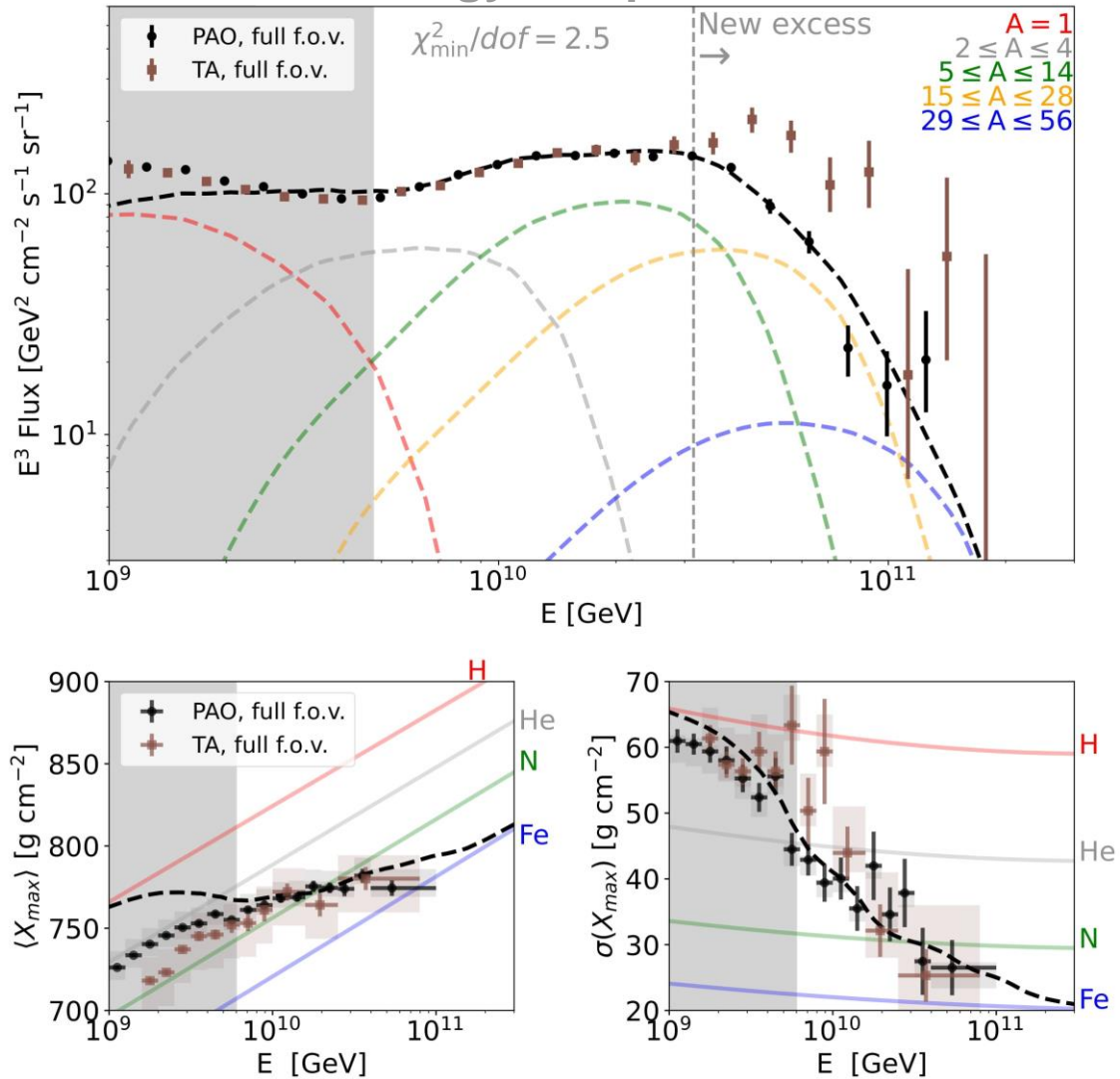
# Propagation including Nuclear Cascade

Heinze et al. (2019)  
github.com/joheinze/PriNCE



## Systematics

### Energy-independent shift

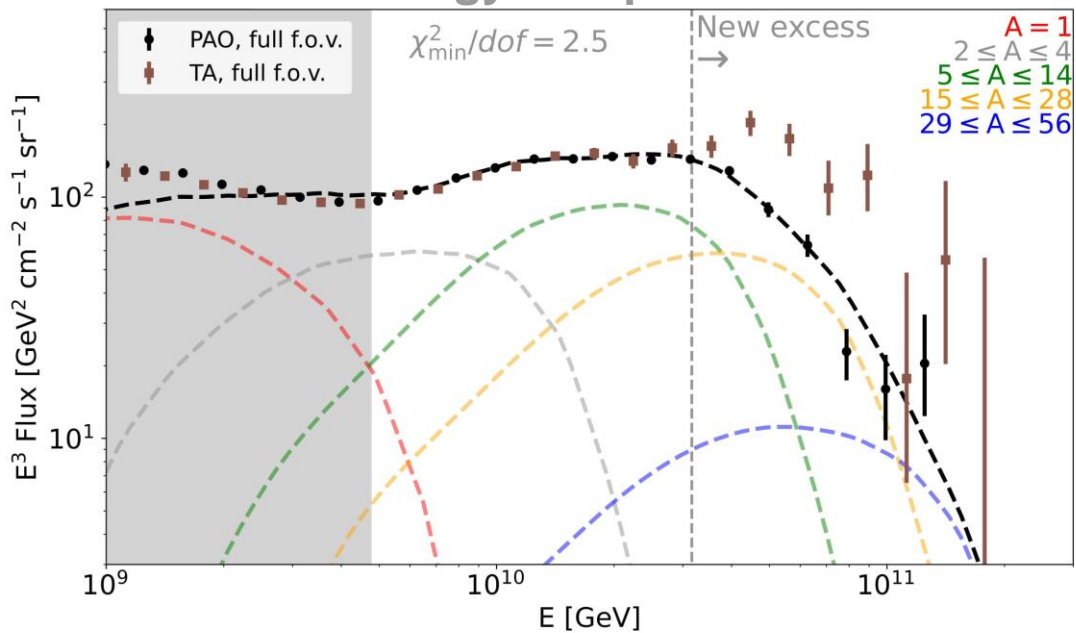




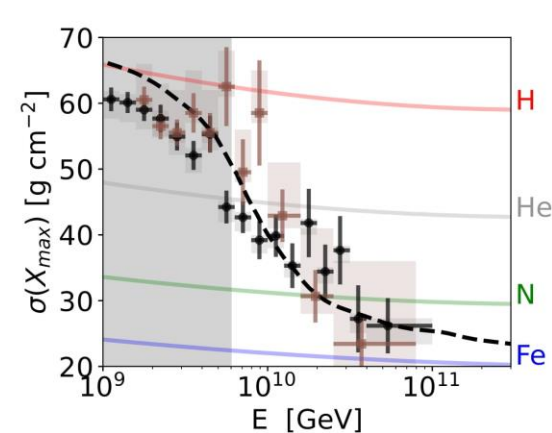
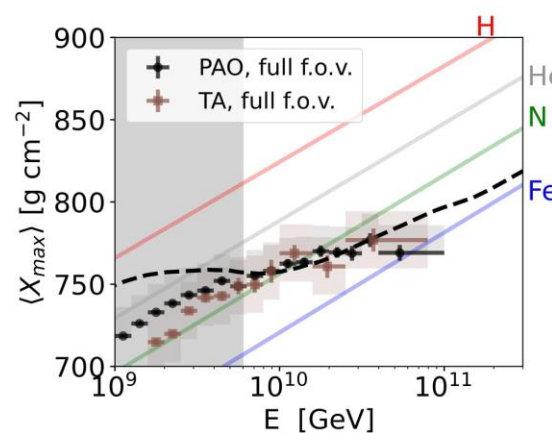
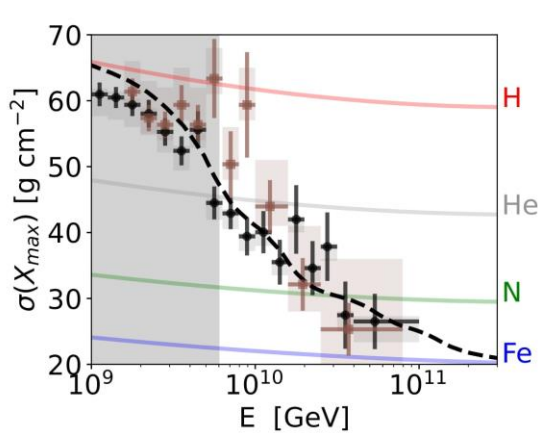
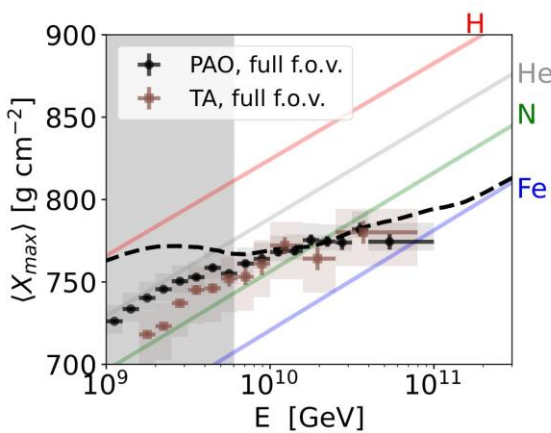
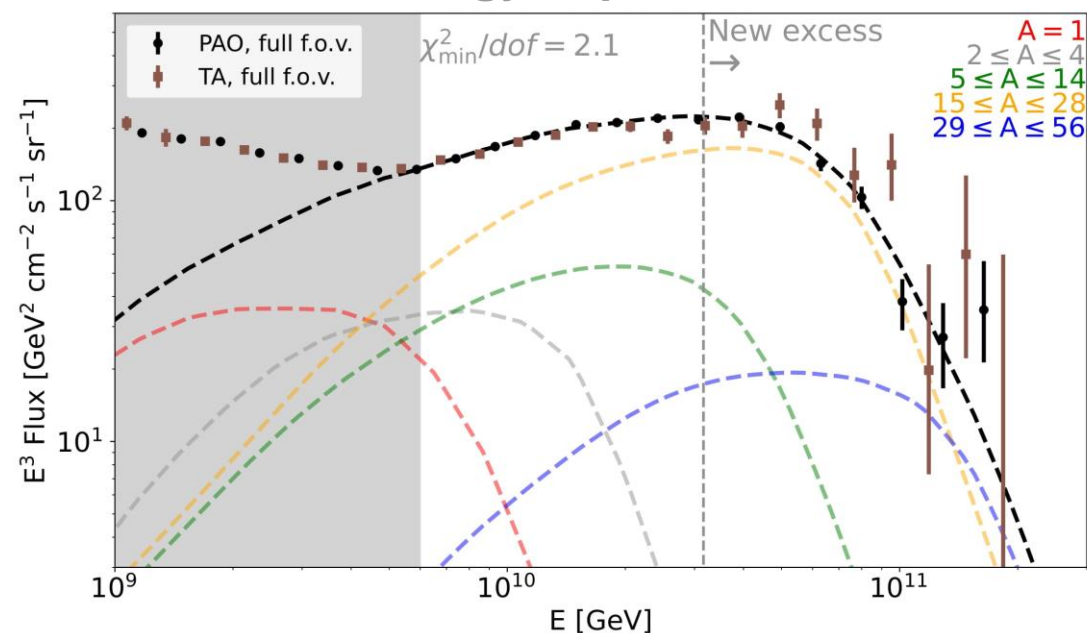
# Results

## Systematics

### Energy-independent shift



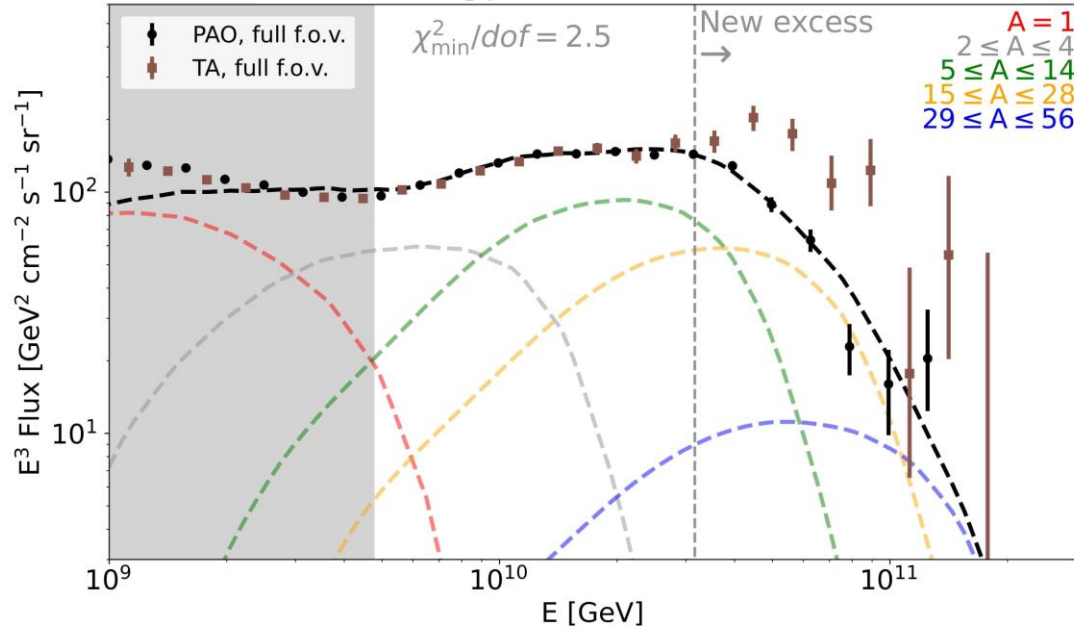
### Energy-dependent shift



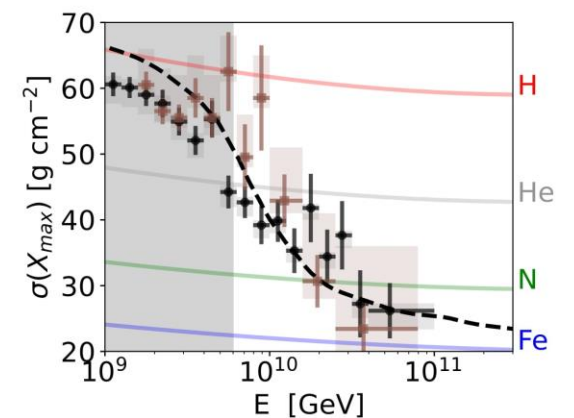
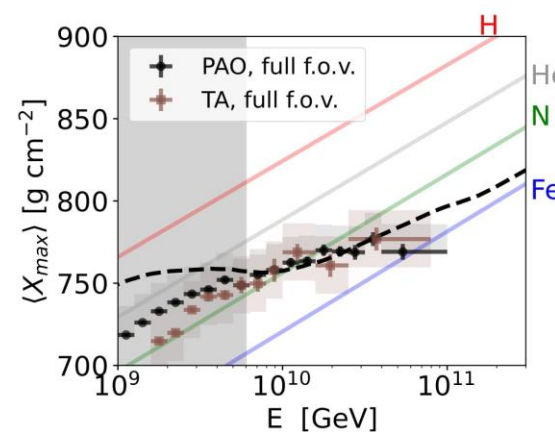
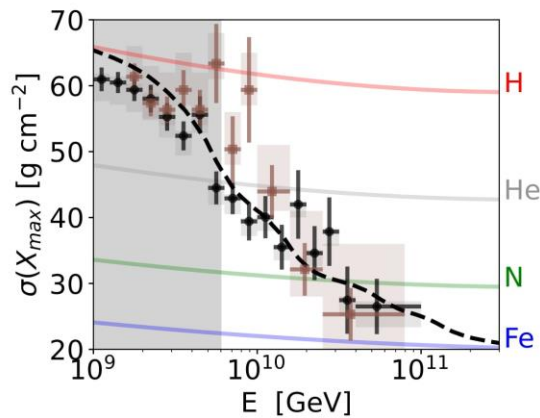
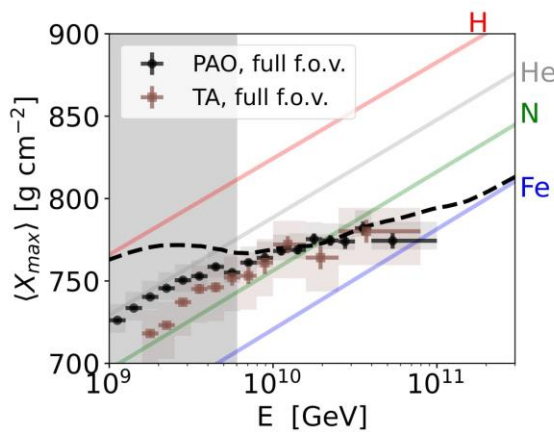
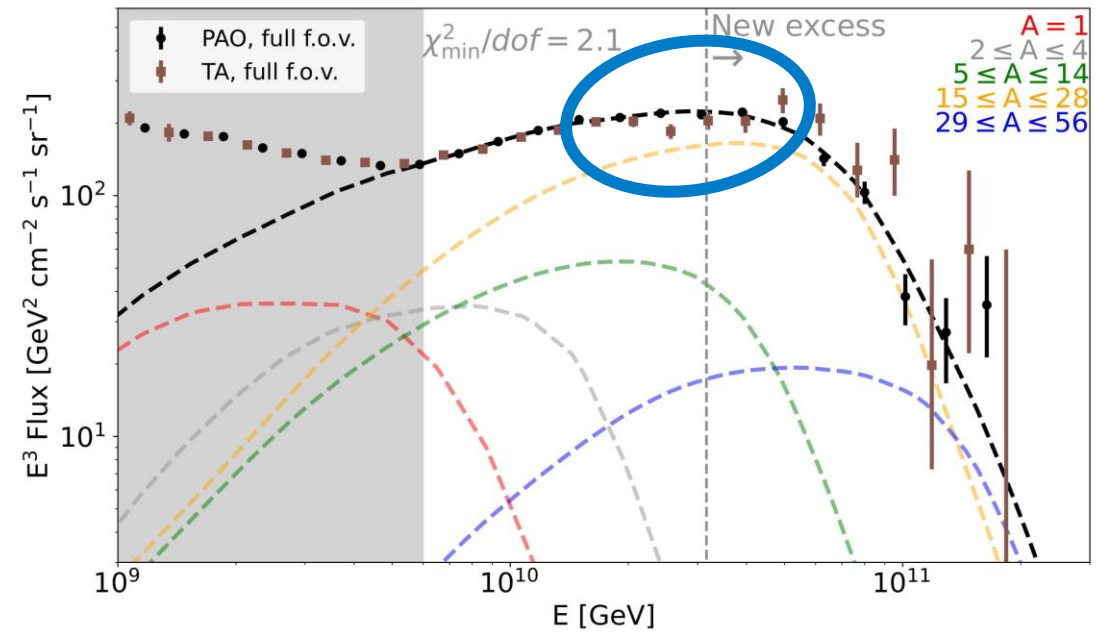
# Results

## Systematics

### Energy-independent shift



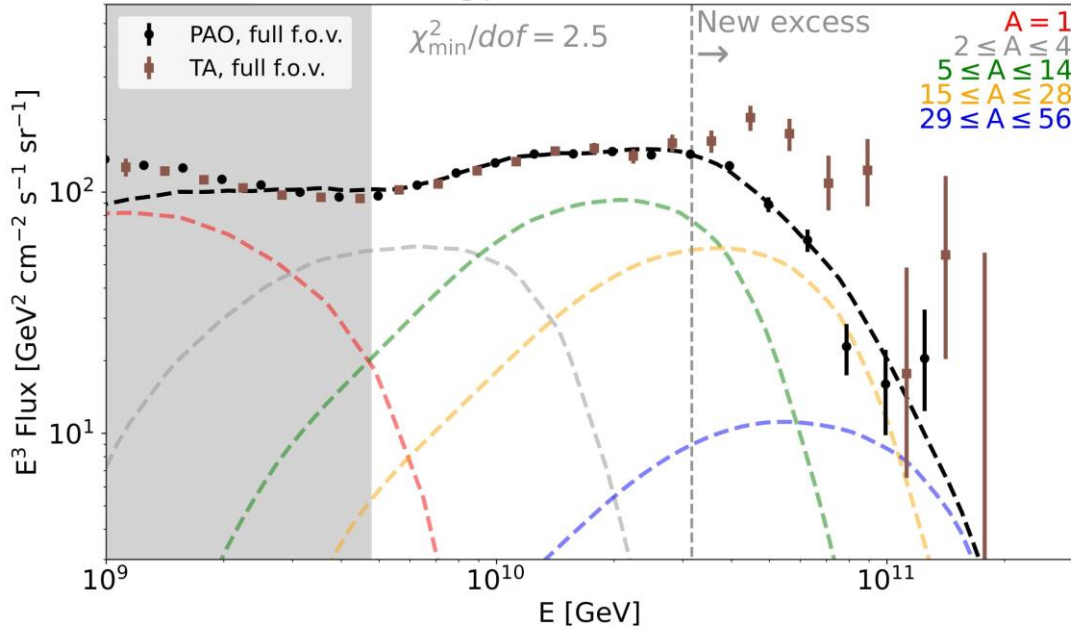
### Energy-dependent shift



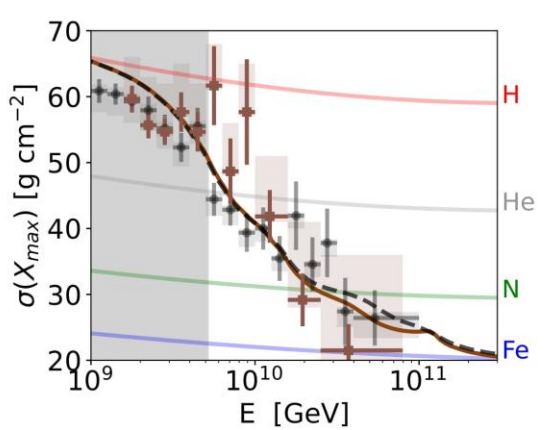
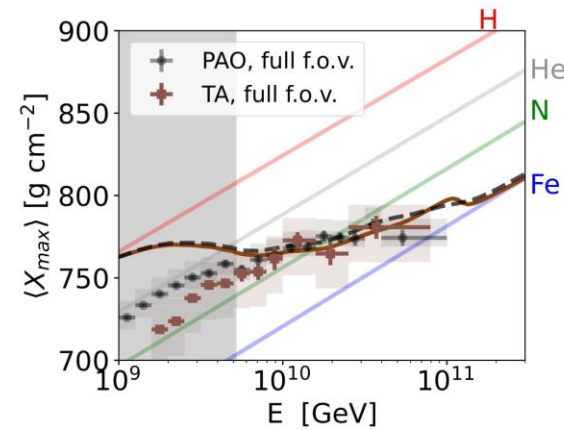
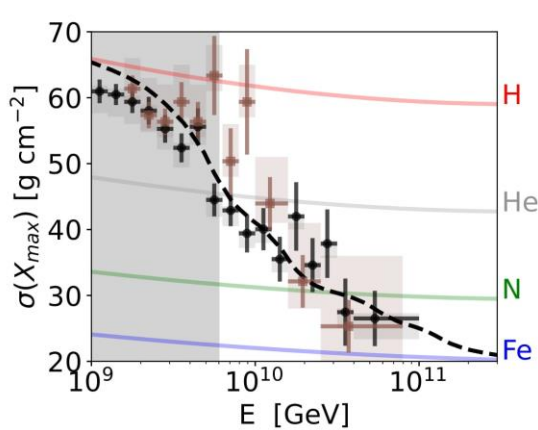
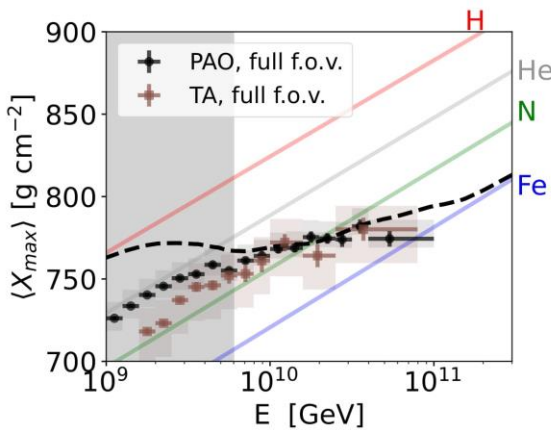
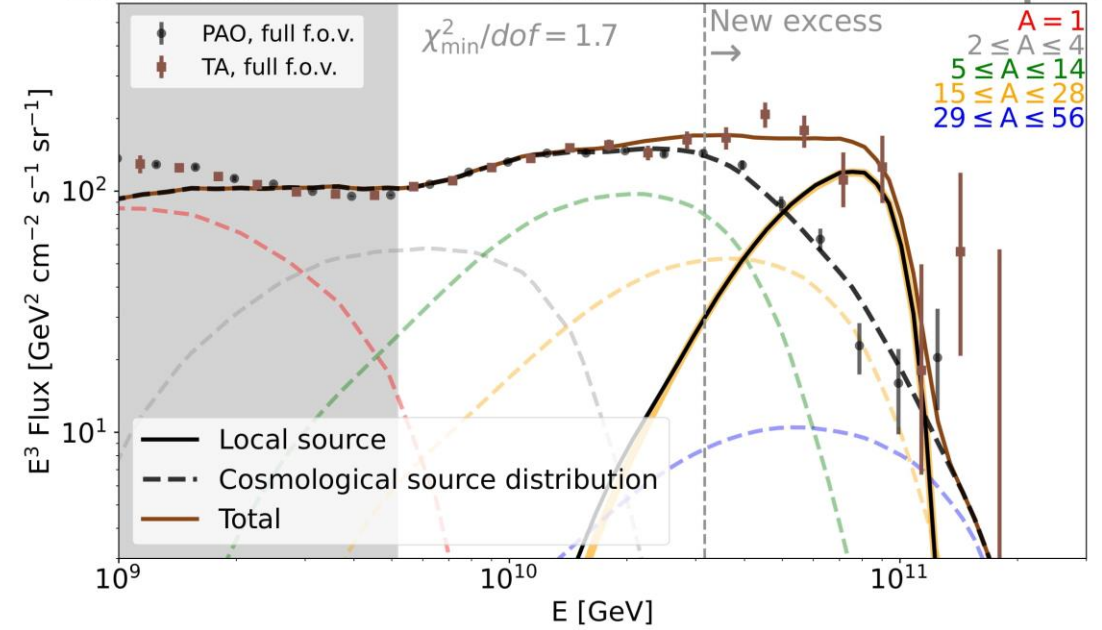
# Results

## Astrophysical origin

### Energy-independent shift

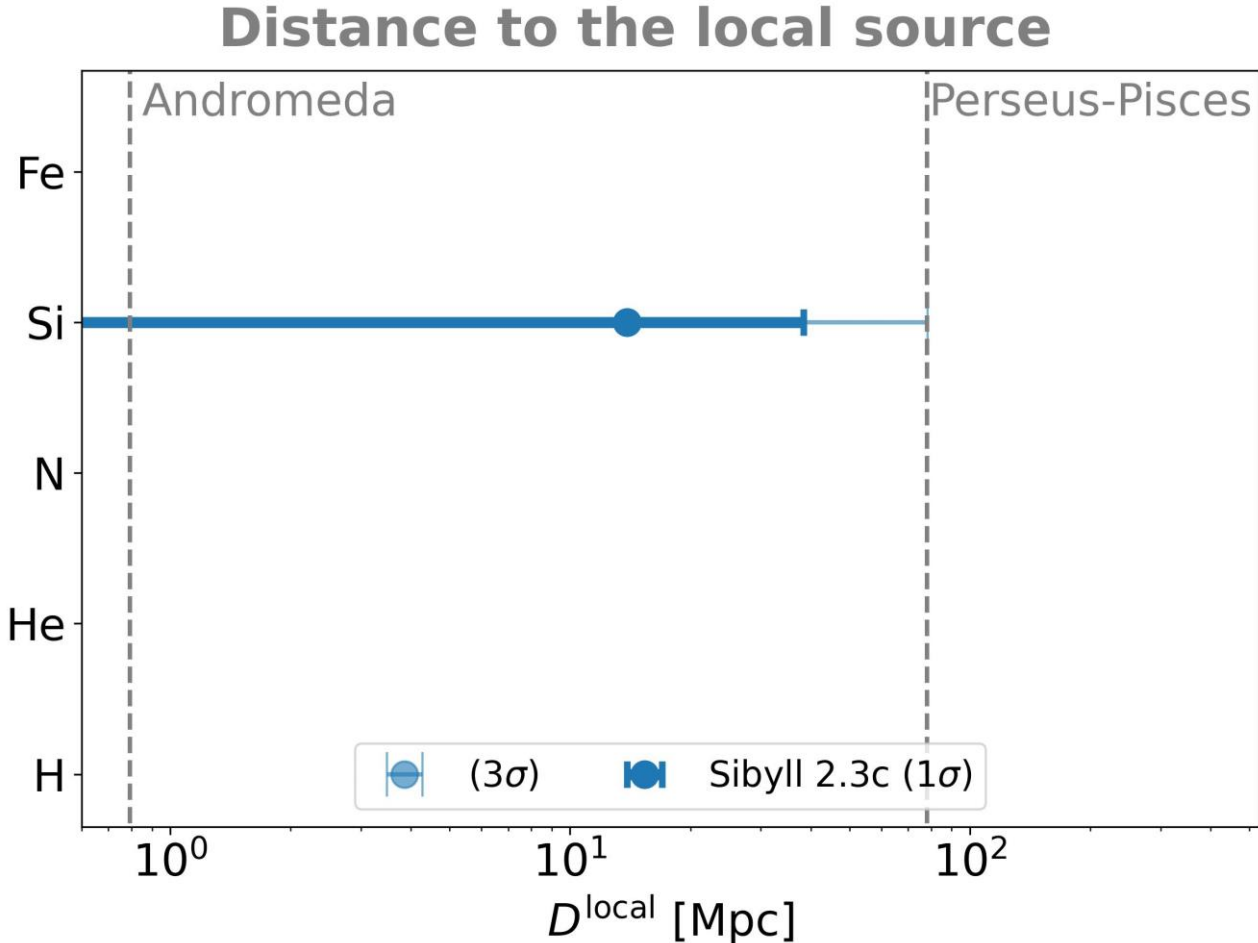


### Local source of silicon-28 in the Northern Hemisphere



# Local source

## Distance



A local source emits only one isotope (silicon-28)

Distance  $13.9^{+9.2}_{-13.9}$  Mpc

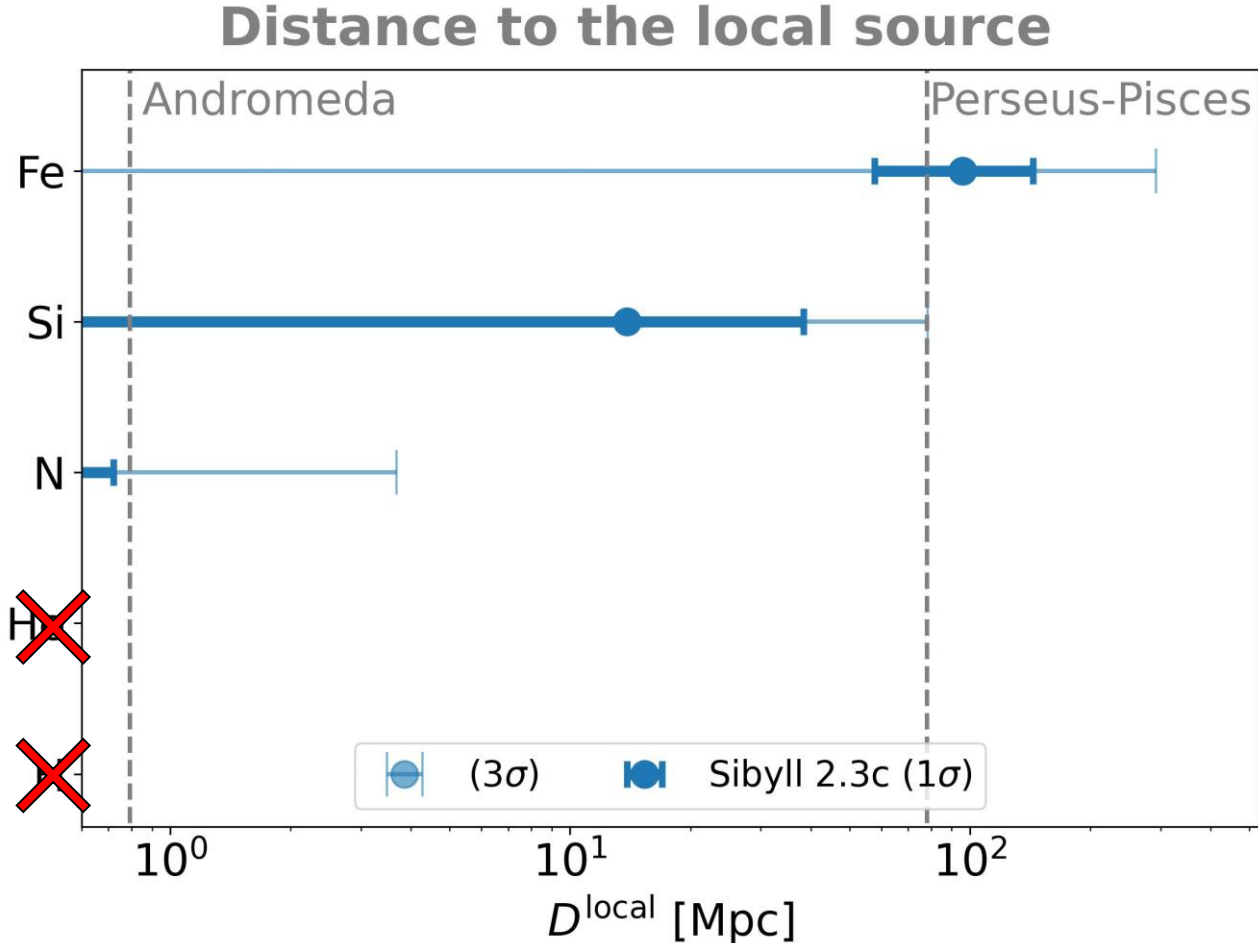
Luminosity  $7.3^{+18.0}_{-7.3} \cdot 10^{41}$  erg/s

The maximum rigidity  $1.3^{+0.2}_{-0.1} \cdot 10^9$  GV



# Local source

## Distance



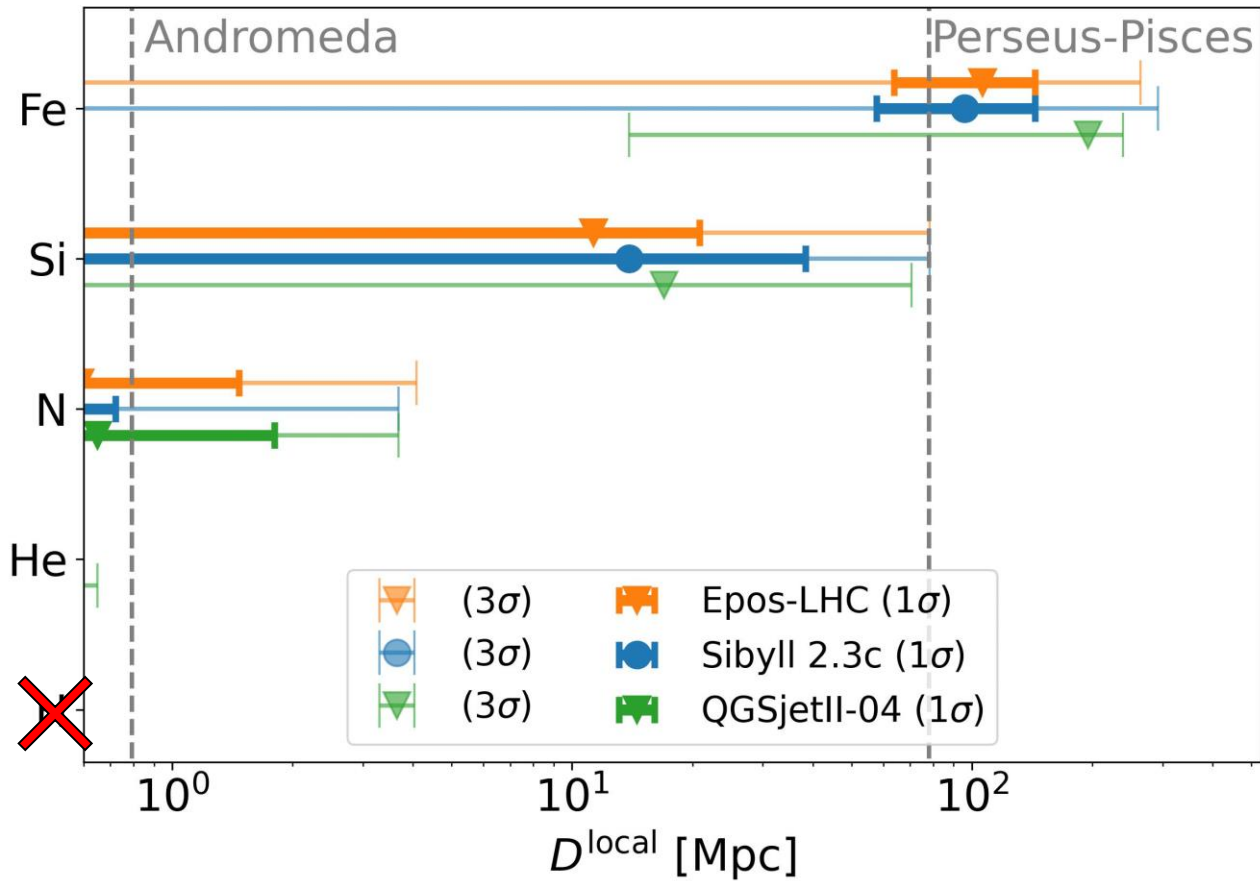
1. Intermediate-mass elements are preferred.
2. He (almost) and H are excluded in 3 sigma.
3. Distance depends on the type of isotope.



# Local source

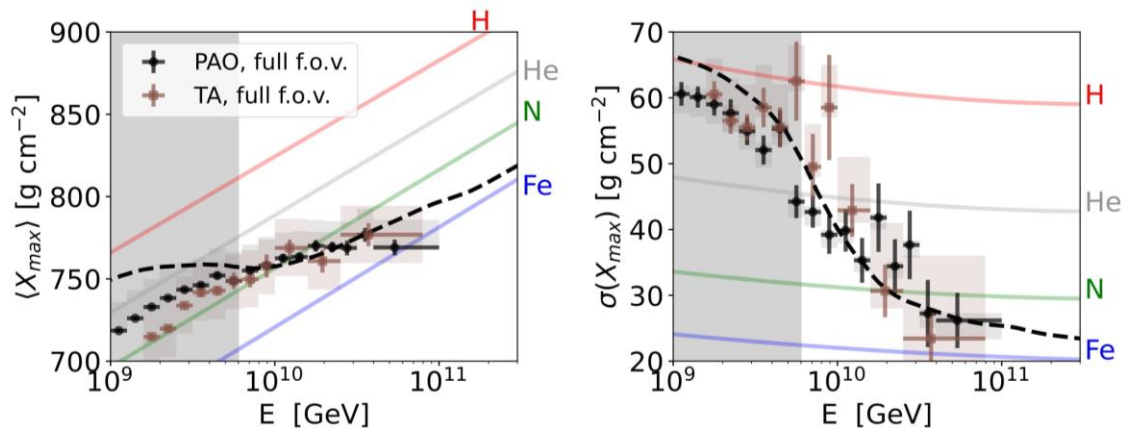
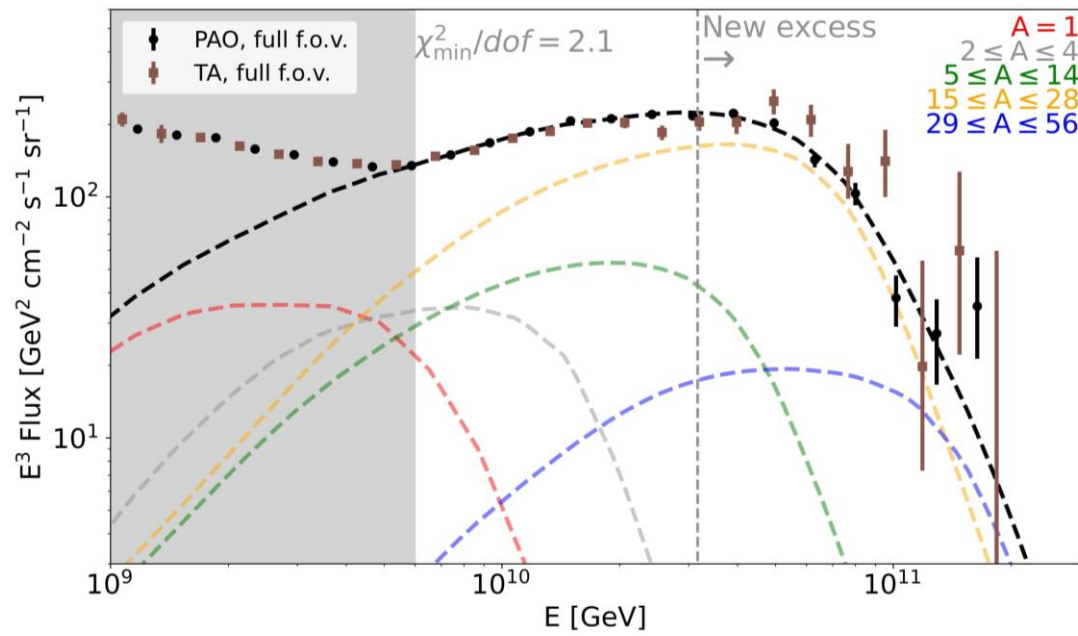
## Distance

### Distance to the local source

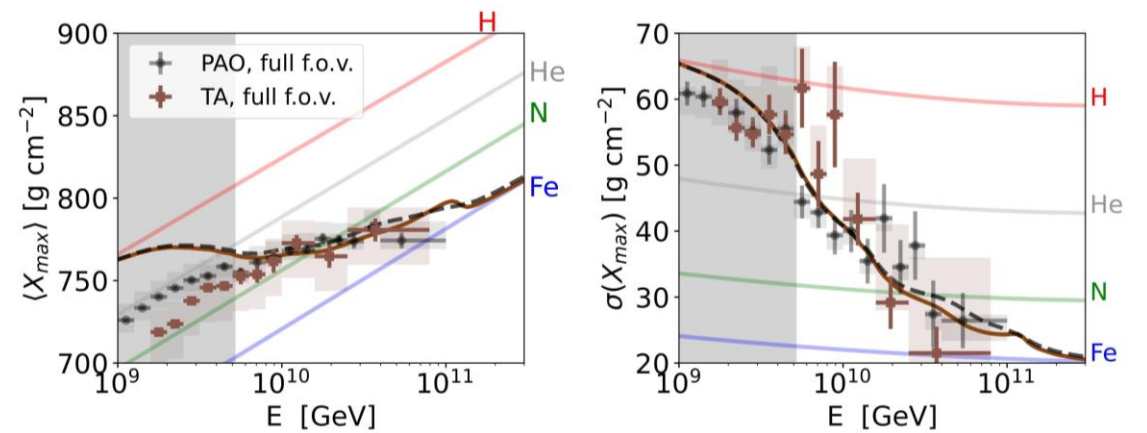
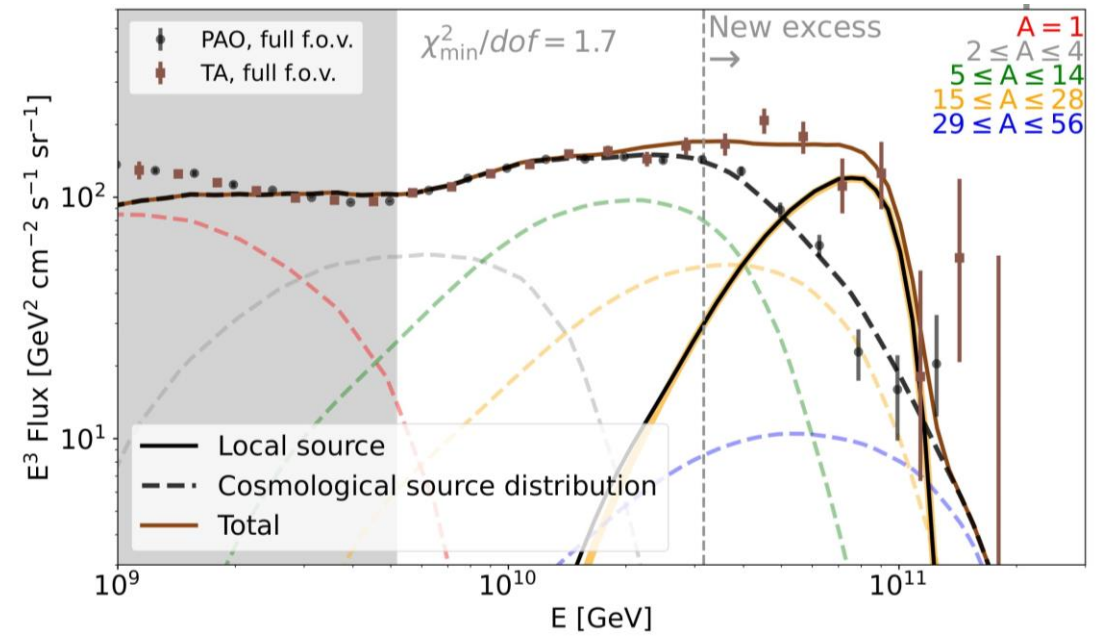


1. Intermediate-mass elements are preferred.
2. He (almost) and H are excluded in 3 sigma.
3. Distance depends on the type of isotope.

# Systematics



# vs Astrophysical origin



# Take-home messages

**Both scenarios could explain the differences**

## Systematics:

1. What about the different declination bands?
2. Physics behind the shift?

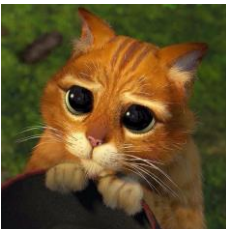
## Astrophysical:

1. What is the local source?
2. Is there a probability to observe it?

Are there any differences between the simple and complex shifts in the common declination band?

© DreamWorks Pictures (7)

More publicly available data for theory :)



# Credits for imagines

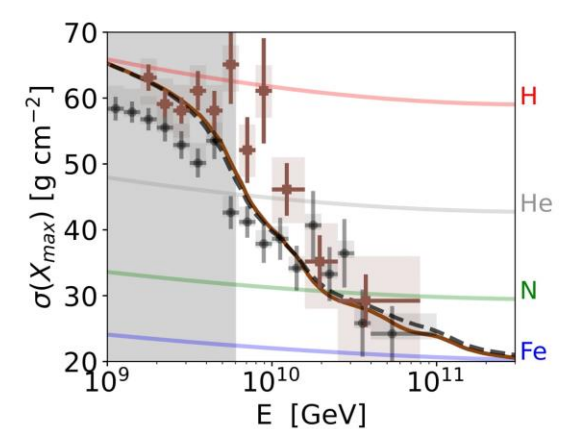
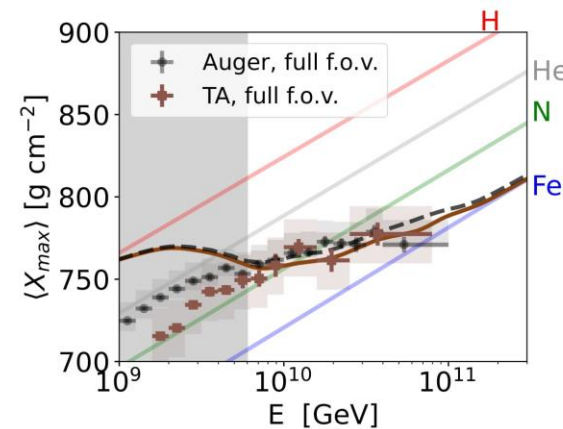
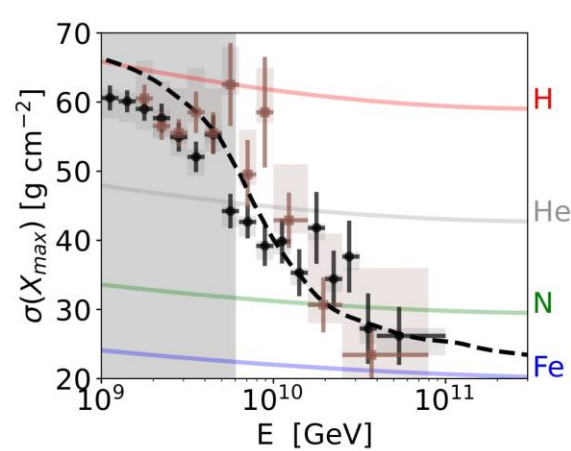
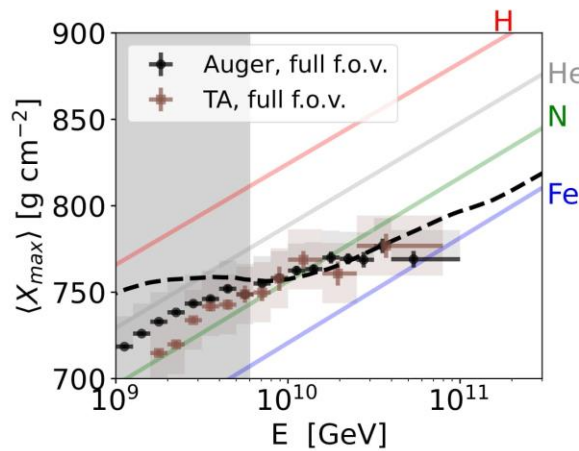
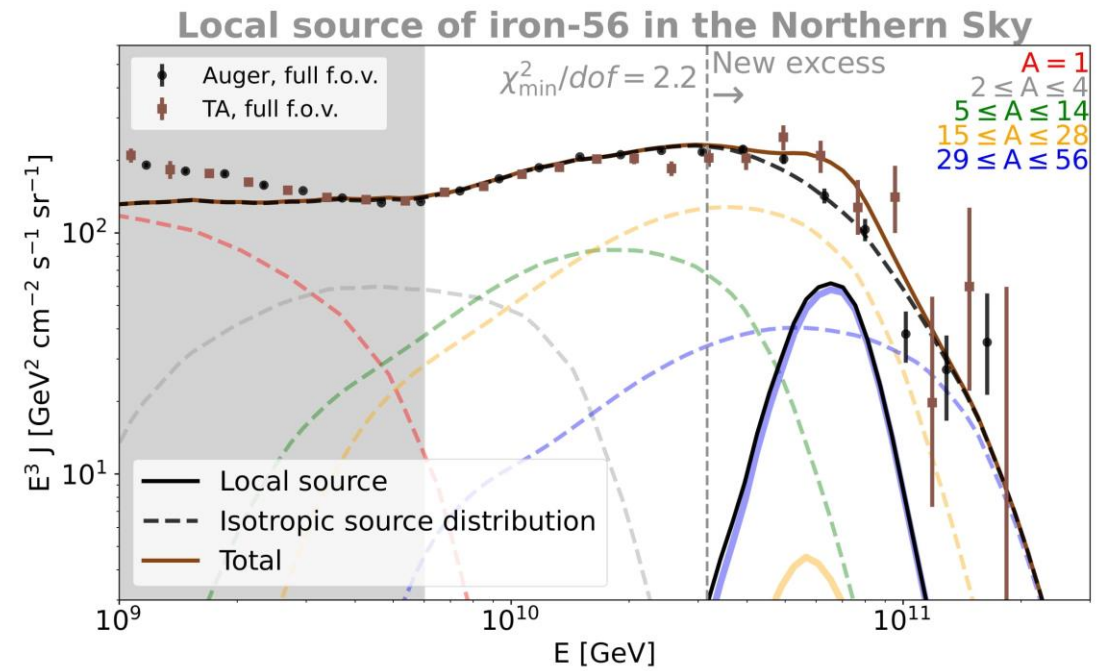
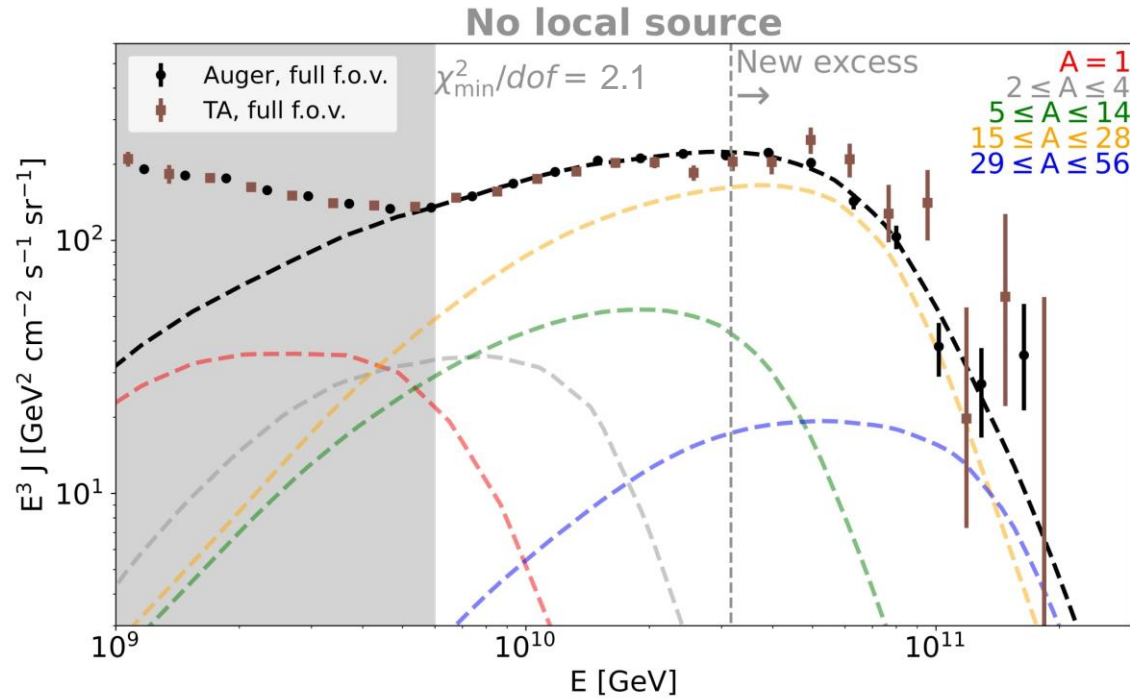
1. “Which Is Better To Launch Your First Mobile App: Android Or iOS?” by Abhijeet Nirmal (<https://www.volumetree.com/which-is-better-for-your-first-mobile-app-to-launch-on-android-or-ios/> )
2. “DC VS. MARVEL: WHO HAS THE HIGHER TOMATOMETER?“ by Mark Hofmeyer (<https://editorial.rottentomatoes.com/article/dc-vs-marvel-who-has-the-higher-tomatometer/> )
3. “Star Wars vs. Star Trek” by Otis Adams (<https://vocal.media/geeks/star-wars-vs-star-trek>)
4. “Spider-Man” 2022 by Sony Pictures Releasing
5. <https://thechoiceisyours.whatisthematrix.com> by WARNER BROS
6. “Shrek” 2001 by DreamWorks Pictures

All images are used for non-profit and educational reasons

# Backup slides



# Energy-dependent shift



# Local source

## Best fit case

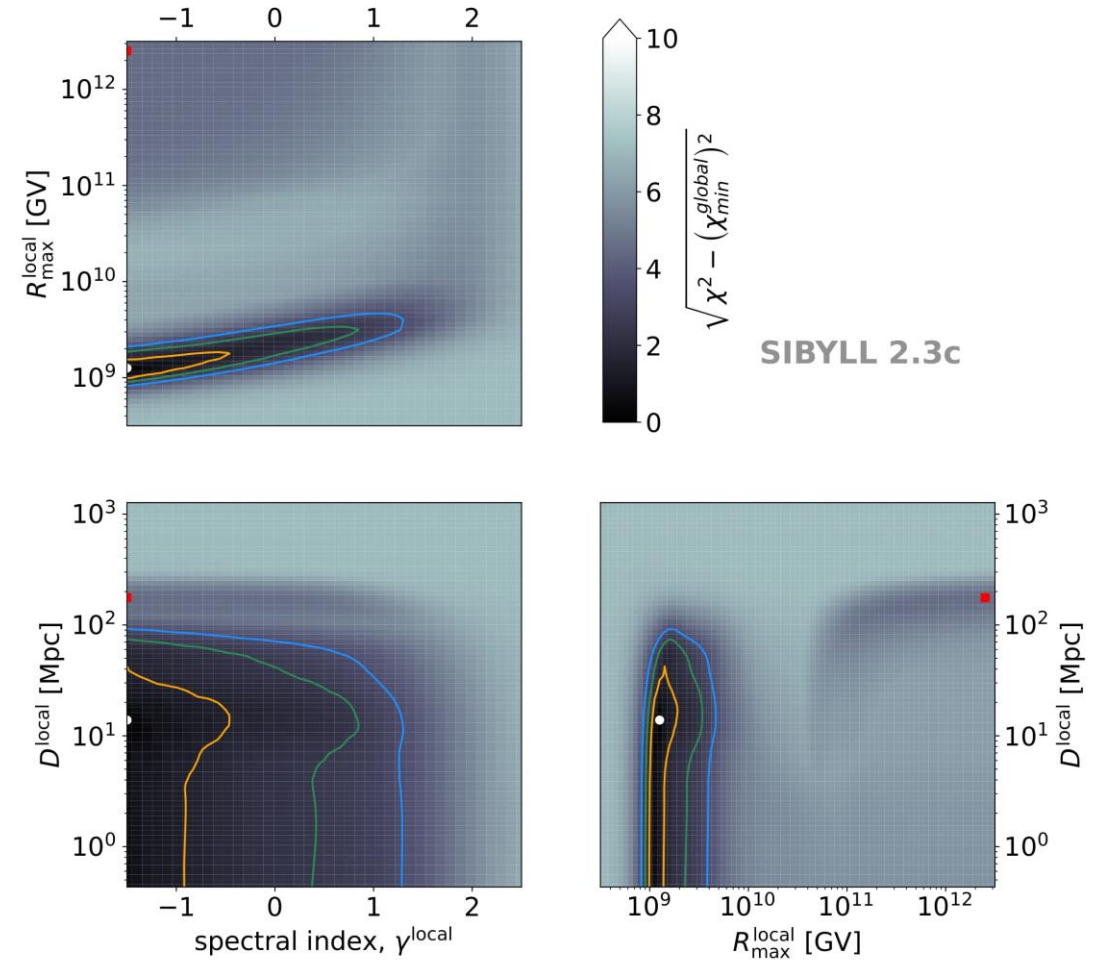
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(silicon-28)

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Luminosity  $7.3^{+18.0}_{-7.3} \cdot 10^{41}$  erg/s

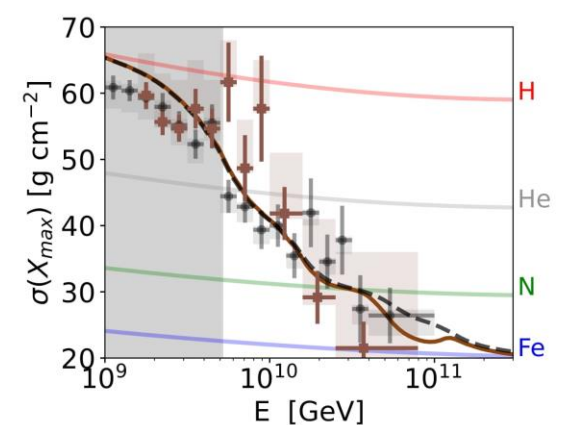
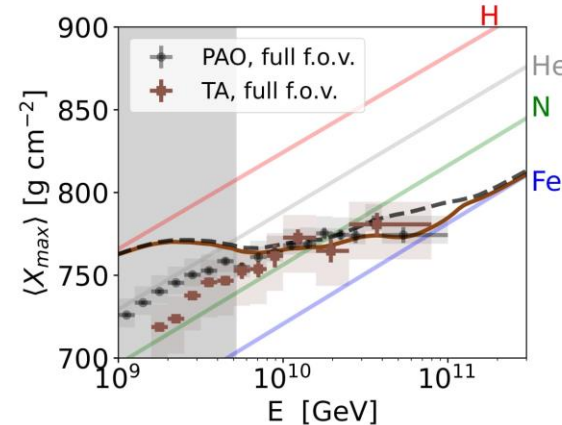
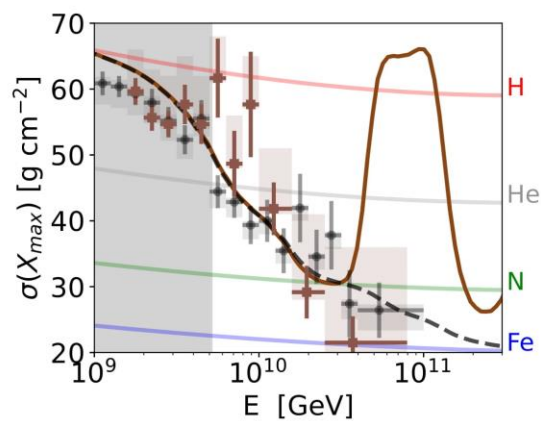
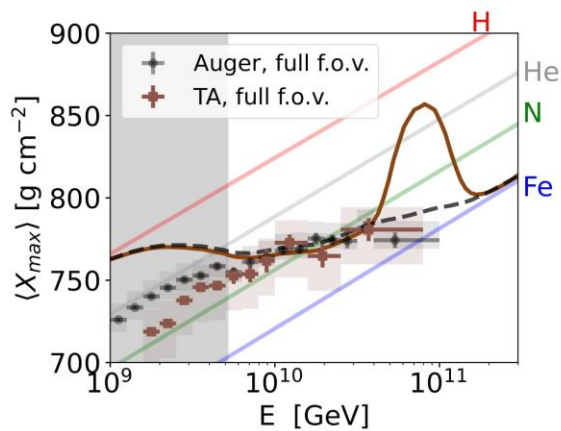
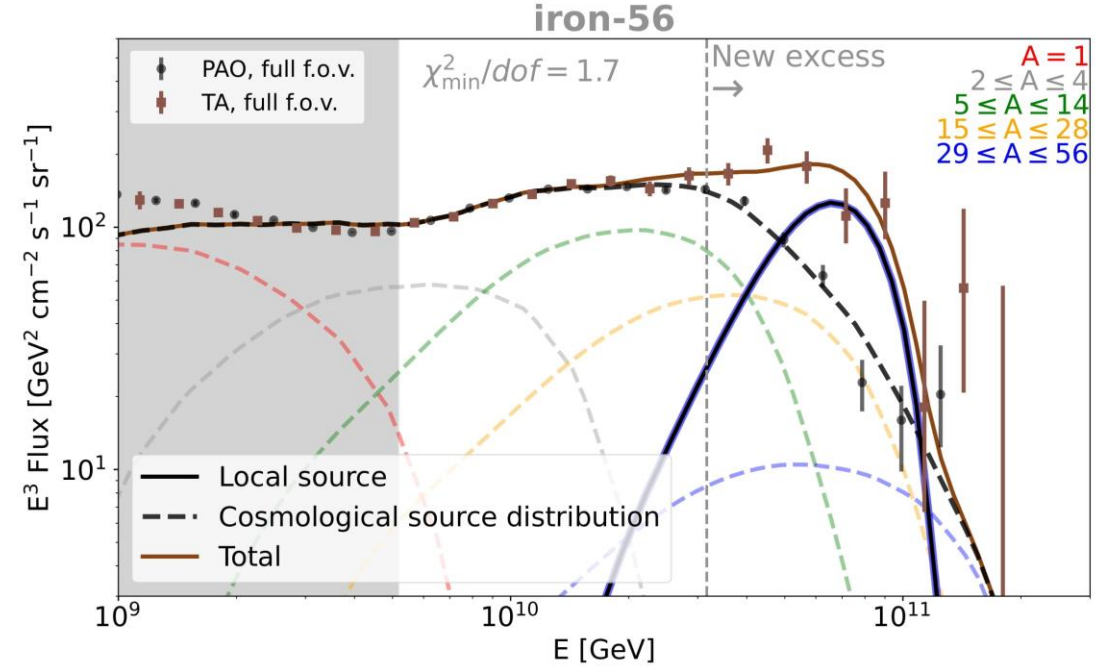
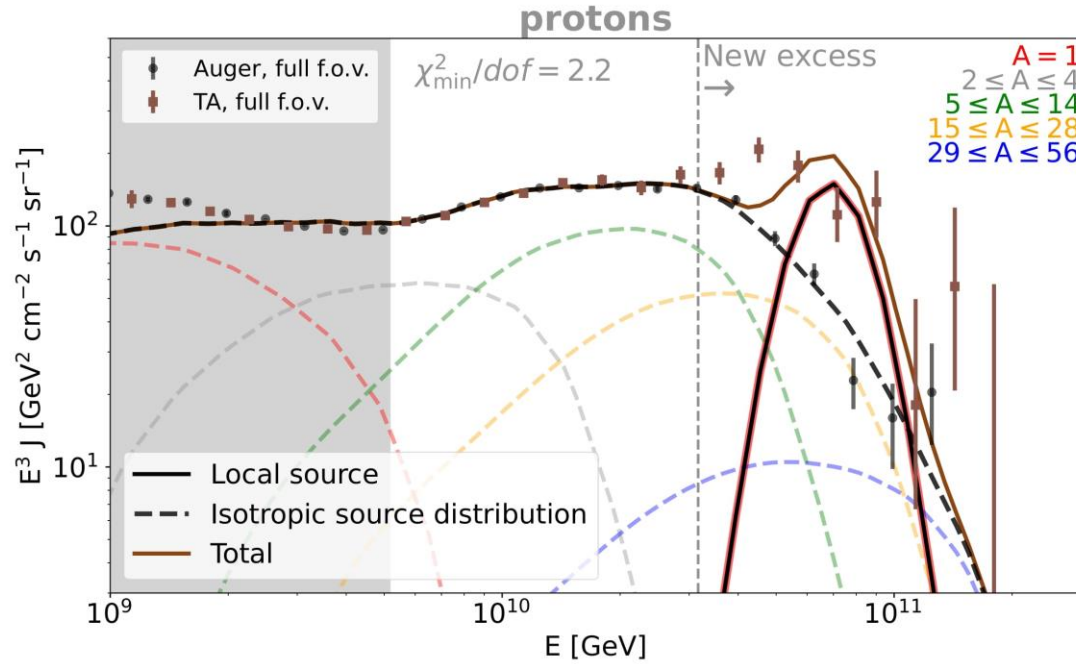
The maximum rigidity  $1.3^{+0.2}_{-0.1} \cdot 10^9$  GV

Local source of silicon-28 in the Northern Sky



# Results

## Astrophysical origin





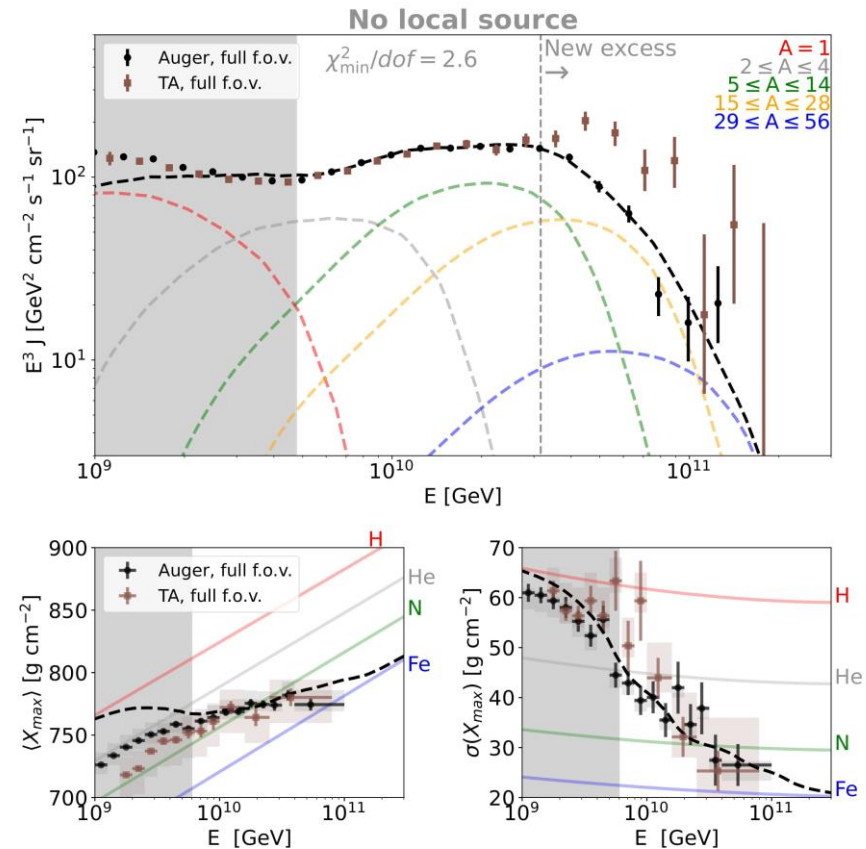
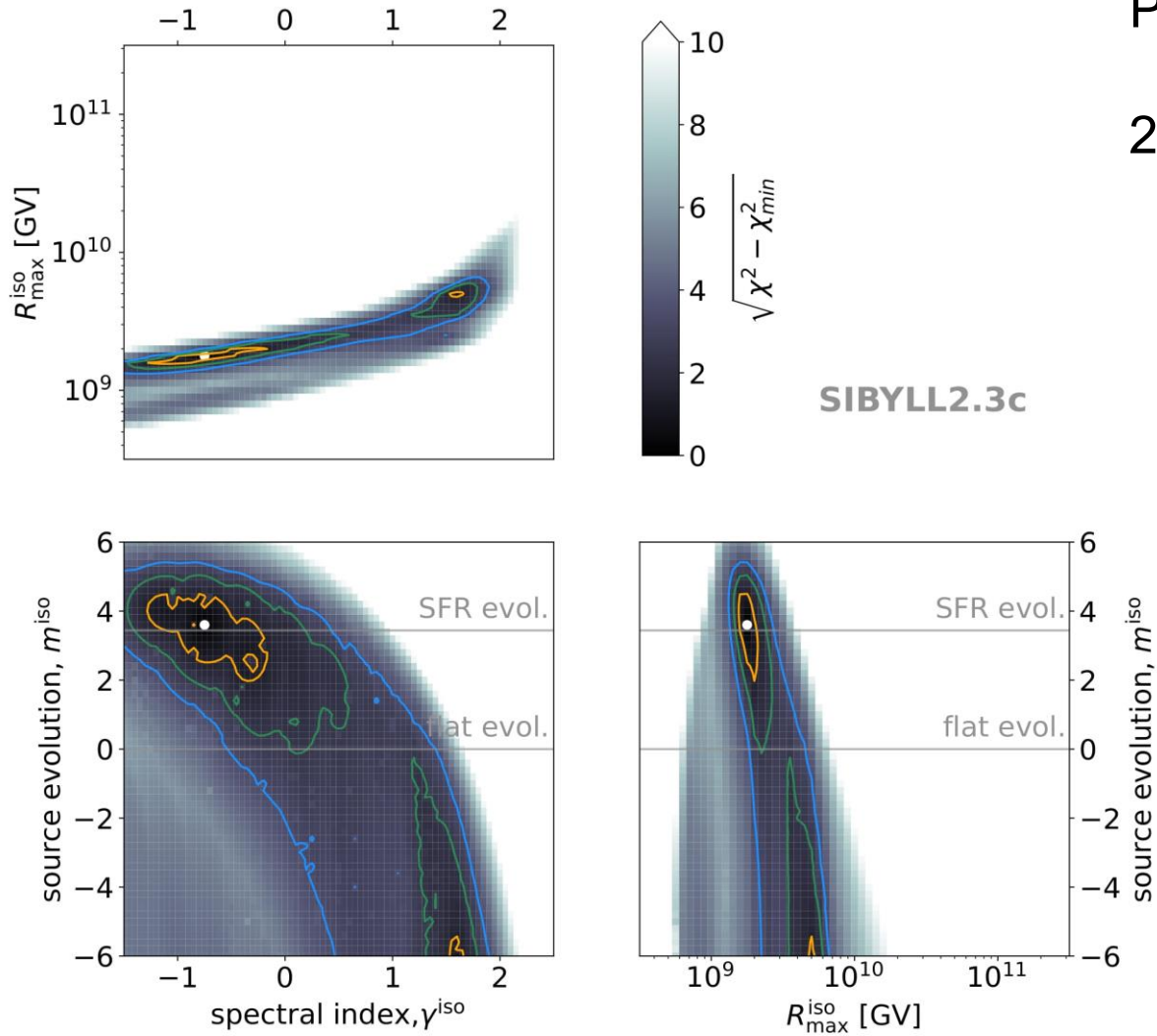
# Results

No Local source

Energy-independent shift

Parameters are similar for all cases

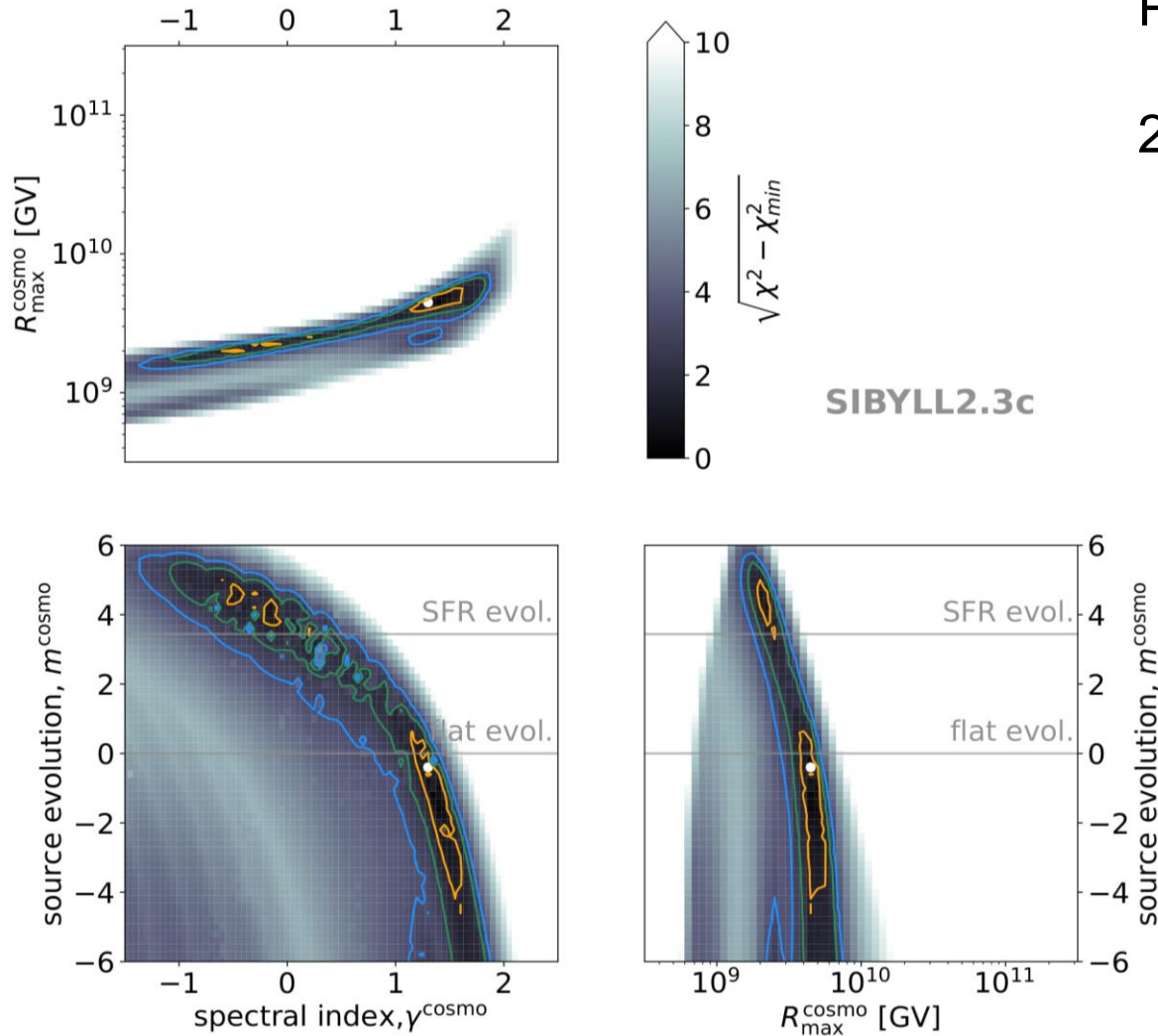
2 local minimum: "PAO" and "TA"



# Results

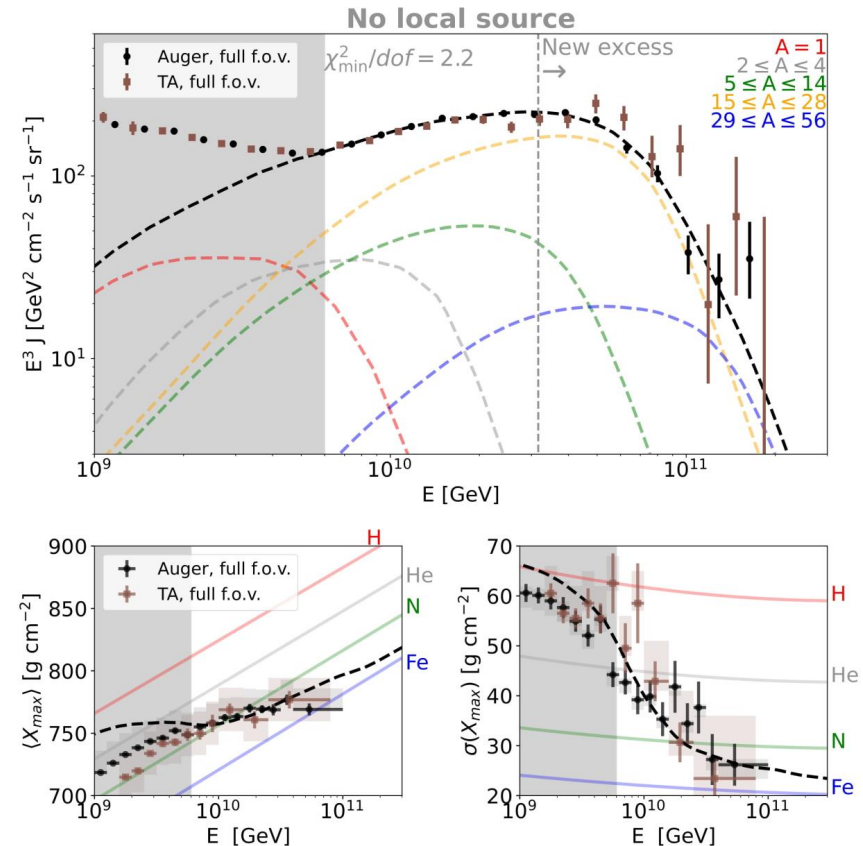
## Systematics

### Energy-dependent shift



Parameters are similar for all cases

2 local minimum: “PAO” and “TA”





# Sources

## Spectrum of injected CR

Choices following *Auger Combined Fit*.  
Simple *Power-law* with *rigidity-dependent* cut-off

$$J_{source\_A}(E) = J_A f_{cut}(E, Z_A, R_{max}) \left( \frac{E}{10^9 \text{GeV}} \right)^{-\gamma}$$

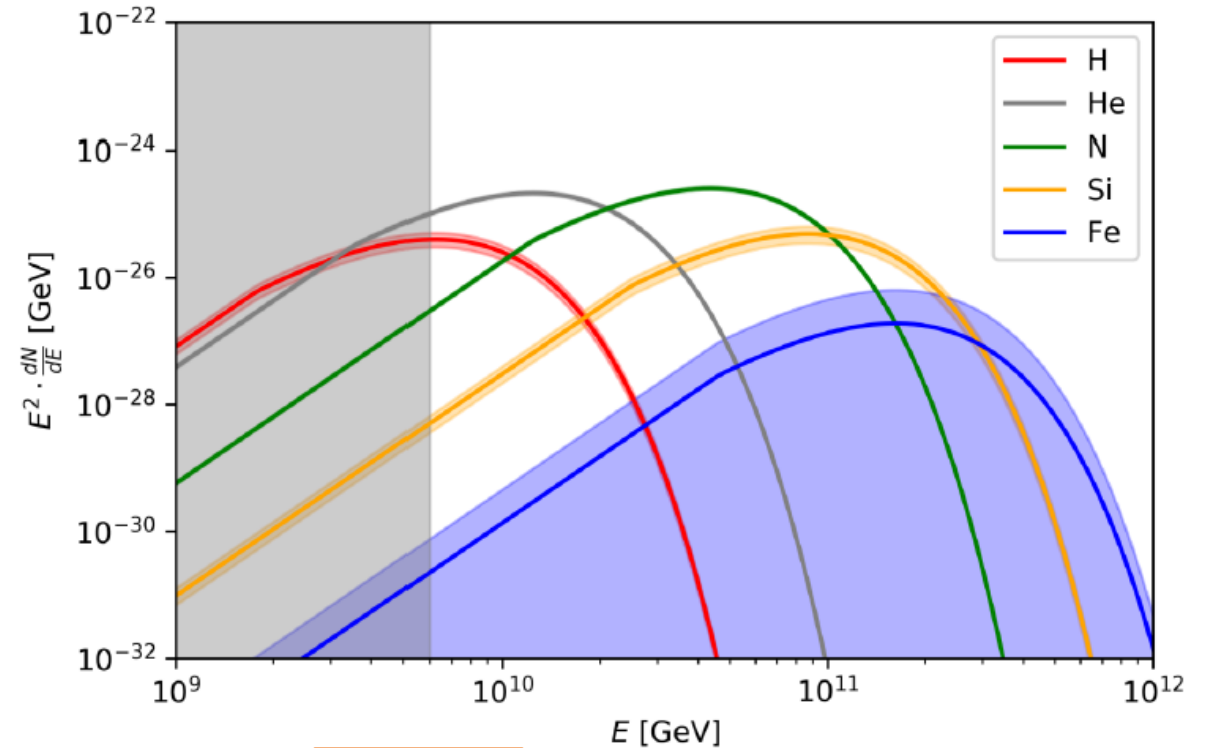
$$f_{cut}(E) = \begin{cases} 1 & , E < Z_A R_{max} \\ \exp\left(1 - \frac{E}{Z_A R_{max}}\right) & , E > Z_A R_{max} \end{cases}$$

Source evolution locally as

$$n_{evol}(z) = (1+z)^m \quad z < 1$$
$$n_{evol}(z) = \delta(z)$$

Only five injection elements:  
*H, He, N, Si, Fe*

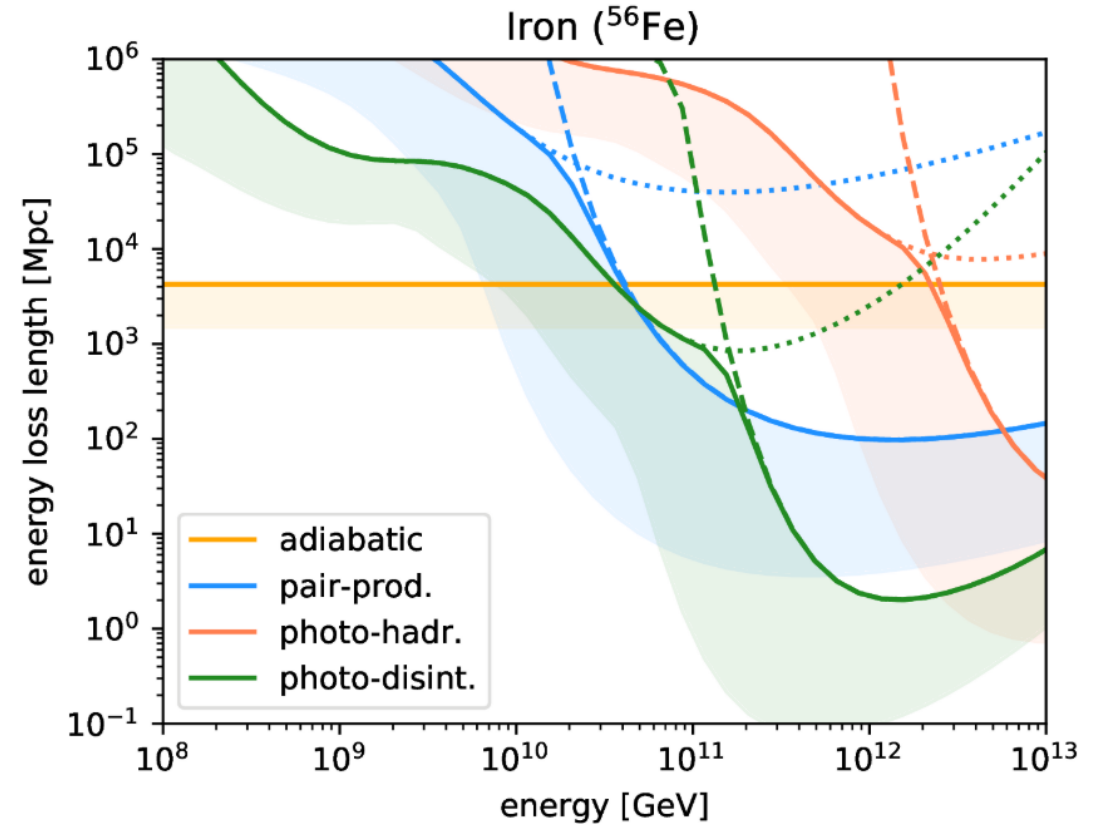
3 main parameters:  
*m(z), gamma, R*



# Propagation

## Transport equation

- About **50 coupled** differential equations
- Speed: **20s – 120s** for single spectrum  
(depending on number of system species)



$$\partial_t Y_i(E, z) = + \underbrace{\partial_E (H E Y_i)}_{\text{adiabatic cooling}} - \underbrace{\partial_E \left( \frac{dE}{dt} Y_i \right)}_{\text{pair - production}} - \underbrace{\Gamma_i Y_i + \sum_j Q_{j \rightarrow i}}_{\substack{\text{photo-hadronic} \\ \text{Photo-disintegration}}} + \underbrace{\mathcal{L}_i}_{\text{Injection}}$$

Jonas Heinze at TeVPA 2018;