LXe eXplore the invisible

XENON - DARWIN Local Group Activities



Cecilia Ferrari - GSSI Science Fair - 14/02/2022



The invisible matters!

We know that we don't know.

- Is it matter?
- How does it interact?
- How much does it weigh?
- Can we directly see it?

But it is for the unknowns that we explore.

- Assume dark matter exists.
- Suppose it "weakly" interacts.
- If m_{DM} > 1 GeV, We may directly see it.





Our invisible eXplorer





We may directly detect DM particles belonging to the Milky Way Halo wind.

- Dual-phase (liquid+gas)
- Energy reconstruction
- 3D event reconstruction
- Fiducialization
- Event discrimination (electronic recoil vs nuclear recoil)









A powerful technique



- <u>arXiv:1705.06655</u> First Dark Matter Search Results from the XENONIT Experiment
- <u>arXiv:1805.12562</u> Dark Matter Search Results from a One Tonne×Year Exposure of XENON1T
- <u>arXiv:1811.12482</u> First results on the scalar WIMP-pion coupling, using the XENON1T experiment
- <u>arXiv:1902.03234</u> Constraining the Spin-Dependent WIMP-Nucleon Cross Sections with XENON1T
- <u>arXiv:1904.11002</u> First observation of two-neutrino double electron capture in 124Xe with XENONIT
- <u>arXiv:1907.11485</u> Light Dark Matter Search with Ionization Signals in XENONIT
- <u>arXiv:1907.12771</u> Search for Light Dark Matter Interactions Enhanced by the Migdal effect or Bremsstrahlung in XENON1T
- <u>arXiv:2006.09721</u> Excess Electronic Recoil Events in XENONIT
- **arXiv:2011.10431** Search for inelastic scattering of WIMP dark matter in XENON1T
- <u>arXiv:2012.02846</u> Search for coherent elastic scattering of solar 8B neutrinos in the XENONIT dark matter experiment

XENONIT Results



Beyond DM Searches





A long story with a sequel



Past Present Image: I

XENON10	XENON100	XENONIT	XENONnT	DARWIN
2005 - 2007	2008 - 2016	2012 - 2018	2019 - 2023	2025
15 x 20 cm	30 x 30 cm	lxlm	1.5 x 1.3 m	2.6 x 2.6 m
14 kg	62 kg	2 tons	5.9 tons	40 tons
~10 ⁻⁴³ cm ²	~10 ⁻⁴⁵ cm ²	~10 ⁻⁴⁷ cm ²	~10 ⁻⁴⁸ cm ²	~10 ⁻⁴⁹ cm ²

Future



Our team of eXplorers



Subateci

C. Ferrari - eXplore the invisible

GSSI Science Fair 2022

NYUAD



The local group



- Walter Fulgione (LNGS Group responsible)
- Marcello Messina (LNGS XENONnT Technical Coordinator)
- Alfredo D. Ferella (UnivAQ Professor)
- **Carla Macolino** (UnivAQ Professor)
- Valerio D'Andrea (UnivAQ PostDoc)
- Riccardo Biondi (LNGS PostDoc)
- Cecilia Ferrari (GSSI PhD Student)
- Andrea Melchiorre (UnivAQ Master Student)

A little group embracing three different institutions.









- Realization and Installation of new Electrodes for the XENONnT TPC
- MugPump for GXe purification









Local group activities on XENONnT - S





Local group activities on DARWIN





- The **ABALONE photosensor** is a modern, scalable, cost-effective, robust alternative to PMT and SiPM
- High radiopurity due to absence of metal
- **High detection performance**: UV sensitivity, low afterpulsing rate, sub-ns timing resolution
- Sensor like G-APD needed to detect photon signal from scintillator

R&D at LNGS and at Stockholm University



Activities on-going!

- Setup optimization
- Data analysis for detector characterization
- GEANT4 simulations
- Characterization in LXe



Join the group!



XENONNT SRO is just finished and data will be unblinded soon. But a **new science run** is coming, and you could be one of the few people so lucky to put their hand on this scientific treasure!

And also, if you are not afraid to get your hand dirty with a little mechanic, electronics, hydraulic ecc... You will have a lot o fun in joining us!





So, if you want to know more about Xenon and Darwin, just **contact us**!!

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It is only in the <u>heart of the</u> <u>mountain</u> that one can see rightly; What is <u>fundamental in physics</u> is invisible to the eye.

Edited from Antoine de Saint-Exupéry, <u>The Little Prince</u>

