\*He recoils in CYGNO 10 L prototype

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Elisabetta Baracchini



# DRECTIONAL DIRECT DARK MATTER SEARCHES INFN

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PRIN:

Prot. 2017T54J9J



# G S Gravitational DM evidences G S I

*"Dark" Matter because it does not interact with light* 

**Galaxy rotation curves** 

# Galaxies motion inside clusters



**Galactic Collisions** 

Cosmic Microwave Background



**Big Bang Nucleosynthesis** 



### Larger scales explored, older times probed



E. Baracchini - CYGNO & INITIUM: Directional Dark Matter searches with optical readouts- GSSI Scientific Fair 2020

#### 2



### G S Directionality as key for unambiguous identification of DM S I

Increasing reliability of any observed signal, increasing difficulty in the experimental technique



Energy dependence: a falling exponential with <u>no peculiar features</u>

### G S Directionality as key for unambiguous identification of DM S I

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Universe 4 (2018) no.11, 116

Energy dependence: a falling exponential with <u>no peculiar features</u> Temporal dependence: <u>a few %</u> annual modulation

### G S Directionality as key for unambiguous identification of DM S I

#### Increasing reliability of any observed signal, increasing difficulty in the experimental technique









**DAMA/LIBRA** Collaboration 1-6 keV Residuals (cpd/kg/keV) DAMA/LIBRA-phase2 ~250 kg (1.13 ton×yr) 0.04 0.02 -0.02 -0.04 -0.06 6250 6500 6750 7000 7250 7500 7750 8000 8250 Time (day)

Universe 4 (2018) no.11, 116



Energy dependence: a falling exponential with <u>no peculiar features</u>

Temporal dependence: <u>a few %</u> annual modulation Directional dependence: an <u>O(1)</u> effect that no background whatsoever can mimic

Directional correlation with an astrophysical source is the only available POSITIVE identification of a DM signal

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#### Capability to reject isotropy down to low threshold, i.e. to fight all backgrounds, including neutral

A. M. Green et. al, Astropart. Phys. 27 (2007) 142



Directional detector can tolerate backgrounds, including neutral

WIMP signal in principle detectable with O(10) 3D events

E. Baracchini - Directional Dark Matter Searches with the CYGNO/INITIUM project - SIF 2019, GSSI, L'Aquila



# Directionality as tool for background rejection, neutrino physics and DM astronomy



#### Capability to reject isotropy down to low threshold, i.e. to fight all backgrounds, including neutral

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Capability to leap beyond the Neutrino Floor and to do Neutrino physics



The Neutrino Floor is an opportunity, not a limit

Sun neutrinos physics



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The Neutrino Floor is an opportunity, not a limit

## Sun neutrinos physics

# Capability to probe DM nature once discovered

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#### Phys.Rev. D98 (2018) no.10, 103006



WIMP & halo properties unbiased constrained with a single measurement

# DM astronomy & DM interactions

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# G SDirectionality: how well preserved in nuclear recoils?G SS IChoose your target material & resolutionS I



50 µm



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# G S Gas TPC concept & features S I



Helium-Fluorine gaseous target for simultaneous Spin Independent & Spin Dependent sensitivity to O(GeV) WIMPs



Energy loss and track topology to efficiently reject background at O(keV) energy threshold



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#### G S S I G S S I







A multi-site Galactic Nuclear Recoil Observatory at the tonscale to probe Dark Matter below the Neutrino Floor and measure <sup>8</sup>B solar Neutrinos <u>with directionality</u>

- Helium/Fluorine-based gaseous TPC for sensitivity to low mass WIMP region for both SI and SD couplings
- Goal of zero background operation after electron/gamma rejection and fiducialization at O(keV)
- Directional and gamma/electron rejection thresholds at O(keV)



#### G S S I CYGNO 3D optical readout: GEM + sCMOS + PMT S I

# **GEM** amplification



#### G S S I CYGNO 3D optical readout: GEM + sCMOS + PMT S I







#### G S S I G S S I

# sCMOS:

#### high granularity X-Y + energy measurements



Market pulled
 Single photon sensitivity
 Decoupled from target
 Large areas with proper optics
 Spectral response matched to CF4







• O(100) um 3D tracking with high quality particle identification (PID)

# G S CYGNO: photographing track S



### G S S I CYGNUS-RD project (2016-2018) G S S I

# JINST 13 (2018) no.05, P05001

## 10 x 10 x 1 cm<sup>3</sup> 0.1 Liters Triple thin GEMs CMOS & PMT on same side





## ORANGE: small prototype Optically ReAdout GEms Camera distance ± 18 cm

### PoS EPS-HEP2017 (2017) 077

24 x 20 x 20 cm<sup>3</sup> 9.6 Liters Triple thin GEMs CMOS & PMT on opposite sides



- 3D printed gas box
- 3D printed field cage with metallic rings
  semi-transparent cathode (wire mesh)
- LEMOn: large prototype Large Elliptical Module Optically readout Camera distance ± 53 cm

ific Fair 2020

Equipped with a suitable large aperture (f/0.95) and <u>E. Baracchini - CYGI</u> a short focal length (25 mm) lens A) Field Cage B) PMT

- (C) Adaptable bellow
- (D) CMOS camera

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#### G S S I CYGNUS-RD project (2016-2018) G S S I

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### He:CF<sub>4</sub> gas mixture with conventional electron drift

# G S He:CF41 atm G S CYGNUS-RD developments toward C/GNO S I

#### Energy threshold @ about 2 keVee



#### JINST 14 (2019) P07011

#### He:CF<sub>4</sub> 1 atm S S G G CYGNUS-RD developments toward CXGNO S

#### Energy threshold @ about 2 keVee



JINST 14 (2019) P07011

Absolute position determination along drift direction (background reduction through fiducialization)



Light transverse profile of both sCMOS images and PMT waveforms sensitive to absolute Z position via fit to diffusion

NIM A 936 (2019) 453-455

# G S He:CF41 atm G S CYGNUS-RD developments toward C/GNO S I

#### **Energy threshold** @ about 2 keV<sub>ee</sub>



#### Absolute position determination along drift direction (background reduction through fiducialization)



Light transverse profile of both sCMOS images and PMT waveforms sensitive to absolute Z position via fit to diffusion

NIM A 936 (2019) 453-455

#### Hint of 2D directionality and sense down to very low thresholds!





#### <u>JINST 14 (2019) P07011</u>

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# G S PHASE O: Mango & Lime



MANGO: a Multipurpose Apparatus for Negative ion studies with GEM Optically readout



#### LIME LIME: Long Imaging ModulE Field Cage (A) PMTs (B) Camera stand (C) В 1 sCMOS + 4 PMT Δ 33 x 33 cm<sup>2</sup> area D 50 cm drift Β 1/18 of CYGNO

- PMMA gas box + copper sheet for EM shielding/gamma shielding/darkening
- Field cage made of copper rings on PMMA supports
  - possibility of testing also DRIFT-like field cage (Kapton foil with copper strips) and **resistive foil**
- 4 PMT for positioning sensitivity with fast readout (center-of-gravity)

# 50 L volume



Under preliminary tests @ LNF, to be installed in underground LNGS by end 2020 to asses performance & measure neutron flux

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To be soon installed in our new overground LNGS lab in HdM



# G S PHASE 1: 1 m<sup>3</sup> detector @ LNGS UNFN



# G S Join us because, above all, is a lot of fun!

# #socialdetector













## https://web.infn.it/cygnus



