

The HERD space mission: Detecting Galactic Cosmic Rays at the Highest Energies

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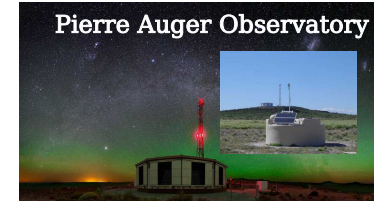
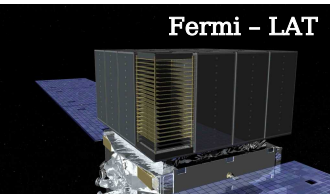
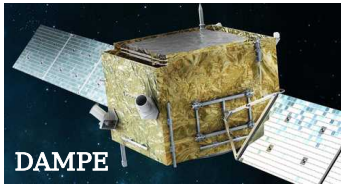
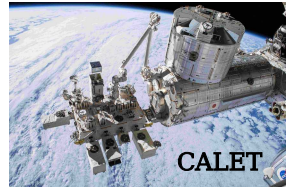
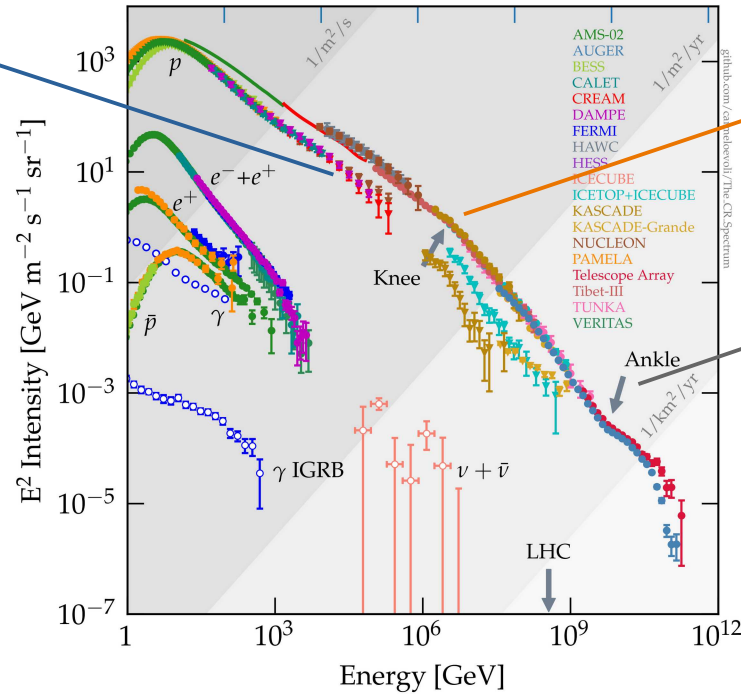
Gran Sasso Science Institute (GSSI) & INFN-LNGS

The Cosmic Ray Landscape

Energies achieved with current space-borne direct CR experiments (~ hundred TeV)

Maximal energies achieved with direct detection CR experiments (~ PeV)

Region covered by indirect CR experiments (~ 10^{20} eV)



Direct CR experiments

- Precise measurement of particle charge and energy
- Small exposure to provide statistically meaningful measurements above few tens of TeV.

Indirect CR experiments

- Huge achievable energies
- Difficulty in making composition studies with small systematics



Experiments with **large acceptances**, operating over **several years** are needed to explore CR spectra at **PeV energies**.

Main scientific objectives

Cosmic Rays: Precise spectra and mass composition up to PeV

Gamma – ray astronomy and transient studies

Electron spectra (and anisotropy) up to tens of TeV

Indirect **Dark Matter searches** with high sensitivity

D. Kyratzis, Il Nuovo Cimento 43C (2020) 117

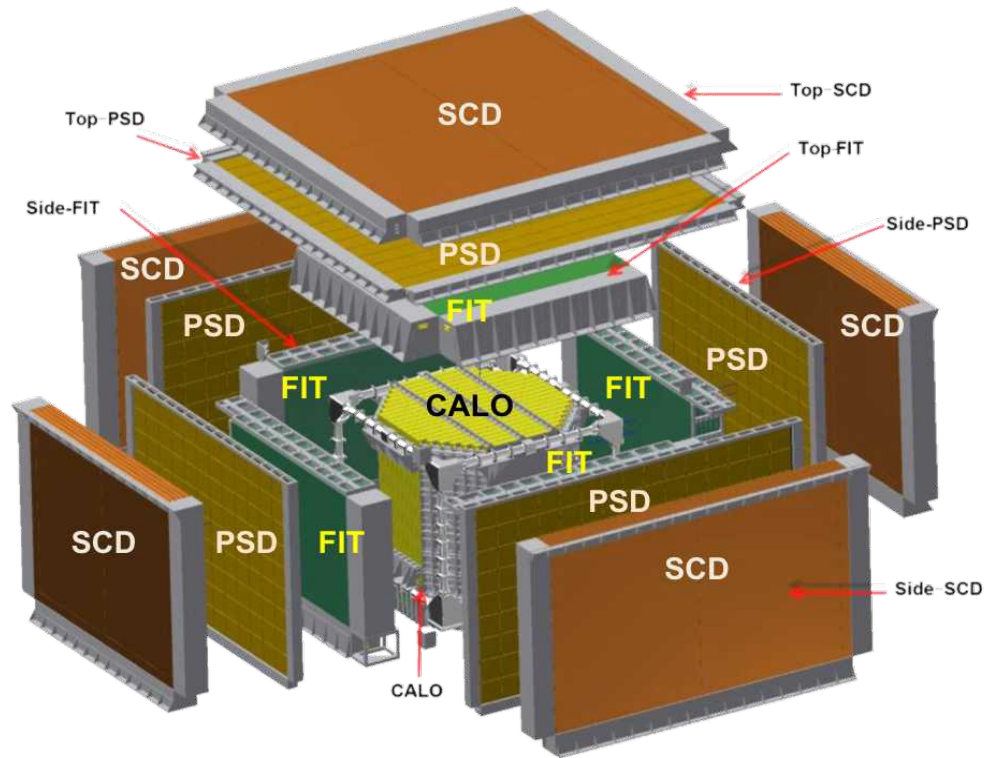
The Collaboration

International synergy between Chinese, Italian, Swiss & Spanish institutes.



HERD: Detector Description

Based on previous experience with AMS-02, FERMI & DAMPE missions



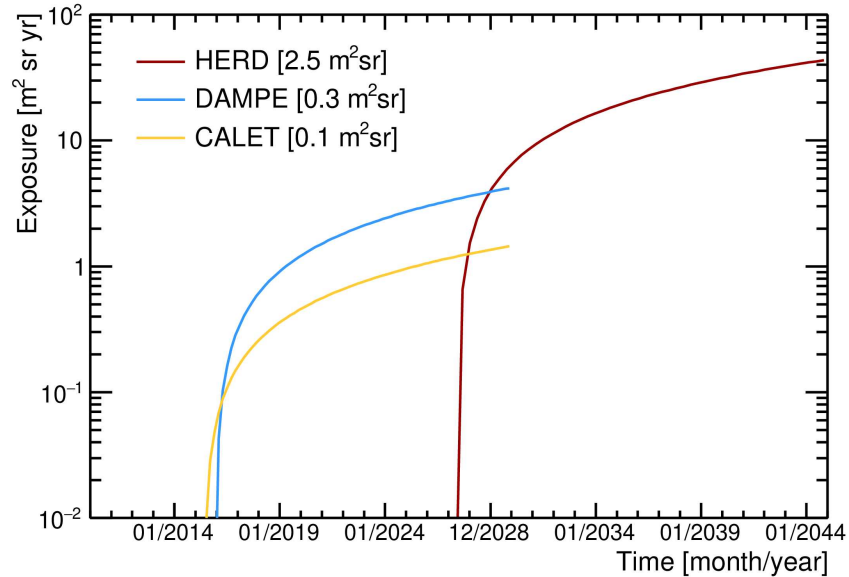
A deep ($\sim 55 X_0, 3 \lambda_I$) 3D cubic calorimeter (CALO), octagonal prism, accurately measuring deposited energy + e/p separation

Fiber Tracker (FIT), determining tracks of impinging particles.

Plastic Scintillator Detector (PSD), providing gamma-ray and charged particle triggers + charge measurement.

Silicon Charge Detector (SCD), ensuring an additional charge measurement

plus **A Transition Radiation Detector (TRD)**, placed on one of the lateral faces, providing energy calibration of nuclei (TeV region).

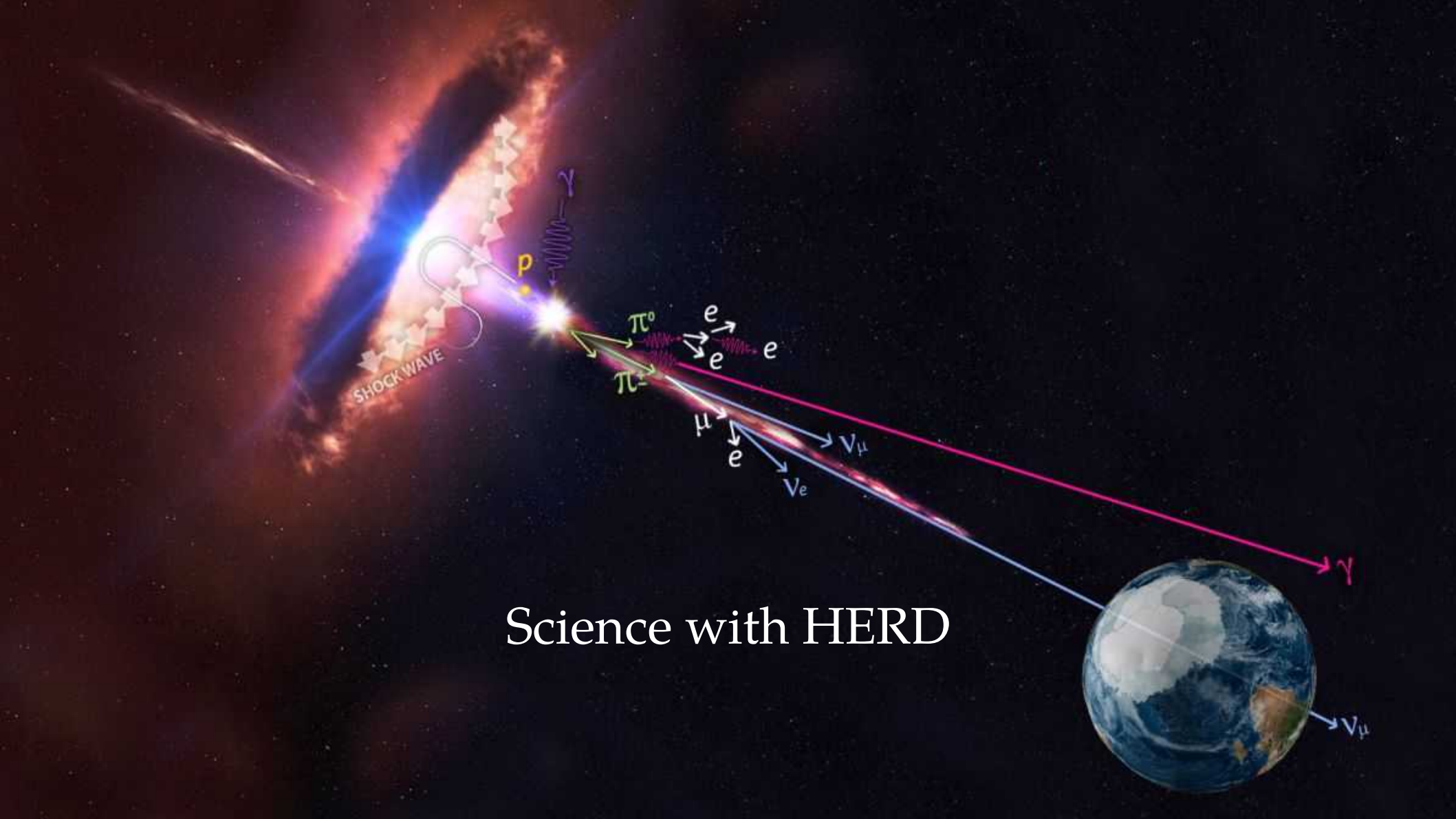


One order of magnitude upgrade in exposure wrt to current gen CR experiments:

15 – 20 $\text{m}^2 \text{sr yr}$

Main Requirements

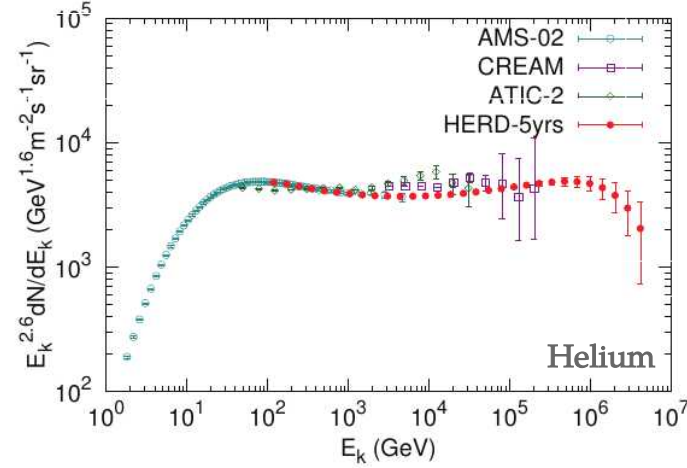
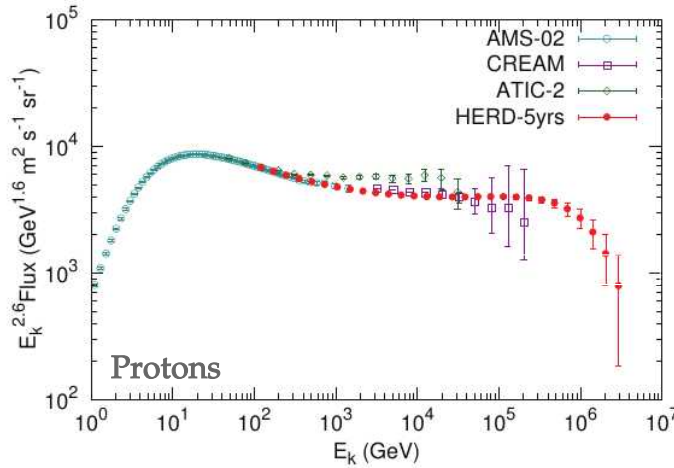
Energy range (e/ γ)	10 GeV-10s TeV(e); > 100 MeV (γ)
Energy range (CRs)	30 GeV – 3 PeV
Angular resolution	0.1 deg. @ 10 GeV
Energy resolution (e/ γ)	1-2% @ 200 GeV
Energy resolution (p)	20-30% @100 GeV – PeV
e/p separation	$\sim 10^{-6}$
Geometric Factor (e)	>3 $\text{m}^2 \text{sr}$ @ 200 GeV
Geometric Factor (p)	>2 $\text{m}^2 \text{sr}$ @ 100 TeV
Pointing	Zenith
Field of View	+/-70 deg
Mass	< 4 tons
Lifetime	~ 10 years



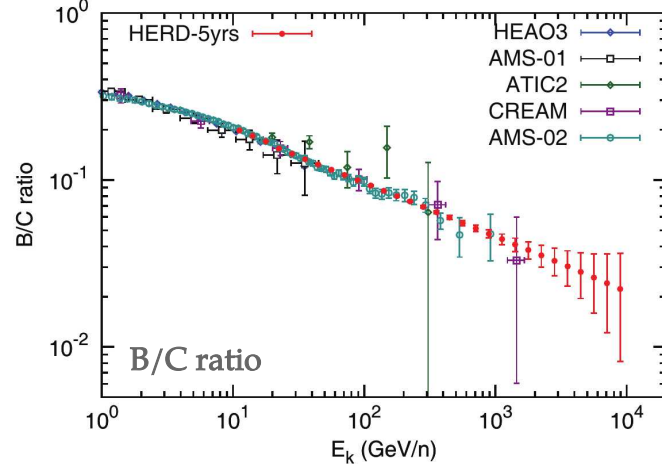
Science with HERD

Cosmic Ray Nucleonic Spectra

Pronounced features such as the “knee” in the CR spectrum can be examined via **direct p & He measurements**



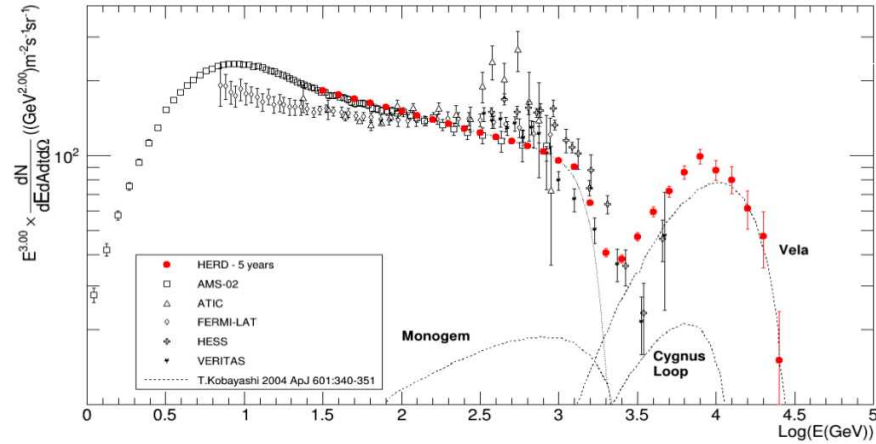
Measuring CR fluxes from protons to iron with direct observations up to the highest energies



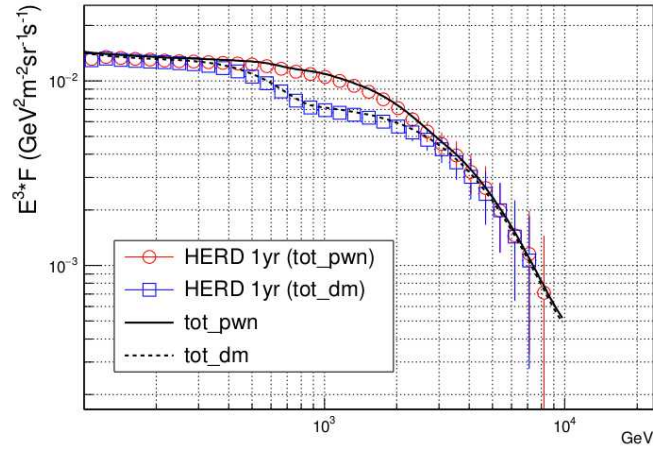
Extending the fluxes of medium mass (BCNO group) to their highest achievable limit with direct measurements

Cosmic Ray Electrons & Gamma – Rays

Expected flux after 5 years of operation



Astrophysical or Dark Matter sources?



HERD will be able to accurately determine the all – electron flux up to several tens of TeV

Multimessenger Era

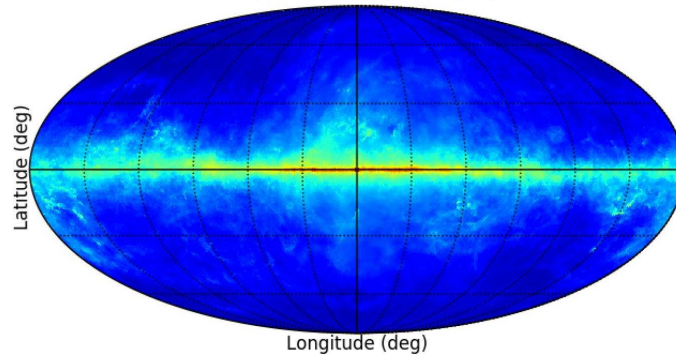
Possible synergy w/ ground-based experiments:

Gamma – rays (CTA, LHAASO)

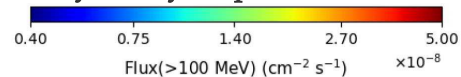
Neutrinos (KM3NeT, IceCube)

Gravitational Waves (Ligo, Virgo)

HERD 5 Years, TS=25, > 10 photons/bin, 4 bin/dec



5-year sky-map above 100 MeV



Owing to its large acceptance & sensitivity, HERD will be able to conduct full gamma-ray sky surveys (> 100 MeV)

L. Farina et al, PoS, ICRC2021(2021) 651

GSSI Group Activities

GSSI is heavily involved (in conjunction with other institutes) in the realization of the HERD PSD, via hardware R&D and MC simulations.



The PSD will operate as **anti-coincidence** providing **charge measurements** of incoming CR nuclei in a range of $Z = 1 - 26$.

Main requirements: **high detection efficiency**, **broad dynamic range** & **good energy resolution**.

Previous, current and upcoming activities concern:

Hardware

Preparation, assembly & performance validation of scintillator bars coupled w/ SiPMs + tested with a multitude of particles and beams (configured at GSSI & LNGS)

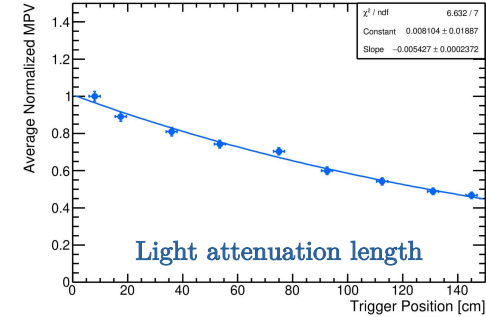
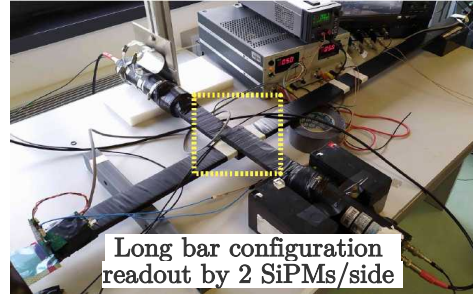
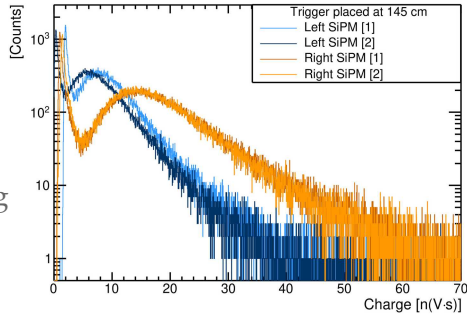
Software

Dedicated GEANT4 - based simulation evaluating performances of various PSD configurations tested in the lab along with their inherent properties

Few results...

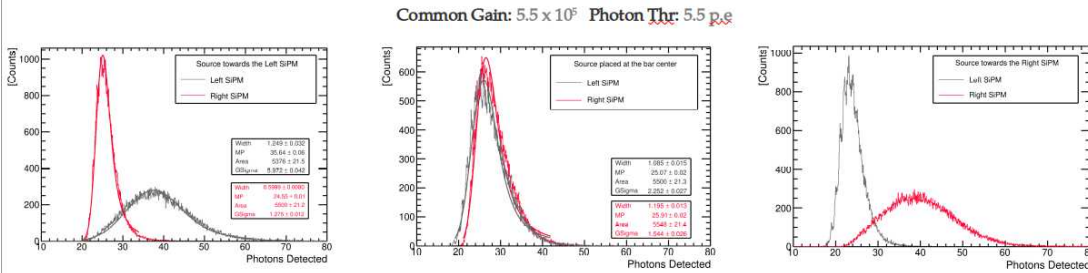
Two scintillator geometries are currently investigated, one based on long bars while the other on square tiles.

Cosmic ray muon charge distributions measured in various trigger positions along the bar



Measurement of light attenuation length extracted from all trigger positions

D. Kyratzis et al, PoS, ICRC2021(2021) 054

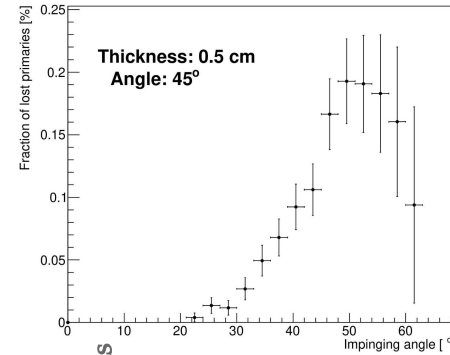


Measurements taken w/ ⁹⁰Sr in 3 positions along the bar

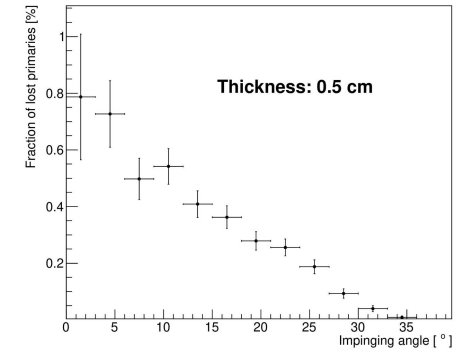


Radioactive source measurements, scanning the PSD configuration along various bar positions

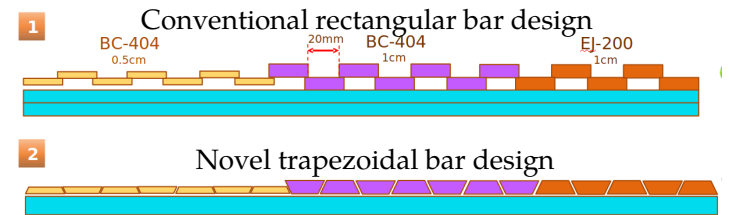
Trapezoidal Bars on GEANT4



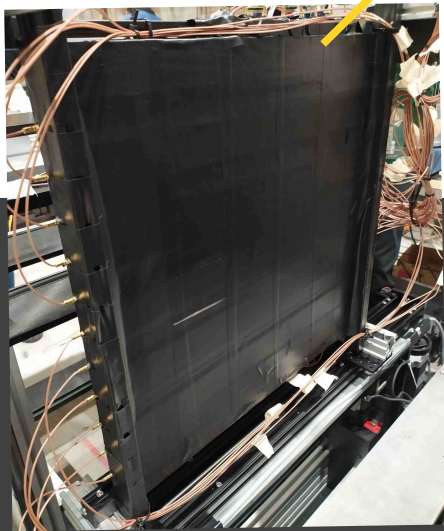
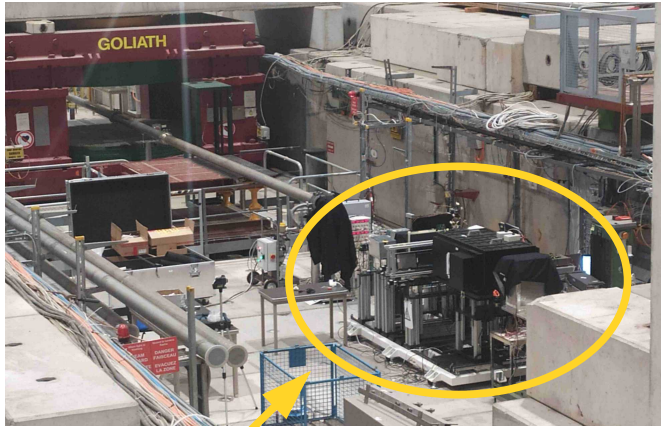
Rectangular Bars on GEANT4



Hermeticity studies

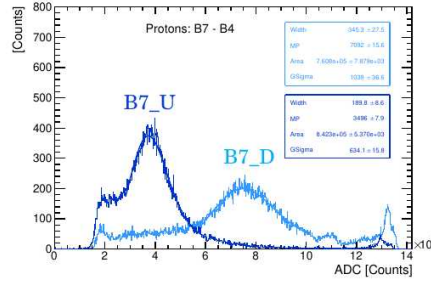
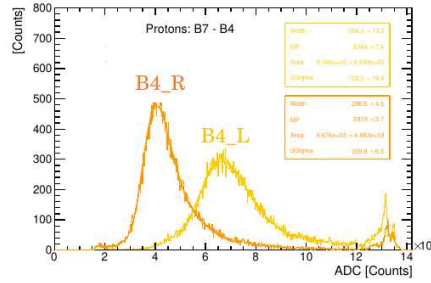


CERN SPS & PS campaigns

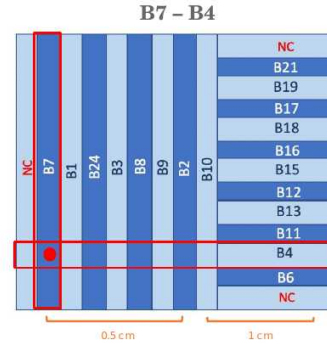
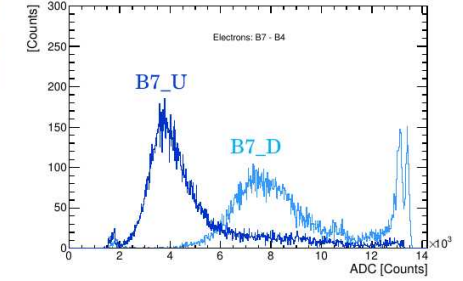
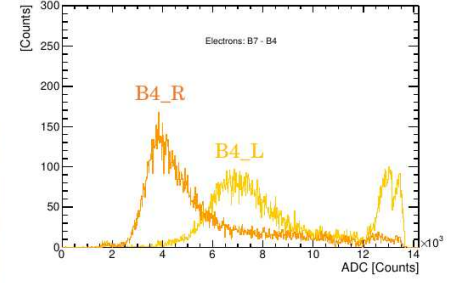


Installation @ CERN - SPS

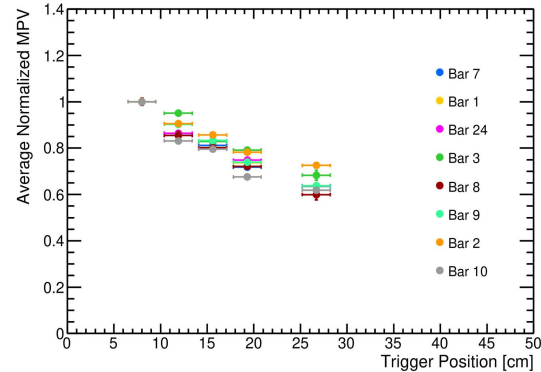
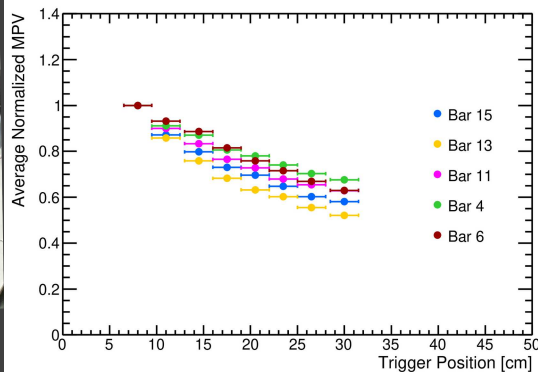
Protons



Electrons



p & e⁻ beam results on specific bars



Evaluation of average MPV behaviour regarding 13 tested bars in various trigger positions

HERD is a novel space – borne detector, to be installed on–board **China's Space Station (CSS)** around 2027, with an **expected lifetime of ~10 years**.

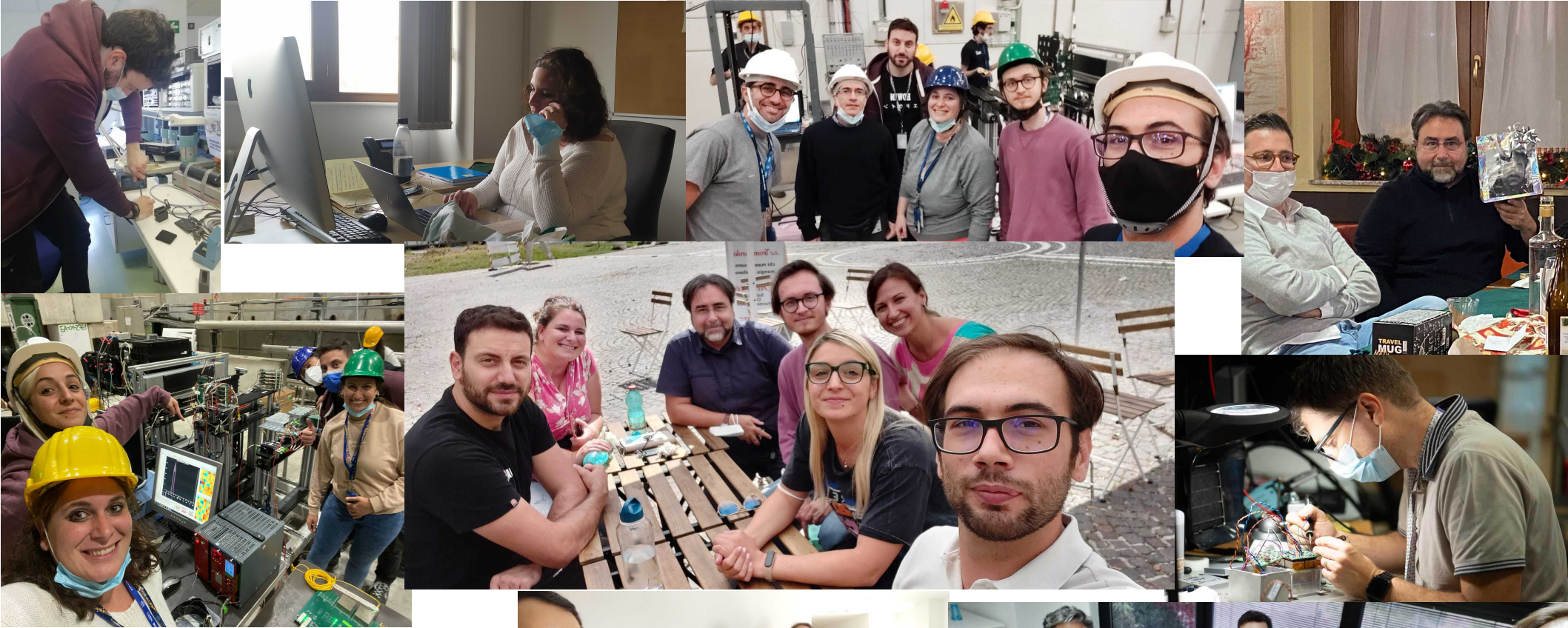
Main scientific objectives revolving around **high energy galactic CRs, gamma – ray astronomy** and indirect **DM searches**.

Novel detector, comprising a deep **3-D calorimeter (CALO)**, a **Fiber Tracker (FIT)**, a **Plastic Scintillator Detector (PSD)** and a **Silicon Charge Detector (SCD)** instrumented on both the **top and lateral sides**, along with a **Transition Radiation Detector (TRD)**.

Coupling **state-of-the-art detector techniques** with a pioneering design in HERD, an **order of magnitude increase in acceptance** can be attained, considering previous & current missions.

Ongoing and future work in HERD foresees continuous development efforts manifested through dedicated **hardware R&D** and **MC simulations** in order to verify its performances.

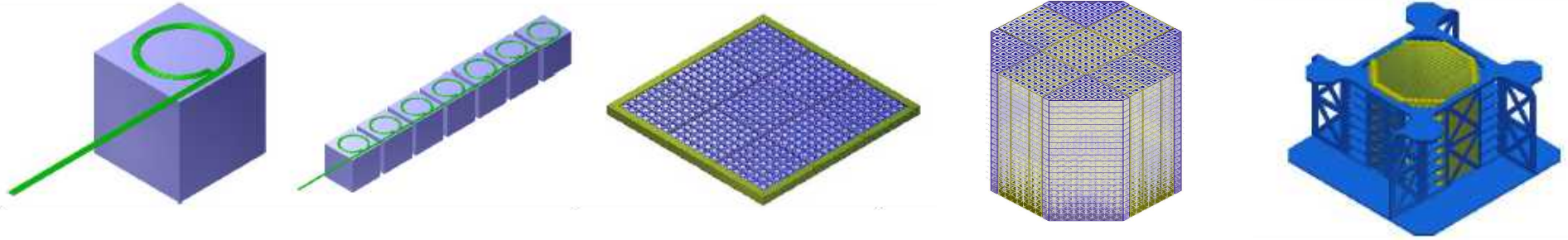




Greetings from the
GSSI Space Group!



HERD sub-detectors: Calorimeter (CALO)



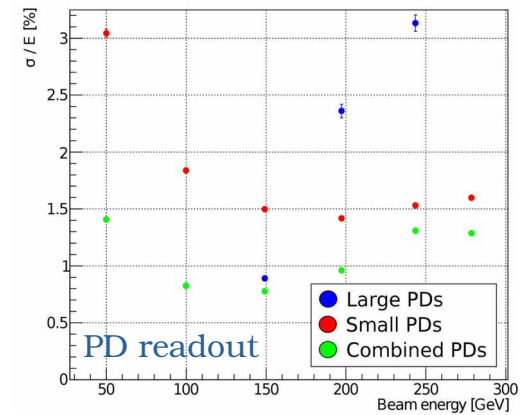
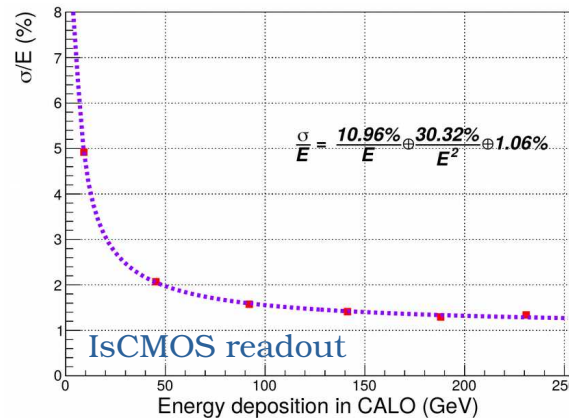
Item	Value
Type of crystal	LYSO
Nuclear interaction length	3 (55 X_0)
Number of crystals	~7500
Crystal dimension	$3 \times 3 \times 3 \text{ cm}^3$

Scintillation light is readout independently by:

- 1) WLS fibers coupled to IsCMOS cameras
- 2) Photodiodes connected to custom frontend electronics

Partial readout of crystals with PhotoDiodes (Calocube) for calibration extended dynamic range & reduced systematics.

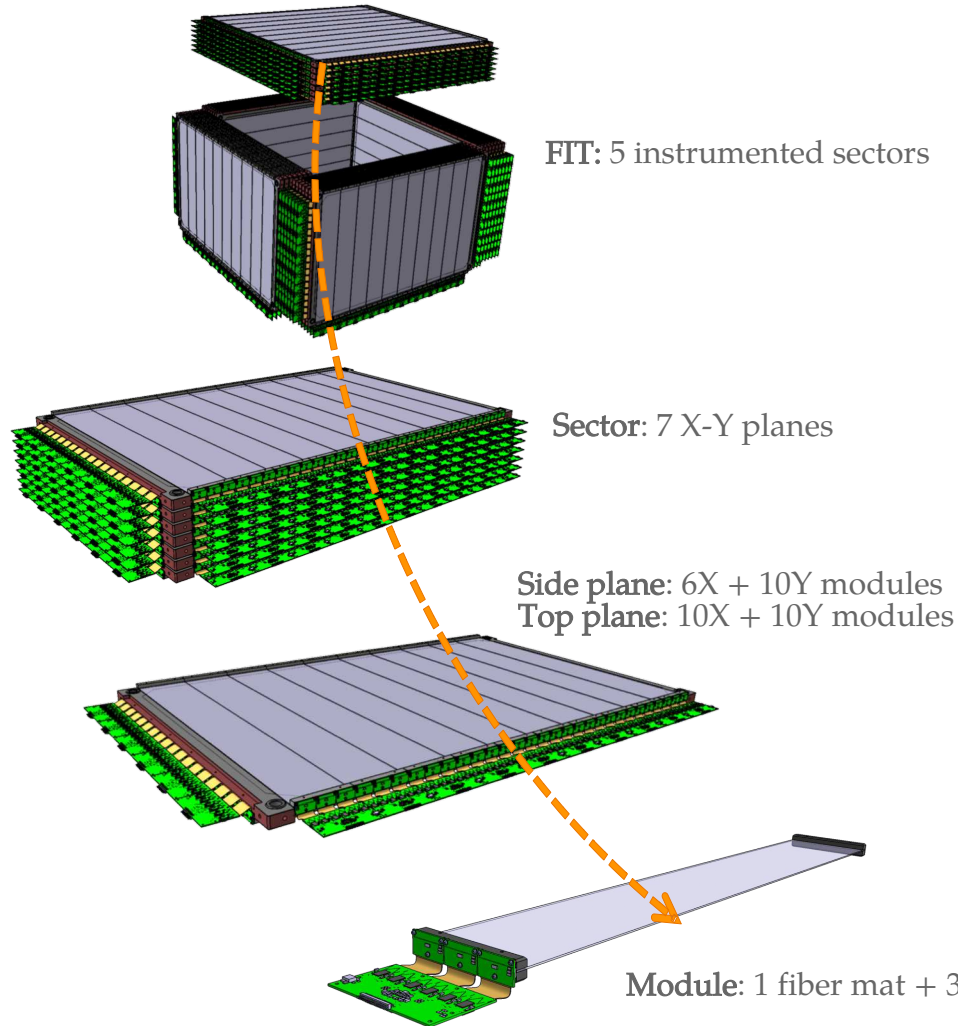
Energy resolution for electrons



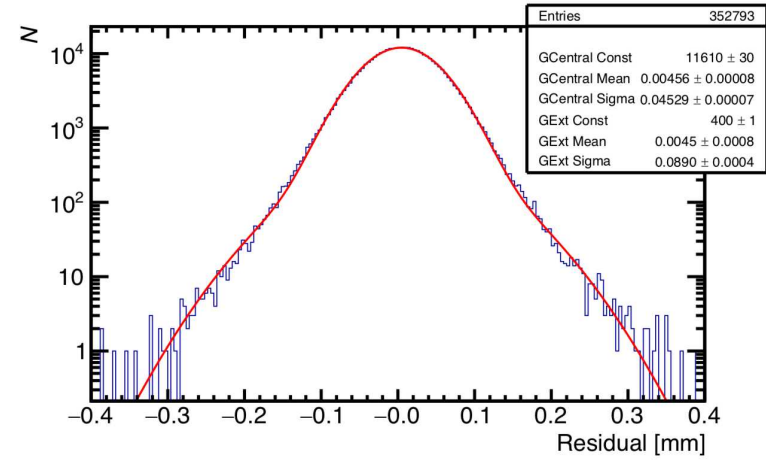
...from beam tests at CERN – SPS

L. Pacini et al, PoS, ICRC2021(2021) 066

HERD sub-detectors: Fiber Tracker (FIT)



Position residual distribution from proton beam tests



Spatial resolution = $(45.0 \pm 0.1) \mu\text{m}$

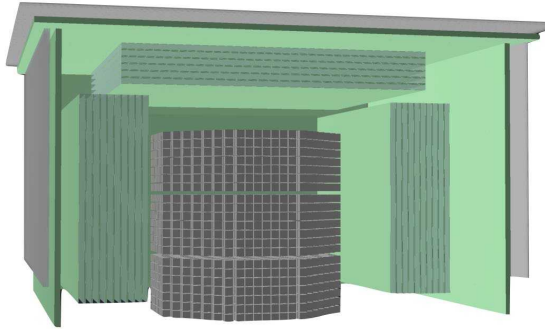


Z	μ_z	σ_z	σ_z/μ_z [%]
2	1.99	0.31	15
3	3.07	0.4	13
4	4.01	0.51	12

Charge resolution for nuclei heavier than protons

HERD sub-detectors: Silicon Charge Detector (SCD)

The SCD is a **silicon micro-strip** detector with the objective of precisely measuring the particle charge



Being the **outermost** detector the SCD aims to avoid early charge – modifying interactions in the PSD

Highly segmented to minimize backplash events moving upward from the CALO

