Gran Sasso Science Institute - February 25, 2020

The High Altitude Nater Cherenkov Observatory





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### Outline

#### Introduction & Motivation

The HAWC Observatory

Recent results

Outlook



#### Duldig, Science 314 (2006) 429-430









# UHECR detection<sub>\*</sub>

Auger Collaboration, Science 357 (2017) 1266-1270

 $E > 8 \text{ EeV}; 45^{\circ} \text{ smoothing}$ 



#### NASA's Fermi telescope resolves supernova remnants at GeV energies



- First resolved TeV γ-ray image of a Shell type SNR (Resolution ~10 arcmin)
- Acceleration source of cosmic rays, but is it evidence of protons?





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Tanaka et al., The Astrophysical Journal 685 (2008) 988

Leptonic

VS.

#### Hadronic

### Scientific Motivation

- Constrain the origin of cosmic rays by measuring gamma-ray spectra to 100 TeV.
- Probe particle acceleration in astrophysical jets with wide field of view, high duty factor observations.
- Explore new physics with an unbiased survey of the TeV sky.

# **Experimental Techniques**

- ✓ Background free
- ✓ Large duty cycle
- ✓ Large aperture

- Small area

Space-based detectors Low energy threshold EGRET, Fermi-LAT



# **Experimental Techniques**

- ✓ Large effective area
- Excellent background rejection

- Small aperture
- Low duty cycle

Imaging Atmospheric Cherenkov Telescopes High sensitivity HESS, MAGIC, VERITAS



# **Experimental Techniques**

- ✓ Large aperture
- Excellent background rejection
- ✓ Large duty cycle



- Moderate area
- Ground array of air-shower particle detectors Large aperture + High duty cycle Milagro, Tibet, ARGO, HAWC



# HAWC

- 2<sup>nd</sup> generation water Cherenkov
- Wide instantaneous field of view (2 sr)
- High duty cycle (> 90%)
- Large area (22,000 m<sup>2</sup>)

A second generation wide-field  $\gamma$ -ray detector

### Main Features

- Most bright Galactic GeV sources
  extend to TeV
- Best instrument for hard spectrum and extended sources



### The HAWC Observatory



300 - 7 m x 5 m steel Water Cherenkov Detectors (a.k.a. tanks) with 4 PMTs at 4,100 m a.s.l. in Mexico



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#### Effect of the laser calibration on the observation of the shadow of the Moon

o deflection matches 2 TeV median energy

- angular resolution < shadow width of 1.2°</p>
- o position verifies pointing



### HAWC site



# HAWC site

#### LMT (4,600 m)



#### Pico de Orizaba (18,500 ft)

### Deployment status



From 2011 to 2015

# Deployment status



From 2011 to 2015

### Design improvements



### Design improvements



Fermi-LAT sky smoothed map E > 50 GeV (Pass 8 - 6 years of data) (courtesy of M. Ajello)

### Design improvements



# Science Kesults

Penn State HAWCers in Mexico (2017)

#### Abeysekara et al [HAWC] ApJ 843 (2017) 39 The Crab



#### Abeysekara et al [HAWC] ApJ 843 (2017) 39 The Crab



Abeysekara et al [HAWC] ApJ 843 (2017) 40

### first HAWC catalog



#### 2HWC J1930+188

- coincident with VER J1930+188
  SNR G54.1+00.3 PSR J1930+1852
- TeV emission was reported to be pointlike and likely from PWN
   nearby molecular CO cloud



VERITAS pt-src upper limit
 ~1.4% of Crab

2HWC J1927+187\*

- associated with 2HWC J1930+188?
- ongoing analysis on spatial morphology







[VERITAS+Fermi-LAT+HAWC!] ApJ 866 (2018) 24

#### Abeysekara et al [HAWC] ApJ 843 (2017) 40



#### Abeysekara et al [HAWC] ApJ 843 (2017) 40



ZHWC J1939+2270

#### Geminga



#### Abeysekara et al [HAWC] Science 6365 (2017) 911-914

Positron excess from nearby pulsars?



Dec. [deg]

#### Abeysekara et al [HAWC] Science 6365 (2017) 911-914





#### Abeysekara et al [HAWC] Science 6365 (2017) 911-914



Estimated positron energy flux at Earth

#### Abeysekara et al [HAWC] Nature 562 (2018), 82-85



#### VHE emission from the jets of a microquasar

#### Abeysekara et al [HAWC] Nature 562 (2018), 82-85



#### VHE emission from the jets of a microquasar

# Recent Developments

Elemen

Penn State HAWC group at home (2018)

#### Abeysekara et al [HAWC] ApJ 881 (2019) 134





Abeysekara et al [HAWC] Phys. Rev. Lett. 124 (2020) 021102

### $\gamma$ -ray sky above 56 TeV







#### Abeysekara et al [HA





#### FIG. 3 of the PRL





# Outriggers



### Outlook: HAWC South







### Conclusions

#### Main results

irsa Minor

Coma Beregifes

HydroTi

30

Sextans

Can's Major

/\_30

Conak-60max

Sculpt

Extended regions, transient events, highest energies

#### Other results

Pisces

- Dark matter, extended regions, • diffuse emission, cosmic rays, ...
- EBL, solar physics, ...

Sagittan.

30

Region C Region B Region A Ursa Major I Wittend Major I PSR 80540+23 & HAVIS[0543+233 Triangulum II Segue II Segue9 0 කො

### Conclusions

#### Main results

Extended regions, transient
 events, highest energies



## Conclusions

#### Main results

 Extended regions, transient events, highest energies

#### Other results

- Dark matter, extended regions, diffuse emission, cosmic rays, ...
- EBL, solar physics, ...

#### Multi-wavelength physics

- MoUs with IceCube, IACTs, etc
- AMON
- HAWC alerts

#### Outlook

- Array of Outriggers
- Southern Observatory





The HAWC Collaboration

# Recent HAWC publications

- "Multiple Galactic Sources with Emission Above 56 TeV Detected by HAWC," Physical Review Letters 124 (2020) 021102
- "Measurement of the Crab Nebula Spectrum Past 100 TeV with HAWC," The Astrophysical Journal 881 (2019) 134
- "Searching for dark matter sub-structure with HAWC," Journal of Cosmology and Astroparticle Physics 07 (2019) 022
- "MAGIC and Fermi-LAT gamma-ray results on unassociated HAWC sources," Monthly Notices of the Royal Astronomical Society 485, 356 (2019)
- "All-sky Measurement of the Anisotropy of Cosmic Rays at 10 TeV and Mapping of the Local Interstellar Magnetic Field," The Astrophysical Journal 871, 96 (2019)
- "Very-high-energy particle acceleration powered by the jets of the microquasar SS 433," Nature 562, 82-85 (2018)
- "Constraints on spin-dependent dark matter scattering with long-lived mediators from TeV observations of the Sun with HAWC," Physical Review D 98, 123012 (2018)