

35th cycle, PhD in Astroparticle Physics - 3rd year admission

Neutron flux measurement at LNGS with CYGNO/INITIUM

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Outline

CYGNO experiment and the neutron background

Nuclear recoils studies

Underground background studies

Conclusions and future perspectives

G S S I The CYGNO/INITIUM project and neutron background in underground experiments



The CYGNO/INITIUM project

- CYGNO: high resolution Time Projection Chamber with optical readout for directional dark matter searches
- He:CF₄ 60/40 gas mixture at atmospheric pressure, triple GEM amplification, sCMOS cameras + PMT
- \rightarrow 3D tracking capability, ~ 1keV energy threshold
- In sinergy with INITIUM project on negative ion drift with the addition of SF₆
- R&D studies with different prototypes



Neutron background at LNGS

- Rare event searches experiments operate underground (reduction of cosmic ray background 10⁻⁶ muon, 10⁻³ neutron @LNGS)
- **Neutron background** is a major issue for many underground experiments (nuclear astrophysics, neutrinoless double beta decay, direct DM detection)
- Mainly produced by spontaneous fission, (α, n) reactions, radioactivity of the set up
- Neutral particles (neutrons, neutrinos) which interact with nuclei can mimic a WIMP-nucleus scattering
 - Neutrinos cannot be shielded – directionality is key
 - Neutrons can be shielded knowledge of spectrum is paramount for ٠ the experiment design

LIME can provide a spectral, directional measurement of the neutron flux underground at LNGS





WIMP

neutron

neutrino

Neutron background at LNGS

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- The neutron flux measurements done so far:
 - Don't agree with each other
 - Have large binning/limited spectral sensitivity
 - Many techniques based on thermal neutron absorption
- With the LIME gaseous TPC we can:
 - Directly measure the **fast** component through NR detection
 - Enhance sensitivity to **thermal** component with the addition of ³He (no change in gas mixture properties), through ³He(n,p)³H reactions
 - **Directional measurement**
- Moreover, the neutron measurement underground can show the feasibility of the technique for DM detection





The LIME prototype



PMT (×4) sCMOS camera ORCA fusion • • Entries Mean 701 Std Dev 278 13% energy 120 resolution at 5.9 keV 7663 Mean Sigma 1007 10000 4000 6000 8000 12000 Cluster light (photons)

50 liters sensitive volume

• He:CF₄ 60/40 atm.pressure

• Triple GEM amplification

33x33 cm² readout area, 50cm drift

• 1sCMOS camera + 4 PMT



- Already commisioned at LNF
- Tested operation stability,
- Energy resolution at 13% at 5.9keV (⁵⁵Fe) throughout the volume

The LIME prototype







Nuclear recoils: tracks simulation and analysis



Low energy NR ionization

- NR from DM (or neutrons) are expected to be in the range 1-100 keV
- A significant fraction of the energy is lost to non-ionizing processes
- The stopping power decreases: ions lose the majority of their energy at the beginning of their path
- Key features to be studied:
 - Ionization quenching factor $QF(E) = \frac{E_{ionization}}{E}$
 - Ionization profile $\frac{dE_{ionization}}{ds}$
- Relevant for energy calibration, track analysis development, direction and head-tail study







SRIM simulations

- SRIM (Stopping and Range of Ions in Matter) to simulate the passage of ions (NR) in He:CF₄ in 60/40 proportion for different energies and atoms (He, C and F)
- Post processing of SRIM simulations was necessary to retrieve the QF and the ionization profile (SRIM is not optimised for the study of the tracks)
- Main issues:
 - Nuclear energy losses (secondary recoils/cascades) are significant but no explicit description is given by SRIM
 - No 3D ionization energy deposits: from the 3D *total* energy distribution I applied a **conversion factor** (dependent on the QF)

Primary ionization
 Secondary ionization

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Quenching factor

I computed the QF from the ionization profile along the shooting direction (x):

$$E_{ionization} = \int_0^{x_{max}} \left[\left(\frac{dE}{dx} \right)_e + \left(\frac{dE}{dx} \right)_n \right] dx \quad \longrightarrow \quad QF(E) = \frac{E_{ionization}}{E}$$

The QF is a *global* feature of the track; but the ionization energy losses of the ion changes while it loses energy



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Ionization profile

- The simulated ionization profile is fundamental for the study of directionality and head-tail effect
- I applied the conversion factor to the 3D energy distribution
- Whenever the energy of a secondary recoil is high enough, I construct a cascade based on expected stopping power of the hit atom (same for tertiary recoil, etc...)
- Results are consistent with SRIM results (both for energy deposit profile and spatial distribution)



Simulation of images

- The 3D ionization profile is the starting point to produce simulated images as seen from the sCMOS camera
- Ionization energy deposit → n° electrons (w-value) → multiplication in GEM avalanche → n° photons
- Positions of light determined by diffusion in drift region and GEM stack
- Light in the image depending of geometrical factors (dimension of detector and camera sensor)
- Sensor noise is added
- I produced a sample based on LIME parameters, I reconstructed the tracks with clustering algorithm* (GAC), and I analysed the results







*E. Baracchini et al 2020 JINST 15 T12003



Track direction: as a *preliminary* approach, it is the inclination of the line which maximizes the variance of the pixel distribution, weighted with their intensity. Further studies are needed

igma angle

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20

10

20

40

60



- Diffusion between 0.5 and 1.2 mm
- F recoils (SRIM) •

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60 Torr, 68.03 percent spread



18.86

Std Dev

• DMTPC results:

Diffusion between

0.5 and 1.2 mm

- Track sense: from head-tail effect identification.
- Asymmetry of the track (light integral difference in the two halves of the track)
- Relative position of intensity peak with respect to centre
 - Max intensity (macro)pixel / peak found in profile





C.Deaconu, PhD Thesis, MIT (2009)

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Tested on data taken with AmBe neutron source with LEMON*:

1200 NR identification: the number of peaks found in the profile 1000 could be used as a ER/NR discriminating variable (in addition to Atul's work on ML) NR identification 500 ER 0.8 400 10cm drift 20cm drift 300 0.6 30cm drift 40cm drift 0.4 100% efficiency at 30keV 0.2 (I will soon test it on an ER sample) Cosmic 140 20 40 80 100 Energy [keV] 120F

*E. Baracchiniet al., Meas. Sci. Technol. 32 025902 (2021)



LIME background study for underground operation



LIME background simulation

- **Background simulation** is fundamental for optimizing the shielding, choice of building materials, development and optimization of data analysis, DAQ set up, estimation of the duration of the data taking, study of the sensitivity reach...
- I simulated with GEANT4 the two background components:
 - Induced by natural gamma and neutron fluxes at LNGS
 - Induced by radioactivity of materials
- LIME will operate through 3 phases:
 - <u>1° phase</u>: external gamma background characterization
 - <u>2° phase</u>: neutron flux measurement
 - <u>3° phase</u>: MC validation and internal background characterization
- I optimized the shielding configurations for the 3 phases given the available space underground



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Shielding optimization

I compared the performance of the use of **copper** or a combination of **copper and lead** as a gamma shielding



<u>Copper was chosen as the shielding best option</u>



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Energy [keV]

 10^{2}

10

 10^{-2}

 10^{-1}

External background simulation



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Internal background simulation

- Background simulation from radioactivity of the detector's components is ongoing - from preliminary results, the rate is O(10⁵ events/yr)
- The only NR background comes from the GEMs could be excluded by fiducialization
- Activity measured at LNGS by M.Laubestein (copper rings and resistors are being measured)

Component	²³⁸ U (²³⁴ <i>m</i> Pa)	²³⁸ U (²²⁶ Ra)	²³⁵ U	²³² Th (²²⁸ Ra)	²³² Th (²²⁸ Th)	⁴⁰ K
Camera body [Bq/pc]	7	1.8	0.4	2.1	2.1	1.9
Camera lens [Bq/pc]	0.9	0.41	0.031	0.08	0.08	11
GEM foil [Bq/ m^2]	< 0.104	0.004	< 0.002	< 0.004	< 0.002	< 0.045
Acrylic [Bq/kg]		0.003		0.005	0.004	0.035

- Single components of a broken BMI express sCMOS camera were measured, primary contribution from sensor and electronic board
 - We could use a custom non-radioactive sensor
 - Electronic board can be moved away from the detector





Spectrum deconvolution

The measured NR spectrum will be the result of the convolution of the neutron spectrum with the detector response





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I'm starting to study an iterative unfolding method based on the bayesian approach

Future perspective and conclusions



Future perspectives

- Analysis of simulated NR will be improved and further developed with a larger and realistic sample
- **Comparison of simulated and measured tracks** will be done to validate the MC simulation and analysis
- A full simulation of the **internal background** will allow the optimization of LIME underground data taking and final choice of materials
- I will continue the R&D work on the MANGO prototype at overground
 LNGS I will study the use of ³He for thermal neutron detection
- I will study the **unfolding method** to retrieve the neutron spectrum
- By the end on 2021 LIME will be installed and commissioned underground, next year the **data taking will start**





Conferences/workshops:

• I presented a talk in the Early Career Plenary in CPAD Instrumentation Frontier Workshop 2021, online, 18/03/21 - 22/03/21. <u>https://indico.fnal.gov/event/46746/contributions/210387/</u>

Scientific publications:

 Directional Dark Matter Searches with CYGNO, CYGNO Collaboration, Particles 2021, 4,343-353. <u>https://doi.org/10.3390/particles4030029</u>

List of attended schools and courses:

- INFN SOUP 2021 International School on Underground Physics, online, 28/06/21 02/07/21
- QSFP 2021 International School on Quantum Sensors for Fundamental Physics, online,06/09/21 - 17/09/21
- GSSI interdisciplinary course: Probabilistic reasoning in inference, forecasting and decision making, G. D'Agostini, 07/06/21 - 09/06/21, 14/06/21 and 16/06/21





backup



SRIM ionization profile



ER rejection factor

- Simulation for 755:5 Torr of He:SF₆ CYGNUS collaboration <u>https://arxiv.org/abs/2008.12587</u>
- R = expected electron recoil rejection factor based on track length selection, 50% efficiency for NR selection

$$R = N_{all}/N_{surv},$$

• If the rate is kept below 10⁵ events/yr above 20keV, a zero background measurement is possible





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