#### ET - Specific on baffling strategy (M. Andrés, M. Martínez)

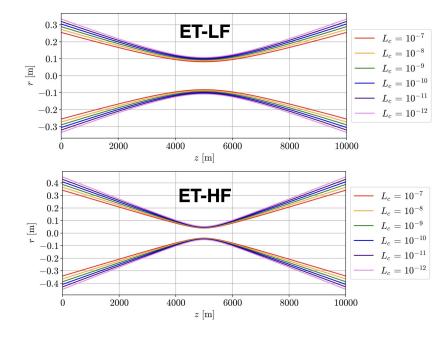
- Stray Light Simulations for ET
  - Beam losses and apertures
  - Baffles layout inside the ET tube
  - Preliminary noise estimations
- New studies
  - First studies on non-ideal configurations
  - Baffle vibrations
  - Tolerances on beam-pipe alignment
- Case for instrumented baffles

Minimum aperture of the tubes

To be able to set the minimum aperture of the vacuum tube to start iterating with the design, the level of Gaussian clipping losses is set to 1e-8. With this choice, and using the expression

$$r(z, L_c) = \frac{w(z)}{\sqrt{2}} \sqrt{\ln\left(\frac{1}{L_c}\right)} + r_{\text{offset}}$$

which allows for an additional beam offset, the minimum apertures of the baffles are of 84cm for ET-HF and 62cm for ET-LF. Assuming a typical baffle height of 8cm, the minimum diameter of the tubes are 1m and 0.8m, respectively.



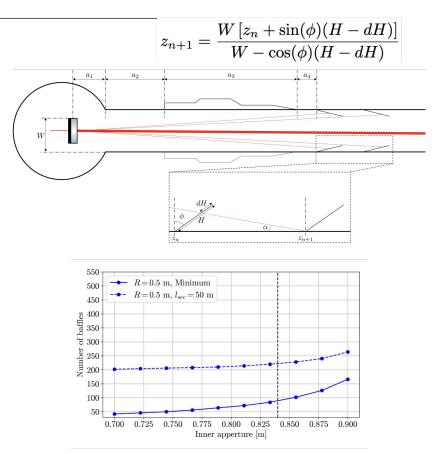
IFAES

Baffle layout inside the main arms

The baffling strategy is that followed both in LIGO and Virgo:

- Using only geometrical arguments, all the pieces of the tube have to be shielded by a baffle.
- In the middle section, whenever zn+1 predicted by the geometrical arguments is larger than the length of a tube section, a baffle is placed at each intersection between sectors.

The number of baffles per FP cavity are 244 (ET-HF) and 222 (ET-LF).







Preliminary noise estimation

#### BACKSCATTERING

We use Thorne and Vinet's formula to compute the backscattering noise from the power distribution that will reach the baffles. This power is obtained using the simulation software SIS and then the noise calculated as

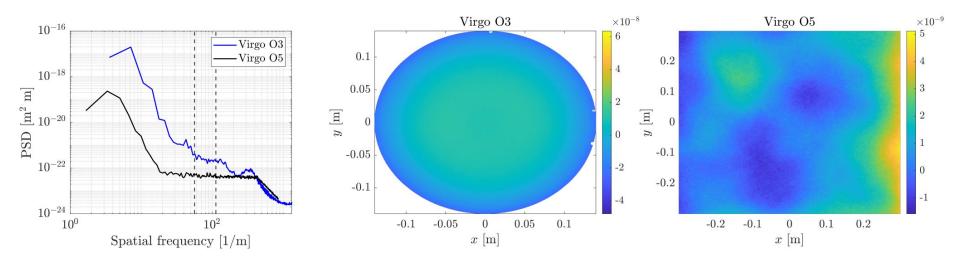
$$d\tilde{h}^{2}(f) = \frac{1}{L^{2}} \left[ \lambda^{2} + \left( \frac{8\Gamma P_{circ}}{cM\pi f^{2}} \right)^{2} \right] \frac{dP}{d\Omega_{bs}} X^{2}(f) dK$$
$$dK = \frac{1}{z^{2}} \left( \frac{dP}{d\Omega_{ms}} \right)^{2} \delta\Omega_{ms}$$

#### DIFFRACTION

We use an analytical method to estimate the diffraction noise caused by randomly serrated baffles.

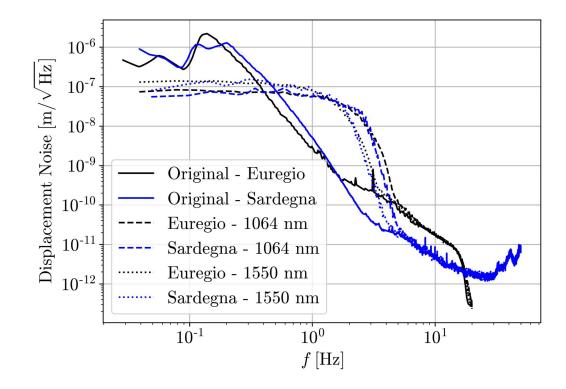
$$\tilde{h}_{\text{diff}}(f) = \sqrt{1 + \left(\frac{8\Gamma P_{circ}}{cM\pi f^2}\right)^2 \frac{1}{\lambda^2}} \\ \times \frac{\kappa\lambda X(f)\sqrt{N_B}}{LR} \left[\frac{\lambda L}{8\pi R\overline{\Delta H}}\right] \left[\frac{\sqrt{\lambda L/4}}{2\pi R}\right]^{1/2}$$

Preliminary noise estimation

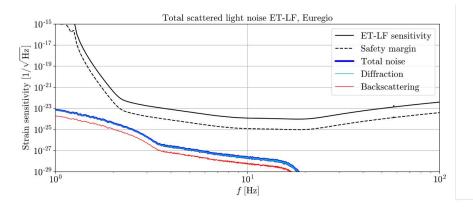


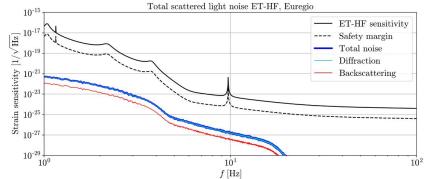


Preliminary noise estimation



Preliminary noise estimation

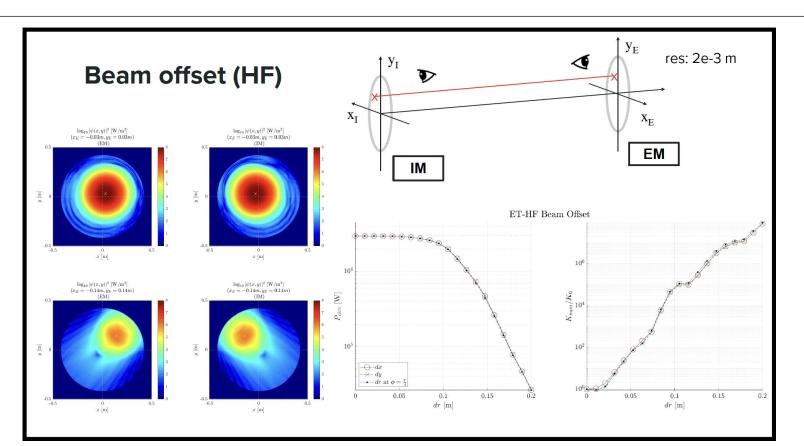




FAS

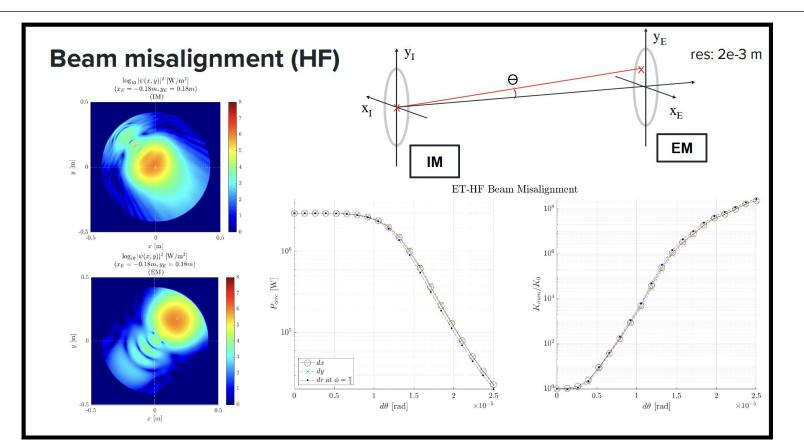
Non-ideal cavities



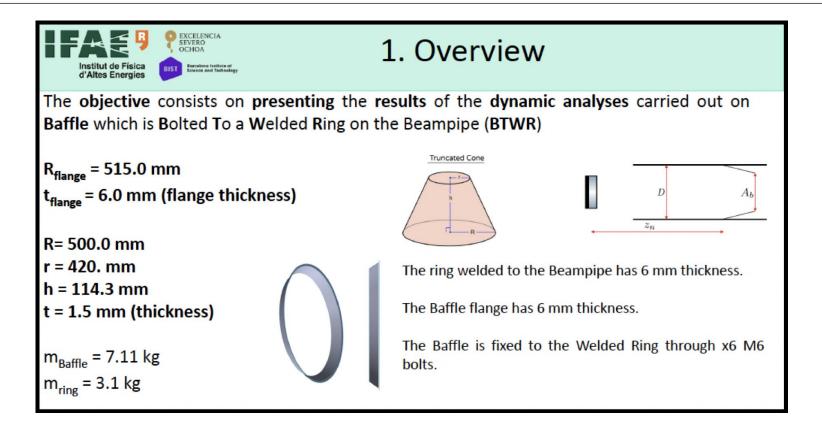


Non-ideal cavities





Vibration analysis



FAS

Vibration analysis



