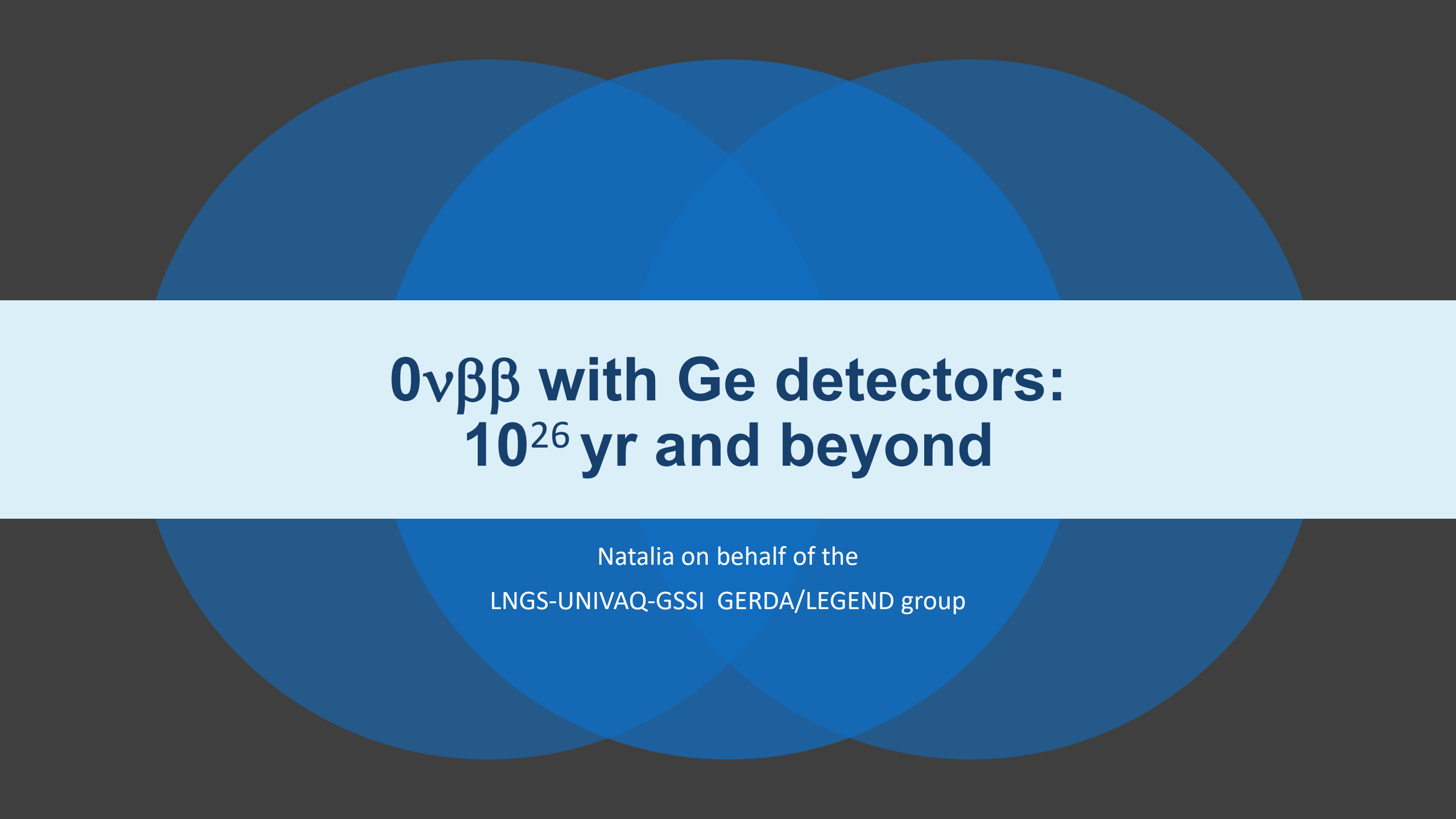
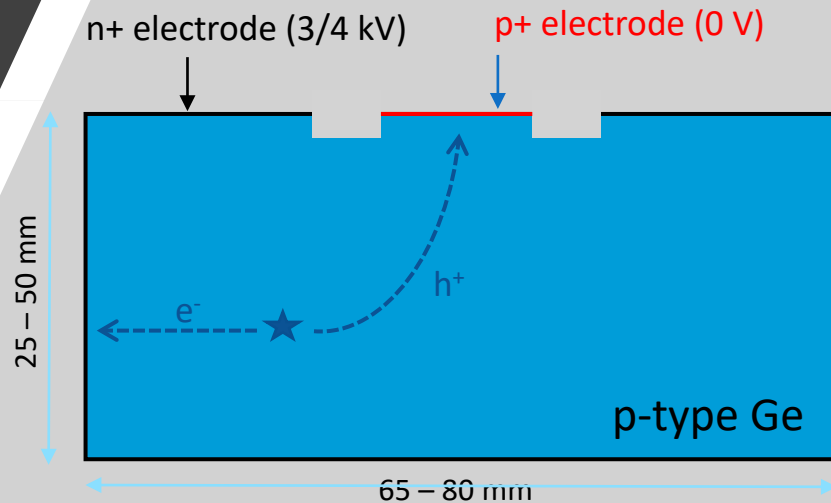




# $0\nu\beta\beta$ with Ge detectors: $10^{26}$ yr and beyond

Natalia on behalf of the  
LNGS-UNIVAQ-GSSI GERDA/LEGEND group

# Searching for $0\nu\beta\beta$ of $^{76}\text{Ge}$

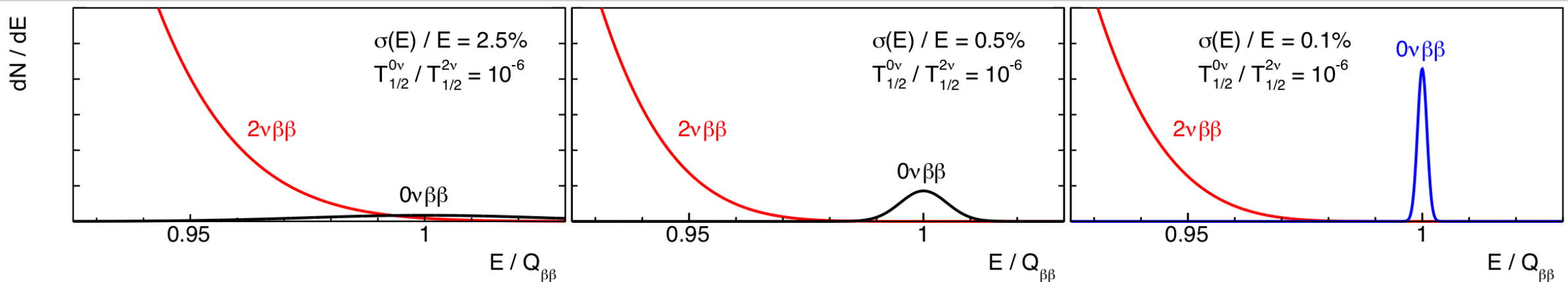


## Why high purity Ge detectors?

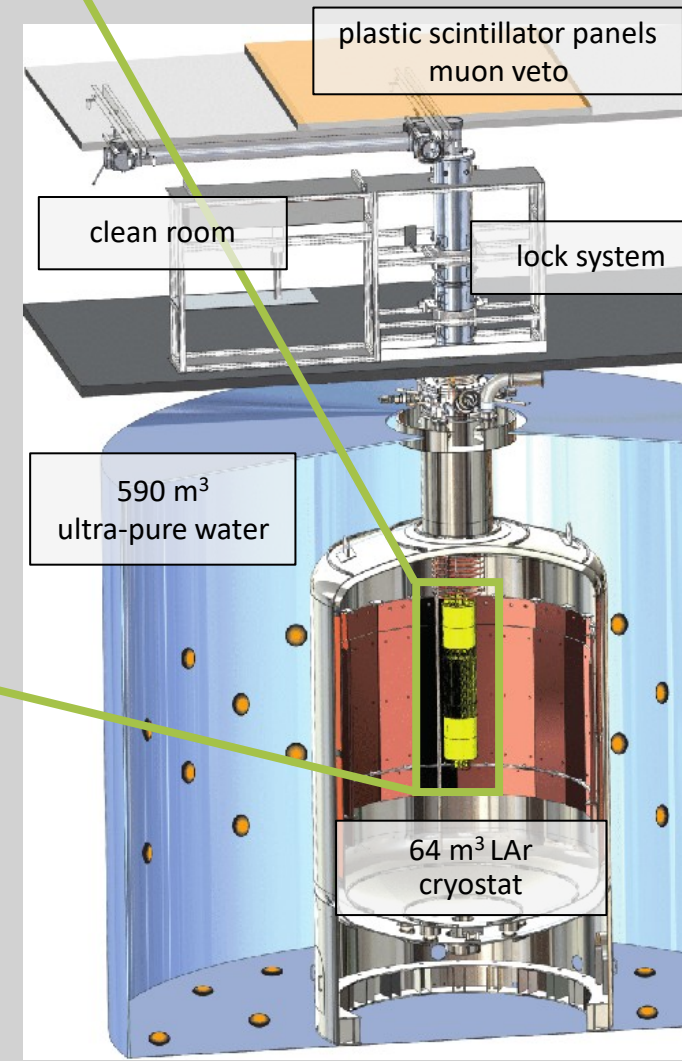
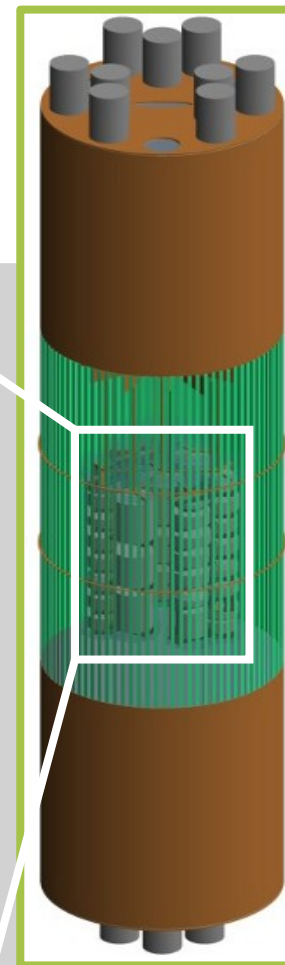
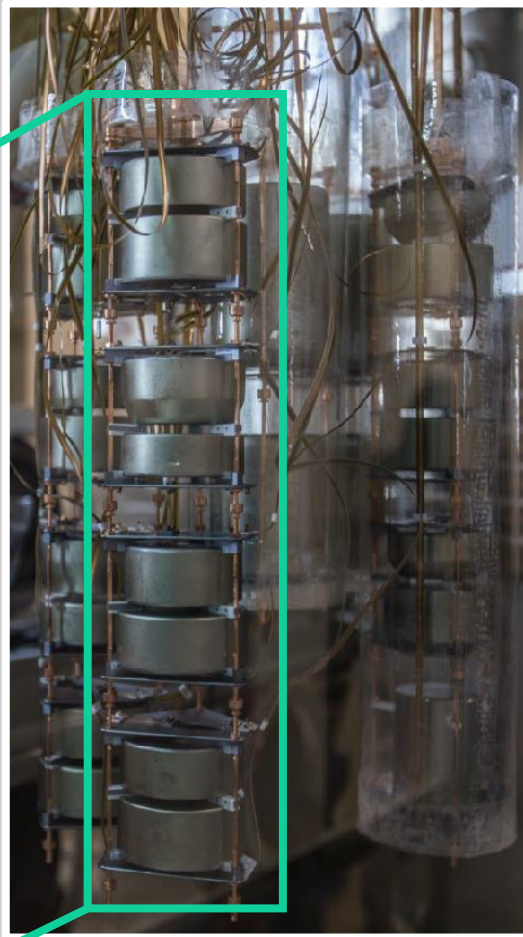
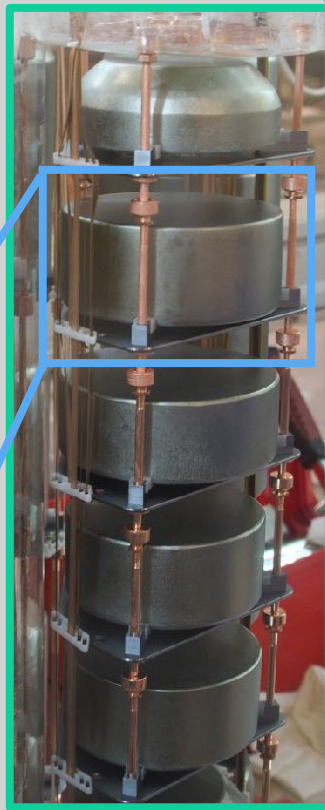
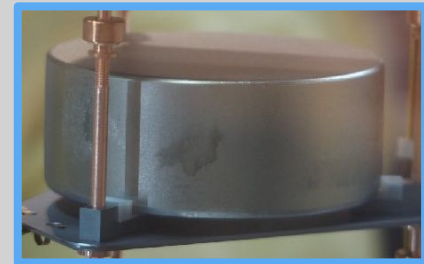
- source = detector → high efficiency
- radio-pure → no intrinsic background [Astropart.Phys. 91 (2017) 15-21]
- high density →  $e^-$  absorbed in 1-2 mm
- semiconductor →  $\sigma(E)/E < 0.1\%$  at  $Q_{\beta\beta}$
- enrichment up to 87% in  $^{76}\text{Ge}$

## $0\nu\beta\beta$ signature:

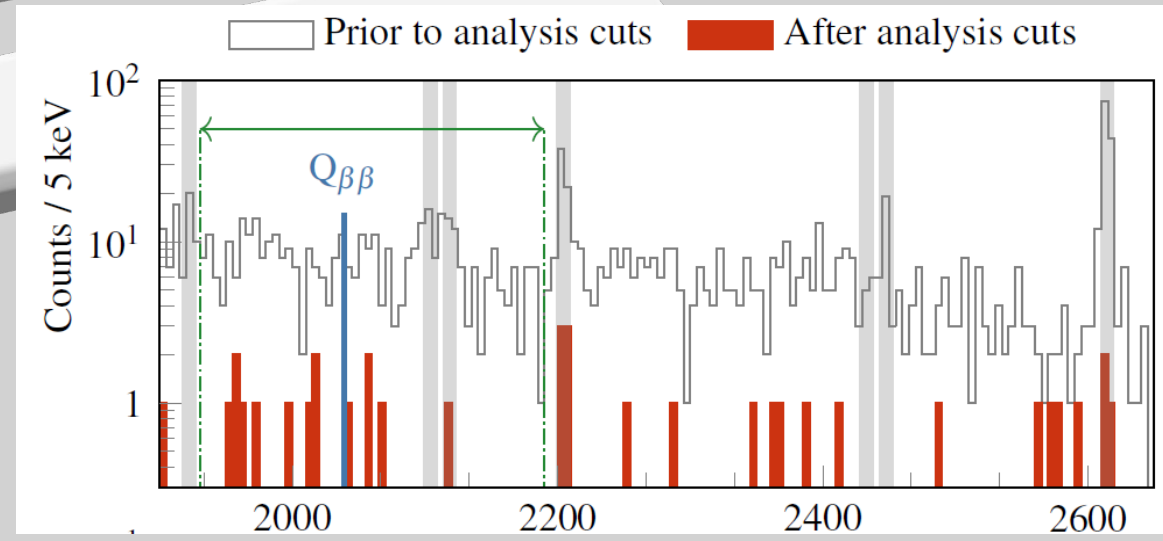
- point-like energy deposition in detector bulk volume
- sharp energy peak at 2039 keV (FWHM = 3-4 keV)



# The GERDA experiment

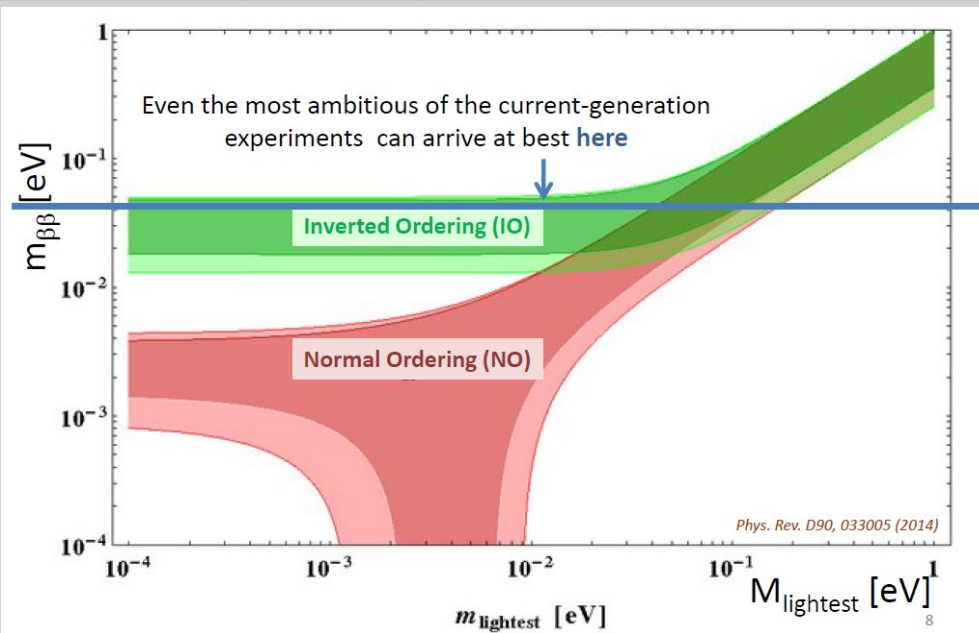


# The GERDA experiment



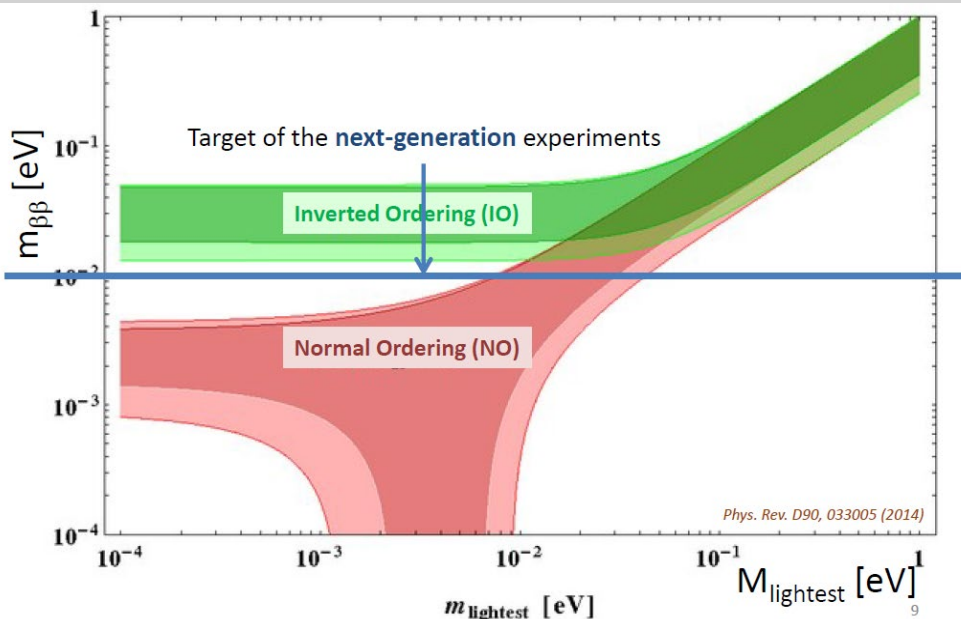
# $10^{26}$ yr and beyond

| EXPERIMENT  | Isotope           | Exposure [kg yr] | $T_{1/2}^{0\nu}$ [ $10^{25}$ yr] | $\langle m_{\beta\beta} \rangle$ [MeV] |
|-------------|-------------------|------------------|----------------------------------|----------------------------------------|
| GERDA       | $^{76}\text{Ge}$  | 127.2            | 18                               | 79-180                                 |
| MAJORANA    | $^{76}\text{Ge}$  | 26               | 2.7                              | 200-433                                |
| KamLAND-zen | $^{136}\text{Xe}$ | 594              | 10.7                             | 61-165                                 |
| EXO         | $^{136}\text{Xe}$ | 234.1            | 3.5                              | 93-286                                 |
| CUORE       | $^{130}\text{Te}$ | 115.9            | 1.5                              | 110-520                                |



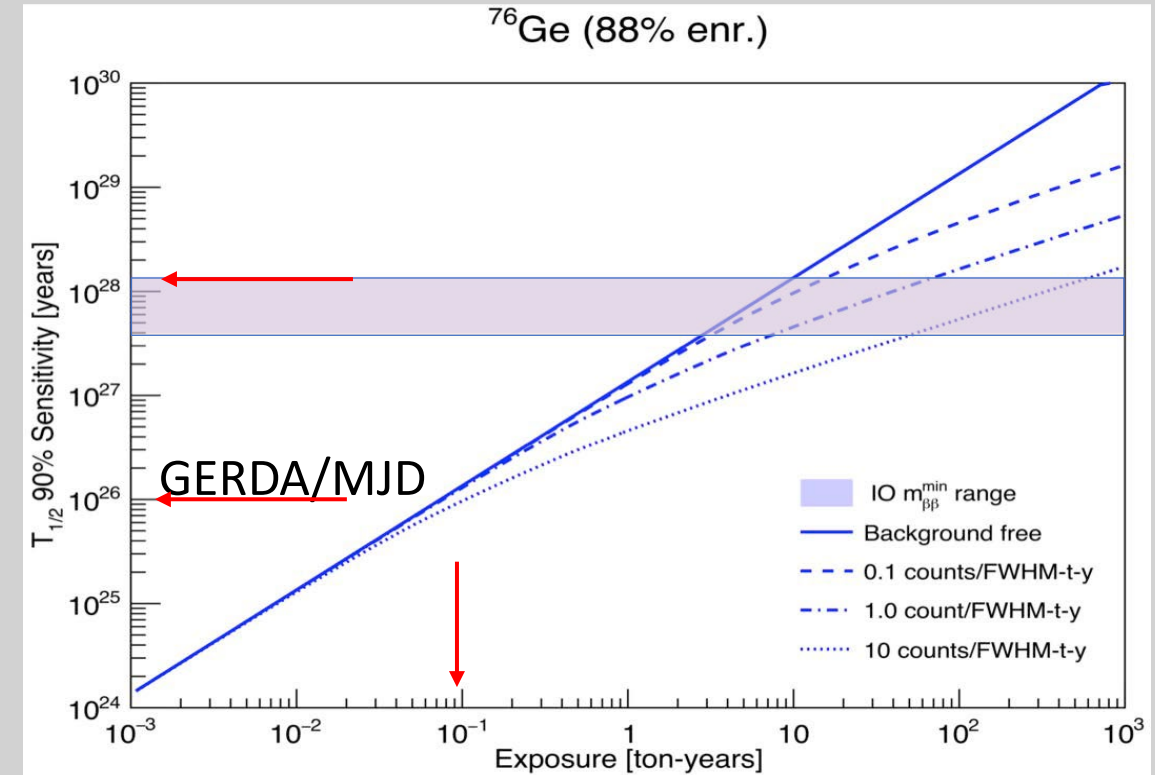
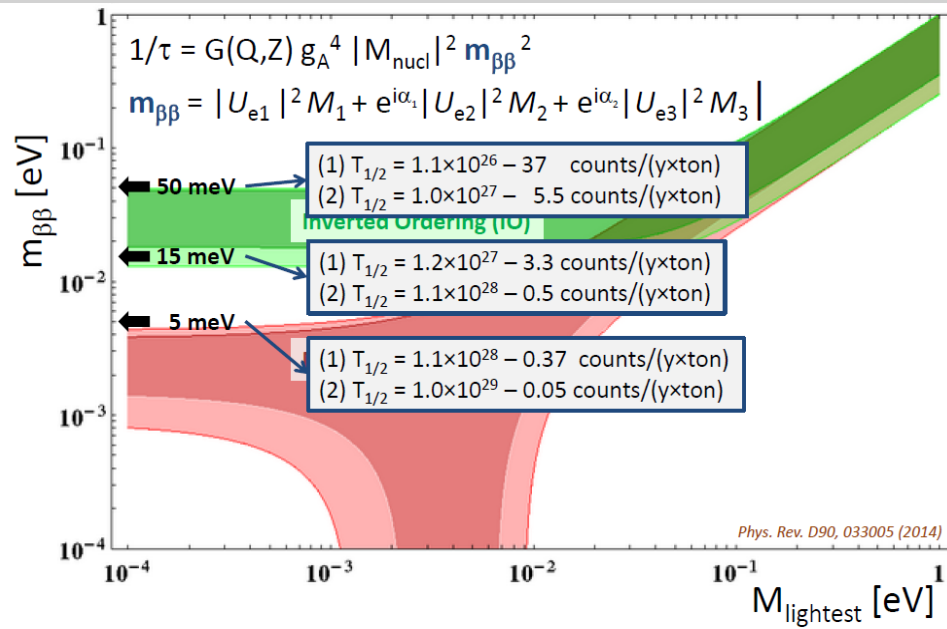
# $10^{26}$ yr and beyond

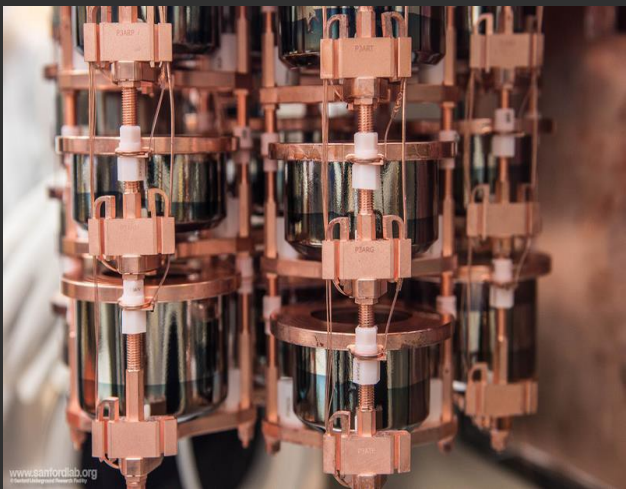
| EXPERIMENT  | Isotope           | Exposure [kg yr] | $T_{1/2}^{0n}$ [ $10^{25}$ yr] | $\langle m_{\beta\beta} \rangle$ [MeV] |
|-------------|-------------------|------------------|--------------------------------|----------------------------------------|
| GERDA       | $^{76}\text{Ge}$  | 127.2            | 18                             | 79-180                                 |
| MAJORANA    | $^{76}\text{Ge}$  | 26               | 2.7                            | 200-433                                |
| KamLAND-zen | $^{136}\text{Xe}$ | 594              | 10.7                           | 61-165                                 |
| EXO         | $^{136}\text{Xe}$ | 234.1            | 3.5                            | 93-286                                 |
| CUORE       | $^{130}\text{Te}$ | 115.9            | 1.5                            | 110-520                                |



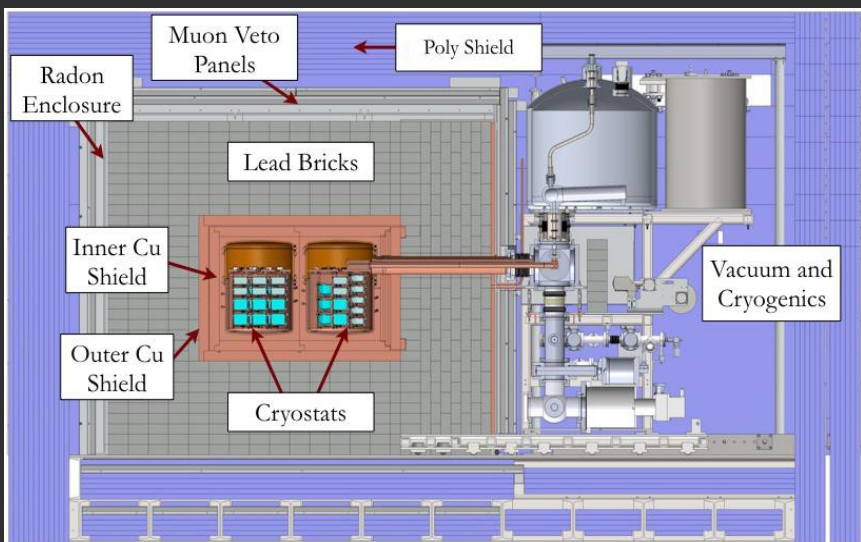
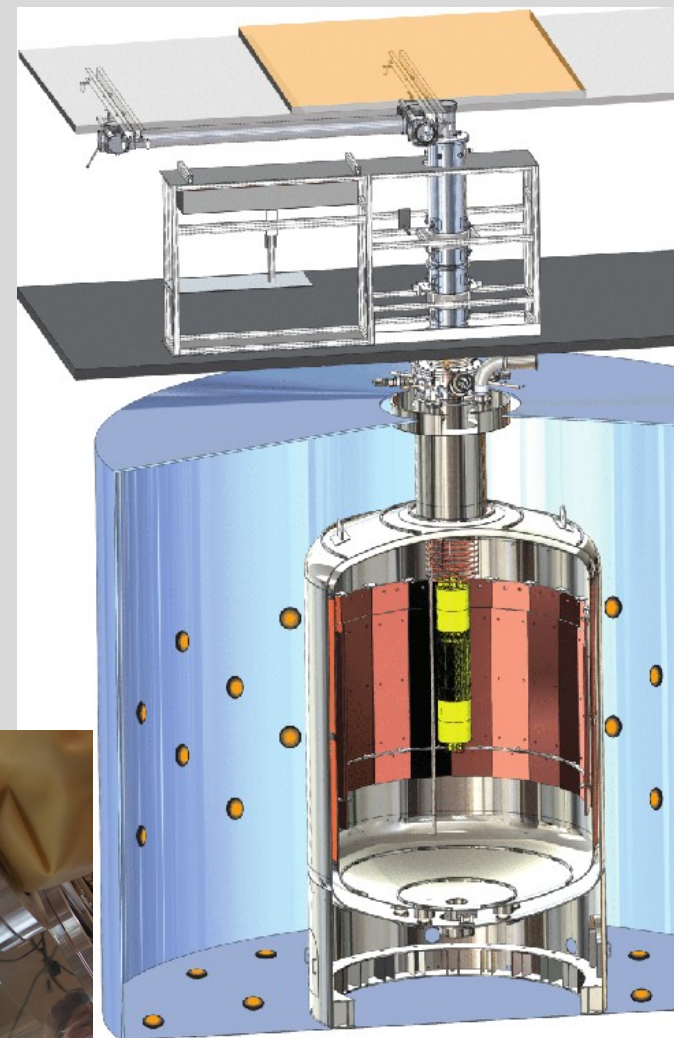
# 10<sup>26</sup> yr and beyond

| EXPERIMENT  | Isotope           | Exposure [kg yr] | $T_{1/2}^{0n}$ [10 <sup>25</sup> yr] | <mbb> [MeV] |
|-------------|-------------------|------------------|--------------------------------------|-------------|
| GERDA       | <sup>76</sup> Ge  | 127.2            | 18                                   | 79-180      |
| MAJORANA    | <sup>76</sup> Ge  | 26               | 2.7                                  | 200-433     |
| KamLAND-zen | <sup>136</sup> Xe | 594              | 10.7                                 | 61-165      |
| EXO         | <sup>136</sup> Xe | 234.1            | 3.5                                  | 93-286      |
| CUORE       | <sup>130</sup> Te | 115.9            | 1.5                                  | 110-520     |





Majorana  
DEMONSTRATOR



## GERDA

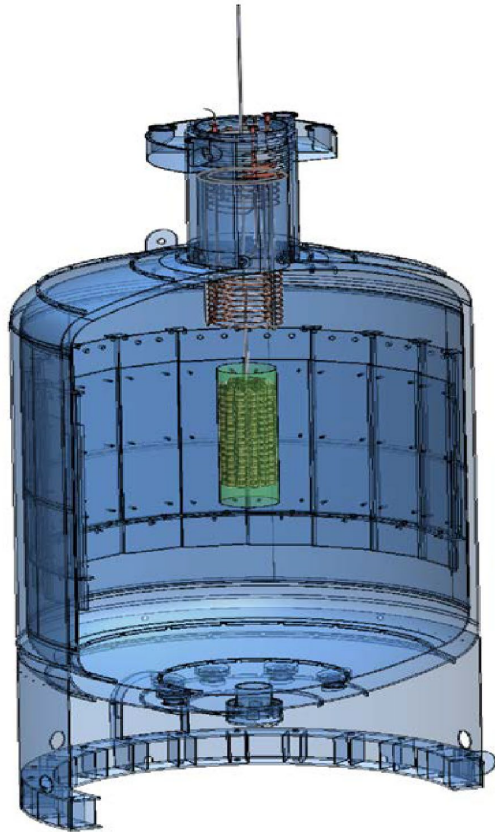
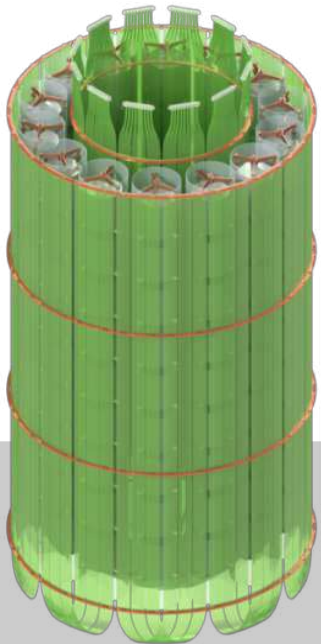
- LAr veto
- Low-A shield, no Pb

## Both

- Clean fabrication techniques
- Control of surface exposure
- Development of large point-contact detectors
- Lowest background and best resolution  $0\nu\beta\beta$  experiments

## MAJORANA

- Radiopurity of nearby parts (FETs, cables, Cu mounts, etc.)
- Low noise electronics improves PSD
- Low energy threshold (helps reject cosmogenic background)



- 70 inverted coax detectors (1.5-2 kg), about 140 kg
  - 28 BEGe's (0.7 kg) about 20 kg
  - 5 ICPC's (2.0 kg) about 10 kg
  - 33 PPC's (0.8 kg) about 28 kg
  - Semi-Coax detectors (either use as is, or recycle) about 15 kg
- Total ~200 kg



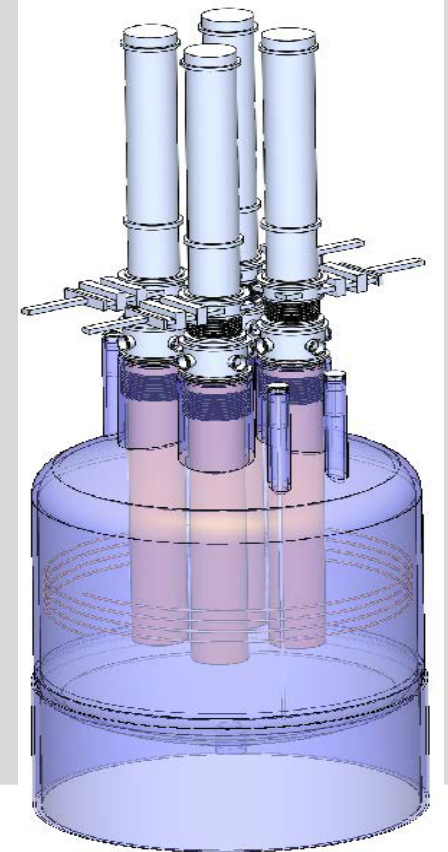
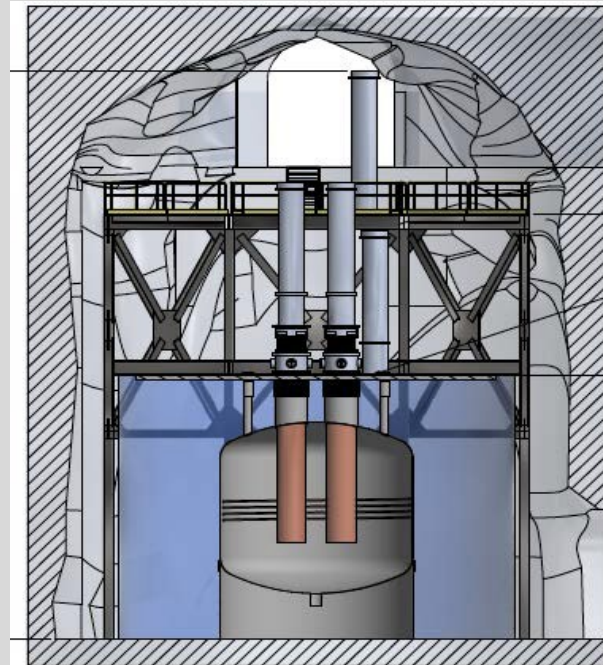


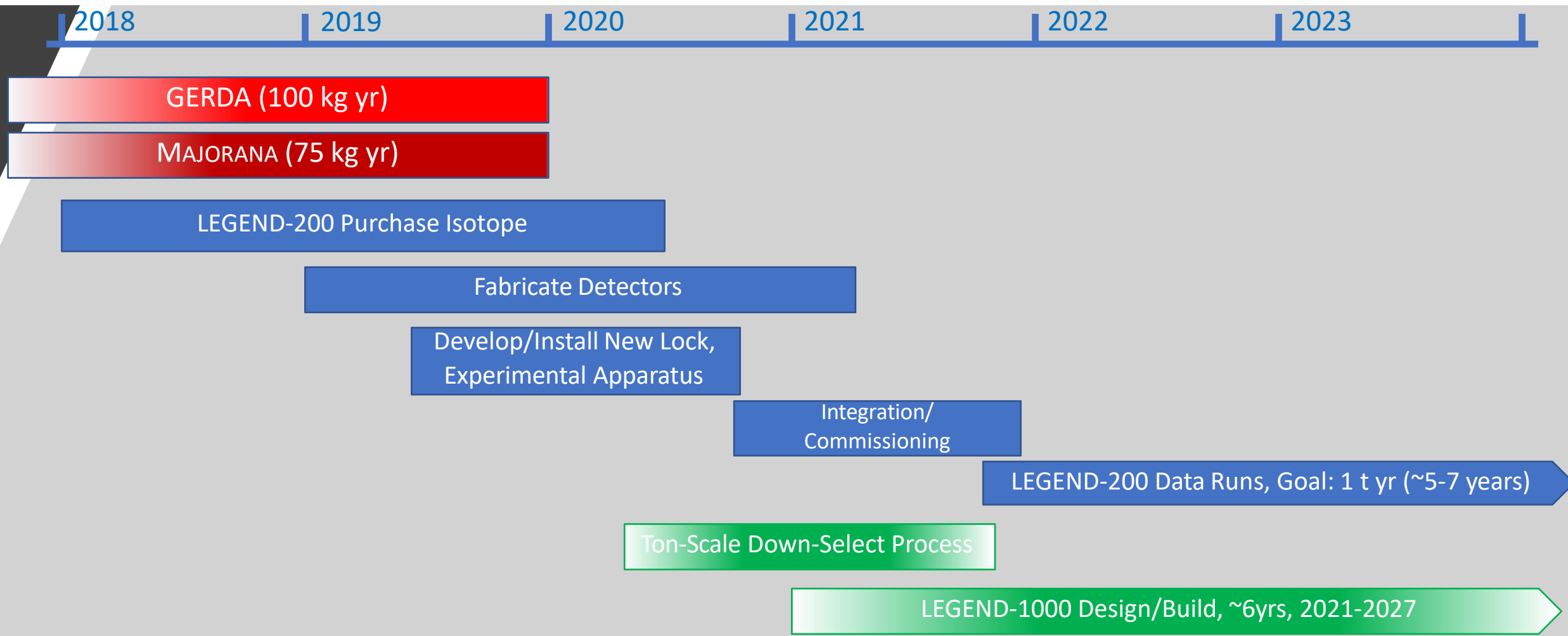
### First phase:

- (up to) 200 kg in upgrade of existing infrastructure at LNGS
- BG goal:  $<0.6 \text{ c}/(\text{FWHM t y})$
- Discovery sensitivity at a half-life of  $10^{27}$  years
- Data start end of 2021

### Subsequent stages:

- 1000 kg, staged via individual payloads
- Background goal  $<0.03 \text{ cts}/(\text{FWHM t yr})$
- Discovery sensitivity at a half-life of  $10^{28}$  years
- Location to be selected





Earliest, and optimistic, LEGEND-1000 Data Start 2025/6

# LEGEND 1000

## BSM Physics program

| Mechanism                           | Signature                                             | Energy range     | Status                | Recent Germanium References                                             |
|-------------------------------------|-------------------------------------------------------|------------------|-----------------------|-------------------------------------------------------------------------|
| Bosonic Dark Matter                 | Peak at $m_b$                                         | 5 — 100keV       | Done in MJD, GERDA    | PRL 118 (2017) 161801<br>PRL 125 (2020) 011801                          |
| Baryon Decay                        | Time Correlation, High Energy                         | 0-10 MeV         | Done in MJD           | PRD 99 (2019) 072004                                                    |
| Fractionally Charged Cosmic rays    | High Multiplicity-coincidence events                  | Few keV          | Done in MJD           | PRL 120 (2018) 211804                                                   |
| WIMP searches                       | Exponential Excess + Annual Modulation. Migdal Effect | < 10 keV         | CDEX/MALBEK/CoGeNT    | PRL 120 (2018) 241301; 1401.6234;<br>Phys.Procedia 61 (2015) 77         |
| Solar axions                        | Peaked Spectra + daily modulation                     | < 10 keV         | Partially Done in MJD | PRL 118 (2017) 161801; Astropart.Phys. 89 (2017) 39, Wiseman PhD Thesis |
| Majoron Emission                    | $2\nu\beta\beta$ spectral distortion                  | $Q_{\beta\beta}$ | Done in GERDA         | Eur. Phys.J. C75 (2015) 416                                             |
| Lorentz Violation                   | $2\nu\beta\beta$ spectral distortion                  | $Q_{\beta\beta}$ | TBD GERDA             | PRD 88 (2013) 071902                                                    |
| Electron Decay                      | Peak at 11.8 keV                                      | ~10 keV          | Done in MJD           | PRL 118 (2017) 161801                                                   |
| Pauli Exclusion Principle Violation | Peak at 10.6 keV                                      | ~ 10 keV         | Done in MJD           | PRL 118 (2017) 161801                                                   |
| BSM physics in Ar                   | Features in Ar Veto spectrum                          |                  | ECEC in Ar36 (GERDA)  | Eur. Phys. J. C75 (2015) 416                                            |

+ Prompt Supernova Neutrinos, SuperWIMPS, Solar Neutrinos, ...

# PEOPLE



Natalia Di Marco

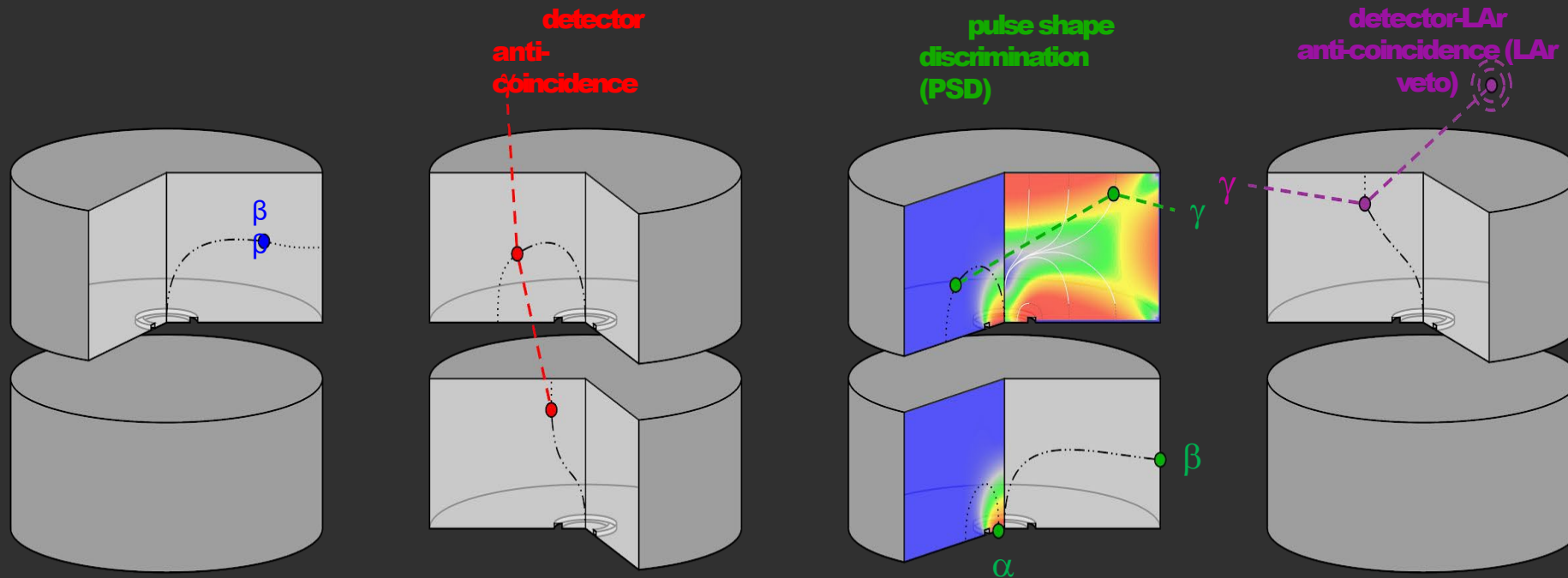


Francesco Salamida  
Carla Macolino  
Valerio D'Andrea



Matthias Laubenstein (Group Leader)  
Chiara Vignoli  
Matthias Junker  
Marco Balata  
Francesco Ferella

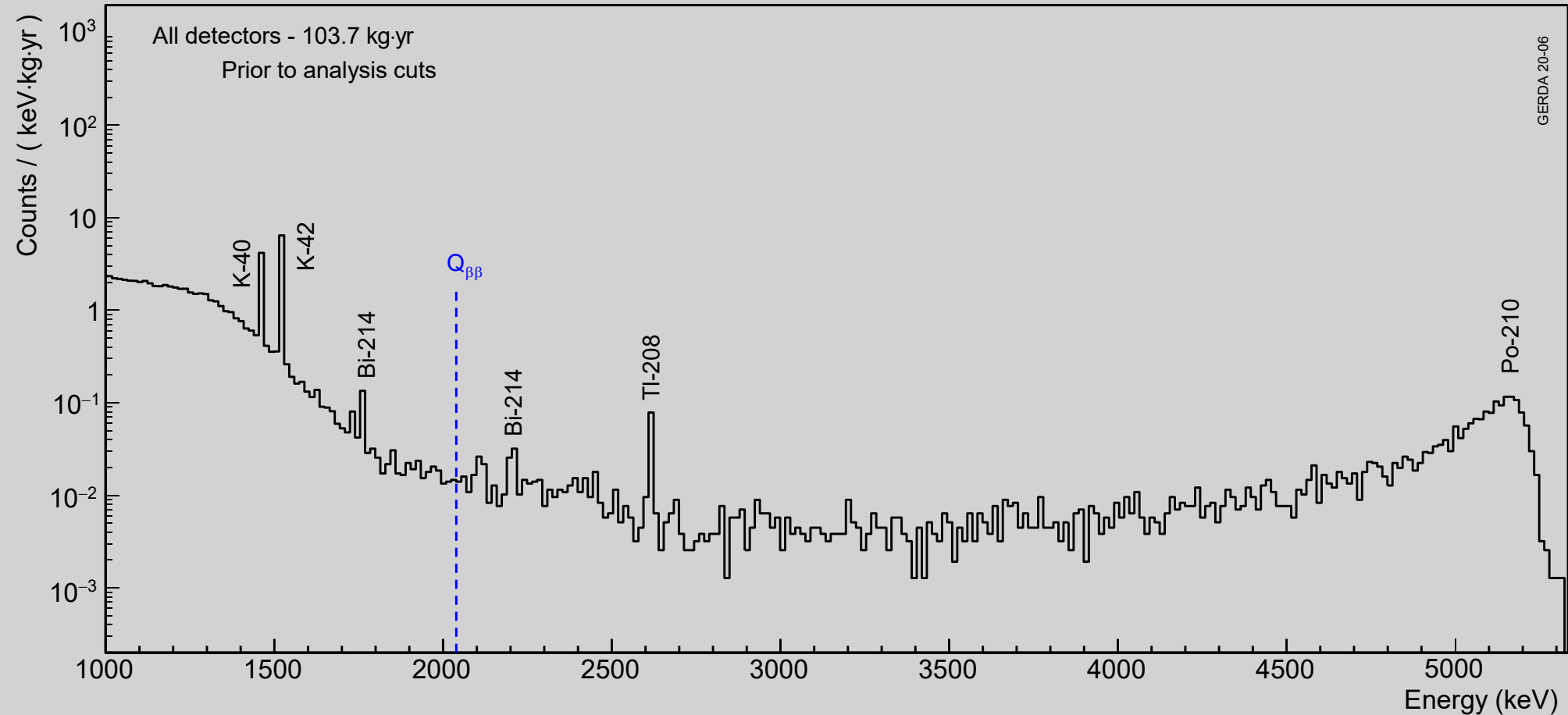
# The GERDA experiment



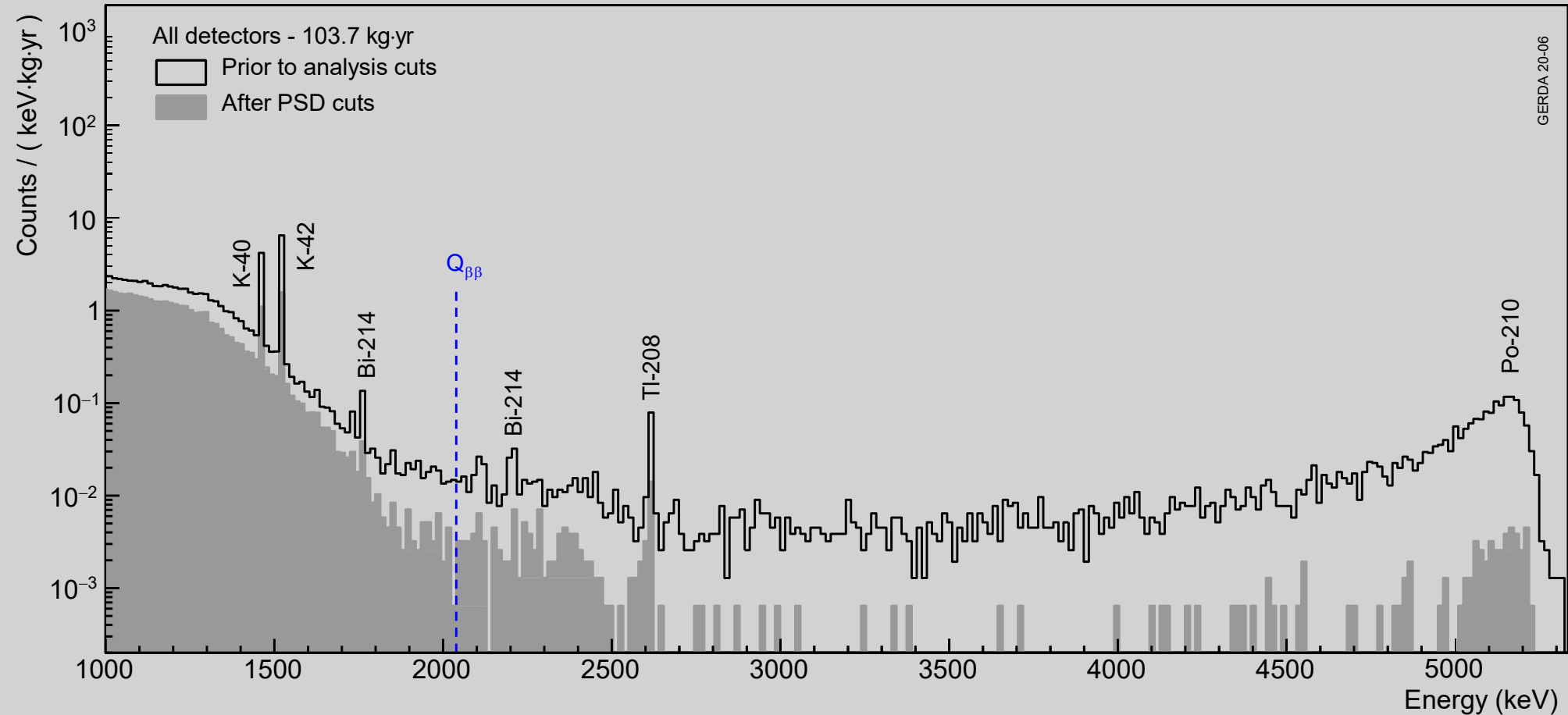
Discriminate **point like** (single site)  $\beta\beta$  topology from:

- **multi-detector** interactions
- **multi-site/surface** interactions
- interactions with **coincident energy**
- **deposition** in surroundings

# The GERDA experiment



# The GERDA experiment



# The GERDA experiment

