



Astroparticle Physics - XXXIII cycle 14/05/2021

Dr. Paolo Gorla



- Majorana neutrinos and Double Beta Decay Physics
- The CUORE experiment & latest results
- The CUORE Background Model & α region studies
- Search of neutrinoless double beta decay of ¹²⁸Te







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Open question in Astroparticle Physics G

Is the Lepton number conserved in nature?

How is the neutrino mass produced?

3) The Big Bang should have created the same amount of matter and anti-matter. Why is there more matter than anti-matter in our Universe?

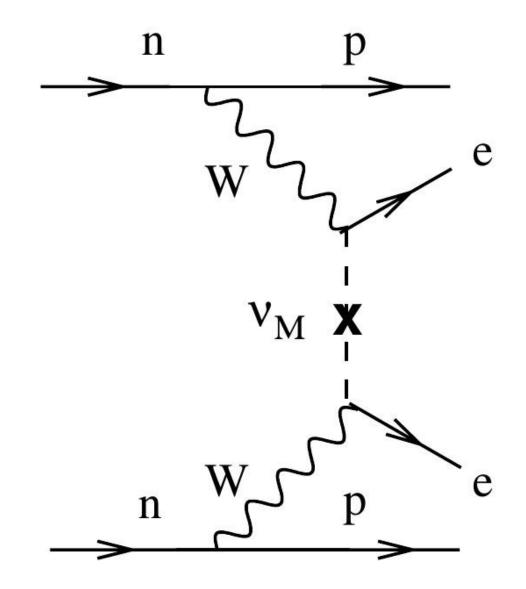
V. Dompè - Search for neutrinoless double beta decay of ¹²⁸Te with the CUORE experiment

1) In the Standard Model, the Lepton number conservation is not associated to a fundamental symmetry.

2) First Standard Model formulation: $m_v = 0$, but neutrino flavor oscillations proved that $m_v \neq 0$.



S Neutrinoless double beta decay (0vββ) G



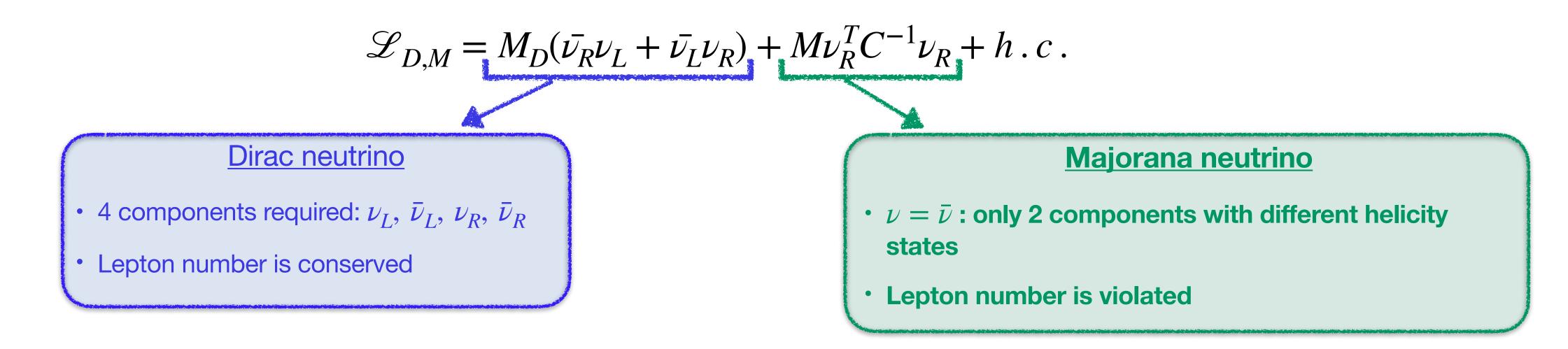
- Only matter is created: first experimental evidence
- Lepton number-violating process: would demonstrate that L is not a symmetry of nature
- Only possible if neutrinos have a Majorana component





S G **Dirac and Majorana Neutrinos**

- antiparticles are coincident ($\nu = \bar{\nu}$)
- New possible mechanism giving rise to ν mass (e.g. see-saw mechanism)
- The general form of the Standard Model Lagrangian for neutrinos:



• Possible explanation of matter-antimatter asymmetry origin via Leptogenesis

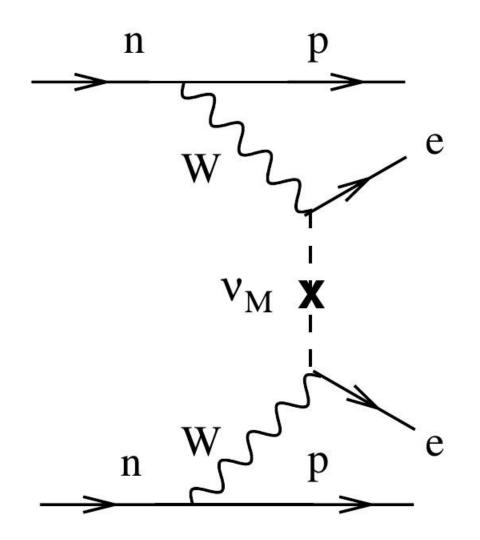
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• 1937 - Ettore Majorana introduced a new theory of fermions: representations of elementary particles and

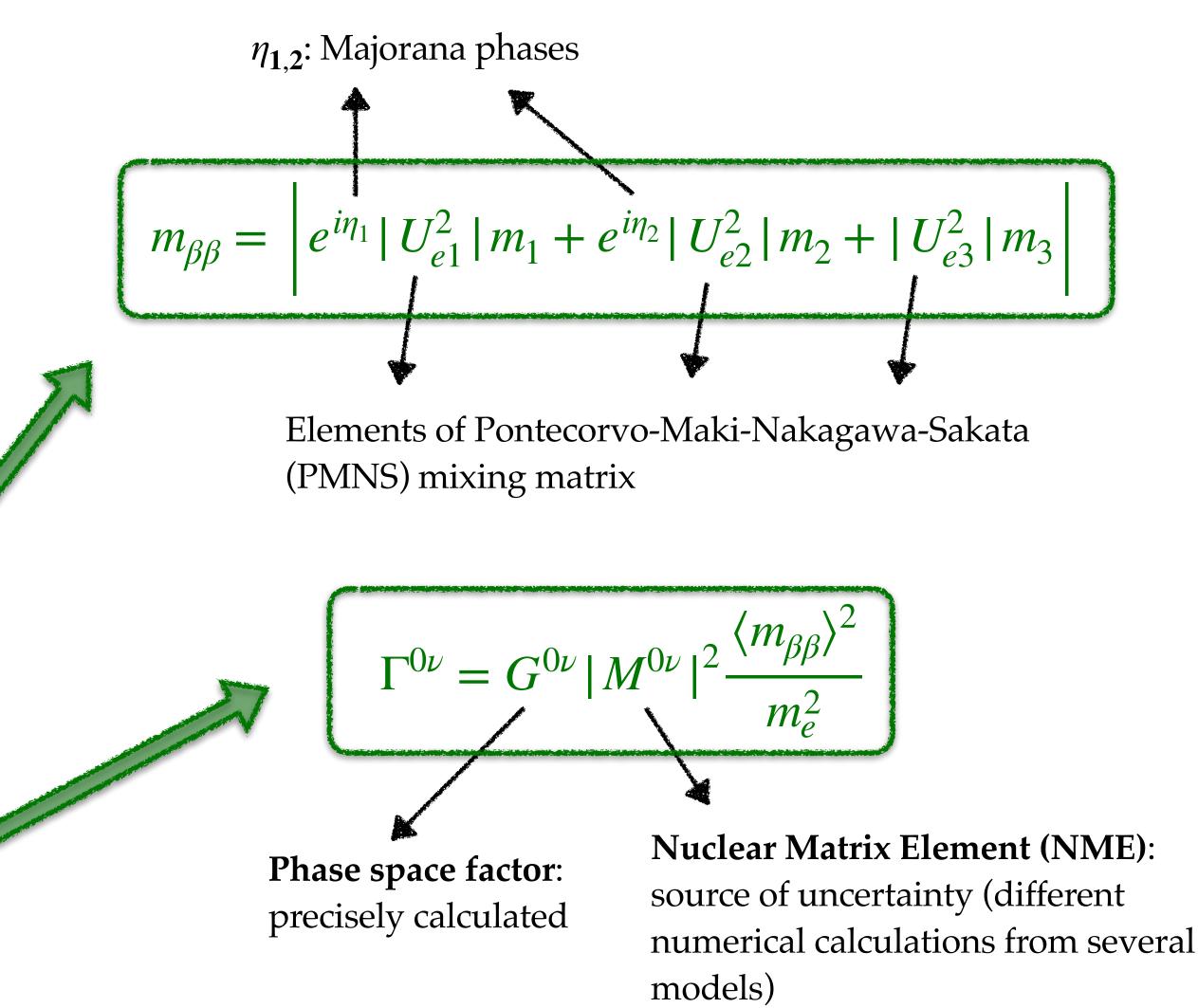




S G Neutrinoless double beta decay



- Light Majorana neutrino exchange mechanism for 0vββ decay
- In this case, we define the Effective Majorana mass $m_{\beta\beta}$
- The $0\nu\beta\beta$ rate is proportional to $m_{\beta\beta}$

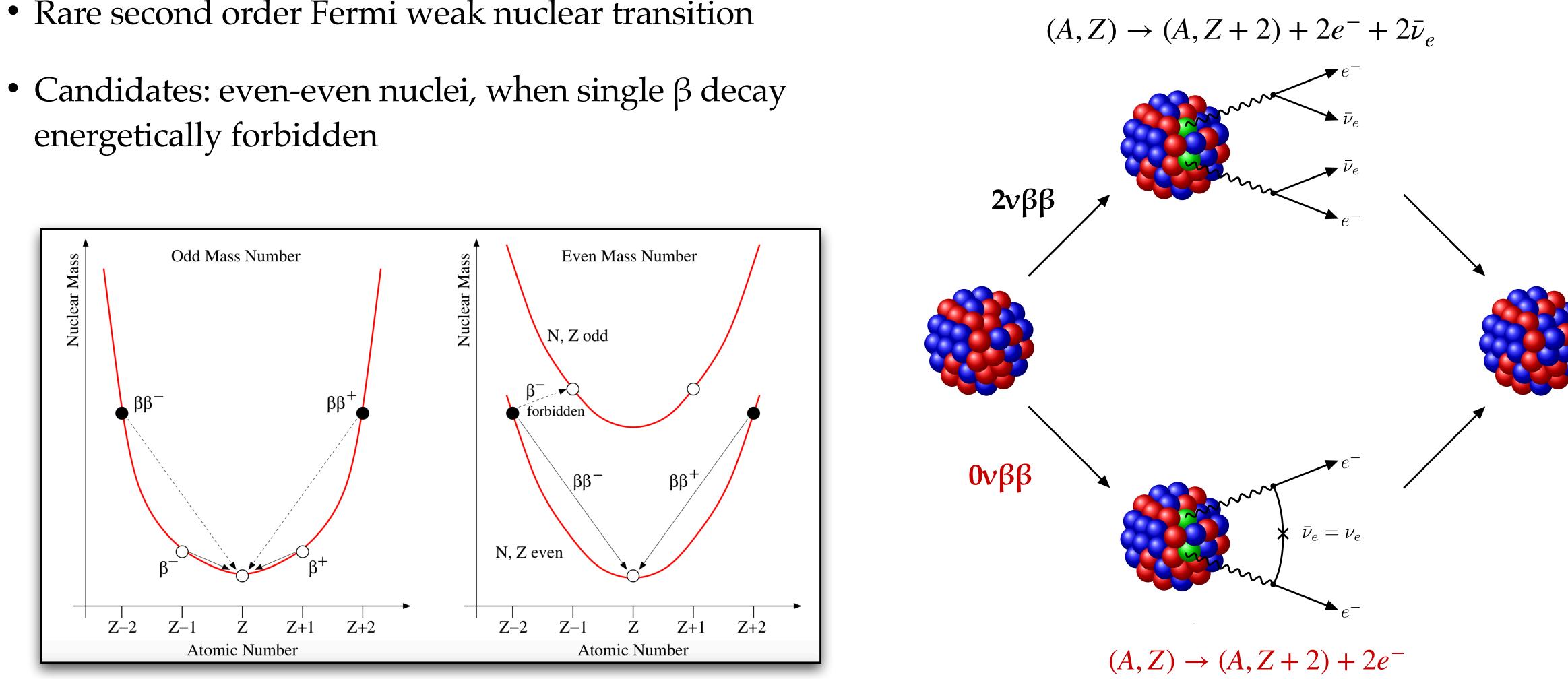






S G Double beta decay: 2νββ and 0νββ S

- Rare second order Fermi weak nuclear transition
- energetically forbidden

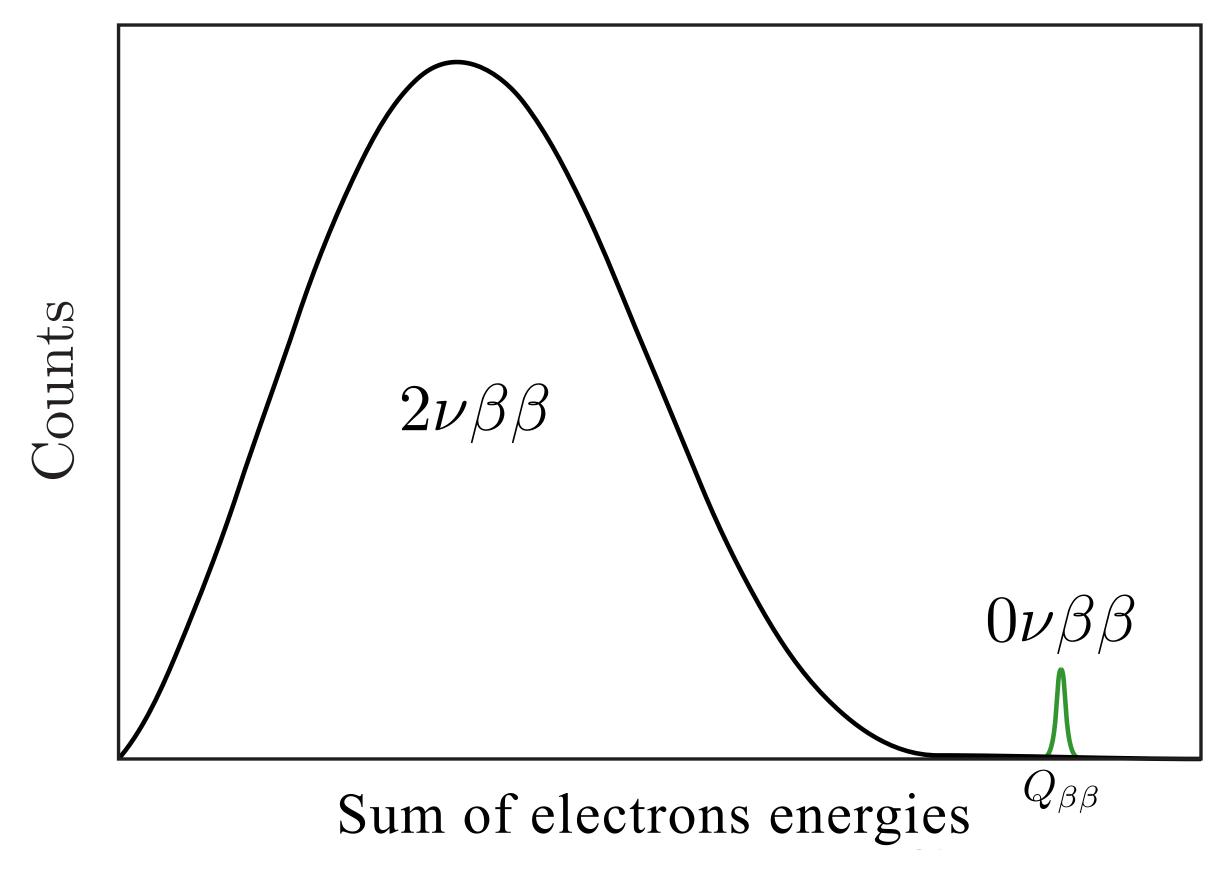








S G Experimental search of 0vββ decay



• Experimental signature of this process is the Q-value of the transition:

$$Q_{\beta\beta} = (M_{parent} - M_{daughter} - 2m_e)c^2$$

- $2\nu\beta\beta$: continuous spectrum from 0 to $Q_{\beta\beta}$
- $0\nu\beta\beta$: monochromatic peak at $Q_{\beta\beta}$





G S I Experimental search of 0vββ decay

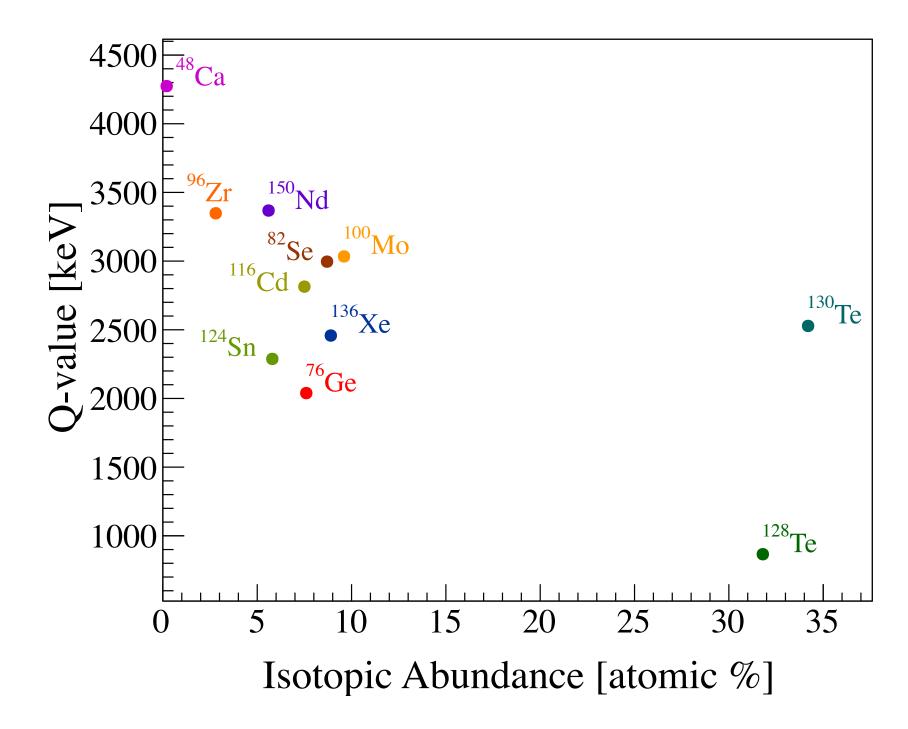
The potential of an experiment to 0vββ decay is evaluated through the *Sensitivity*

Sensitivity:

the half life corresponding to the minimum number of observable signal events above the background, at a given statistical significance n_{σ}

$$S^{0\nu}(n_{\sigma}) = \frac{\ln 2}{n_{\sigma}} \frac{N_a}{A} \sqrt{\frac{M\Delta t}{b\Delta E}}$$

- Detection efficiency
- Isotopic abundance of 0vββ emitter
- Exposure (large detector mass and long live time)
- Background level
- Energy resolution





10



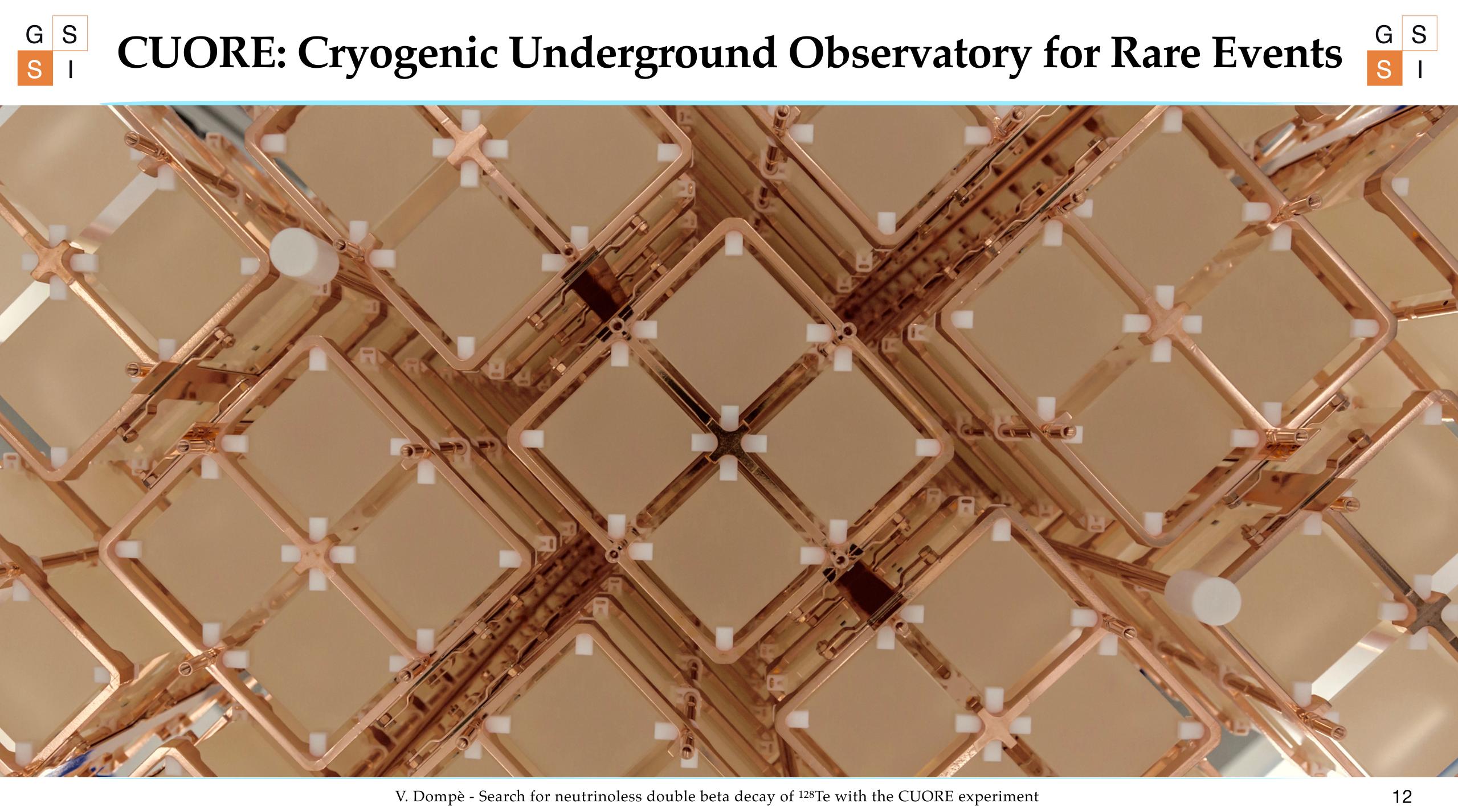
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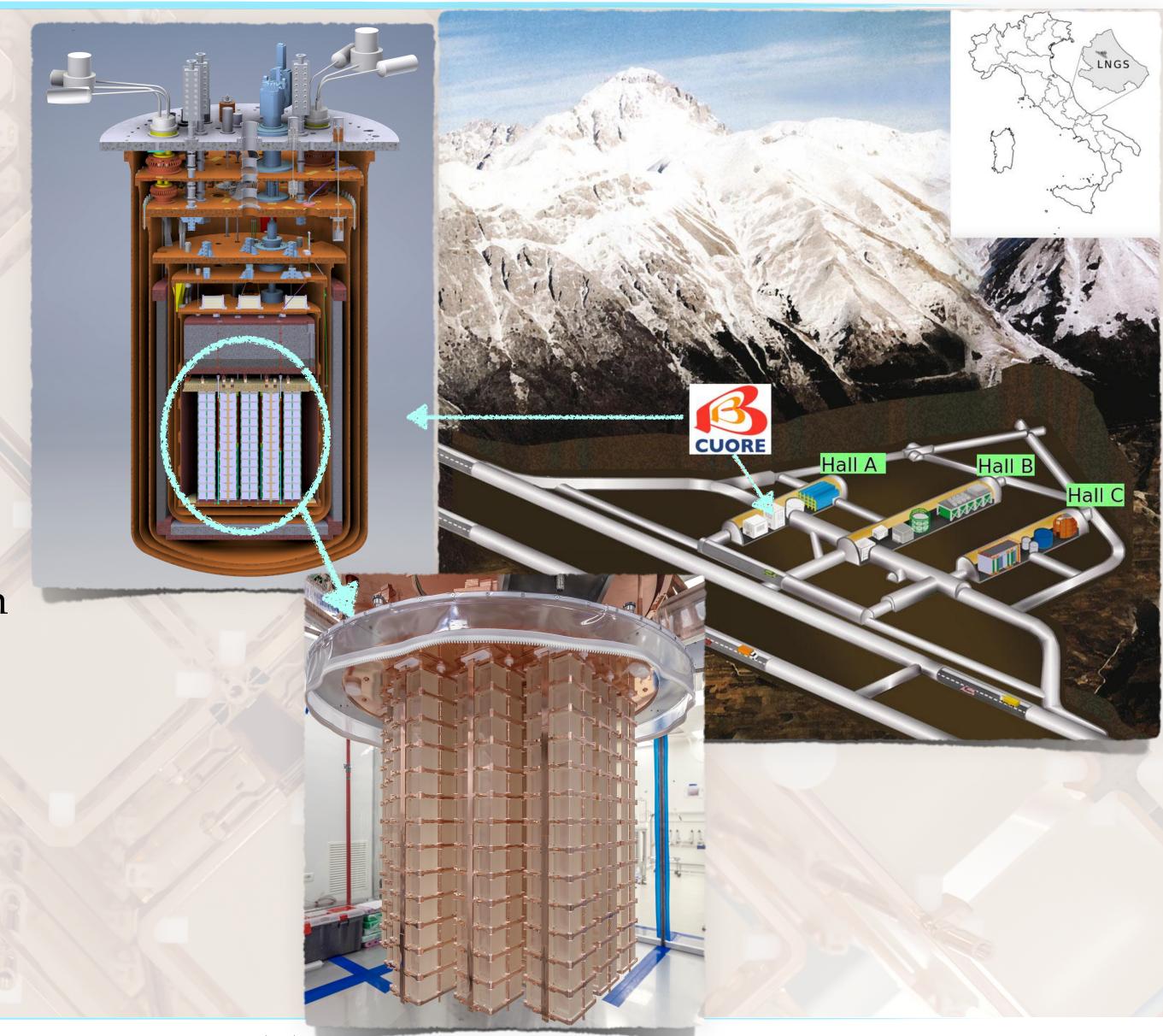
11





CUORE: Cryogenic Underground Observatory for Rare Events

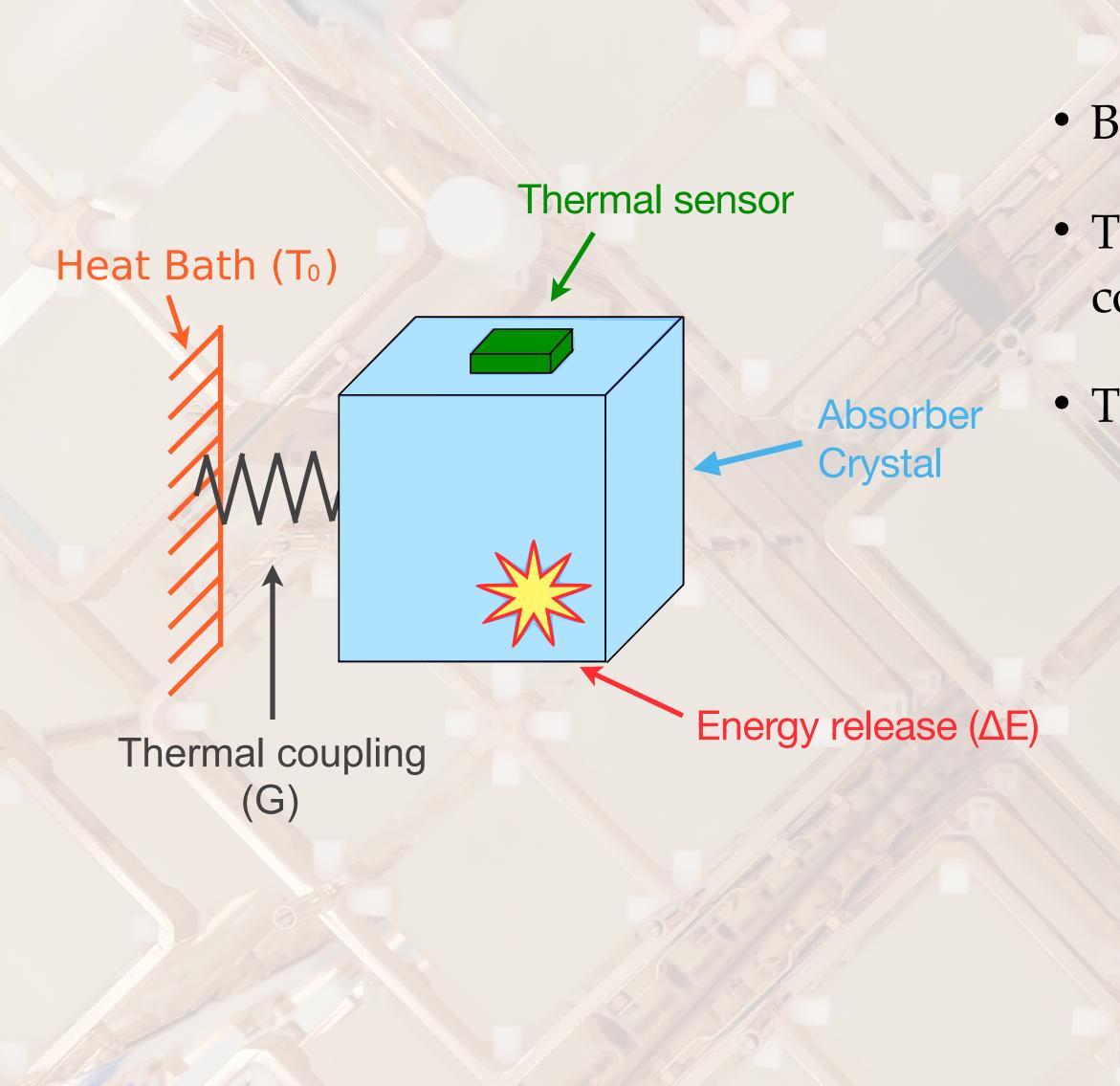
- Main Physics goal: search for 0vββ decay of ¹³⁰Te
- Located at the underground Laboratori Nazionali del Gran Sasso of INFN: 3650 m.w.e. of rock coverage to suppress the cosmic radiation
- 988 natural TeO₂ crystals (742 kg of TeO₂, 206 kg of ¹³⁰Te) arranged in 19 towers
- ¹³⁰Te embedded in the detector itself: ~90% detection efficiency
- Crystals operated as bolometers at ~10 mK







S G **CUORE - The Bolometric technique**



- Bolometer: solid state detector working as calorimeter
- The energy of a particle interacting with the absorber is converted into phonons
- The temperature variation is measured by the thermal sensor

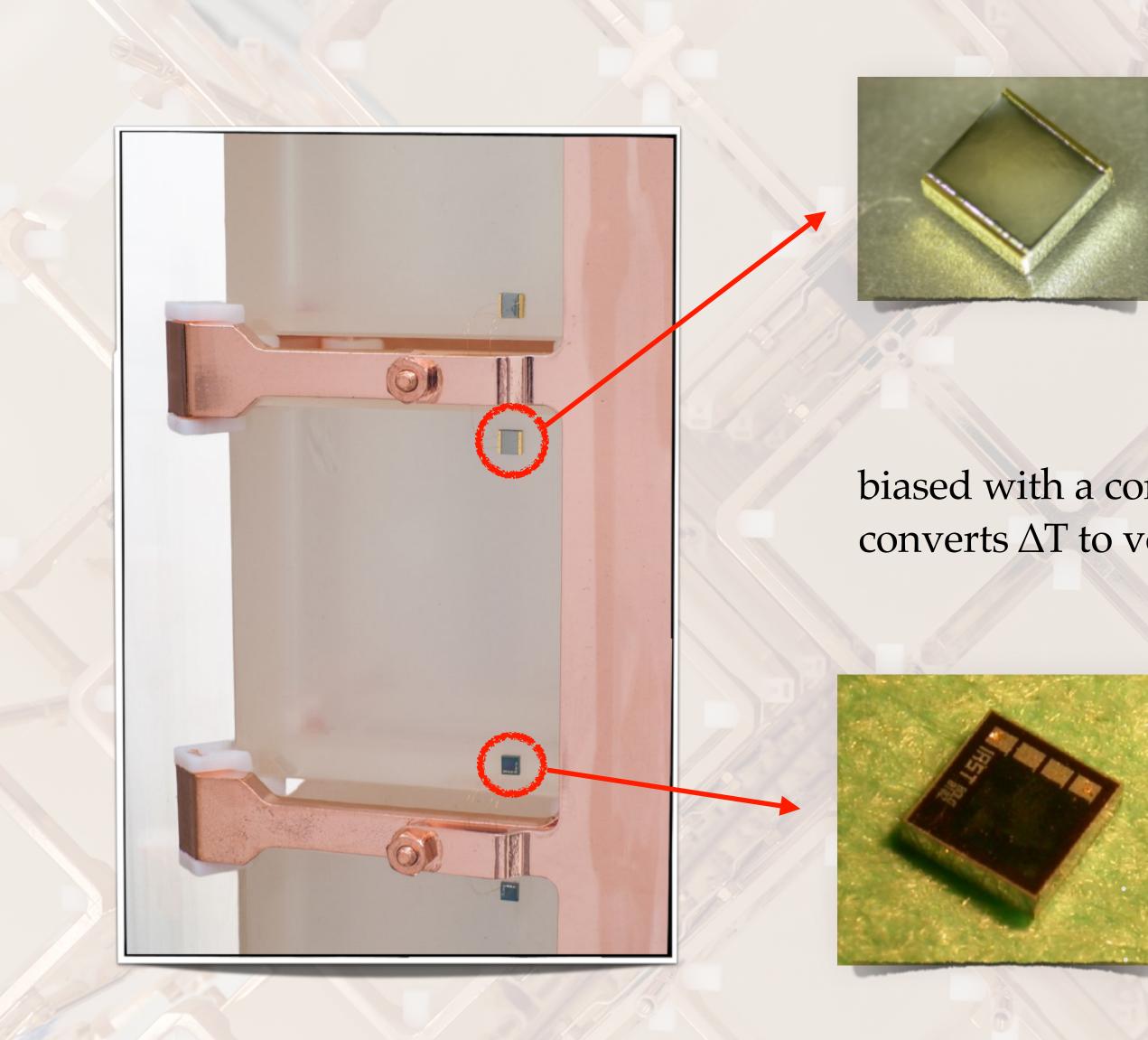
$$\Delta T = \frac{\Delta E}{C}$$
 heat capacity: $C_{TeO_2} \propto T$

 $T_0 = 300 \text{ K}: \quad \Delta E = 1 \text{ MeV} \longrightarrow \Delta T \sim 10^{-18} \text{ - } 10^{-15} \text{ K}$

 $T_0 = 10 \text{ mK}: \Delta E = 1 \text{ MeV} \rightarrow \Delta T \sim 0.1 \text{ mK}$



S G **CUORE - Thermal Sensor and Heater**

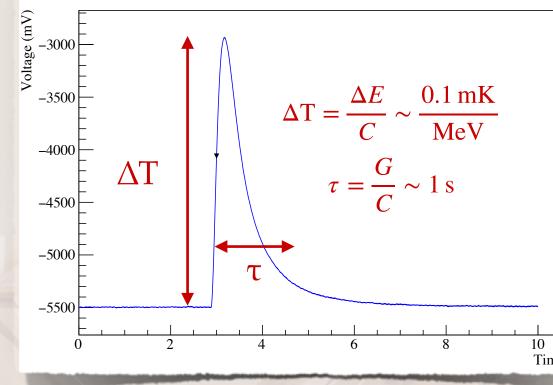


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NTD = Neutron Transmutation **Doped Ge thermistor**

 $R(T) = R_0 e^{\sqrt{\frac{T_0}{T}}}$

biased with a constant current, converts ΔT to voltage signal

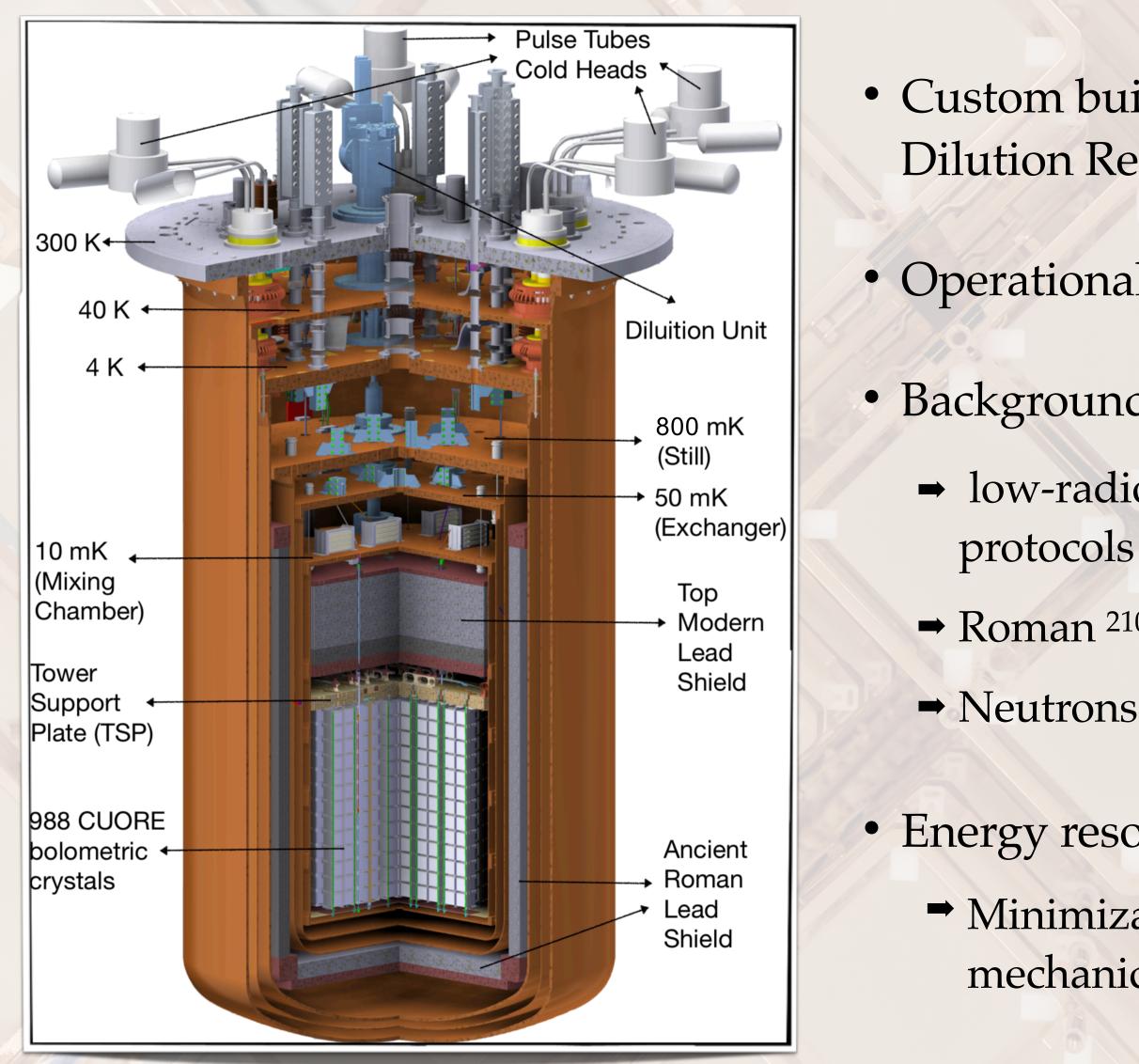


Silicon heater

Periodically fires a pulse at a fixed energy to measure the detector's gain and correct the effect of thermal instabilities (*stabilization*)



S G The CUORE Cryostat

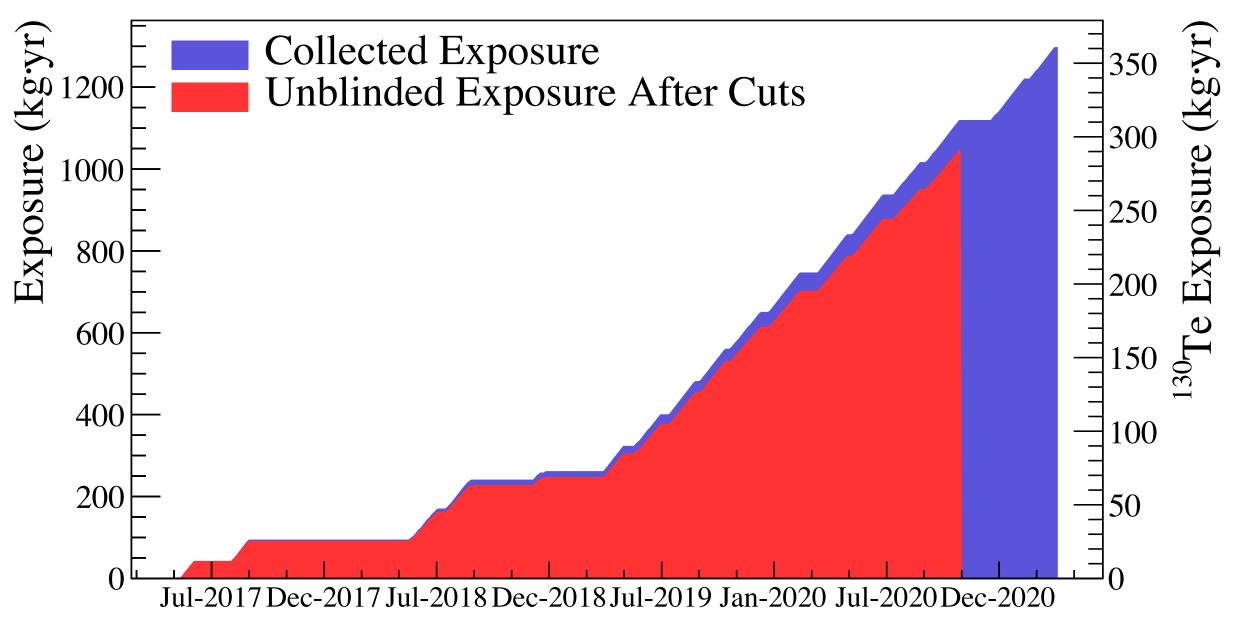


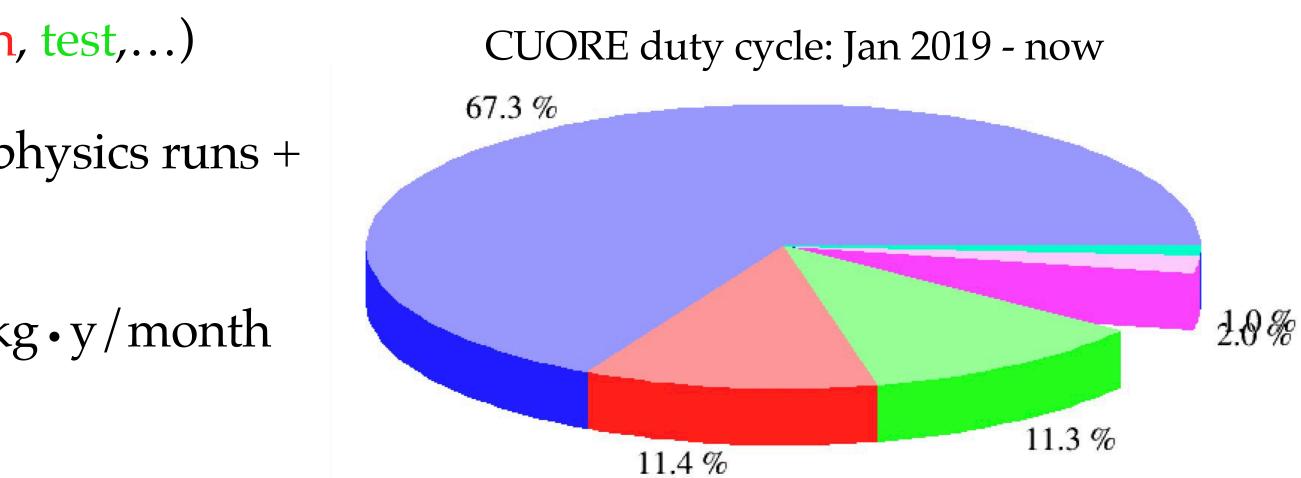
- Custom built cryogen-free structure: 5 pulse tubes + ³He/⁴He **Dilution Refrigerator**
- Operational T ~ 10 mK stable over years
- Background level goal of 10^{-2} counts/(keV·kg·y):
 - → low-radioactivity materials choice, strict cleaning and assembling
 - Roman ²¹⁰Pb-depleted + modern lead shields
 - Neutrons shield: external polyethylene layer with boric acid panels
- Energy resolution < 8 keV at ¹³⁰Te Q_{ββ}:
 - Minimization of vibrational noise: external support structure mechanically decouples the detectors from the cryostat



S G The CUORE data collection

- Data taking organization: *runs* (physics, calibration, test,...)
- 1 *Dataset* (40 60 days) = initial calibration runs + physics runs + final calibration runs
- CUORE data taking is proceeding smoothly (~69 kg y / month since spring 2019)





2021: 1-ton · y analyzed exposure milestone!

- Total collected exposure: 1221.06 kg-y
- Total analyzed exposure: 1038.4 kg·y





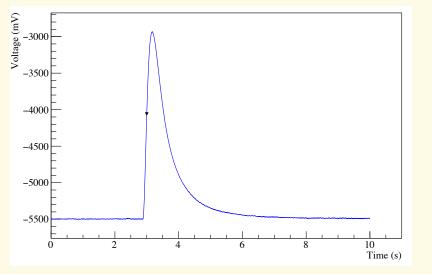
S G The CUORE data processing

From the raw triggered waveform to the energy spectrum:

First basic quantities:

Preprocess

- Baseline, BaselineRMS, BaselineSlope (bolometer working conditions before particle interaction)
- Waveform parameters: SingleTrigger, NumberOfPulses (pile up parameters)



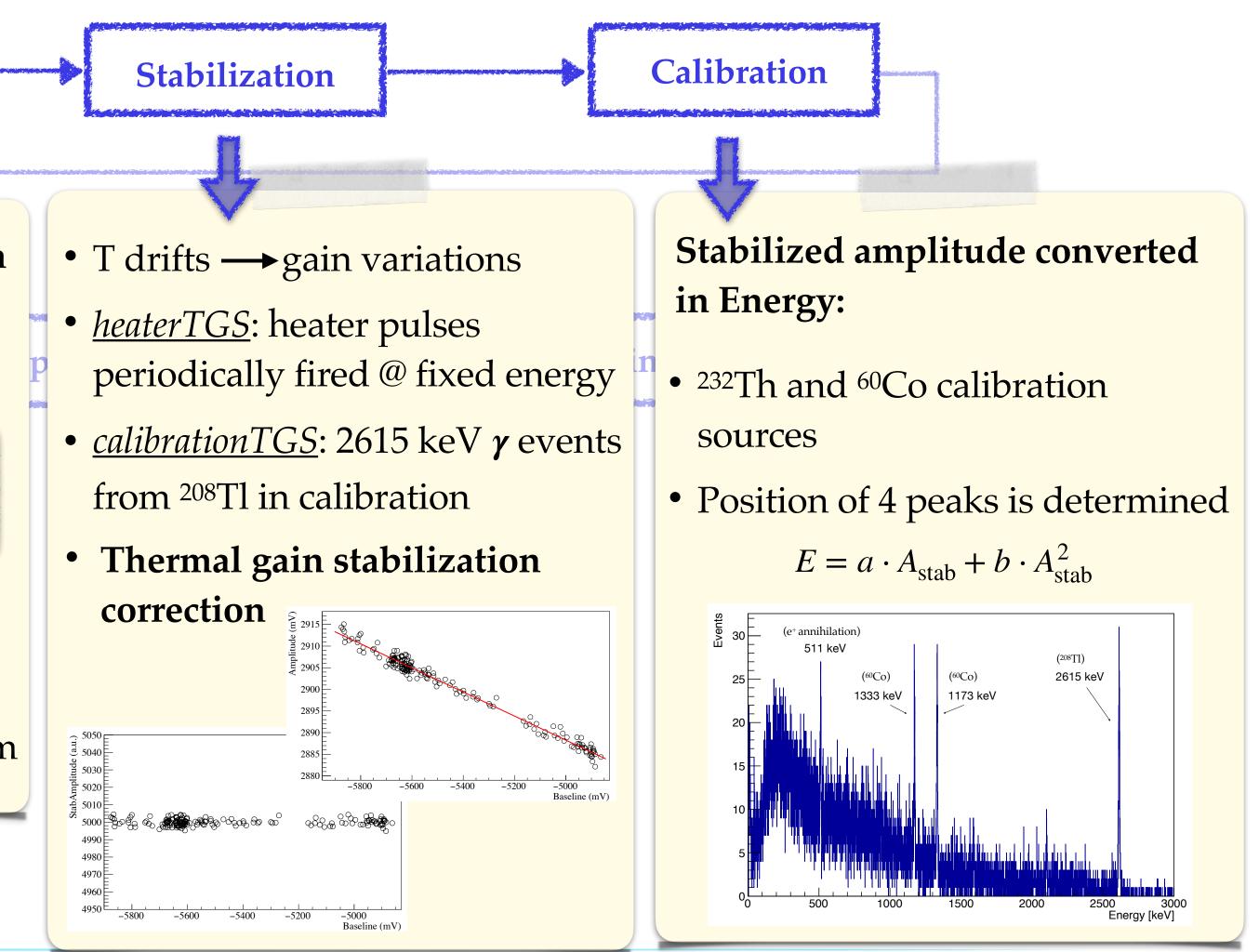
Signal amplitude evaluation

Amplitude

• Optimum Filter: SNR is maximized

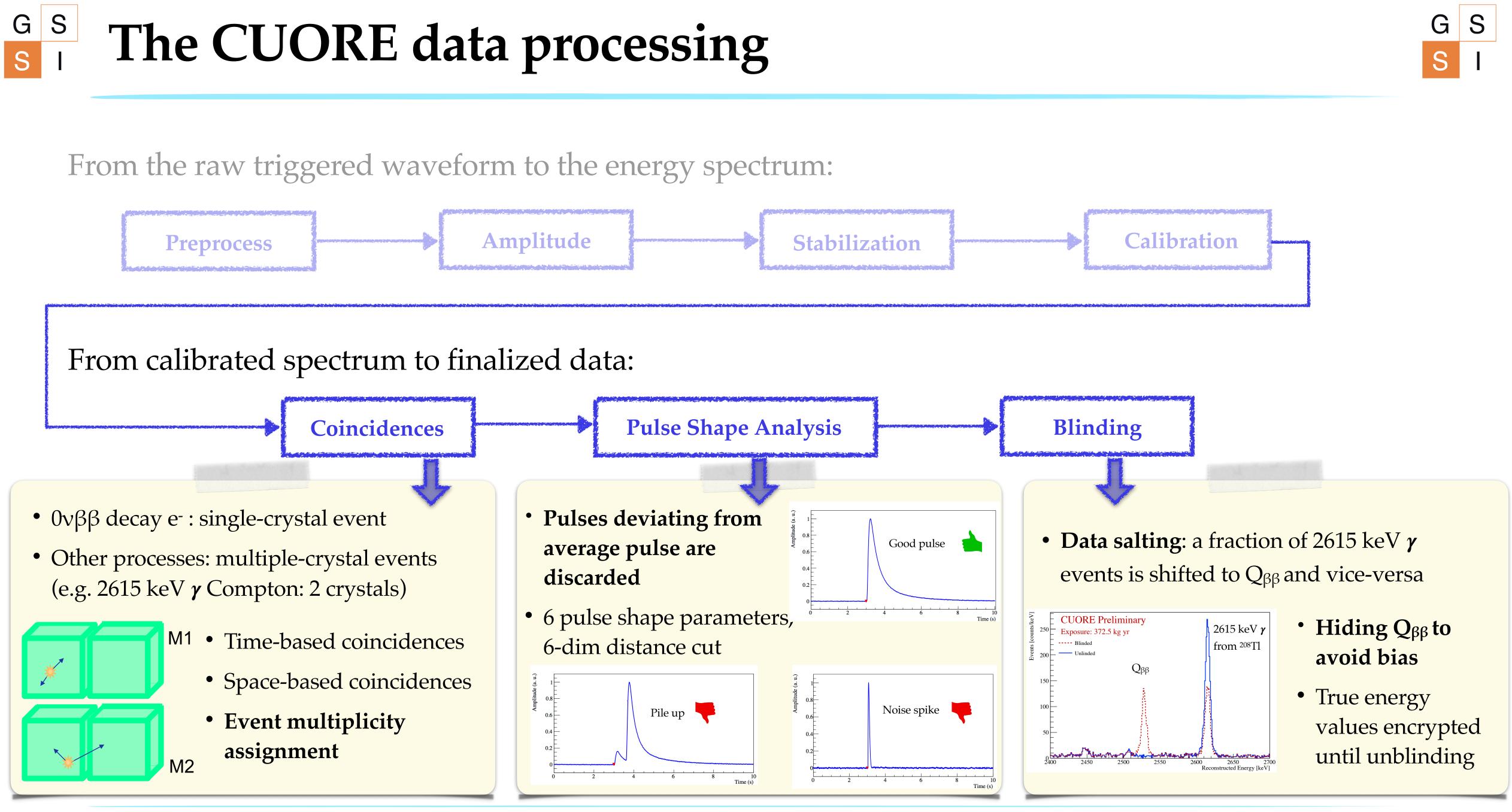
$$V_i^{OF}(\omega) \propto e^{i\omega t_M} \frac{S_i^*(\omega)}{N_i(\omega)} V_i(\omega)$$

- $V_i(\omega) =$ non-filtered signal
- $S_i(\omega)$ = response function
- $N_i(\omega)$ = noise power spectrum







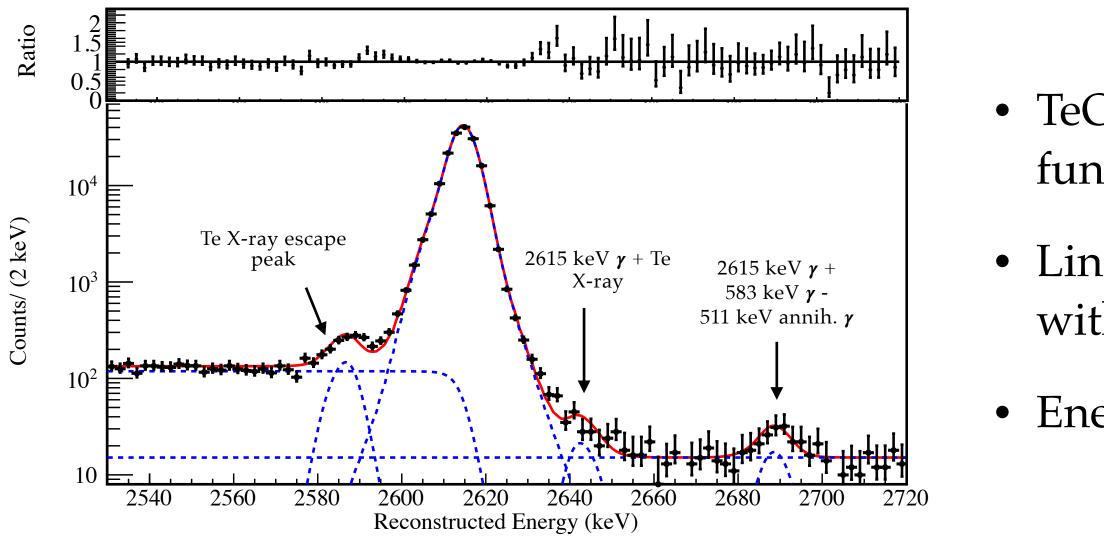


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19

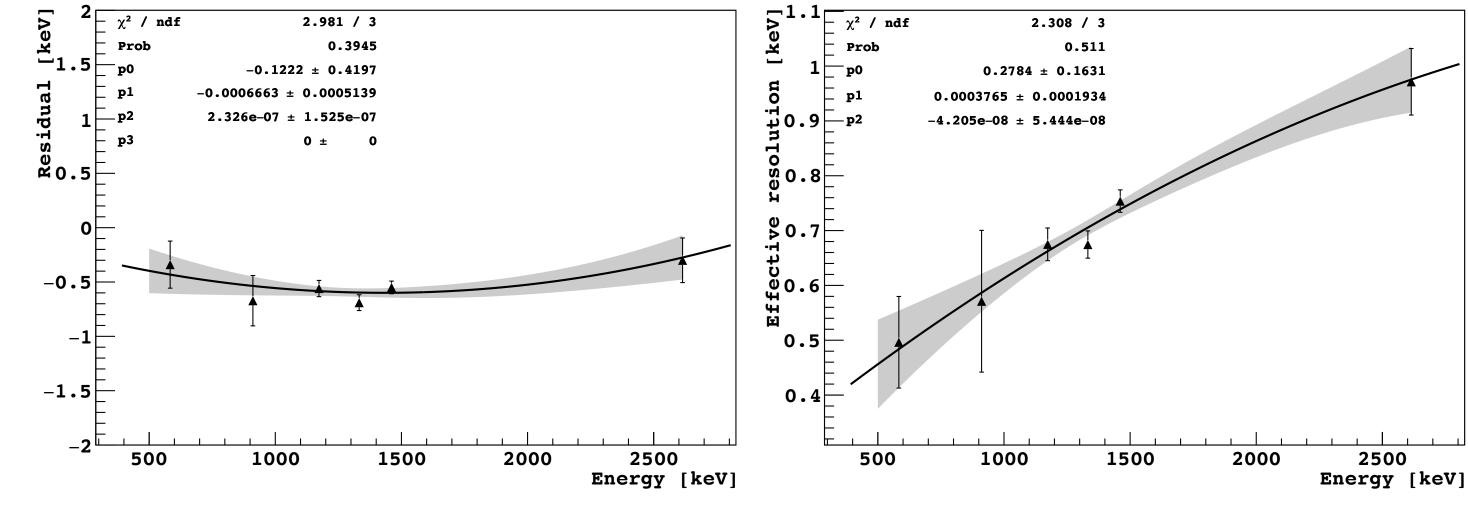
S The CUORE detector response function (lineshape)



• Lineshape in physics data: most prominent peaks fitted

G

- Resolution appears energy dependent, small bias on energy reconstruction
- 2nd order polynomial fit to extract the resolution and bias energy dependence

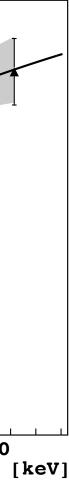


• TeO₂ bolometers exhibit a slightly non-gaussian response function

• Lineshape evaluated on the 2615 keV line in calibration: fit with 3 Gaussian for each detector-dataset

• Energy resolution in calibration is extracted (~7.7 keV)



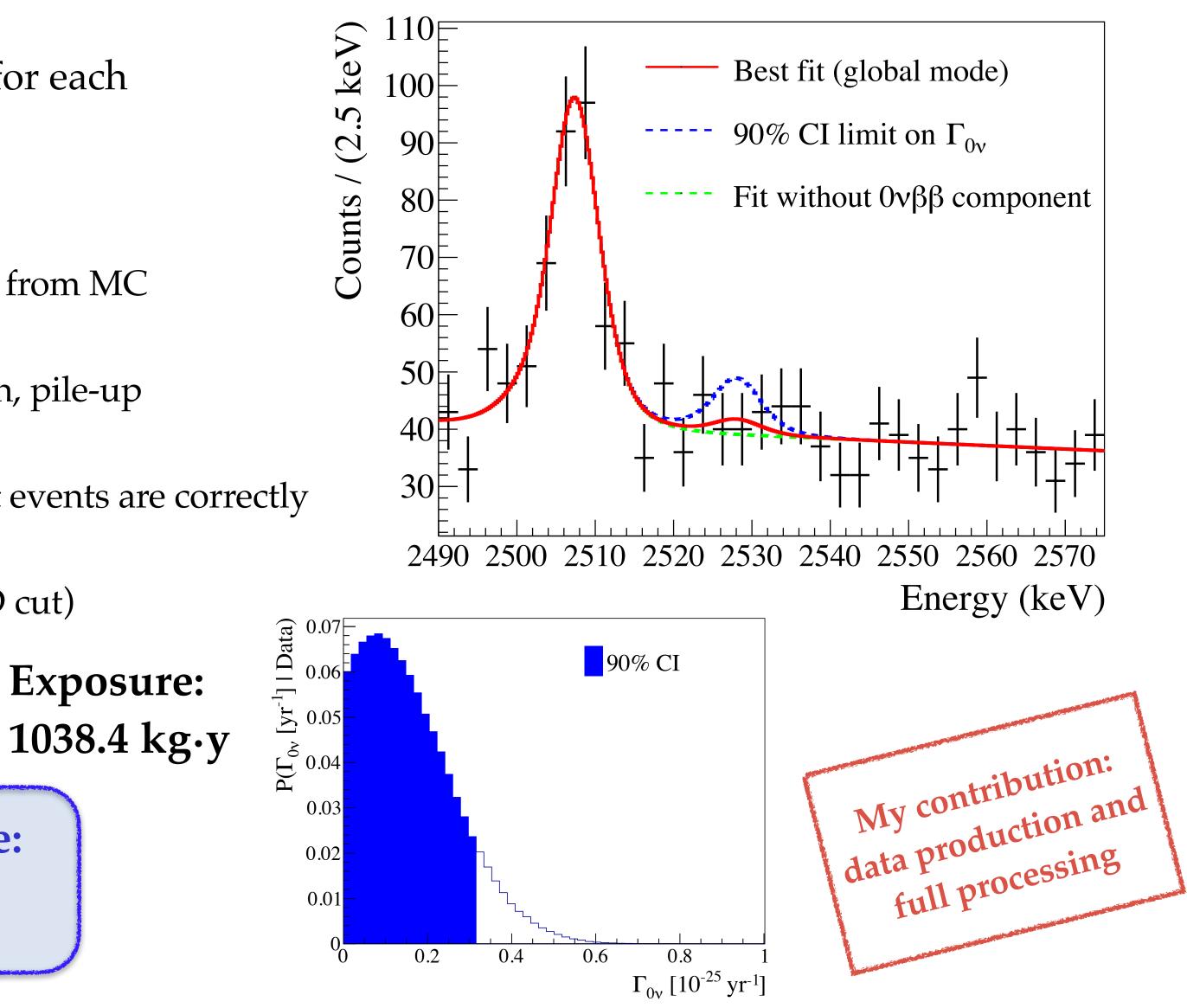




S G CUORE latest results on ¹³⁰Te 0νββ decay

- Unbinned Bayesian fit simultaneously performed for each detector-dataset
- Efficiencies: \bullet
 - <u>Containment efficiency</u> ($0\nu\beta\beta$ decay in single crystal): from MC simulations
 - <u>Reconstruction efficiency</u> (trigger, event reconstruction, pile-up identification)
 - <u>Anti-coincidence efficiency</u> (probability that single-hit events are correctly identified)
 - ➡ <u>PSD efficiency</u> (fraction of events that survive the PSD cut)
- No evidence of ¹³⁰Te $0\nu\beta\beta$ decay is observed

90% C.I. limit on ¹³⁰Te 0νββ decay half-life: $T_{1/2}^{0\nu} > 2.2 \cdot 10^{25} \text{ y}$









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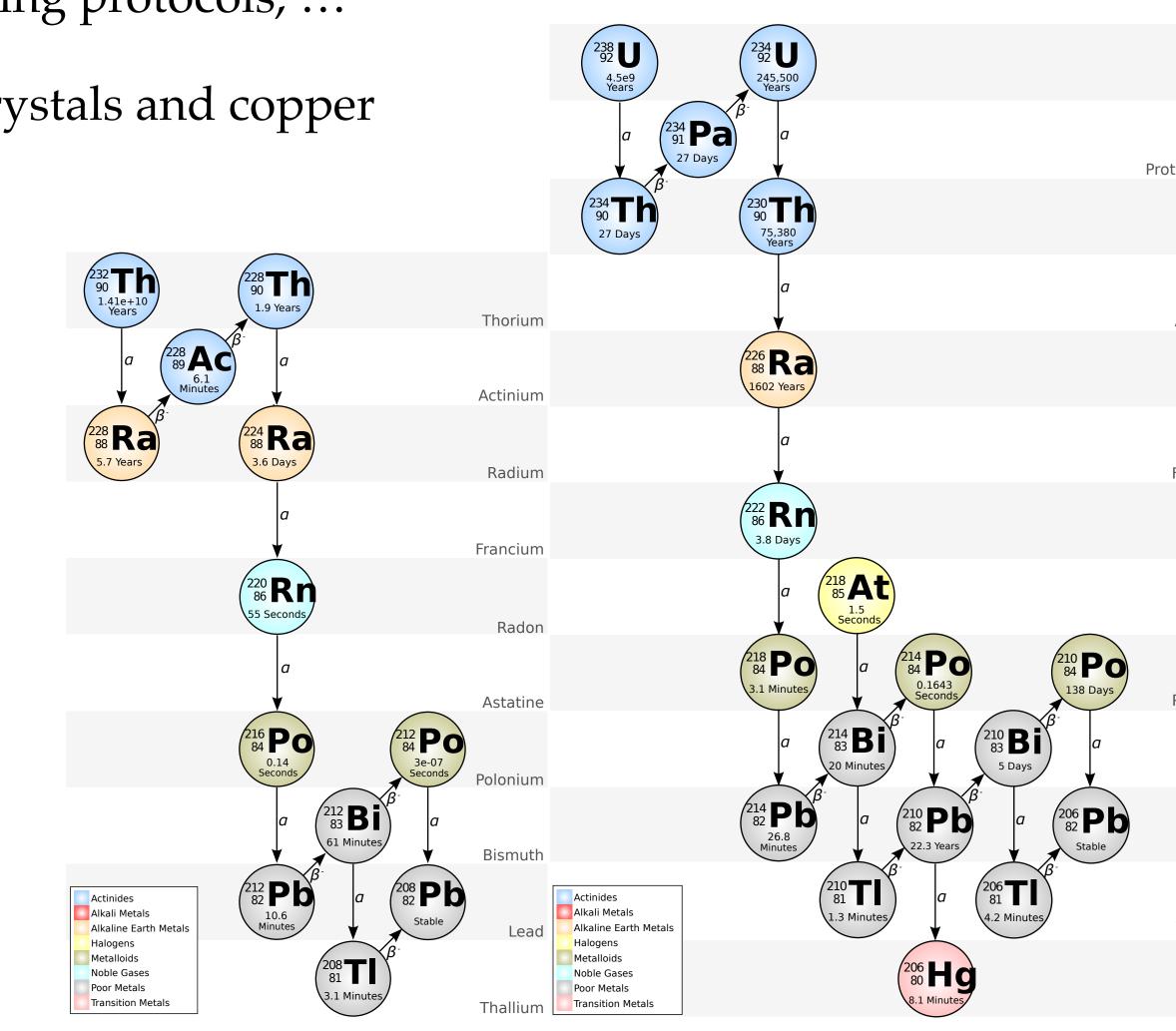
S G **Radioactive sources**

- Radiopure material selection, cleaning and assembling protocols, ...
- ... but: dominant background sources come from crystals and copper residual contamination

```
<sup>232</sup>Th chain nuclei
<sup>238</sup>U chain nuclei
<sup>40</sup>K (environment)
<sup>60</sup>Co, <sup>54</sup>Mn (Cu cosmogenic activation)
<sup>190</sup>Pt (crystal bulk)
. . . . . .
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• Quantitative evaluation of the physics background contributing to the observed spectrum:

BACKGROUND MODEL



Credits:



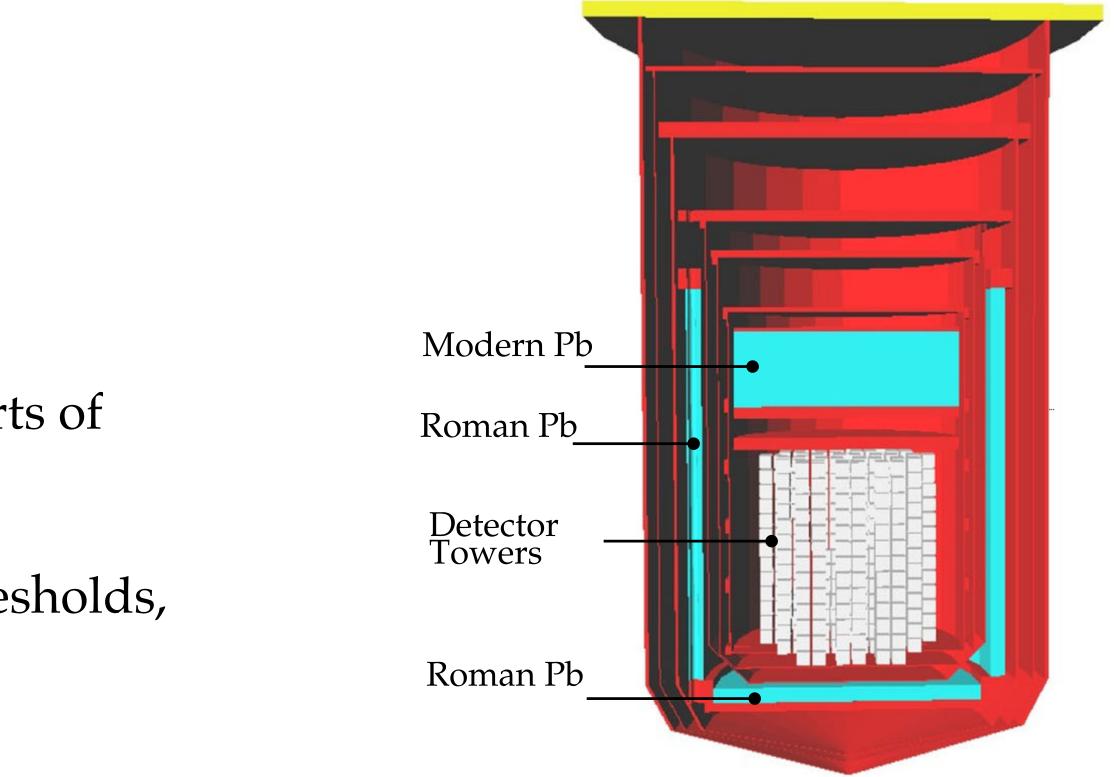




S G The CUORE Background Model (BM)

BACKGROUND MODEL - Detailed Monte Carlo simulations:

- Software based on GEANT4
- Geometry of the CUORE setup
- Radioactive sources and their locations
- Radiation interactions with the different parts of the detector
- Detector response, instrumental effects (thresholds, resolution,...)

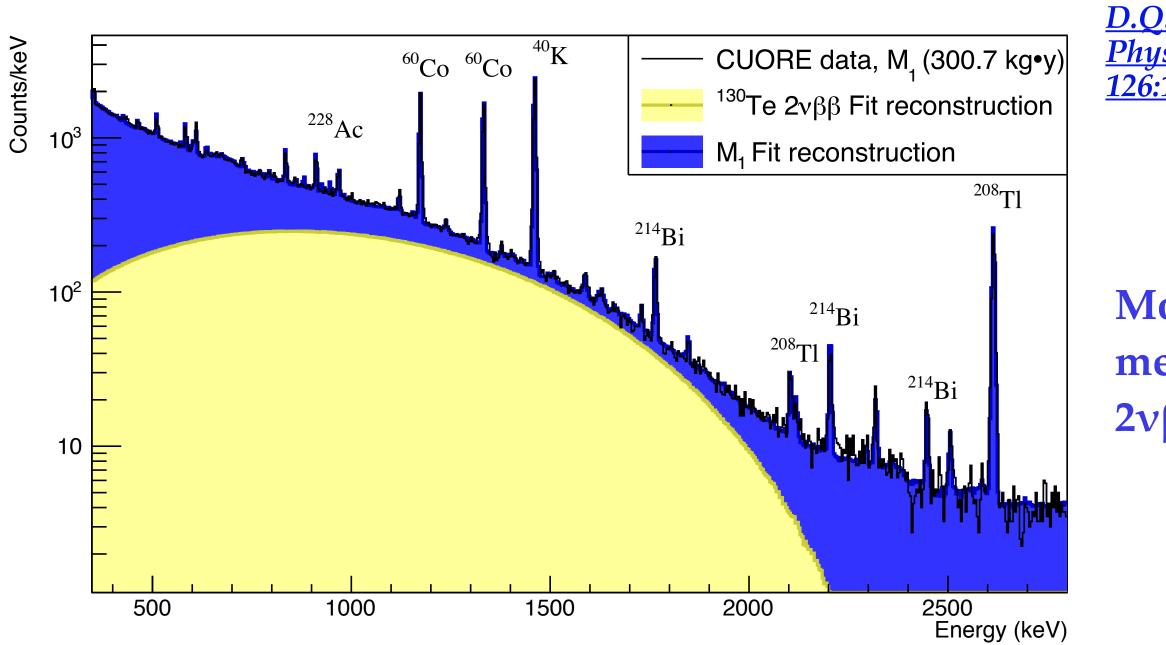


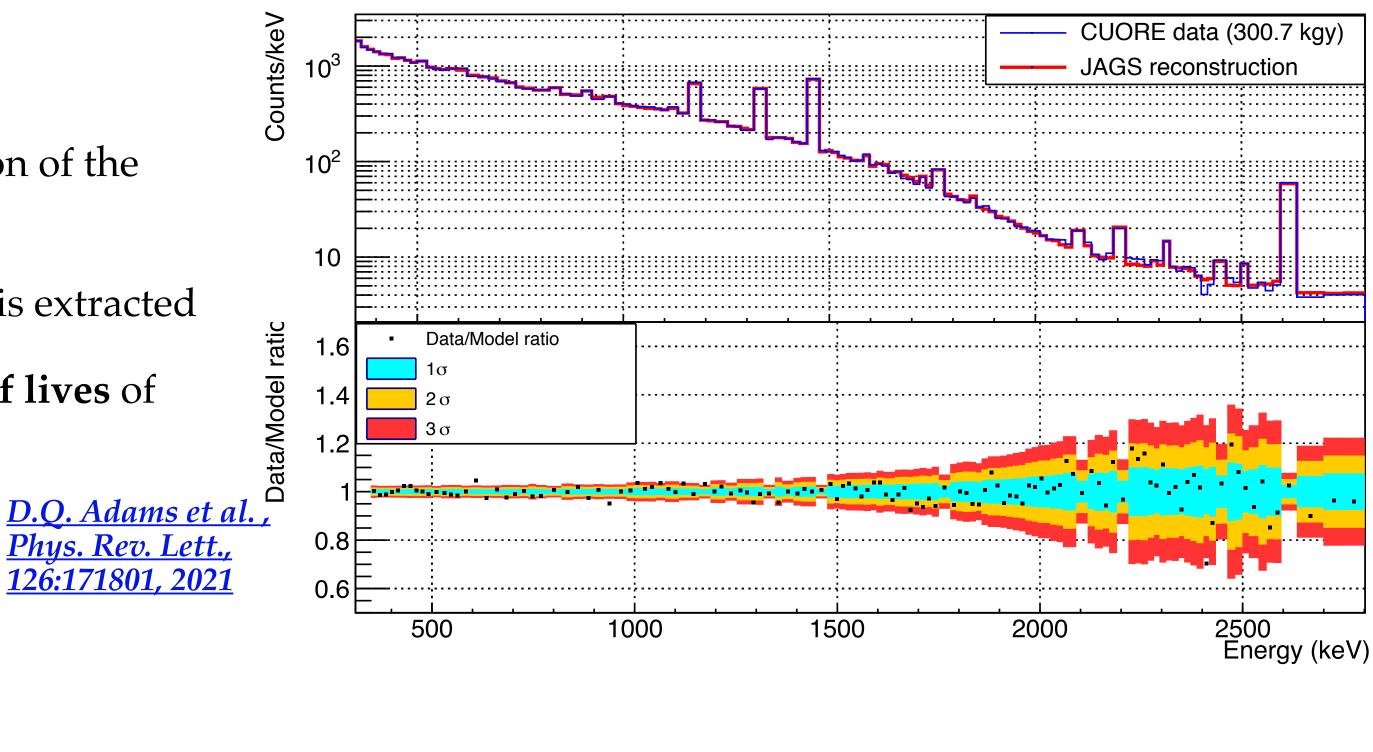




S G The CUORE Background Model (BM): 2νββ decay of ¹³⁰Te

- Identified background contributions are simulated
- Bayesian fit on experimental data with a linear combination of the MC simulations
- Fit parameters: a normalization factor N_i for each source j is extracted
- N_i used to extract the **activity** of the contaminants and **half lives** of processes (e.g. $2\nu\beta\beta$ decay $T_{1/2}$)





Most precise measurement of ¹³⁰Te **2v**ββ **decay** half-life:

$$T_{1/2}^{2\nu} = 7.71_{-0.06}^{+0.08} (\text{stat})_{-0.15}^{+0.12} (\text{syst}) \cdot 10^{20}$$

Faithful reconstruction of the γ region (up to 2.8 MeV)







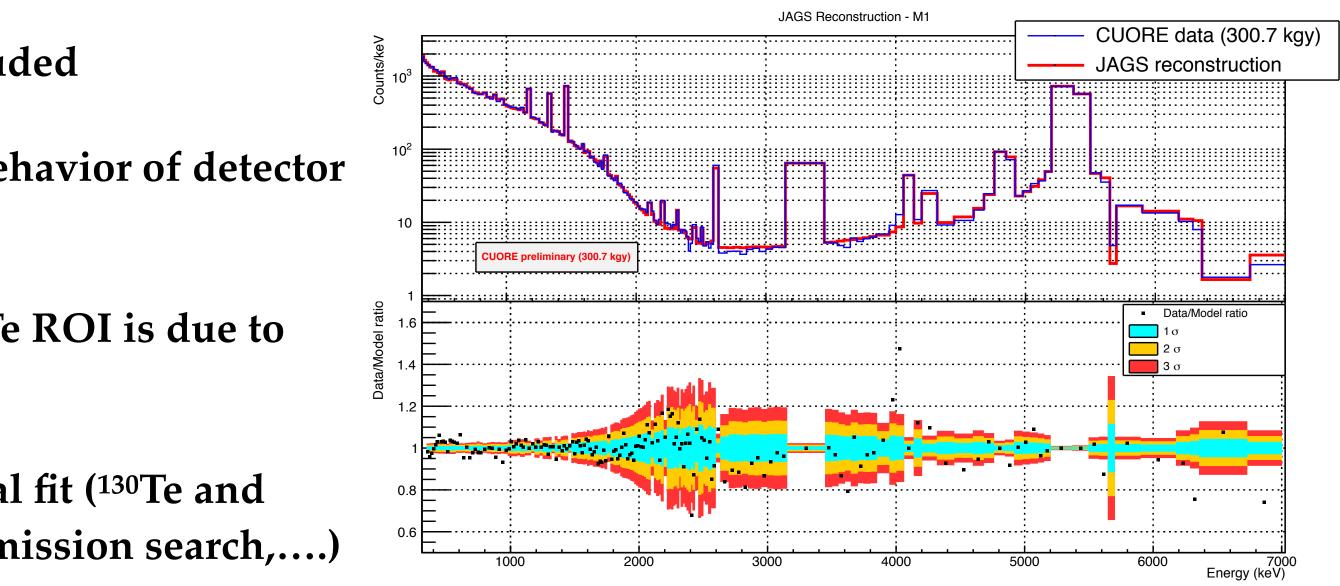


Improving the CUORE Background Model S G

- Fit range: 350 keV 2.8 MeV the α region is excluded
- The α region reconstruction is challenging (non-ideal behavior of detector response, surface contamination profile modeling, ...)
- Important to investigate since >90% of events in the ¹³⁰Te ROI is due to degraded alphas
- Fundamental for rare events analysis requiring a spectral fit (¹³⁰Te and ¹²⁸Te $2\nu\beta\beta$ half life, CPT violation signature, Majoron emission search,....)

Possible improvements to the α **region:** more detailed knowledge of the sources

- Contaminant position in the detector
- α particle quenching in TeO₂ crystals
- Breaking points in the radioactive chains



New inputs to the Background Model

Fit quality improvement, more precise interpretation of the data







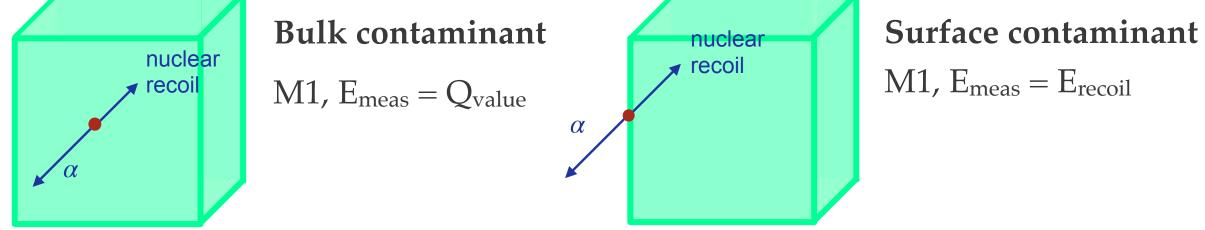


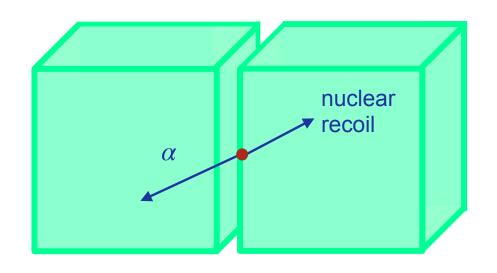
S G **CUORE alpha region - signatures**

- α particles travel short distances (5 MeV α : 10 μ m in Cu)
- Only visible if the contaminant is in the crystal's bulk/ surface or close to it
- Event coincidences study gives information on the source position
- Bulk/surface contaminants produce different signatures in terms of released energy and event multiplicity

$$\alpha$$
 decay: $Q_{value} = E_{\alpha} + E_{recoil}$

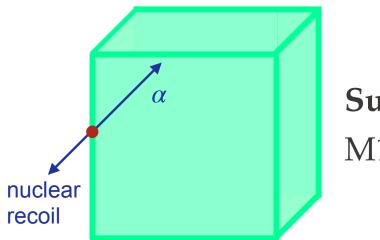






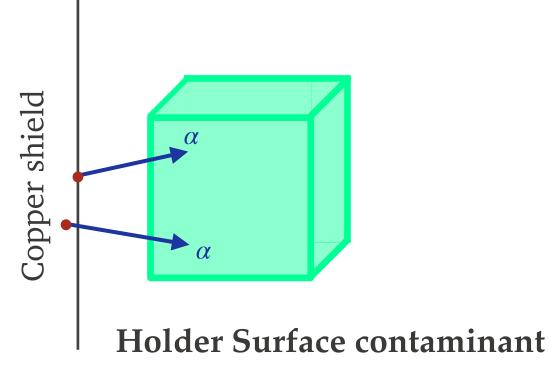
Surface contaminant

M2 single energy: $E^{(1)}_{meas} = E_{\alpha}$, $E^{(2)}_{meas} = E_{recoil}$ M2 total energy: $E_{meas} = Q_{value}$



Surface contaminant

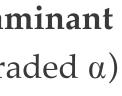
M1, $E_{meas} = E_{\alpha}$



M1, $E_{\text{meas}} \leq E_{\alpha}$ (degraded α)

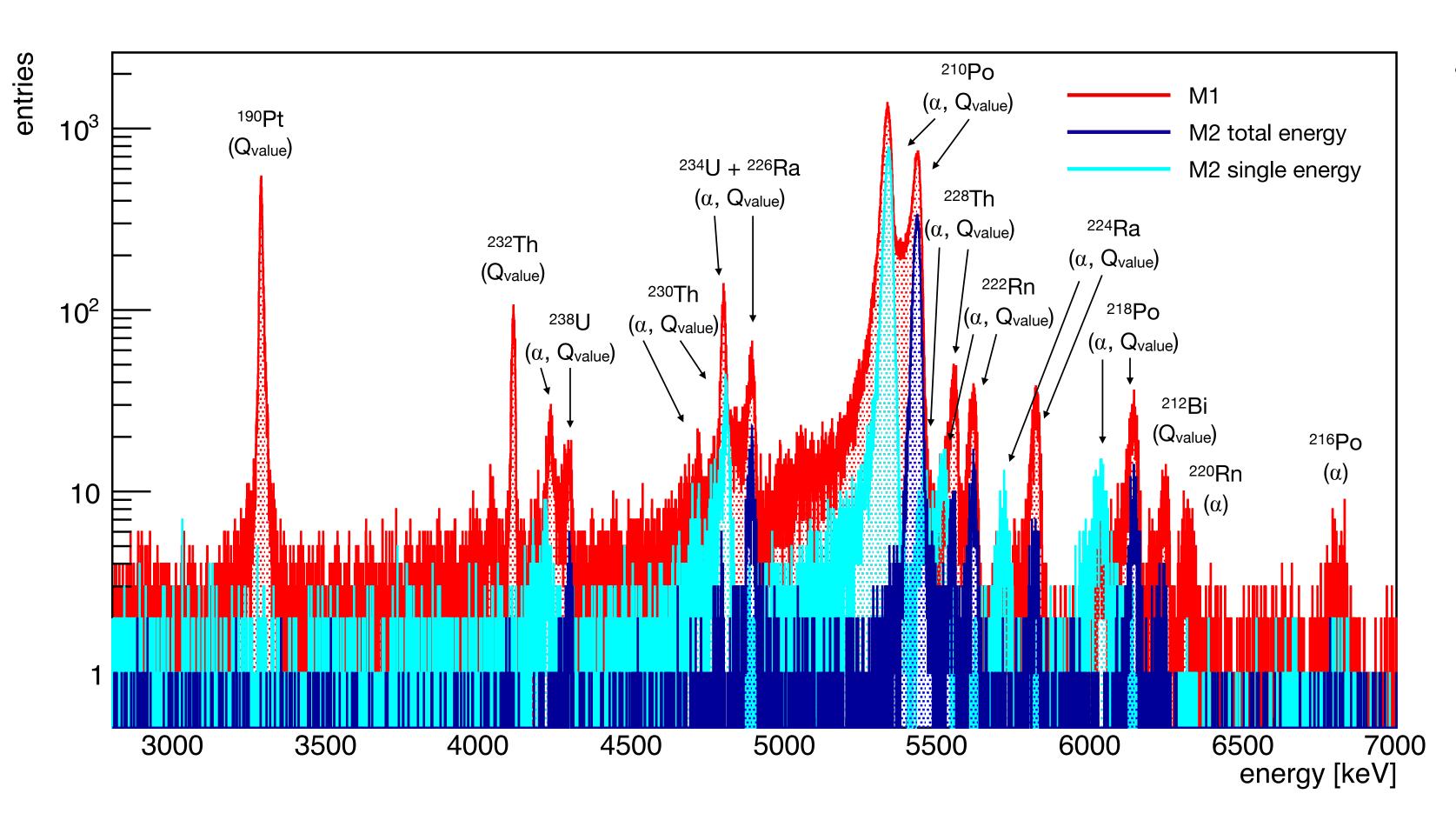








S G **CUORE** alpha region - contributions S



• 238U chain:

Q_{value}: M1 and M2sum <u>α</u>: M1 and M2

crystals' bulk crystals' surface holder's surface

• ²³²Th chain:

Q_{value}: M1 (all elem.) M2sum (not all elem.) $\underline{\alpha}$: M1 and M2 (not all elem.) dominant; contribution

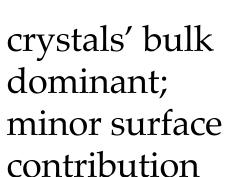
• 190Pt:

<u>Qvalue</u>: M1 only $\underline{\alpha}$: no line



crystals' bulk only







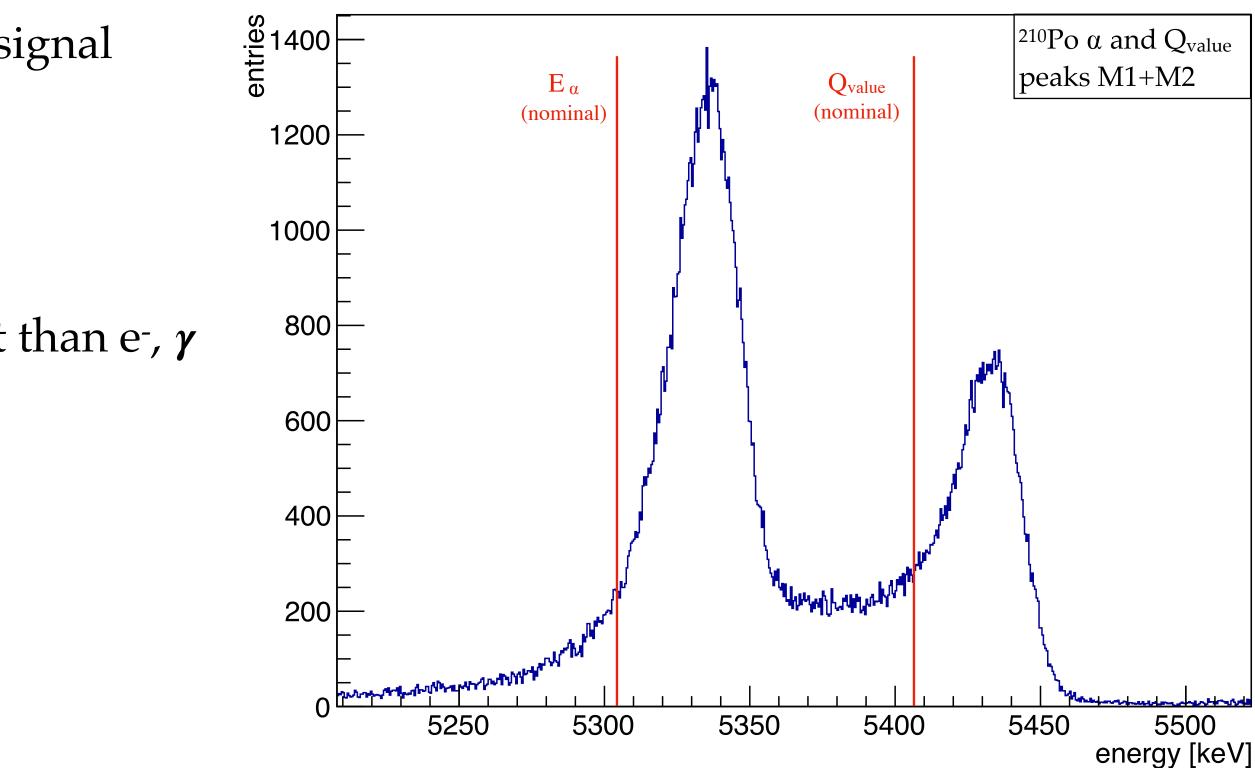
S G **CUORE** alpha region - peaks position in energy

An α particle interacting with TeO₂ crystals generates a signal higher than expected from *γ* calibration

- Measure *electron-equivalent energy* \bullet
- TeO₂ response to α particles appears to be different than e⁻, γ lacksquare

QUENCHING FACTOR:
$$\frac{E_{measured}}{E_{nominal}} \equiv QF$$

GOAL: quantify the QF in CUORE to include the appropriate correction in the Background Model MC simulations



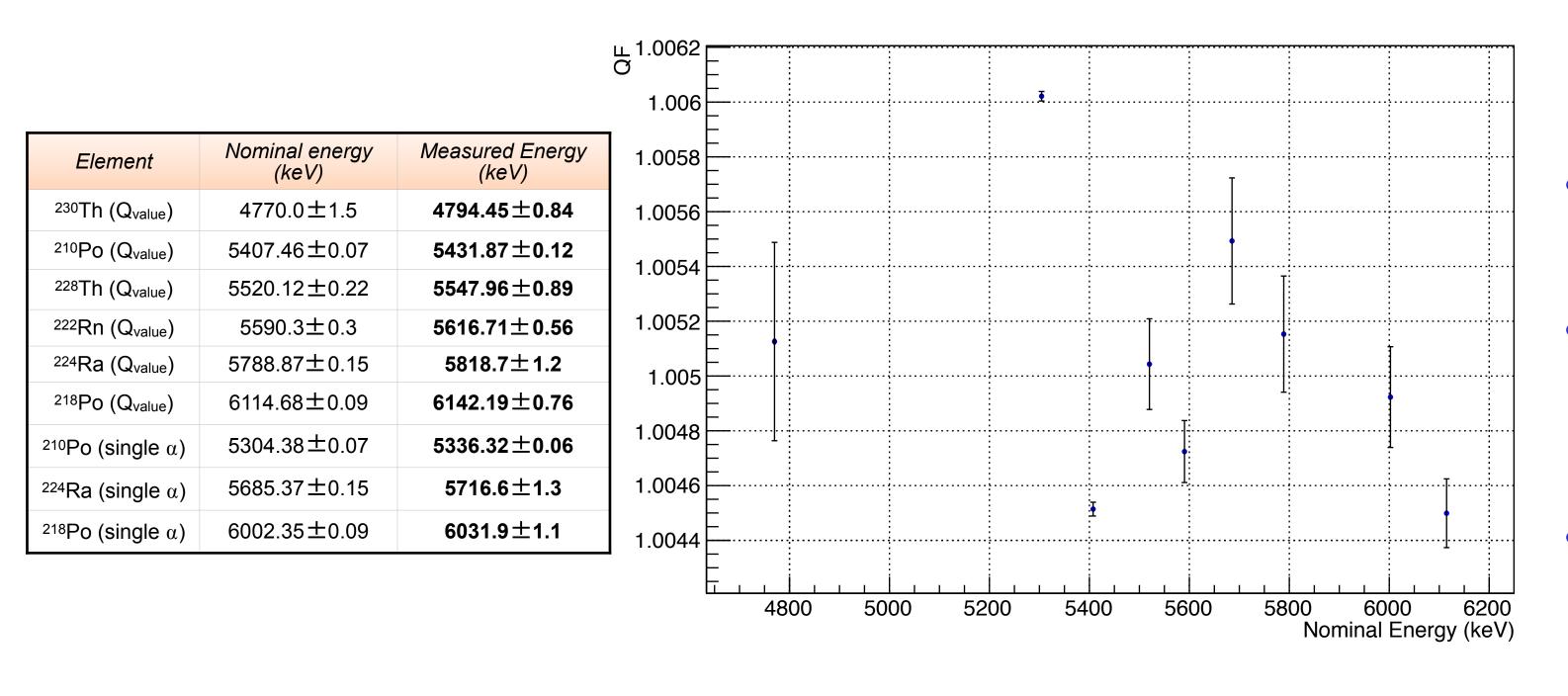


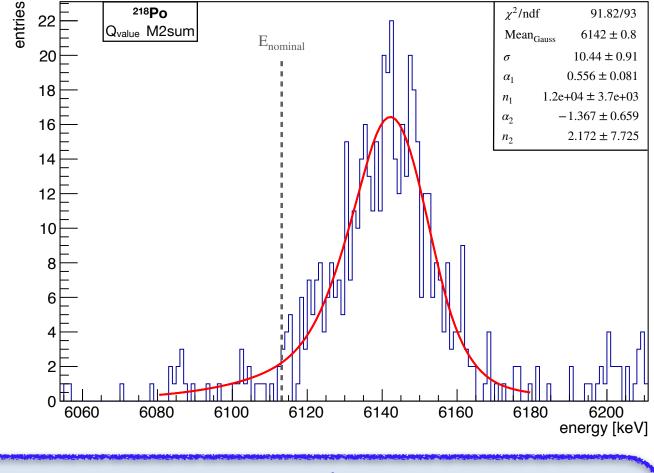




S G **CUORE** alpha region - peaks position in energy

- Estimation of the peaks position in energy fit function: combination of 2 Crystal Ball functions sharing the gaussian parameters
- Crystal Ball function = Gaussian function with a smoothly joined power law tail
- **E**_{measured} estimator: gaussian mean extracted from the fit





Results

- CUORE-0: QF ~ 1.007 **CUORE (this work): QF ~ 1.005**
- Systematic uncertainties under investigation: fit model, choice of E_{measured} estimator, calibration function
- A different QF in CUORE-0 and CUORE could be an effect of the different NTDs bias settings



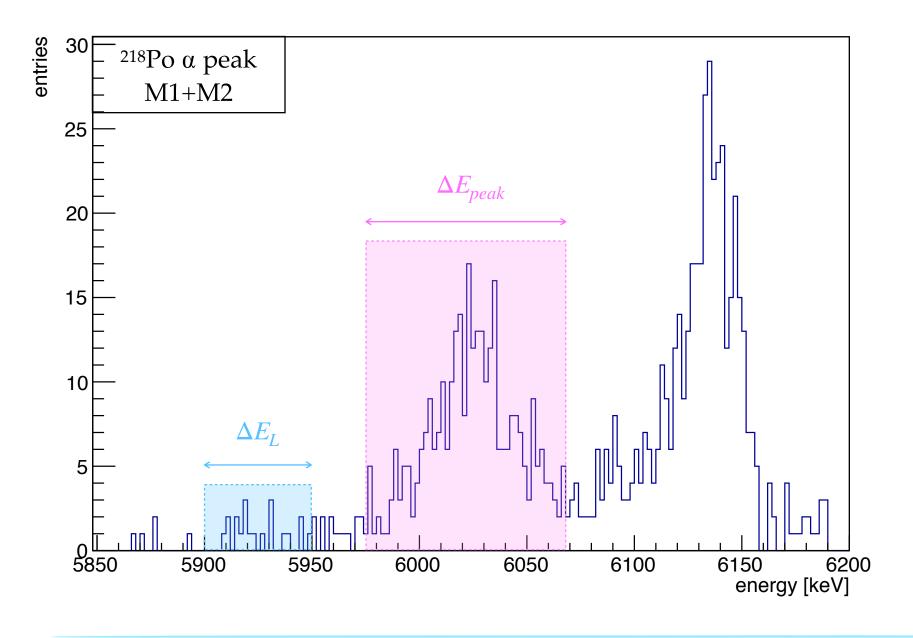


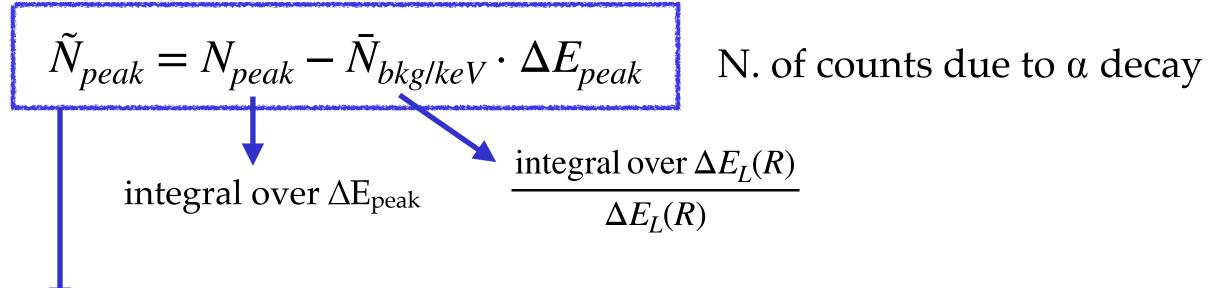




S G CUORE alpha region - secular equilibrium of ²³⁸U chain

- Evaluation of the activity of the ²³⁸U chain isotopes (not enough statistics for ²³²Th)
- Time dependence of the α peaks intensity over a period of 2 years
- Decay rates evaluated by integration of the peaks on spectra for each dataset
- Background subtraction from the peak integral: \bullet





- Scaled by dataset exposure: decay rate in counts/kg/day
- Time at half of the dataset associated



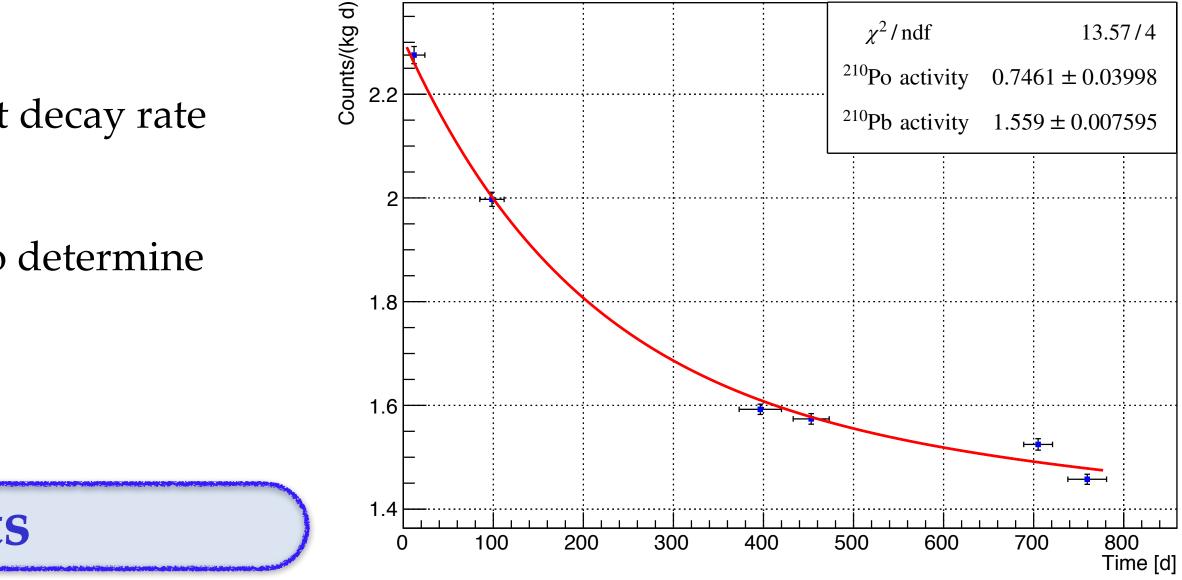


S G CUORE alpha region - secular equilibrium

I determined the activity of ²³⁸U nuclei from their rate distributions over $\Delta t = 2$ years

- ²³⁸U, ²³⁰Th: $T_{1/2} >> \Delta t$, constant rate is expected
- ²¹⁸Po: $T_{1/2} \ll \Delta t$, constant rate is expected, equal to parent decay rate if secular equilibrium occurs
- $^{210}Po + ^{210}Pb$: fit with 2 exponentials on the $^{210}Po \alpha$ peak to determine the two contributions
- Extracted activities scaled by the decay branching ratio

			Results
Element	Half life	Activity (10 ⁻² counts/kg/day)	• 229T T 220TT 1 219T
238	4.47 • 10 ⁹ y	1.72 ± 0.08	• ²³⁸ U, ²³⁰ Th, ²¹⁸ Po
²³⁰ Th	7.54 • 10 ⁴ y	1.31 ± 0.08	 equilibrium do
²¹⁸ Po	3.1 mins	1.48 ± 0.05	• ²¹⁰ Pb and ²¹⁰ Po:
²¹⁰ Pb	22.3 y	155.9 ± 0.8	
²¹⁰ Po	138.4 days	74.6 ± 3.9	 ²¹⁰Po activity ~2
			implanted on su



- o activities are constant and are very similar: secular own to ²¹⁸Po
- two breaking points of the chain
- 2 times lower than ²¹⁰Pb: **secular equilibrium between** ²¹⁰Pb surfaces and daughters not reinstated yet







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- ¹²⁸Te is another $\beta\beta$ emitting tellurium isotope: ¹²⁸Te \rightarrow ¹²⁸Xe
- High natural abundance: 31.75%
- $Q_{\beta\beta} = (866.6 \pm 0.9) \text{ keV}$

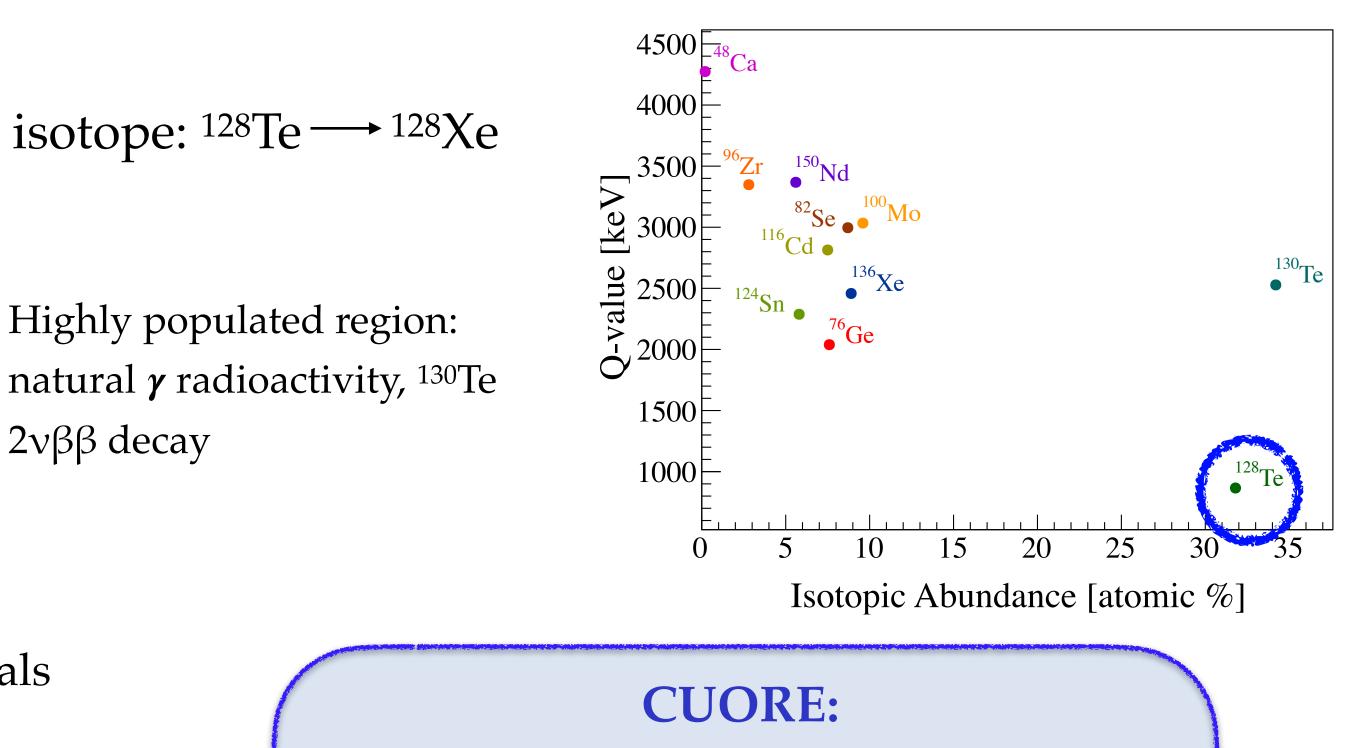
 $2\nu\beta\beta$ decay

- Latest ¹²⁸Te 0νββ decay results:
 - → From direct experiments (MiDBD in 2003, 6.8 kg of TeO₂, 2 crystals) enriched in ¹²⁸Te at 82.3%):

 $T_{1/2}^{0\nu} > 1.1 \cdot 10^{23} \,\mathrm{y}$

➡ From geochemical experiments: (refers to the sum of 2v and 0v modes)

 $T_{1/2}^{128Te} = (2.0 \pm 0.3) \cdot 10^{24} \, \text{y}$



a factor ~10 higher sensitivity is expected, competitive with geochemical results

- N. of ¹²⁸Te $\beta\beta$ emitters in CUORE: $9.519 \cdot 10^{26}$
- ¹²⁸Te mass in CUORE: 188.5 kg

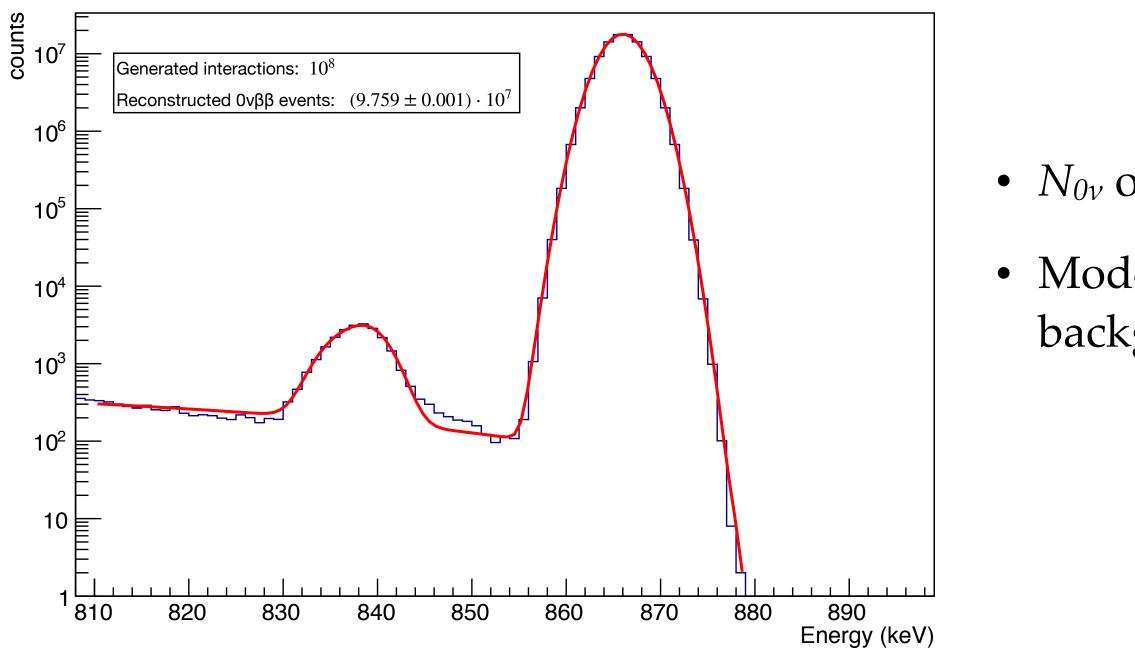
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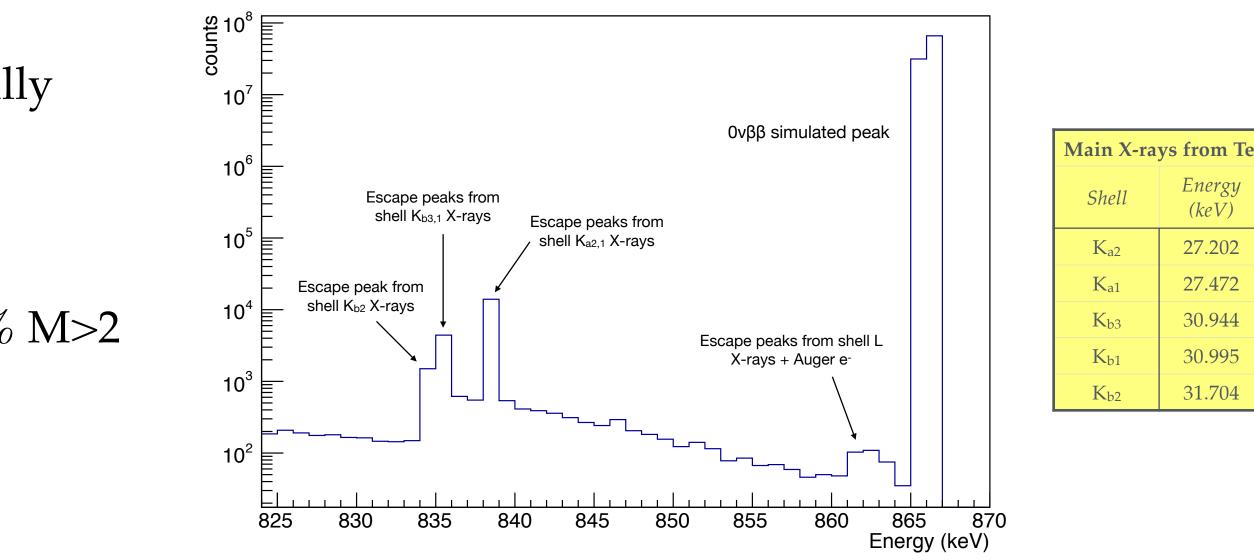


S G Full Containment Efficiency (ϵ_{MC})

- Estimation of the fraction of events where the e- are fully absorbed by the same CUORE crystal
- MC simulation of $N_{MC} = 10^{8} \, {}^{128}\text{Te} \, 0\nu\beta\beta$ decays
- Multiplicities distribution: **98.26**% **M1** 1.24% M2, 0.5% M>2 ullet



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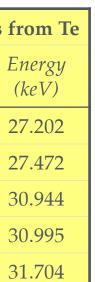
• $N_{0\nu}$ of events reconstructed at $Q_{\beta\beta}$ peak: fit on M1 simulated spectrum

• Model function: 3 X-ray escape peaks + $0\nu\beta\beta$ peak + linear background

Full containment efficiency:

$$\epsilon_{MC} = \frac{N_{0\nu}}{N_{MC}} = (97.59 \pm 0.01)\%$$





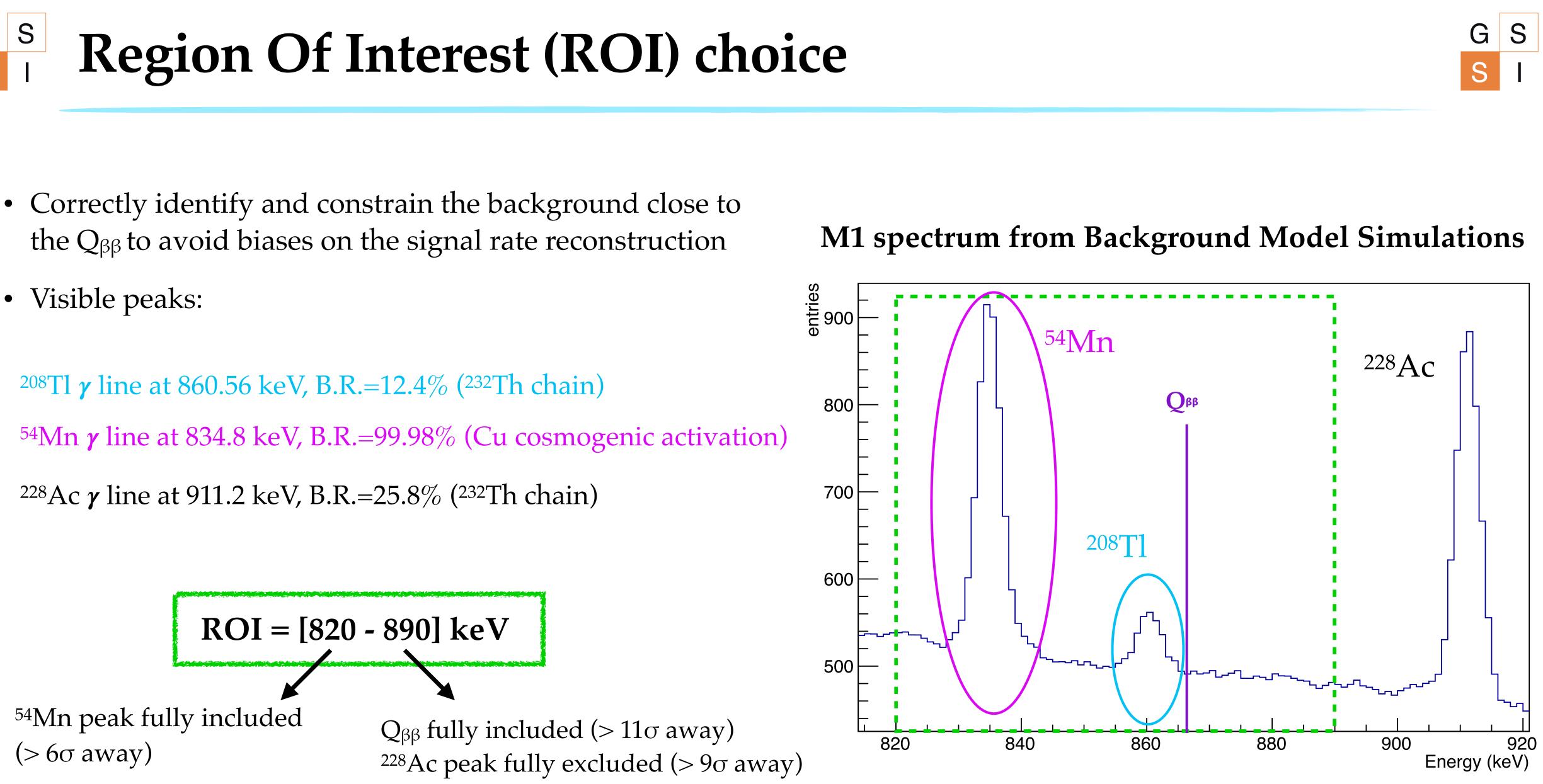




S G

- Visible peaks:

²⁰⁸Tl γ line at 860.56 keV, B.R.=12.4% (²³²Th chain)







- **Bayes' theorem**: it is possible to make an inference on an observable by evaluating its posterior probability distribution from the product of the likelihood function and the prior probabilities

$$P(\overrightarrow{\theta} | D, M) = \frac{1}{\int_{\Theta} \mathcal{D}}$$

BAT software: sampling from the posterior distribution of all parameters with a Markov Chain Monte Carlo

• Bayesian binned fit simultaneously performed on 5 datasets — Analyzed exposure: 309.33 kg·y

 $\frac{\mathscr{L}(D \mid \overrightarrow{\theta}, M) \pi(\overrightarrow{\theta})}{\mathscr{L}(D \mid \overrightarrow{\theta}, M) \pi(\overrightarrow{\theta}) d \overrightarrow{\theta}}$

- *Global mode*: parameter values corresponding to the maximum of the full posterior distribution

- *Marginalized posterior*: posterior distribution integrated over all the parameters with exception of one







Bayesian Fit: model G

- Binned likelihood for each dataset is product of Poissonian terms
- the center of the bin (small bin-width approximation):

$$\mathscr{L} = \prod_{ds} \prod_{i}^{N_{bins}} \frac{\mu_{i}^{n_{i}} e^{-\mu_{i}}}{n_{i}!}$$

$$\mu_{i} = S \cdot f_{S}^{ds}(i) + C_{Mn} \cdot f_{Mn}^{ds}(i) + C_{Tl} \cdot f_{Tl}^{ds}(i) + C_{b} \cdot f_{flat}(i) + f_{linear}(i)$$
Signal peak ⁵⁴Mn peak ²⁰⁸Tl peak Continuous background

• The expected number of counts μ_i for the i-th bin is given by the value of the model function at





S G **Bayesian Fit: parameters**

Modeled with the detector
response function (lineshape)

$$\begin{split}
\frac{C_{h}}{\Delta E_{ROI}} + \frac{m_{ds}}{\Delta E_{ROI}} (E_{i} - E_{1/2}) \\
\mu_{i} &= S \cdot f_{S}^{ds}(i) + C_{Mn} \cdot f_{Mn}^{ds}(i) + C_{Tl} \cdot f_{Tl}^{ds}(i) + C_{b} \cdot f_{flat}(i) + f_{linear}(i) \\
\text{Signal peak} & \text{Summary statements} \\
\text{Signal peak} & S = \Gamma^{0\nu} \cdot \frac{N_{A}}{A_{TeO_{2}}} \cdot \eta_{128} \cdot (M\Delta t)_{ds} \cdot \epsilon_{ds}^{cut} \cdot \epsilon_{MC} \\
\text{sunts:} & (T_{1/2}^{Mn} = 312.2 \ d) & C_{Mn} = \Gamma_{Mn} \cdot e^{-\frac{t_{ds}}{c_{Mn}}} \cdot (M\Delta t)_{ds} \cdot \epsilon_{ds}^{cut} \\
\text{sunts:} & (T_{1/2}^{Tl} = 3.1 \ min) & C_{Tl} = \Gamma_{Tl} \cdot (M\Delta t)_{ds} \cdot \epsilon_{ds}^{cut} \\
\text{subs background counts:} & C_{b} = Bl_{ds} \cdot size_{bin} \cdot (M\Delta t)_{ds} \\
\text{subs background counts:} & m_{ds}
\end{split}$$

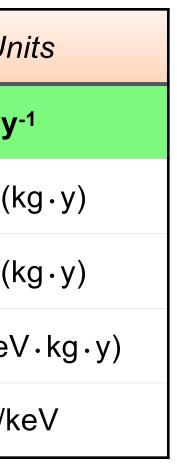
• N. of signal co

- N. of ⁵⁴Mn cou
- N. of ²⁰⁸Tl cou
- N. of continuc
- Linear backgro

Uniform priors on all the parameters (no independent measurements available)

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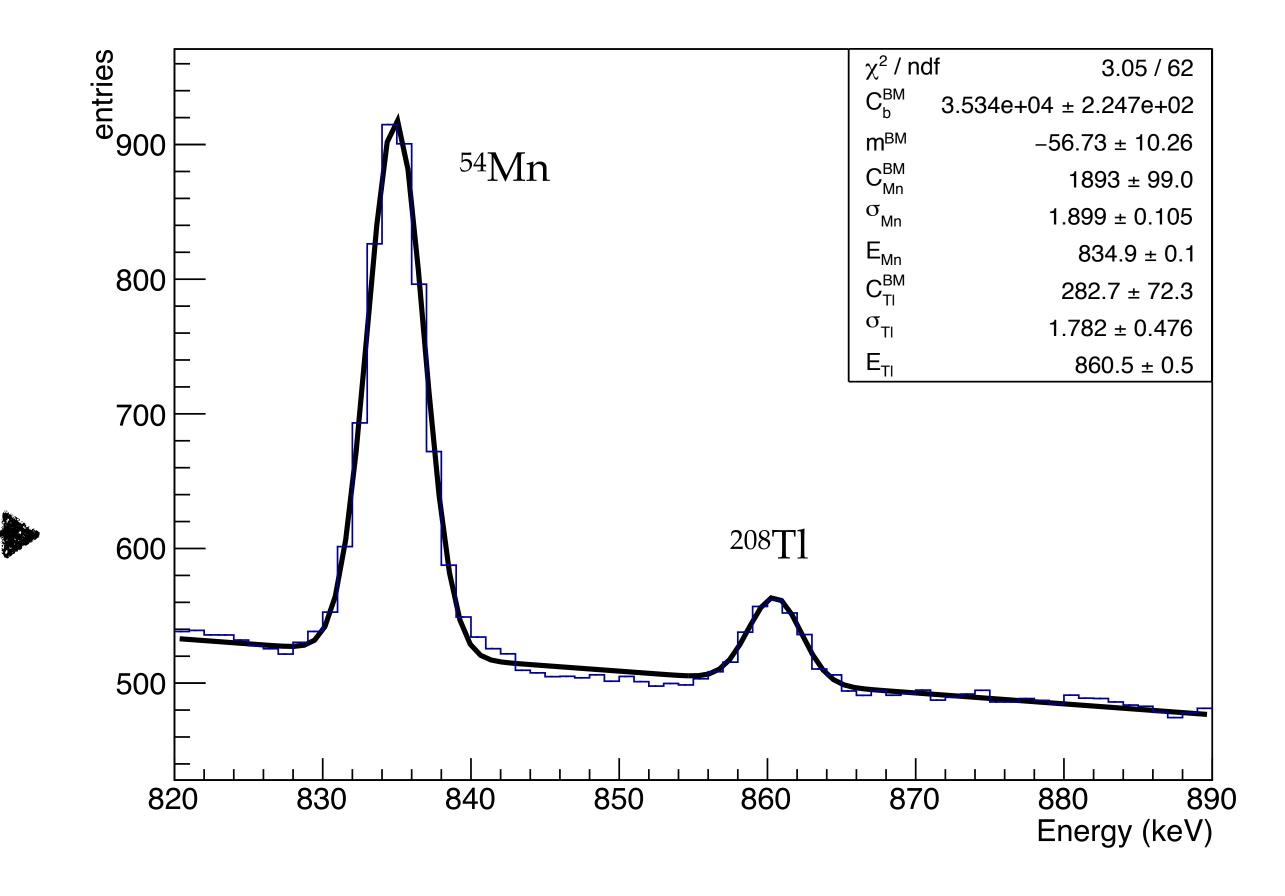
/-1



S G Fit validation: toyMC generation

- Fit model validated on toyMC simulations
- I developed a dedicated code to simulate the various components of the ROI spectrum
- I extracted the parameters to produce the toyMC from a fit on the BM simulations with 2 gaussians + linear background

Parameter	Units	Value for toyMC
$\Gamma^{ ext{toy}}_{Mn}$	cts/(kg · y)	16.27
$\Gamma^{ ext{toy}}_{Tl}$	cts/(kg·y)	0.95
BI ^{toy}	cts/(keV⋅kg⋅y)	1.68
m ^{toy}	1/keV	-0.4







Fit validation: Background reconstruction test

- I generated 10⁴ toyMC spectra with bkg components only (⁵⁴Mn and ²⁰⁸Tl peaks, linear background) according to
- Bayesian fit with signal + background components independently run on each toyMC ($\Gamma^{0\nu} > 0$ only allowed)
- Reconstructed VS generated parameters: the distributions of the global modes was built for each parameter
- The distributions are expected to be centered at the generated ones.

1st test: verify that the fit correctly reconstructs the simulated background contributions

Parameter	Units	Value for toy
$\Gamma^{ m toy}_{Mn}$	cts/(kg · y)	16.27
$\Gamma^{ m toy}_{Tl}$	cts/(kg · y)	0.95
BI ^{toy}	cts/(keV · kg · y)	1.68
m ^{toy}	1/keV	-0.4

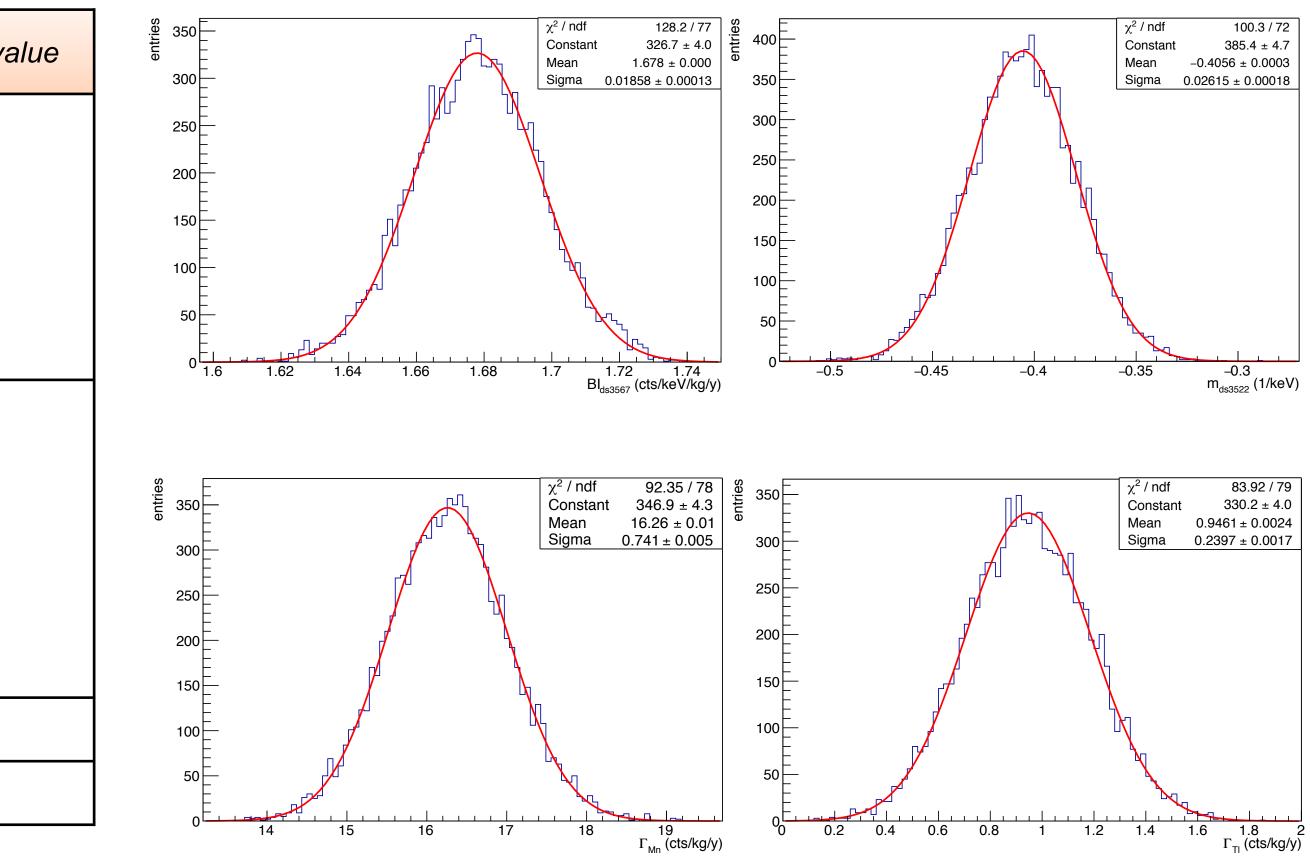






S G Fit validation: Background reconstruction test S

Parameter	Units	Mean Reconstructed Value	Generated va
			Jeneraleu va
Bl _{ds3522} cour	nts/(keV • kg • y)	1.6775±0.0002	
Bl _{ds3552} cour	nts/(keV • kg • y)	1.6782 ± 0.0002	
Bl _{ds3555} cour	nts/(keV • kg • y)	1.6781 ± 0.0002	1.68
Bl _{ds3564} cour	nts/(keV • kg • y)	1.6776 ± 0.0002	
Bl _{ds3567} cour	nts/(keV • kg • y)	1.6779 ± 0.0002	
M _{ds3522}	1/keV	-0.4056 ± 0.0003	
M _{ds3552}	1/keV	-0.4064 ± 0.0003	
M _{ds3555}	1/keV	-0.4064 ± 0.0003	-0.4
Mds3564	1/keV	-0.4060 ± 0.0003	
Mds3567	1/keV	-0.4055 ± 0.0003	
Γ_{Mn}	1/(kg • y)	16.259 ± 0.007	16.27
Γ _{Tl}	1/(kg • y)	0.946 ± 0.002	0.95

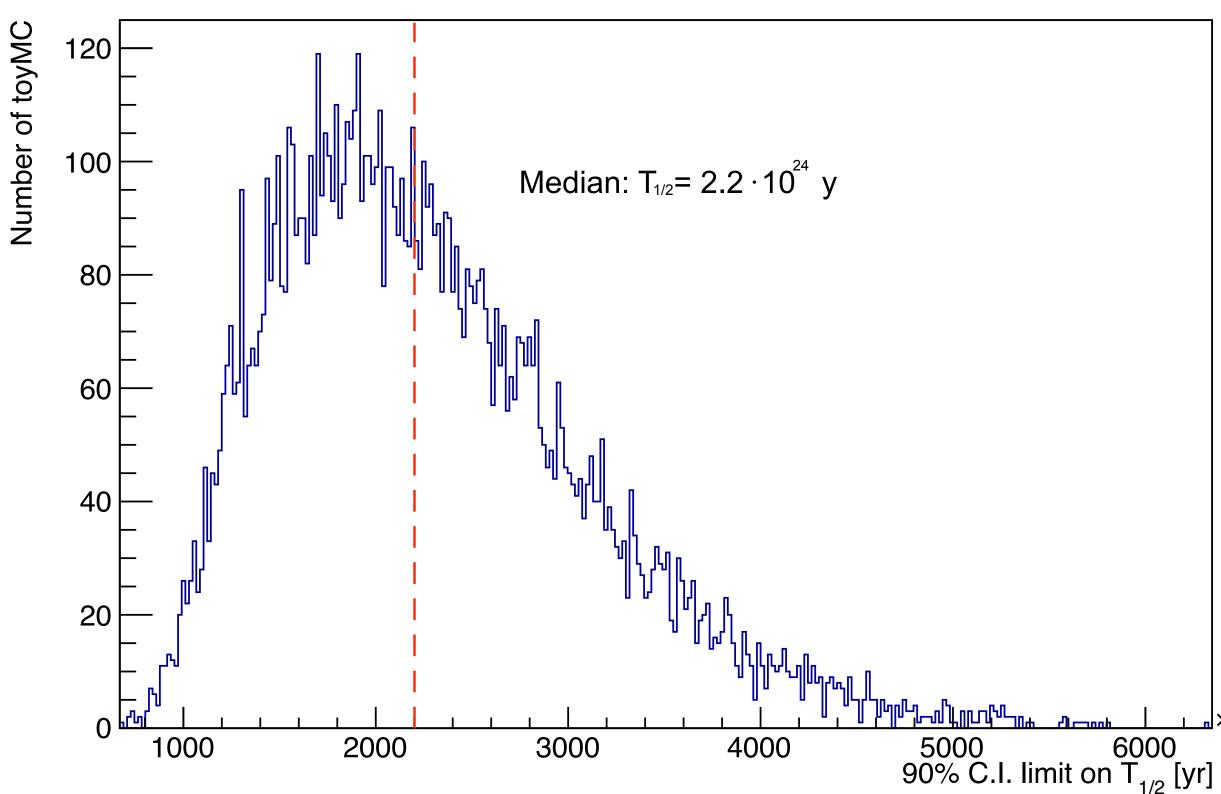








2nd test: CUORE Exclusion Sensitivity (or Limit Setting Sensitivity) extraction procedure median of the distribution of the 90% C.I. limits on $T_{1/2}$



Extraction of the 90% C.I. half life limit from each of the 10⁴ Bayesian fits on the toyMC with BM parameters

Distribution of the half life limits and extraction of the median

CUORE exclusion sensitivity test: $T_{1/2} = 2.2 \cdot 10^{24} y$

This value provides an indication of the CUORE exclusion sensitivity extracted from the measured data.

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×10<sup>21</sup>
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V. Dompè - Search for neutrinoless double beta decay of ¹²⁸Te with the CUORE experiment







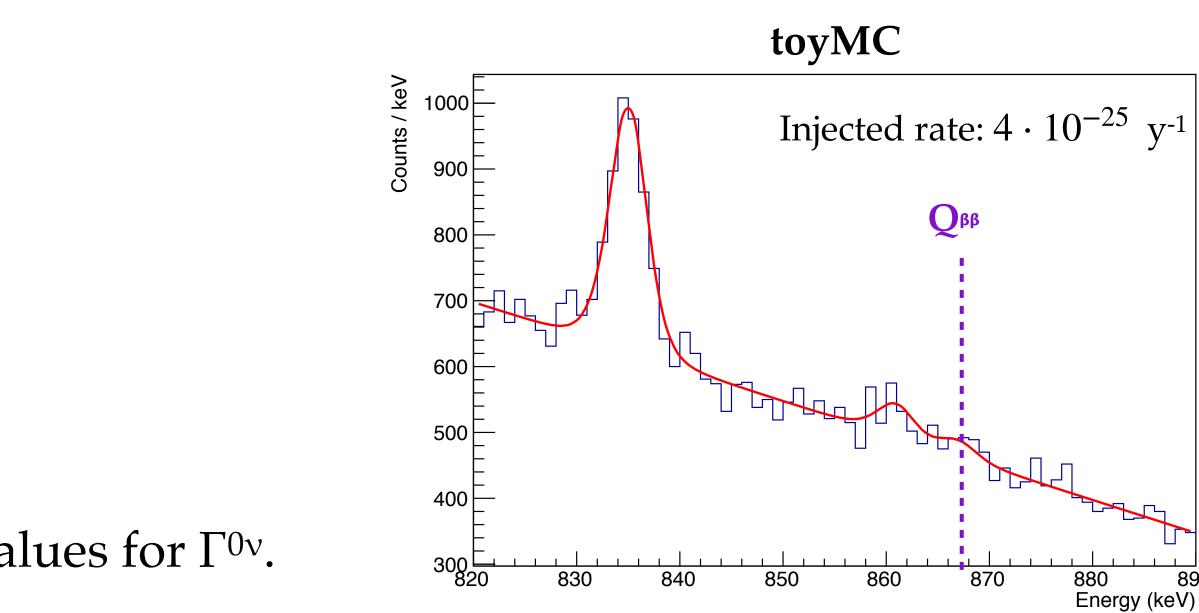
S G Signal Rate reconstruction test

3rd test: study of possible bias in the signal reconstruction

- I introduced the signal component: I tested the Bayesian fit for 5 different signal rates
- I took the test median sensitivity as reference:

$$T_{1/2} = 2.2 \cdot 10^{24} \,\mathrm{y} \longrightarrow \Gamma^{0\nu} = 3.2 \cdot 10^{-25} \,\mathrm{y}^{-1}$$

- Five rates: $(2, 4, 6, 8) \cdot 10^{-25}$, 10^{-24} y⁻¹
- 2000 toyMC with for each rate
- All the toys were fitted allowing also negative values for $\Gamma^{0\nu}$.









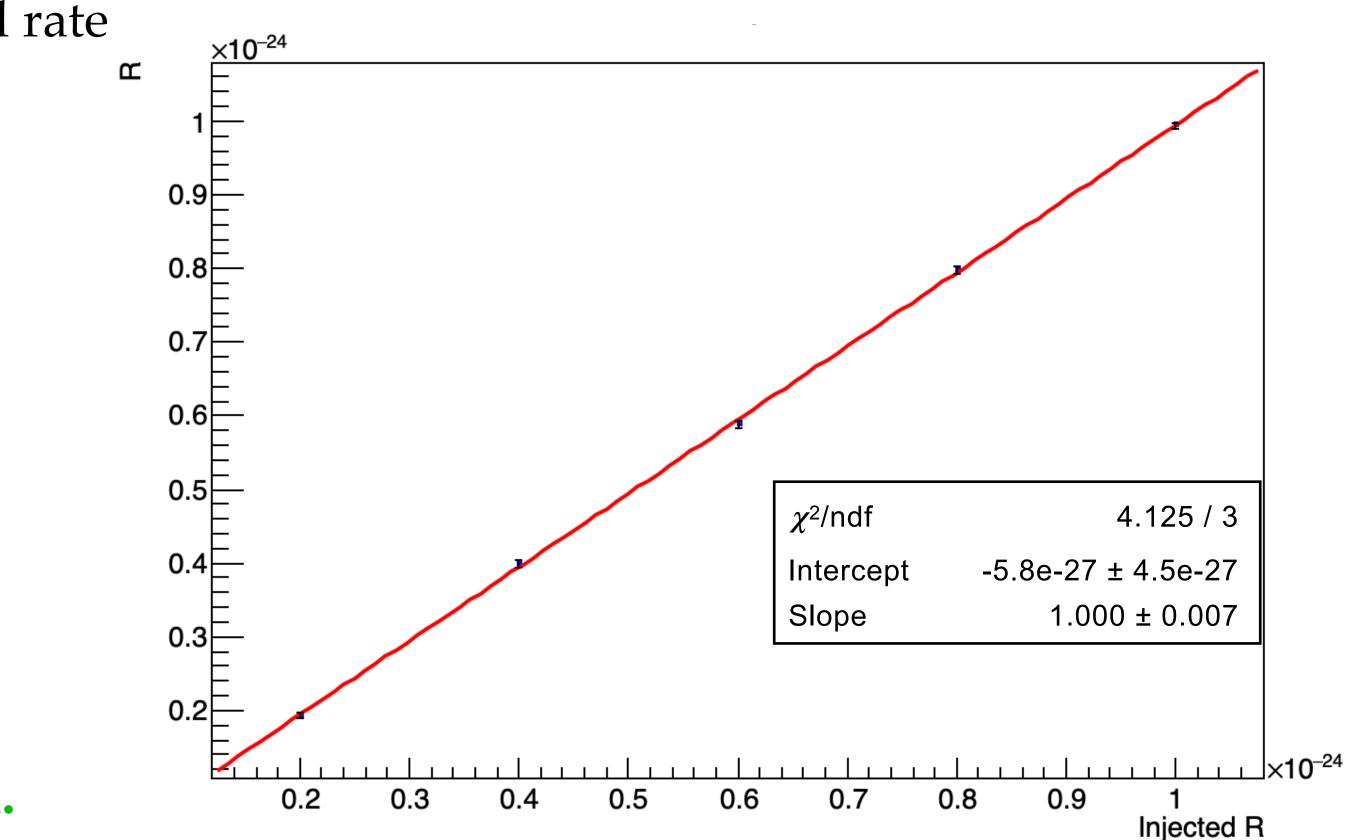
S G Signal Rate reconstruction test

- Distributions of the $\Gamma^{0\nu}$ global modes built for each of the five injected rates
- Mean of each distribution: mean reconstructed rate
- Mean reconstructed VS injected signal rates:

Results of the linear fit:

Intercept: compatible with 0 at 1.3σ Slope: compatible with 1 at < 1σ

No bias is introduced on $\Gamma^{0\nu}$ reconstruction.







S G S Results on 0vββ decay of ¹²⁸Te

Results

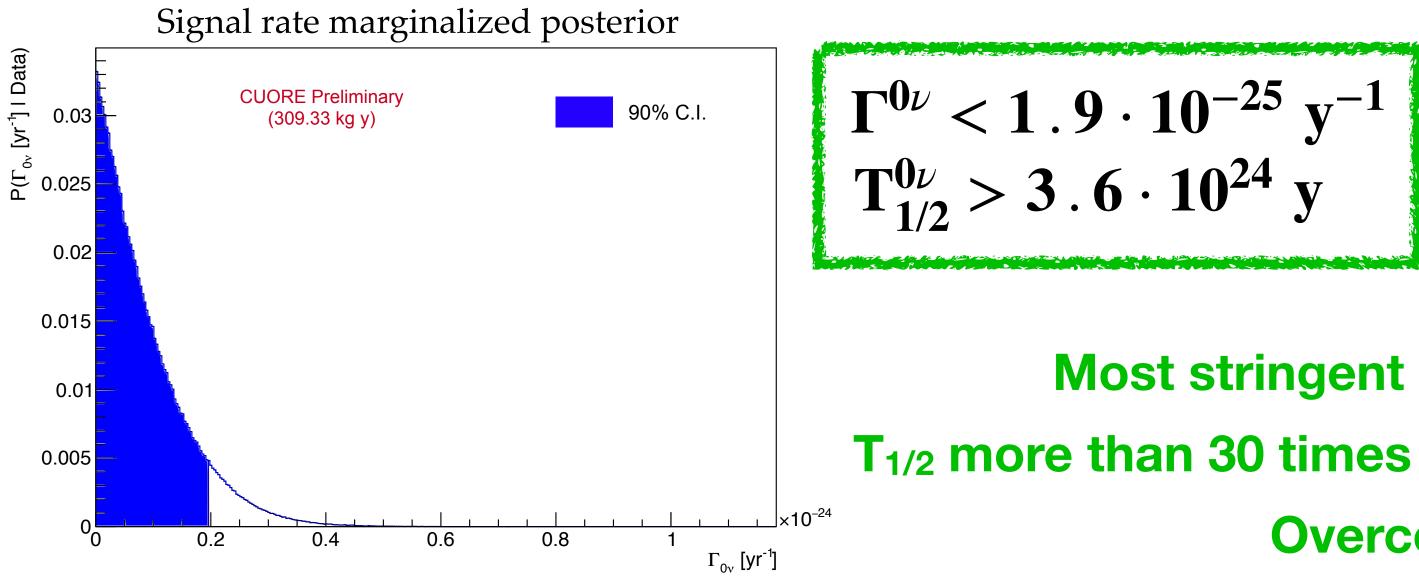




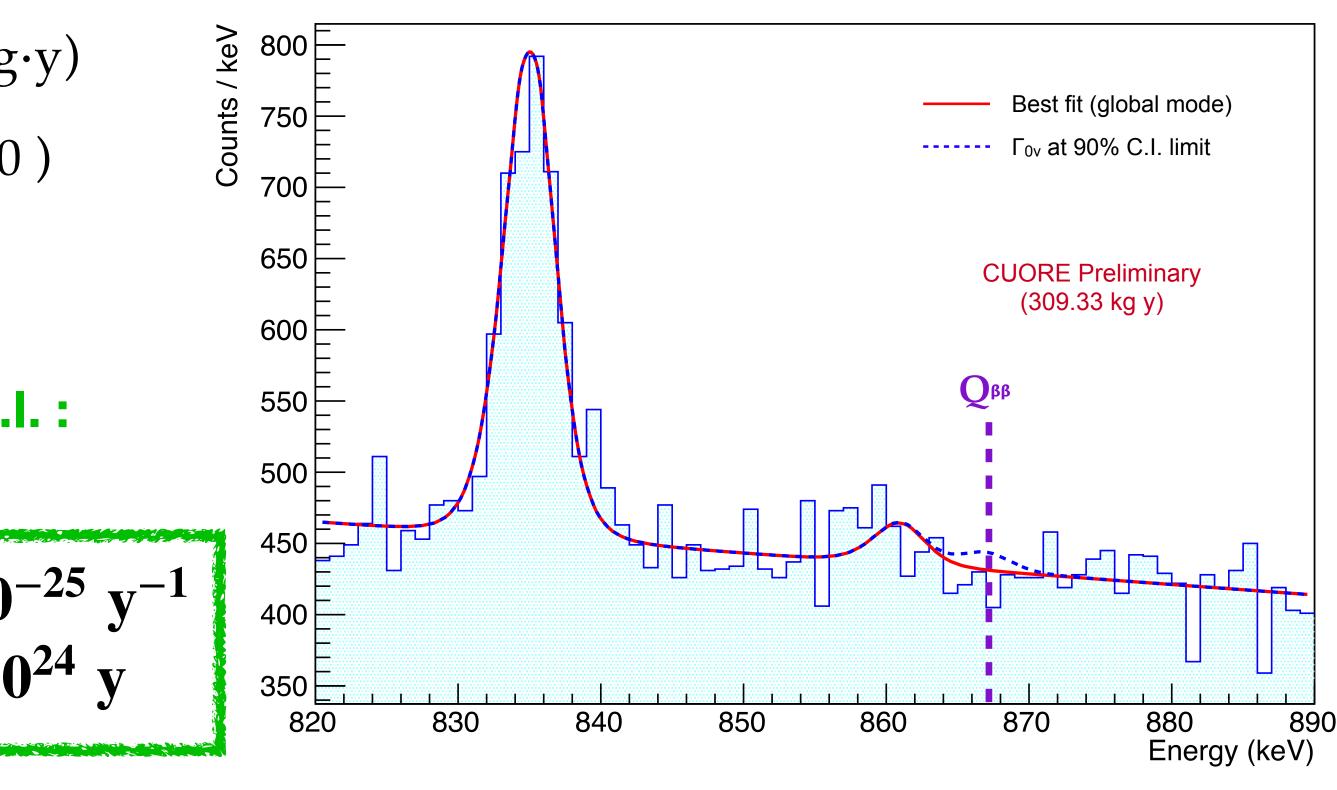
S G Results on 0v\bbeta\beta decay of 128Te

- Bayesian fit on CUORE data (exposure: $309.33 \text{ kg} \cdot \text{y}$)
- $\Gamma^{0\nu}$ prior restricted to physical values only ($\Gamma^{0\nu}>0$)
- No evidence of ¹²⁸Te 0νββ decay is observed.

Limits from marginalized posterior at 90% C.I. :



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Most stringent limits on 0vββ decay of ¹²⁸Te to date $T_{1/2}$ more than 30 times higher than the last direct experiment result **Overcoming geochemical results**

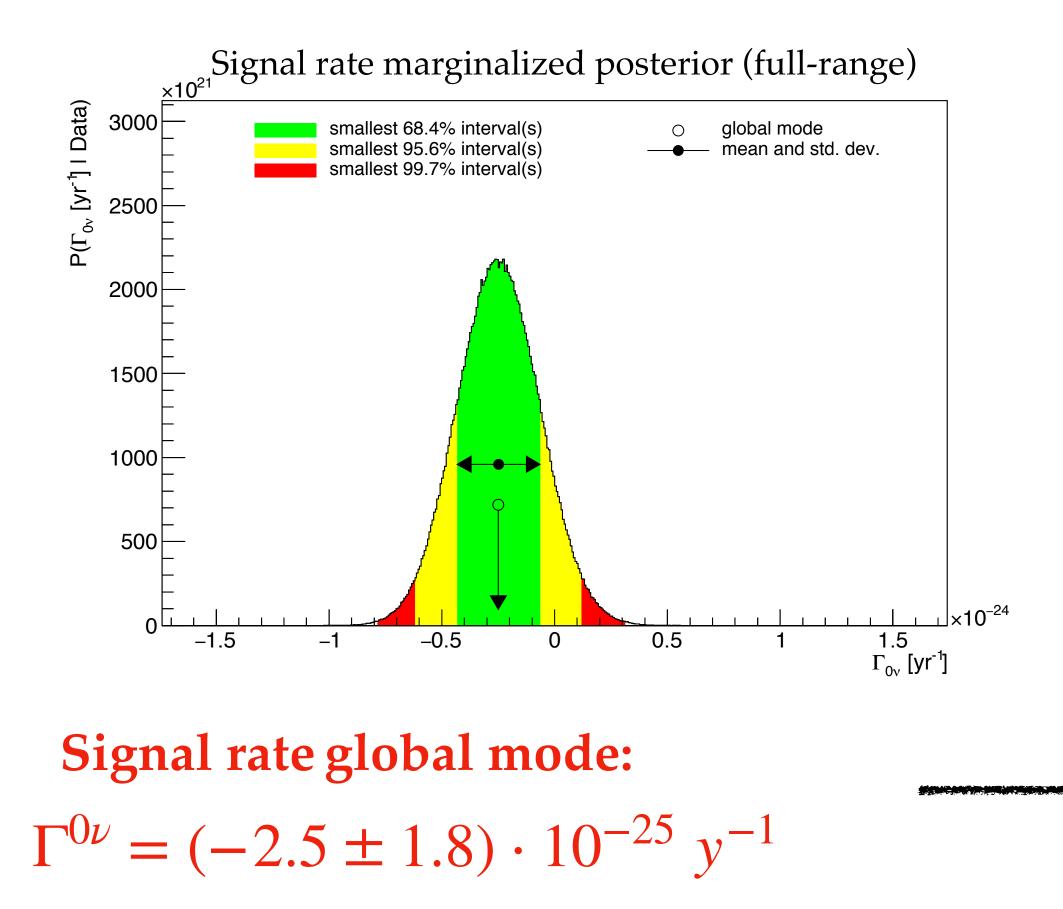


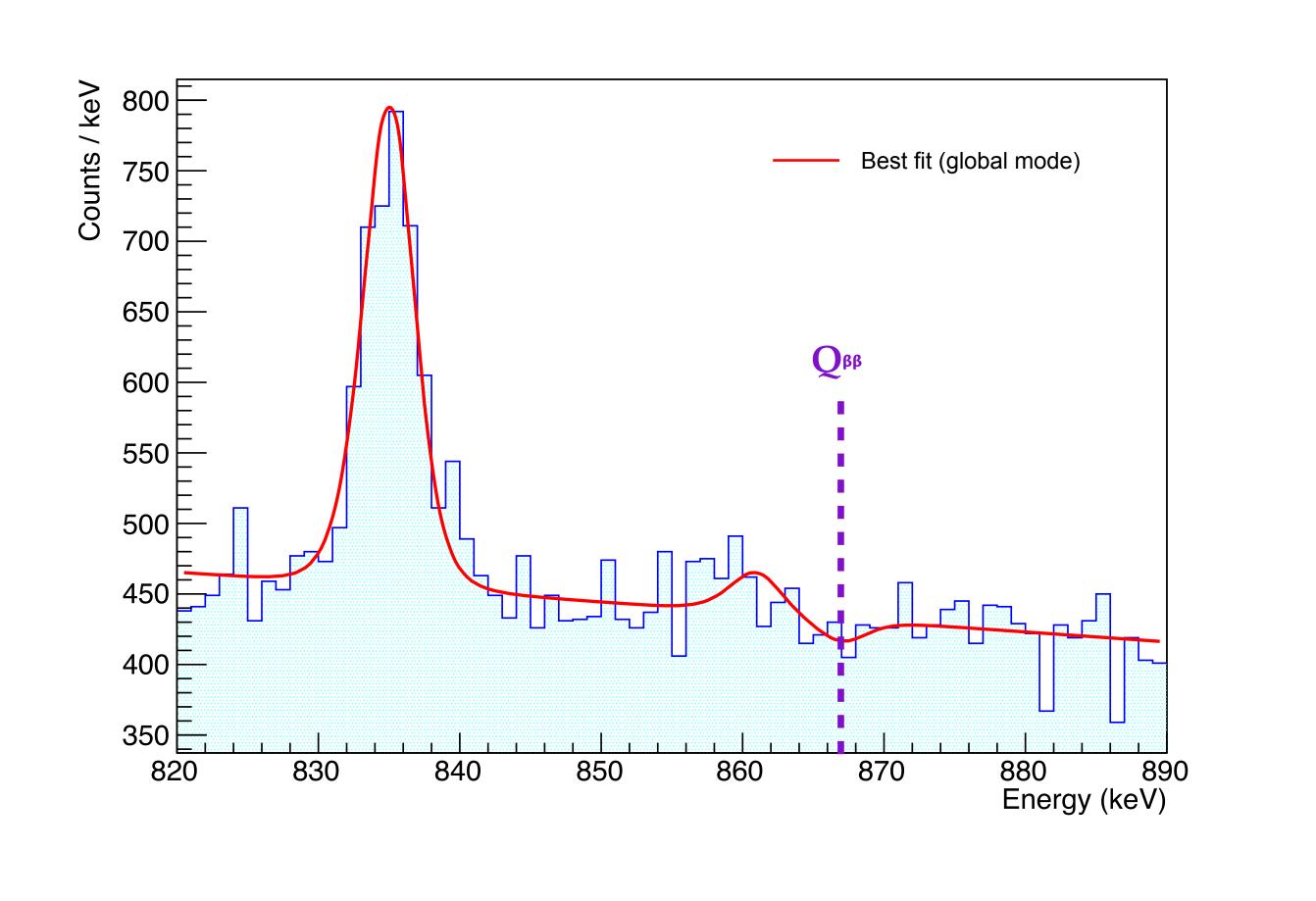




S G Results on 0vββ decay of ¹²⁸Te S

Fit repeated with $\Gamma_{0\nu}$ prior allowed to assume negative values





Signal rate underfluctuation of $\sim 1.4\sigma$ significance

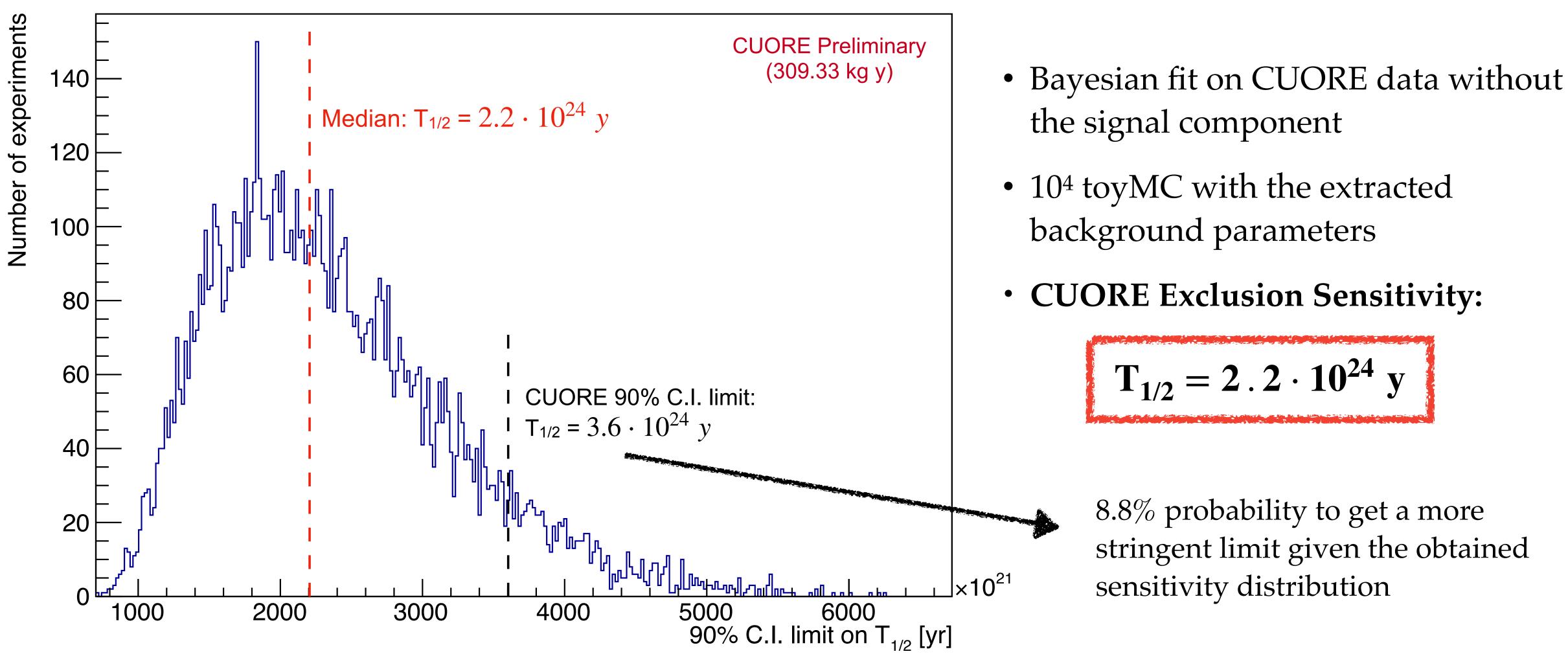






S G **Results on 0vββ decay of ¹²⁸Te**

CUORE Exclusion Sensitivity to 0vββ decay of ¹²⁸Te









S G 0vββ decay of ¹²⁸Te: perspectives

- This analysis has been approved by the CUORE Collaboration
- The paper is in preparation
- Evaluation of the systematics is ongoing and will be included in the final publication.







S G Summary & Conclusions

<u>Analysis of ¹²⁸Te 0vββ decay in CUORE with 309.33 kg · y exposure:</u>

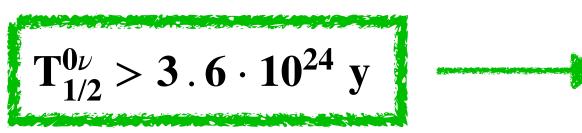
- No evidence of $0\nu\beta\beta$ decay
- NEW Bayesian 90% C.I. lower limit on ¹²⁸Te 0v $\beta\beta$ decay: $T_{1/2}^{0\nu} > 3.6 \cdot 10^{24} \text{ y}$
- More than a factor 30 improvement with respect to the previous direct limit (MiDBD)
- Overcoming the limits obtained from geochemical experiments for the first time
- CUORE exclusion sensitivity: $T_{1/2} = 2 \cdot 2 \cdot 10^{24} \text{ y} = 8.8\%$ probability of a more

<u>CUORE Background Model - α region:</u>

- Identification of the contaminations and their position (surface, bulk)
- *α* **quenching in CUORE**: investigation and quantification
- Secular equilibrium and breaking points in radioactive chains (²³⁸U)

Main CUORE results on ¹³⁰Te 0νββ and 2νββ decay:

- **Data production**: online data monitoring and full reprocessing analysis sequence
- **Detector optimization**: vibrational noise reduction by tuning the Pulse Tubes phases



stringent limit

Will be included in the **CUORE Background Model**

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MOST STRINGENT IN LITERATURE







Backup slides





S G Systematics in ¹²⁸Te 0vββ decay fit

The study of the following systematics is currently ongoing:

¹²⁸Te $Q_{\beta\beta}$ energy Containment efficiency (ϵ_{MC}) ¹²⁸Te Isotopic abundance (η) **Cut Efficiency** do Lineshape parameters t0

- ϵ_{MC} , η , ϵ_{cut} : nuisance parameters in the fit
- $Q_{\beta\beta}$, Lineshape parameters: I repeat the fit varying the value of the systematic parameters, then I merge the posteriors weighting on the respective prior

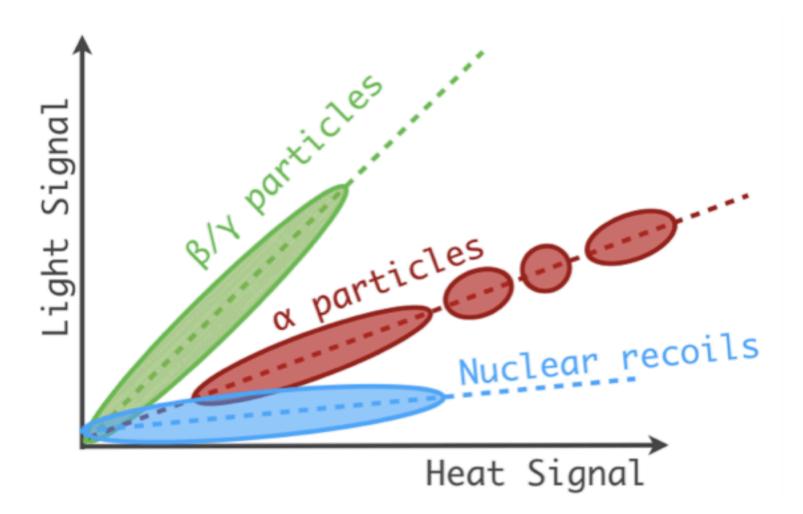
Fit model	90% C.I. limit on T _{1/2} (y)
Stat. Only	$> 3.63 \cdot 10^{24}$
w/ ¹²⁸ Te Q _{ββ}	$> 3.20 \cdot 10^{24}$
w/ ¹²⁸ Te isotopic abundance	$> 3.58 \cdot 10^{24}$
w/ Containment Efficiency	$> 3.63 \cdot 10^{24}$



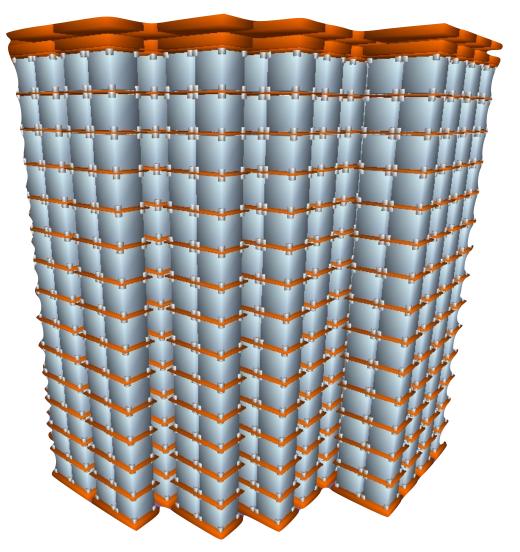


S G **CUPID: Cuore Upgrade with Particle Identification**

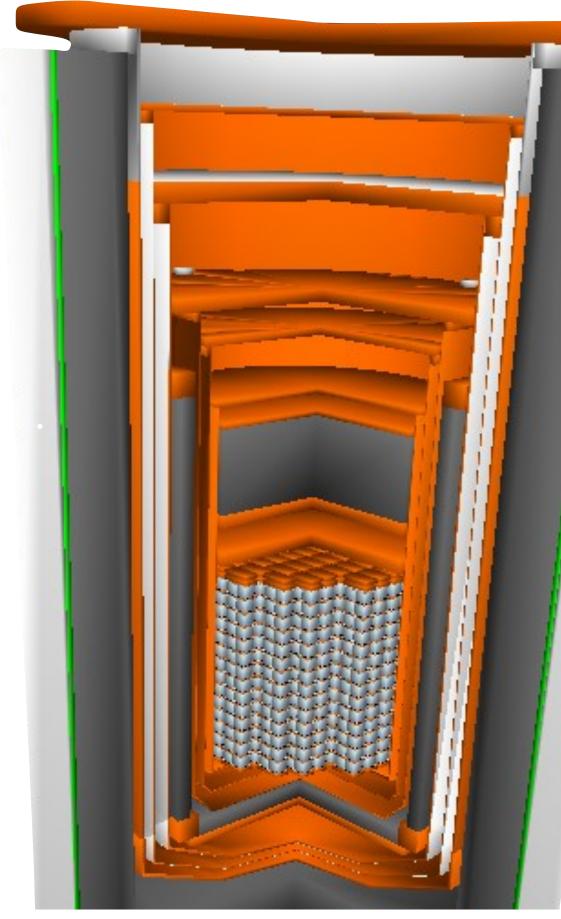
- $0\nu\beta\beta$ of ¹⁰⁰Mo, $Q_{\beta\beta} = (3034.40 \pm 0.17)$ keV ----- above γ radioactivity
- Ton-scale array of ~1500 Li₂¹⁰⁰MoO₄ scintillating bolometers (240 kg of ¹⁰⁰Mo) + light detectors
- Will be hosted by the CUORE cryostat
- **Background-free technology**: simultaneous read-out of thermal signal + scintillation light allows to reject α background



CUPID towers



CUORE cryostat









G Lineshape function in Binned Fit

- In the unbinned fit, the lineshape response function is defined as a continuous function of energy
- A lineshape function is evaluated for each detector-dataset pair, and its value is computed for each event.
- The binned fit does not operate on each event, but on histograms: one histogram for each dataset
- In the binned fit, the lineshape should be applied to each bin
- An average lineshape function *f*^{*ds*} mediated over all the detectors *c* is defined for each dataset:

$$f_{j}^{ds}(i) = \frac{\sum_{c} f_{j}^{c-ds}(i) \cdot exposure_{c-ds} \cdot size_{bin}}{\sum_{ch} exposure_{c-ds}}$$

$$j = S, Mn, Tl$$





S G See-saw mechanism

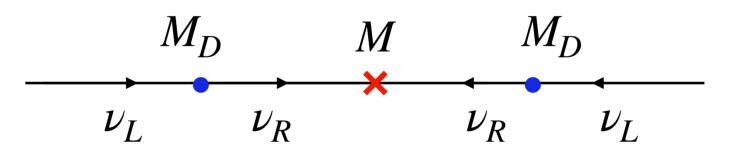
Dirac mass term: $M_D \sim mass of charged$ leptons

M_{I}	ſ

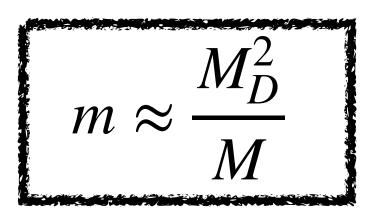
 ν_R

 u_L

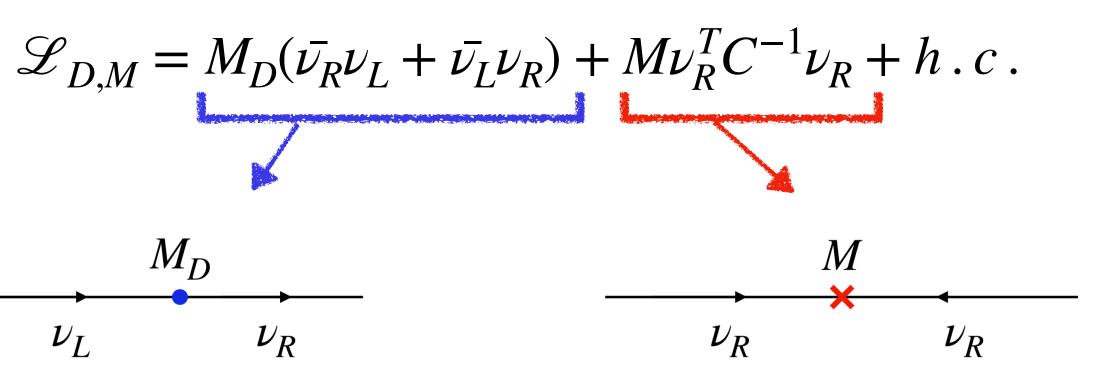
Combining the two vertexes:



Mass induced for v_L :

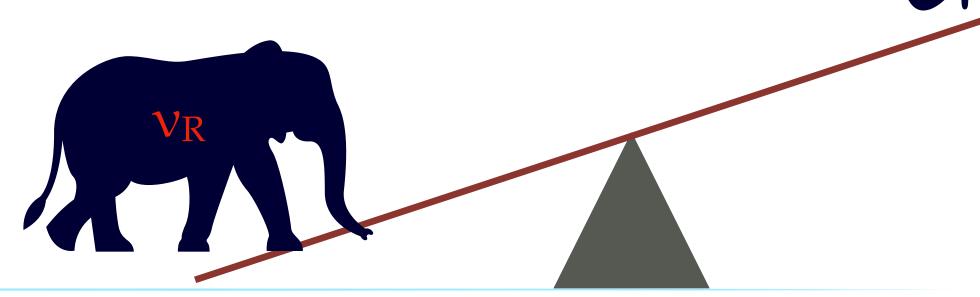


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Majorana mass term: $M = \text{mass of } v_{\text{R}}$

Can be interpreted as a Majorana mass term for v_L









Leptogenesis

- The observation of a lepton number violating process would suggest that leptons played an important role in the creation of matter - antimatter asymmetry in the Universe
- The asymmetry could have been generated in leptons and possibly propagated to the baryonic sector

Sakharov conditions:

- 1. Lepton number is violated
- 2. C and CP are violated
- 3. Universe not in thermal equilibrium





G S J Zero-background counts condition

$$N_{S} = \frac{ln2}{T_{1/2}} \epsilon \eta \frac{N_{A}}{A} (N_{B} = b \Delta E (M\Delta t))$$

Scenario limited by background statistical fluctu

$$N_B \gg 1$$
 : signal is observable if $N_S \ge \sqrt{N_B}$
 $S^{0\nu}(n_{\sigma}) = \frac{\ln 2}{n_{\sigma}} \epsilon \eta \frac{N_a}{A} \sqrt{\frac{M\Delta t}{b\Delta E}}$

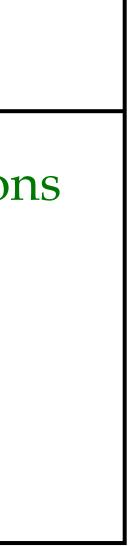
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$M\Delta t$) N. of signal counts

N. of background counts

uations	Zero-background condition		
B	$N_B \leq \mathcal{O}(1)$: no background fluctuation $S_{0-bkg}^{0\nu} = \frac{\ln 2}{n_\sigma} \epsilon \eta \frac{N_a}{A} M \Delta t$		







S G **Conversion to effective Majorana mass** S

$$m_{\beta\beta} = \left| e^{i\eta_1} | U_{e1}^2 | m_1 + e^{i\eta_2} | U_{e2}^2 | m_2 + | U_{e3}^2 | m_3 \right|$$

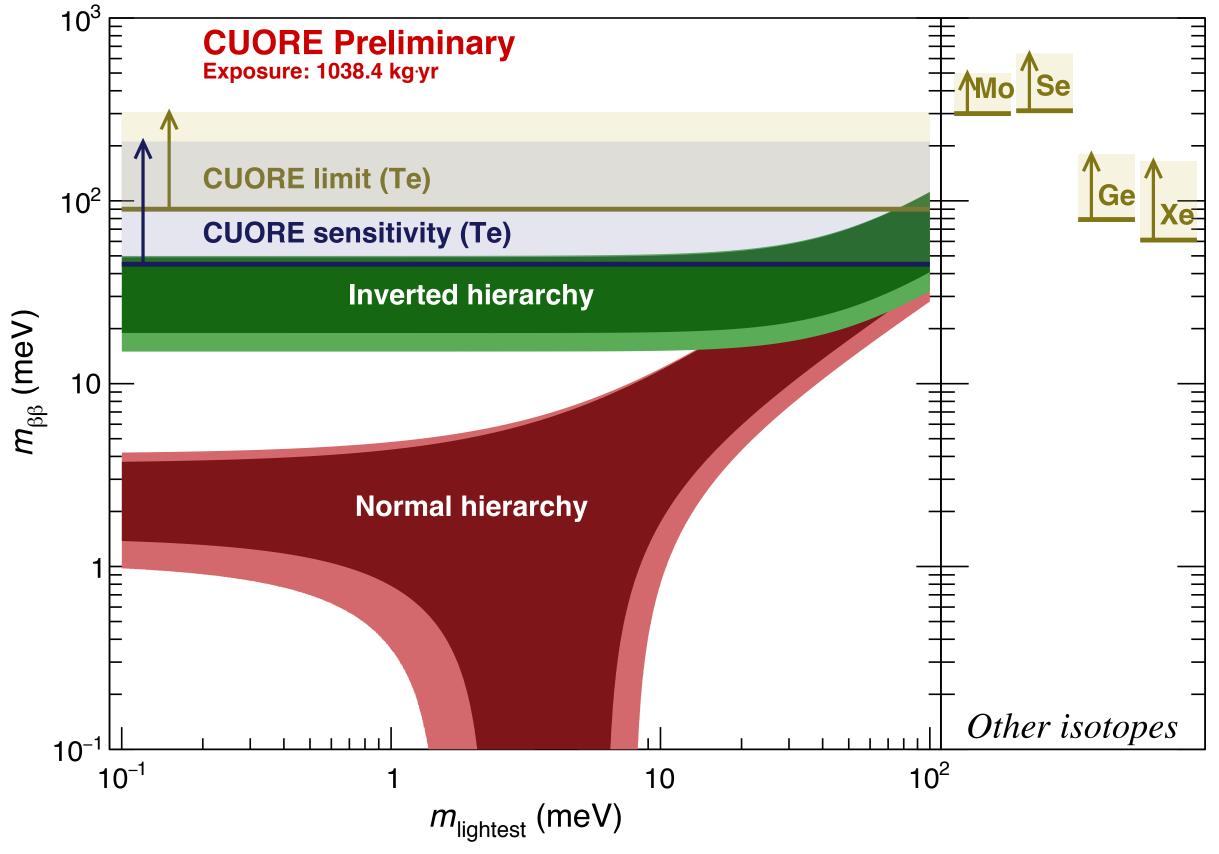
$$\Gamma^{0\nu} = G^{0\nu} |M^{0\nu}|^2 \frac{\langle m_{\beta\beta} \rangle^2}{m_e^2}$$

¹³⁰Te :
$$m_{\beta\beta} < (90 - 305) \text{ meV}$$

¹²⁸Te : $m_{\beta\beta} < (1.1 - 3.0) \text{ eV}$

Ref. NME

[1] Menendez et al. NPA 818 (2009)130	ISM(-StMa)
[2] Neacsu and Horoi PHYS. REV. C 91, 024309 (2015)	ISM(-CMU)
[3] Faessler et al. JoP G: Nucl. Part. Phys. 39 (2012) 124006	QRPA-T
[4] Fang et al. PRC 83 (2011) 034320	
[5] Simkovic et al. PHYS. REV. C 87, 045501 (2013)	
[6] Suhonen et al. JoP G: Nucl. Part. Phys. 39 (2012) 124005	QRPA-J
[7] Hyva rinen and Suhonen PHYS. REV. C 91, 024613 (2015)	
[8] Iachello et al. PRC 87 (2013) 014315	IBM
[9] Barea et al. PHYS. REV. C 91 034304 (2015)	
[10] P.K. Rath et al. PHYS. REV. C 82 064310 (2010)	PHFB
[11] T.R. Rodriguez et al. Phys. Rev. Lett. 105 252503 (2010)	GCM



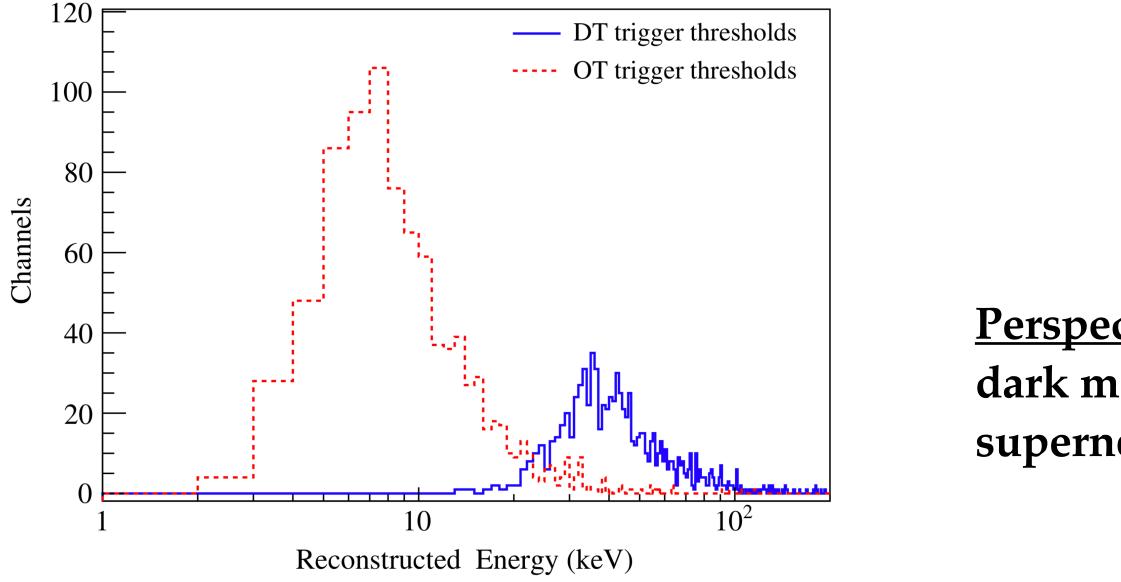




S G **Derivative and Optimum Trigger**

Derivative trigger:

- Fires when the baseline slope is above threshold for a certain number of samples (no info about signal pulse shape)
- Threshold: tens of keV \bullet



Optimum trigger:

- Threshold trigger applied on waveform previously filtered with OF
- Filtered data have higher SNR
- Distinguishes between physical and non-physical pulses
- Threshold: few keV

Perspectives of lowering CUORE thresholds: dark matter searches (WIMP - Te/O nuclei scattering), solar axions, supernova neutrinos





S G Results on 0v\beta\beta decay of 128Te S

 $(\Gamma_{0\nu} > 0)$

Parameter	Units	Global mode	Marginalized mode
L ₀ ^	1/y	1.5 ·10 ⁻³² ± 6.8 · 10 ⁻²⁶	$2^{+103}_{-2}\cdot 10^{-27}$
BI _{ds3522}	cts/(keV.kg.y)	1.48 ± 0.02	$1.48^{+0.02}_{-0.02}$
Bl _{ds3552}	cts/(keV.kg.y)	1.43 ± 0.02	$1.43^{+0.02}_{-0.02}$
Bl _{ds3555}	cts/(keV.kg.y)	1.49 ± 0.02	$1.49^{+0.02}_{-0.02}$
Bl _{ds3564}	cts/(keV.kg.y)	1.48 ± 0.02	$1.48^{+0.02}_{-0.02}$
Bl _{ds3567}	cts/(keV.kg.y)	1.26 ± 0.02	$1.26^{+0.02}_{-0.02}$
Mds3522	1/keV	-0.07 ± 0.03	$-0.06^{+0.02}_{-0.03}$
Mds3552	1/keV	-0.06 ± 0.03	$-0.06^{+0.03}_{-0.03}$
Mds3555	1/keV	-0.08 ± 0.03	$-0.08^{+0.03}_{-0.03}$
Mds3564	1/keV	-0.04 ± 0.03	$-0.04^{+0.03}_{-0.02}$
Mds3567	1/keV	-0.12 ± 0.03	$-0.12^{+0.03}_{-0.03}$
۲Mn	cts/(kg.y)	15.3 ± 0.7	$15.3^{+0.7}_{-0.6}$
ΓΠ	cts/(kg.y)	0.5 ± 0.2	$0.5^{+0.2}_{-0.2}$









S G Results on 0v\bbeta\beta decay of 128Te S

Parameter	Units	Global mode	Marginalized mode
L ₀ ^	1/y	(-2.5 ± 1.8) · 10 ⁻²⁵	$-2.3^{+1.6}_{-2.1}\cdot 10^{-25}$
BI _{ds3522}	cts/(keV.kg.y)	1.48 ± 0.02	$1.48^{+0.02}_{-0.02}$
BI _{ds3552}	cts/(keV⋅kg⋅y)	1.44 ± 0.02	$1.43^{+0.02}_{-0.02}$
BI _{ds3555}	cts/(keV⋅kg⋅y)	1.50 ± 0.02	$1.50^{+0.02}_{-0.02}$
BI _{ds3564}	cts/(keV.kg.y)	1.48 ± 0.02	$1.48^{+0.02}_{-0.02}$
BI _{ds3567}	cts/(keV.kg.y)	1.26 ± 0.02	$1.26^{+0.02}_{-0.02}$
M _{ds3522}	1/keV	-0.06 ± 0.03	$-0.06^{+0.03}_{-0.03}$
m _{ds3552}	1/keV	-0.06 ± 0.03	$-0.06^{+0.03}_{-0.03}$
Mds3555	1/keV	-0.08 ± 0.03	$-0.08^{+0.03}_{-0.03}$
Mds3564	1/keV	-0.04 ± 0.03	$-0.04^{+0.03}_{-0.03}$
Mds3567	1/keV	-0.12 ± 0.03	$-0.12^{+0.03}_{-0.03}$
۲Mn	cts/(kg.y)	15.3 ± 0.7	$15.3^{+0.7}_{-0.7}$
ΓΠ	cts/(kg.y)	0.5 ± 0.2	$0.5^{+0.2}_{-0.2}$

Bayesian fit signal + background on CUORE data - $\Gamma_{0\nu}$ prior allowed to non-physical values





S G Results on 0vββ decay of ¹²⁸Te S

Bayesian fit background-only on CUORE data

Parameter	Units	Global mode	Marginalized mode
BI _{ds3522}	cts/keV/kg/y	1.48 ± 0.02	$1.48^{+0.02}_{-0.02}$
Bl _{ds3552}	cts/keV/kg/y	1.43 ± 0.02	$1.43^{+0.02}_{-0.02}$
BI _{ds3555}	cts/keV/kg/y	1.49 ± 0.02	$1.49^{+0.02}_{-0.02}$
BI _{ds3564}	cts/keV/kg/y	1.48 ± 0.02	$1.48^{+0.02}_{-0.02}$
Bl _{ds3567}	cts/keV/kg/y	1.26 ± 0.02	$1.26^{+0.01}_{-0.02}$
m _{ds3522}	1/keV	-0.07 ± 0.03	$-0.06^{+0.02}_{-0.03}$
Mds3552	1/keV	-0.06 ± 0.03	$-0.06^{+0.03}_{-0.03}$
Mds3555	1/keV	-0.08 ± 0.03	$-0.08^{+0.03}_{-0.03}$
Mds3564	1/keV	-0.04 ± 0.03	$-0.04^{+0.03}_{-0.03}$
Mds3567	1/keV	-0.12 ± 0.03	$-0.12^{+0.03}_{-0.03}$
۲Mn	cts/kg/y	15.3 ± 0.7	$15.3^{+0.07}_{-0.07}$
Гл	cts/(kg y)	0.5 ± 0.2	$0.5^{+0.2}_{-0.2}$



